

Superheavy Elements

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2026 Jun 23





- Motivation & Background
 - History and status of heavy & superheavy element (SHE) production
 - Why are they interesting??
- SHE Recoil Separators
 - General methods & approaches
 - Current status of BGS
- Recent successes
 - Synthesis of ^{290}Lv ($Z=116$)
 - Mass Identification



Cat's Eye Nebula



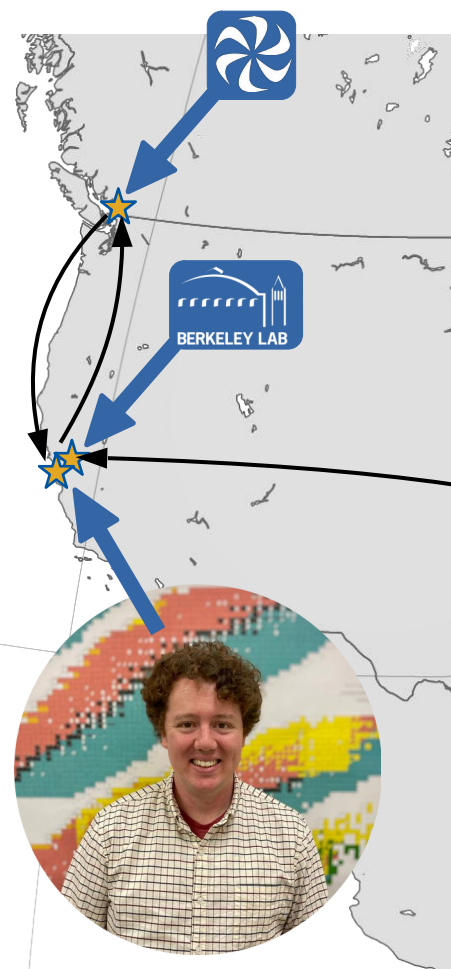
About Me



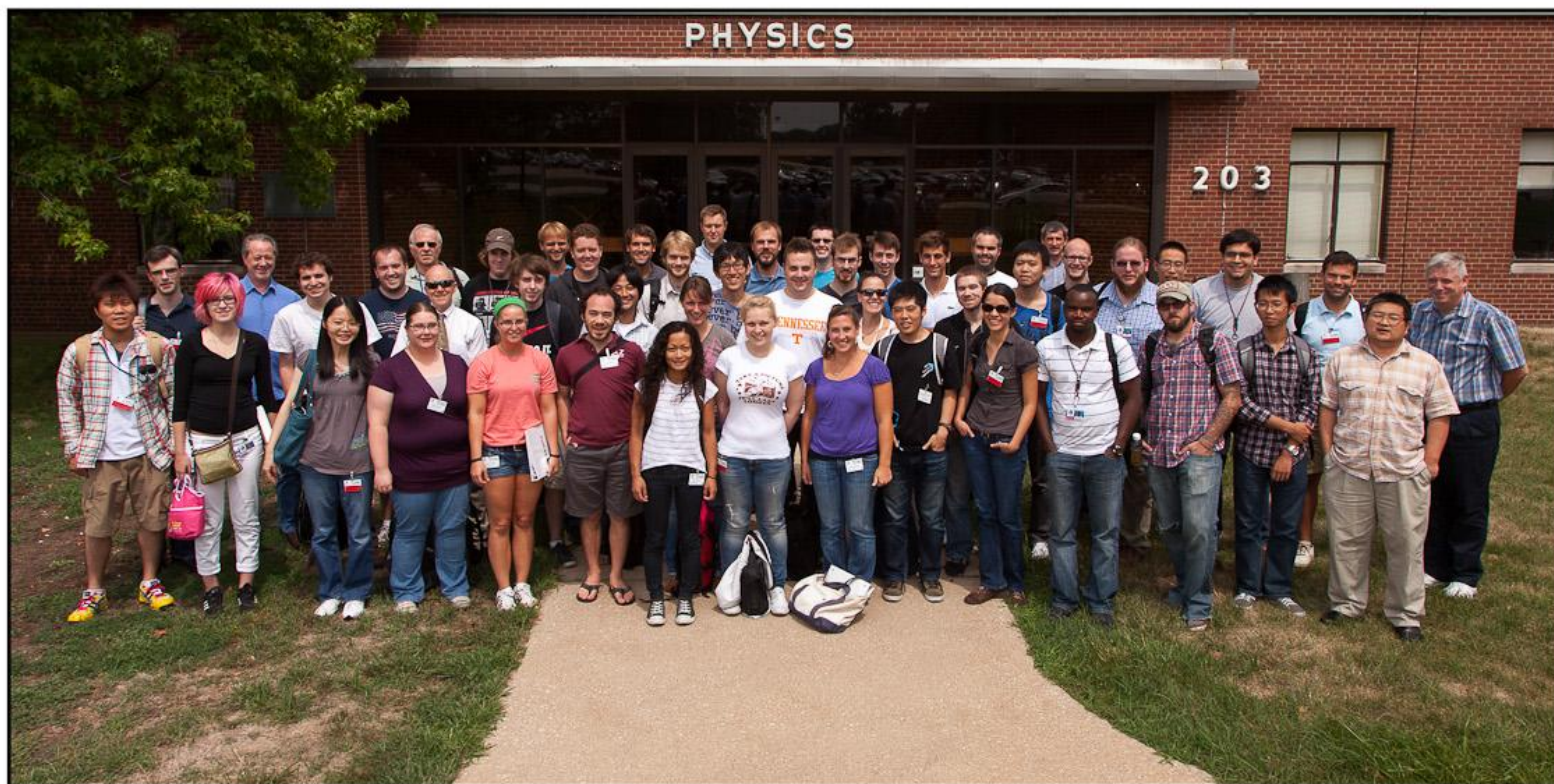
- Undergrad: Florida Atlantic U. (4 yrs)
 - First generation college student
 - No schooling / training in nuclear, but loved science, especially chem & physics
- Graduate: UC Berkeley (6.5 yrs)
 - Worked on super-heavy element reactions
 - EBSS 2012
- Post-doc: TRIUMF (2.5 yrs)
 - EMMA: nuclear structure & astrophysics
- Current: SJSU (6 yrs)
 - Teach nuclear chemistry and research targetry!
 - Goal: support student activity in nuclear science
 - Continue collaborating with the LBNL Heavy Elements group



About Me



EBSS 2026



E B S S 2 0 1 2
A U G U S T 5 - 1 0 , 2 0 1 2



The Elements

1 H																	2 He
3 Li	4 Be											5 B	6 C	7 N	8 O	9 F	10 Ne
11 Na	12 Mg											13 Al	14 Si	15 P	16 S	17 Cl	18 Ar
19 K	20 Ca	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	35 Br	36 Kr
37 Rb	38 Sr	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	52 Te	53 I	54 Xe
55 Cs	56 Ba	57 La ^{>}	72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	84 Po	85 At	86 Rn
87 Fr	88 Ra	89 Ac*	104 Rf	105 Db	106 Sg	107 Bh	108 Hs	109 Mt	110 Ds	111 Rg	112 Cn	113 Nh	114 Fl	115 Mc	116 Lv	117 Ts	118 Og

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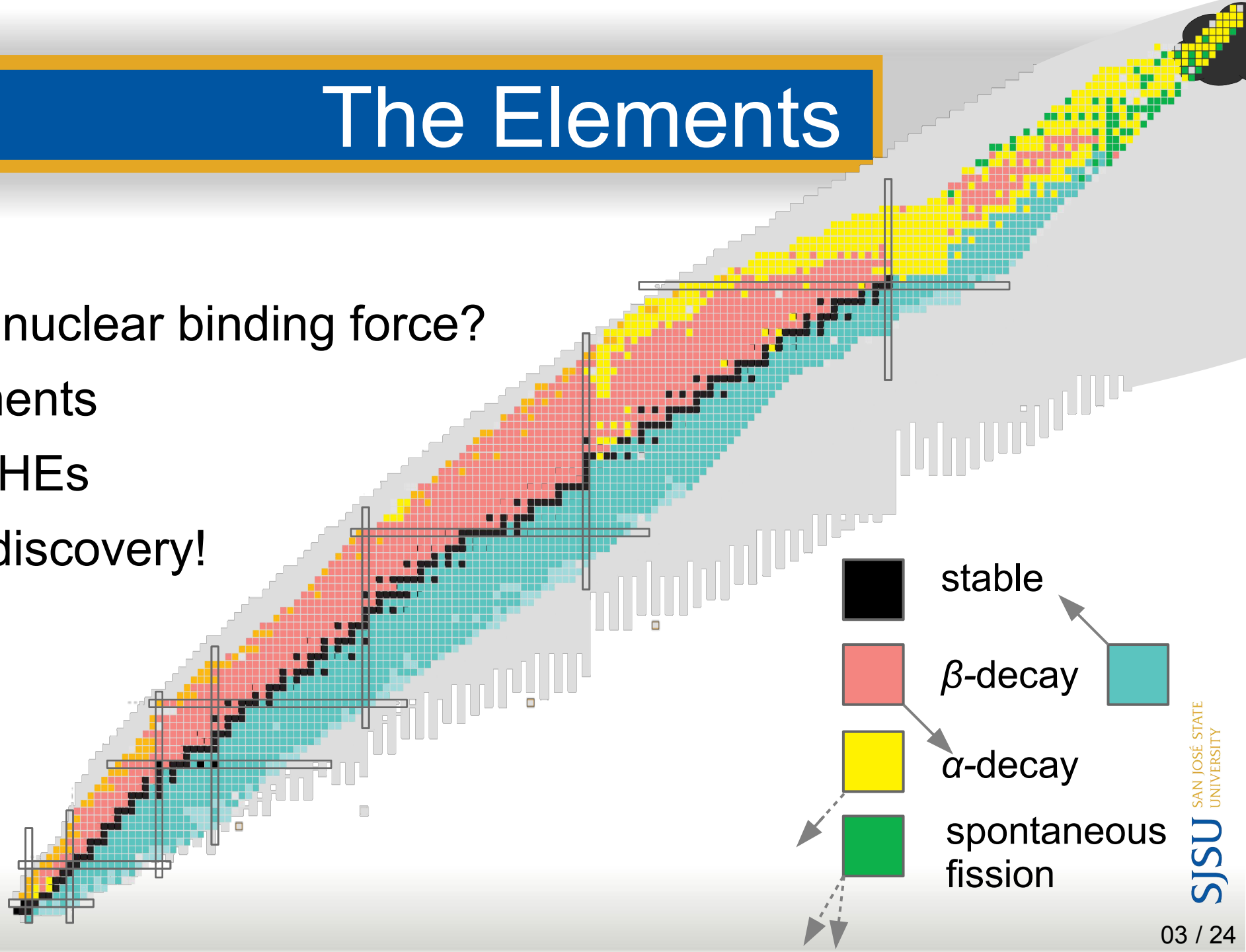
58 Ce	59 Pr	60 Nd	73 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy	67 Ho	68 Er	69 Tm	70 Yb	71 Lu
90 Th	91 Pa	92 U	93 Np	94 Pu	95 Am	96 Cm	97 Bk	98 Cf	99 Es	100 Fm	101 Md	102 No	103 Lr



The Elements

Motivations:

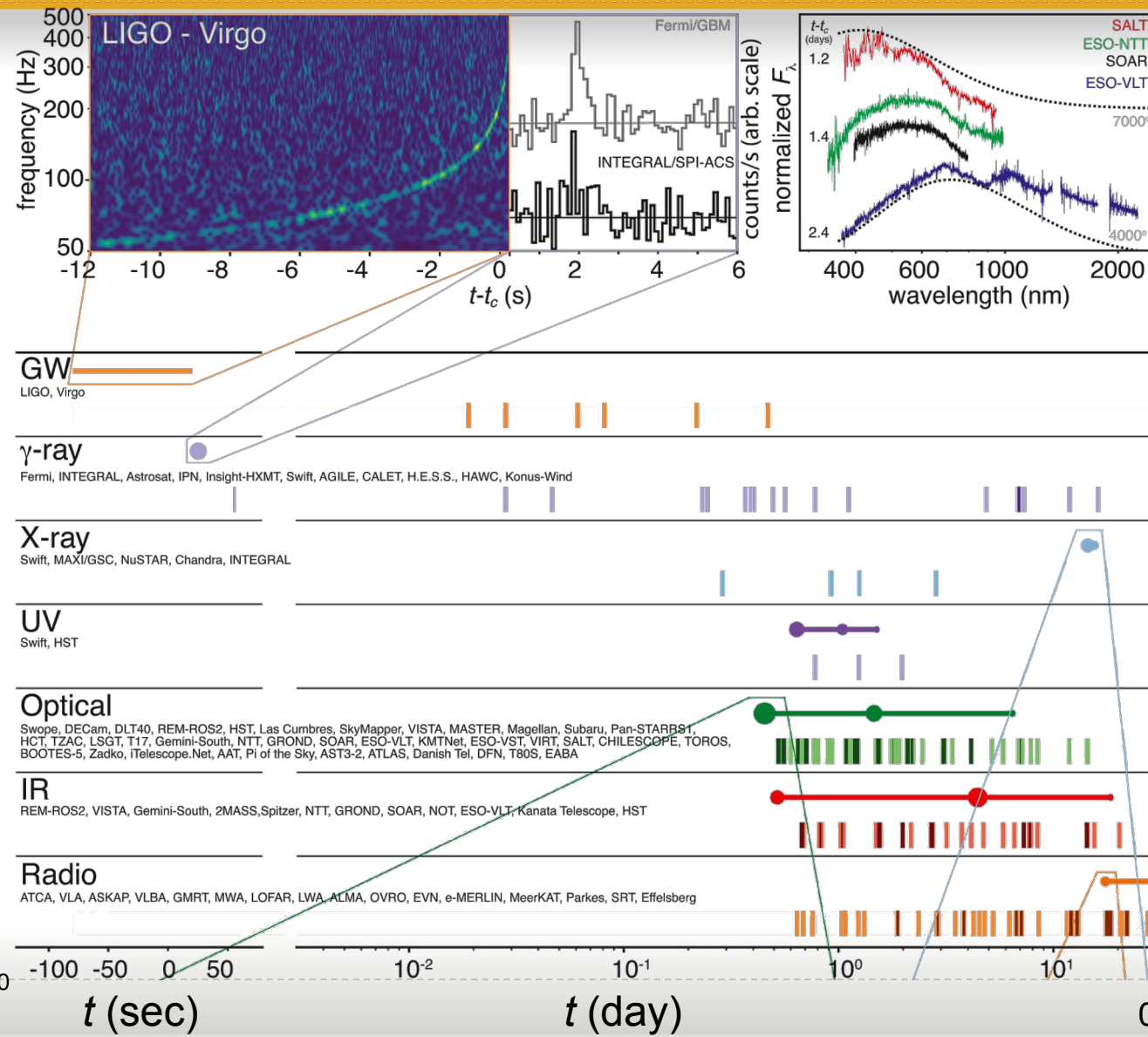
- Upper limits of nuclear binding force?
- Origins of elements
- Chemistry of SHEs
- Excitement of discovery!





Origin of the elements

- Evidence for r -process in neutron star mergers
- Start of the “age of multimessenger astronomy!”
- To study these radioactive nuclei, they must first be produced in the lab

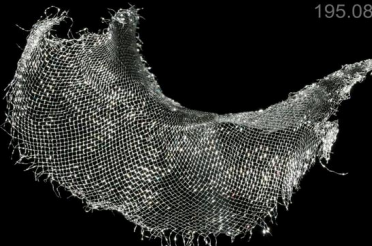
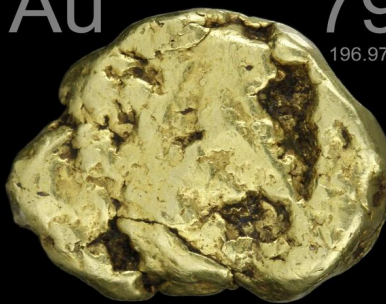


The Astrophysical Journal Letters, 848:L12(59pp), 2017 October 20



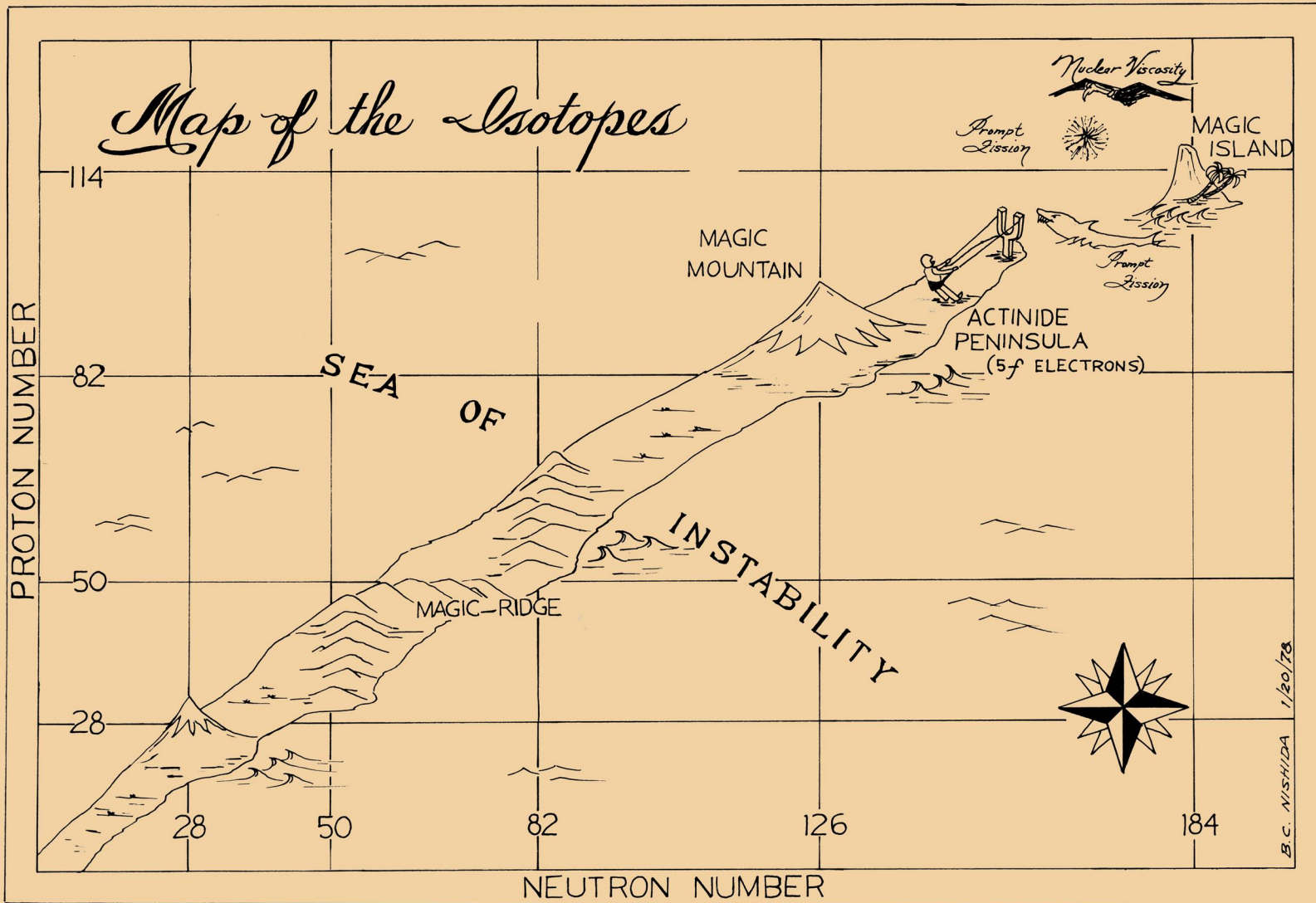
Chemical Behaviors of SHEs

- SHE == breakdown in periodic trends?
 - Relativistic shifts in electronic orbitals

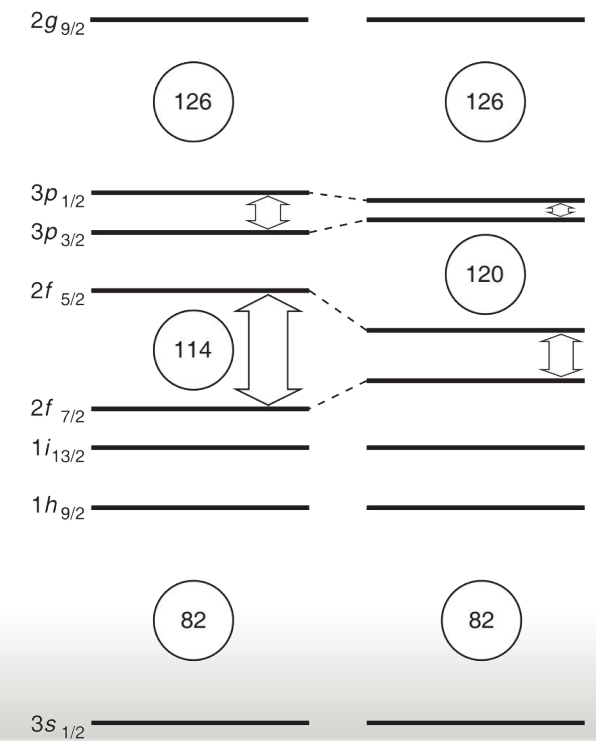
Rh 45 102.91  Rhodium	Pd 46 106.42  Palladium	Ag 47 107.87  Silver	Cd 48 112.41  Cadmium
Ir 77 192.22  Iridium	Pt 78 195.08  Platinum	Au 79 196.97  Gold	Hg 80 200.59  Mercury



Historical Definition of “SHE”

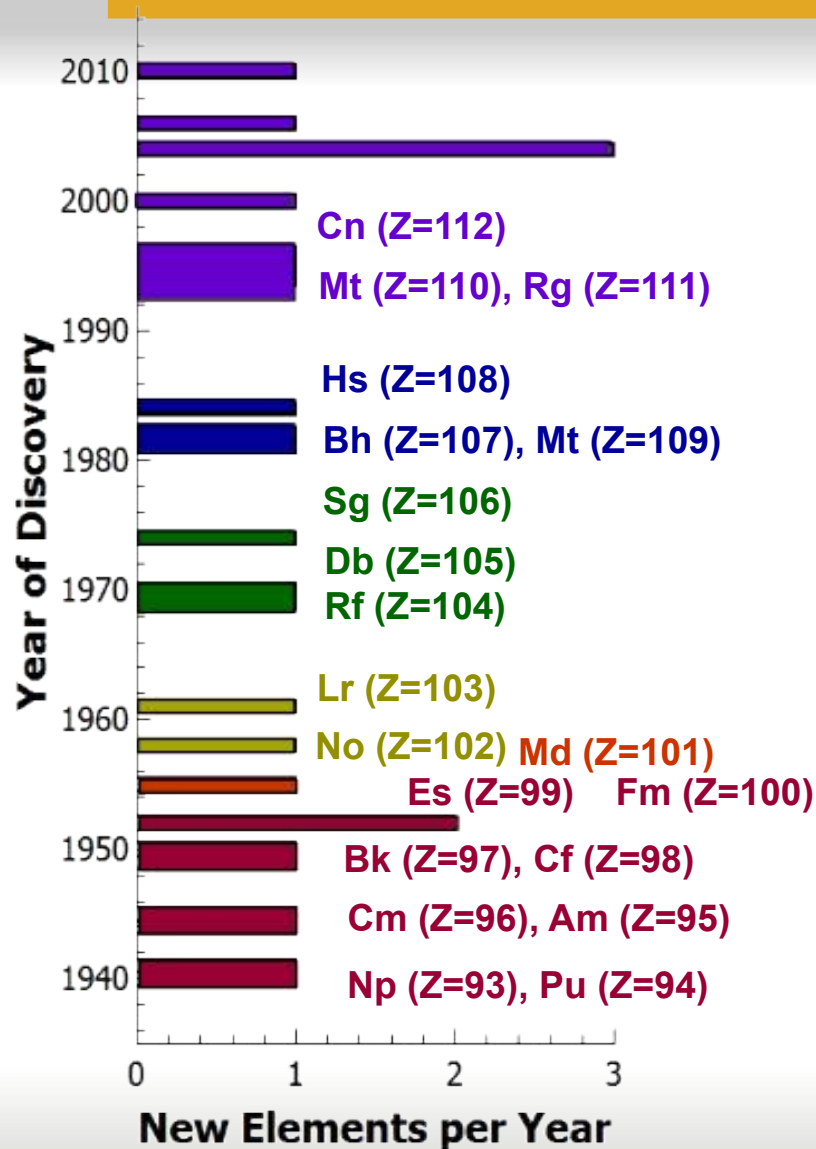


- Search for stability
 - Pre-1967 picture
 - Need to “jump” from Rf → FI , E120





Element Discovery



- IUPAC / IUPAP Definition:

“The discovery of a chemical elements is the experimental demonstration beyond reasonable doubt, of the existence of a nuclide with atomic number Z not identified before, existing for at least 10^{-14} s.”

- Nomenclature:

- Transuranic – element beyond uranium ($Z \geq 93$)
- Transactinide – element beyond lawrencium ($Z \geq 104$)
- Superheavy element – ...



Element Discovery

- Six elements discovered in past two decades
 - $Z = 113$ Nihonium (Nh): 2004 at RIKEN (1 atom in several months)
 - $Z = 114$ Flerovium (Fl): 1998 at JINR (2 atoms in 1 year)
 - $Z = 115$ Moscovium (Mc): 2003 at JINR (4 atoms in 1 month)
 - $Z = 116$ Livermorium (Lv): 2000 at JINR (3 atoms in 1 year)
 - $Z = 117$ Tennessine (Ts): 2010 at JINR (6 atoms in several months)
 - $Z = 118$ Oganesson (Og): 2002 at JINR (3 atoms in several months)

For Release 8 June 2016



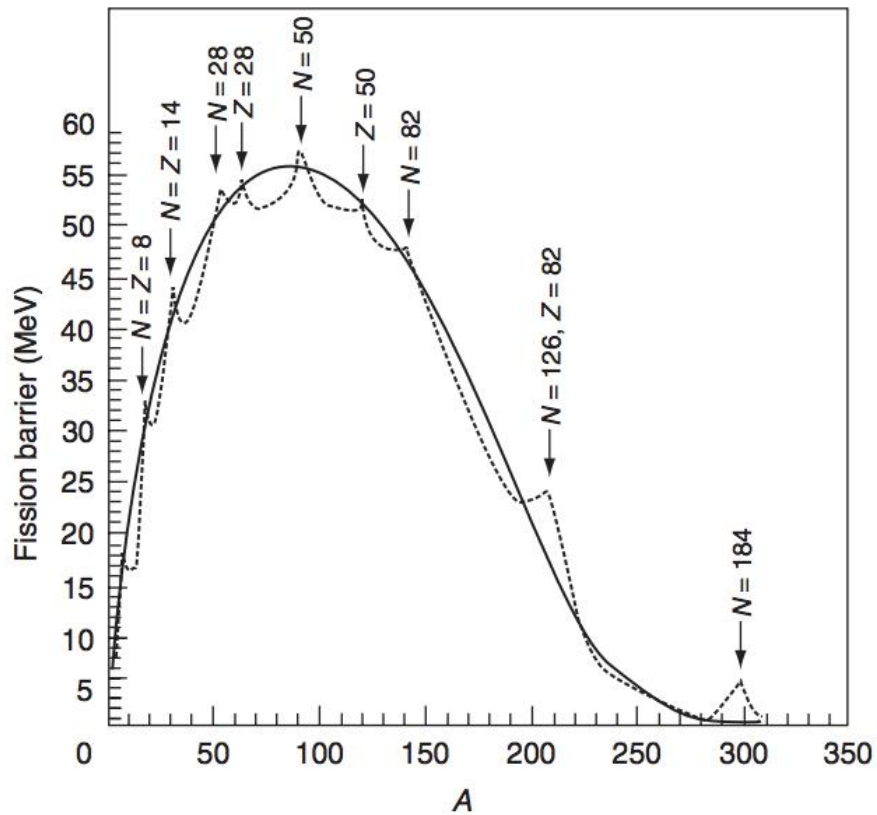
INTERNATIONAL UNION OF
PURE AND APPLIED CHEMISTRY

**IUPAC is naming the four new elements nihonium, moscovium,
tennessine, and oganesson**

Following earlier reports that the claims for discovery of these elements have been fulfilled [1, 2], the discoverers have been invited to propose names and the following are now disclosed for public review:



Stability of SHEs

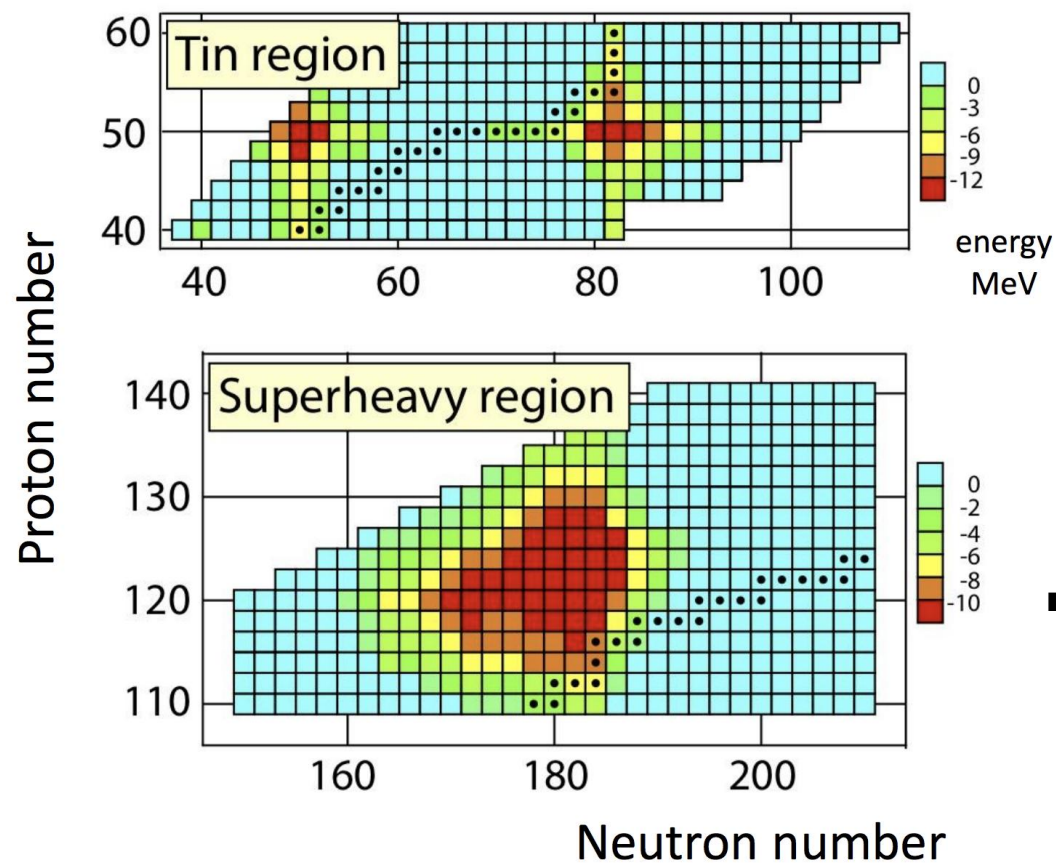


Fission-Barrier heights as a function of A . Smooth line calculated with semi-empirical mass formula and the dotted line illustrates the effects of nuclear structure.

- SHEs exist at edge of fission barriers
 - B_f drops rapidly for heavy nuclei due to increasing repulsive forces.
 - Favorable shell structure can increase a nucleus's resistance to fission.
- Modern SM calculations have diffuse shell stabilization
 - Z closure could be 114 – 120



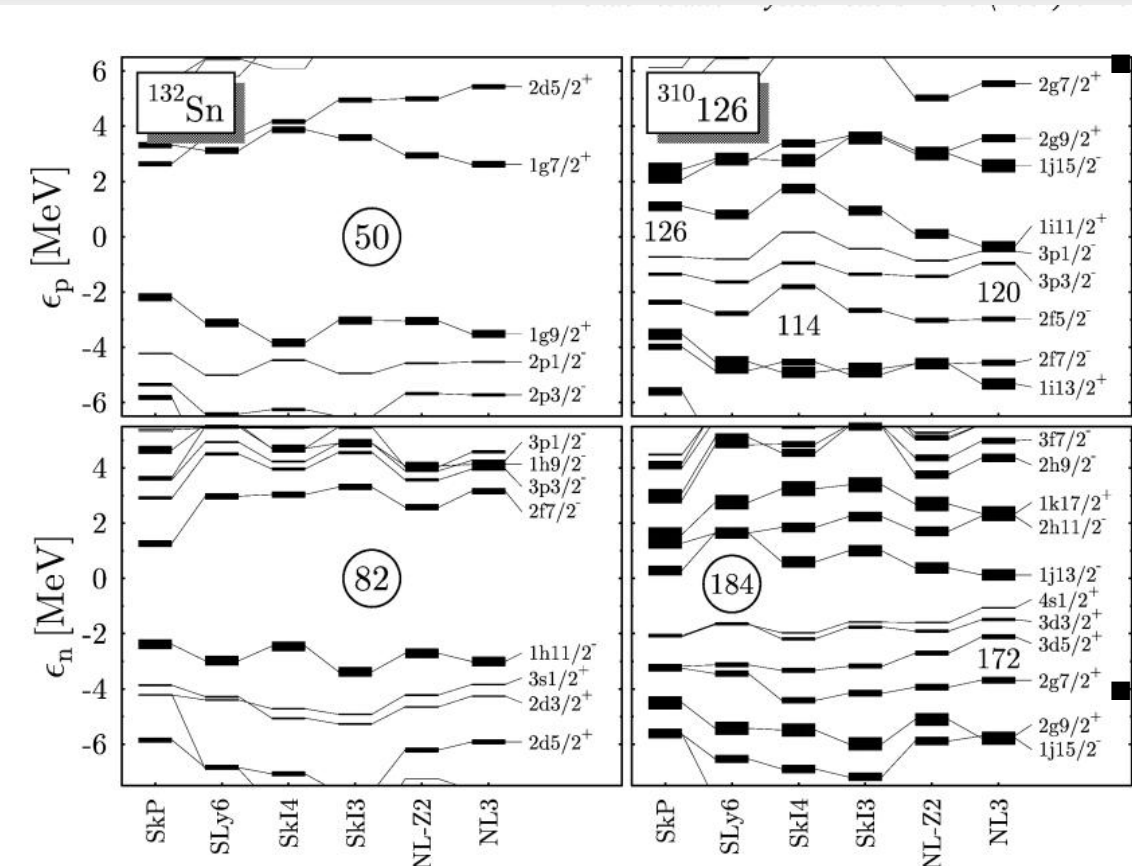
Stability of SHEs



- SHEs exist at edge of fission barriers
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Stability of SHEs



Calculated single-particle spectra for ^{132}Sn and $^{310}126$.
Energies are relative to the Fermi surface.

SHEs exist at edge of fission barriers

- B_f drops rapidly for heavy nuclei due to increasing repulsive forces.
- Favorable shell structure can increase a nucleus's resistance to fission.

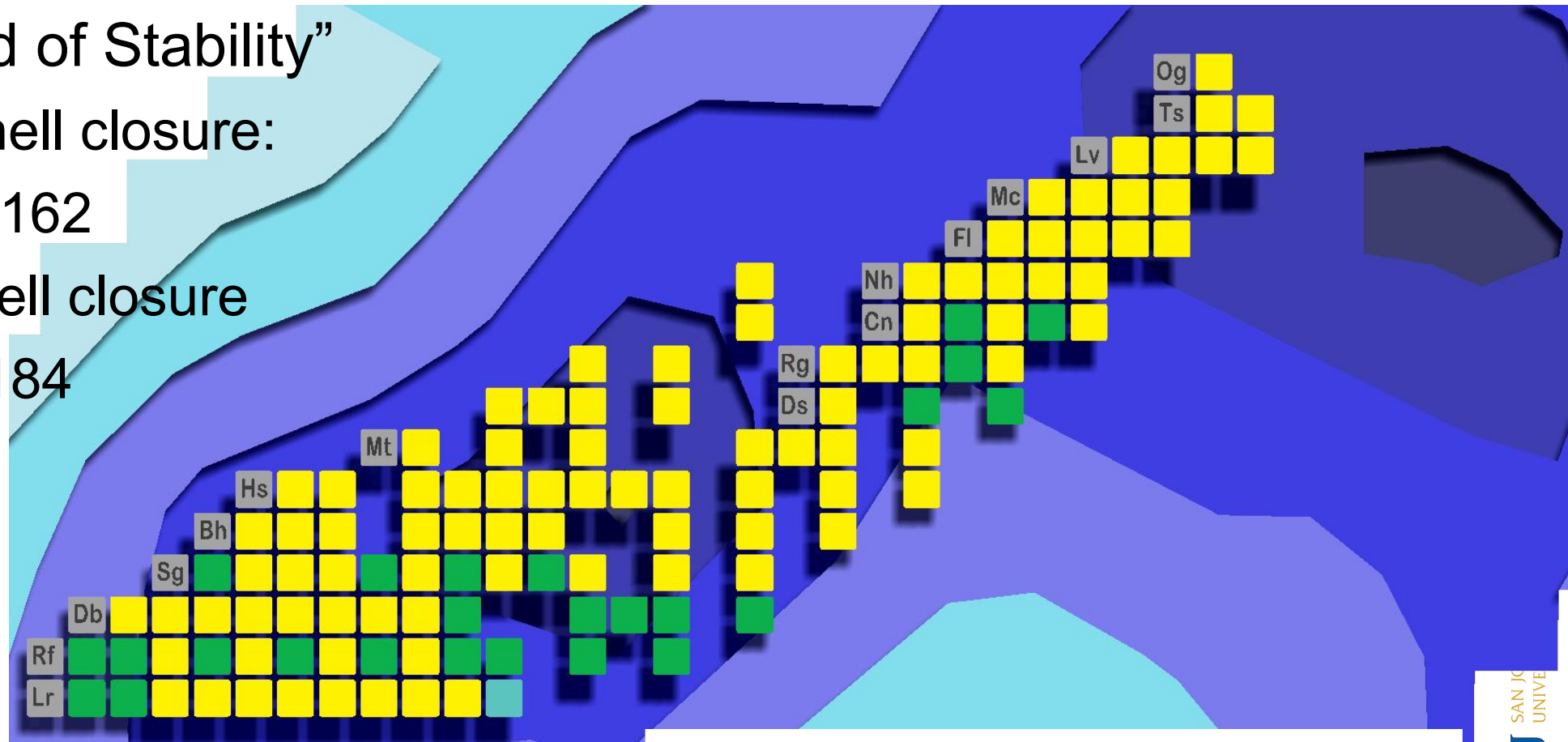
Modern SM calculations have diffuse shell stabilization

- Z closure could be 114 – 120



Modern SHEs

- Edge of “Island of Stability”
 - Deformed shell closure:
 $Z \sim 108$, $N \sim 162$
 - Spherical shell closure
 $Z \sim 114$, $N \sim 184$



Shell effects from Sobczewski et al: Phys. Rev. C 63 (2001) 034306



Element Synthesis

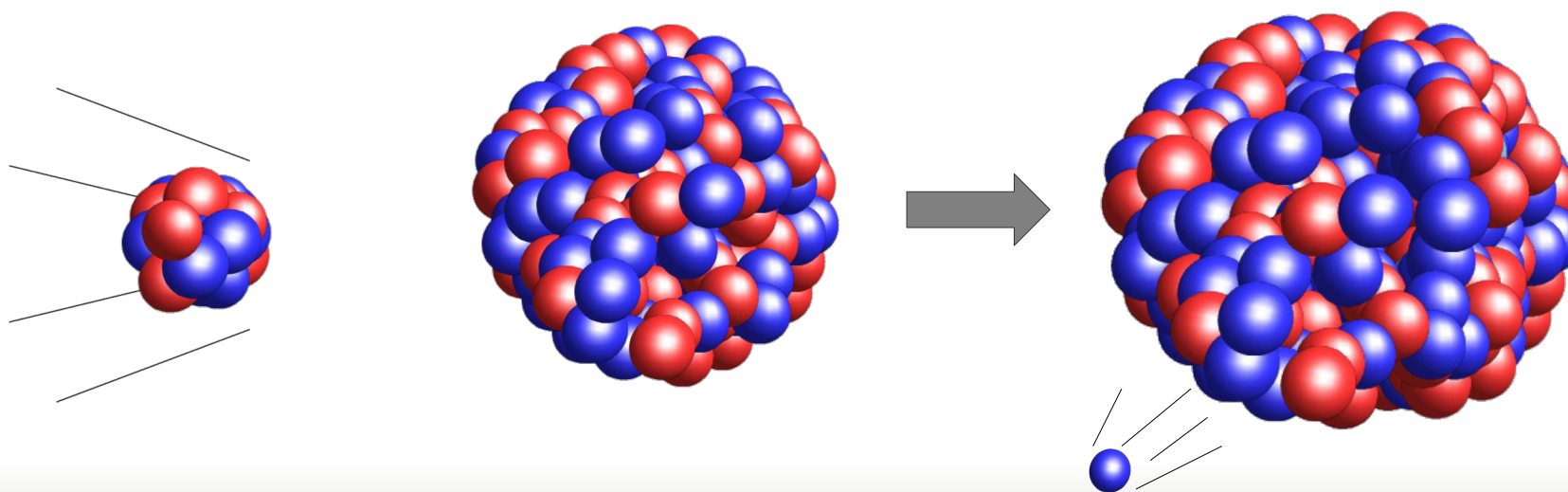
- Techniques for element synthesis

- (1) Reactor-based

- n capture + β^- decay
 - Similar to s-process

- (2) Accelerator-based reactions

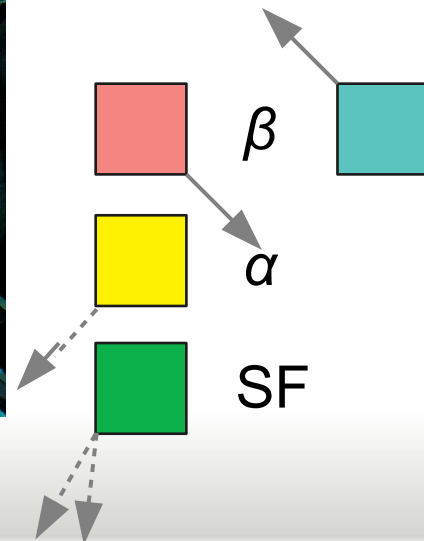
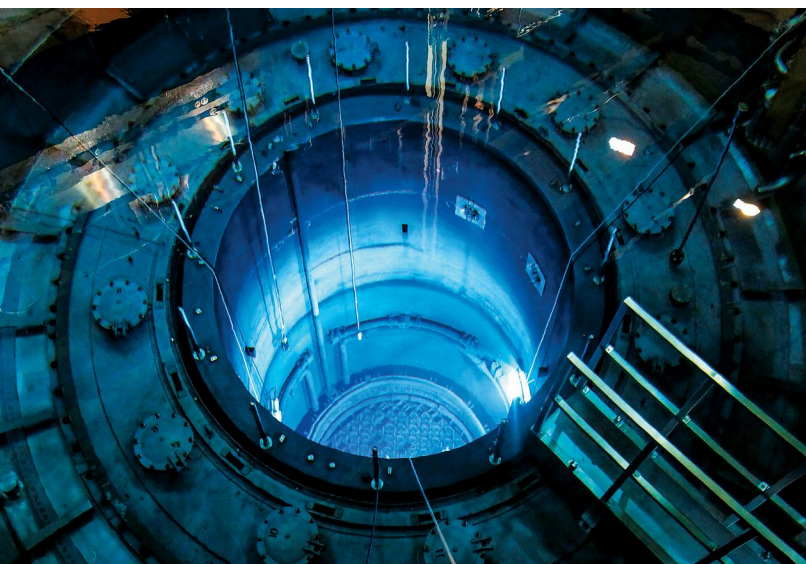
- Compound Nuclear
 - Multi-nucleon transfer





Heavy Element Production

- Reactor Breeding
 - Oldest technique, since 1940s
 - n capture must compete with decay kinetics



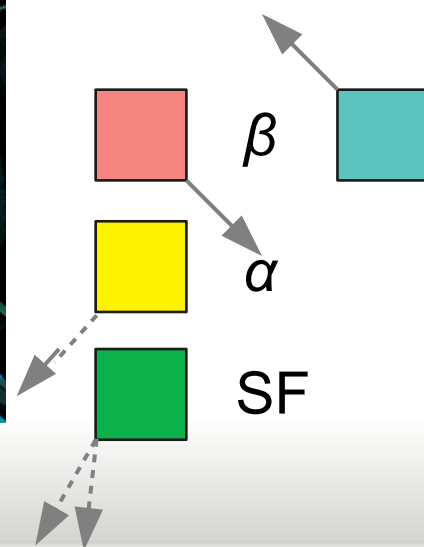
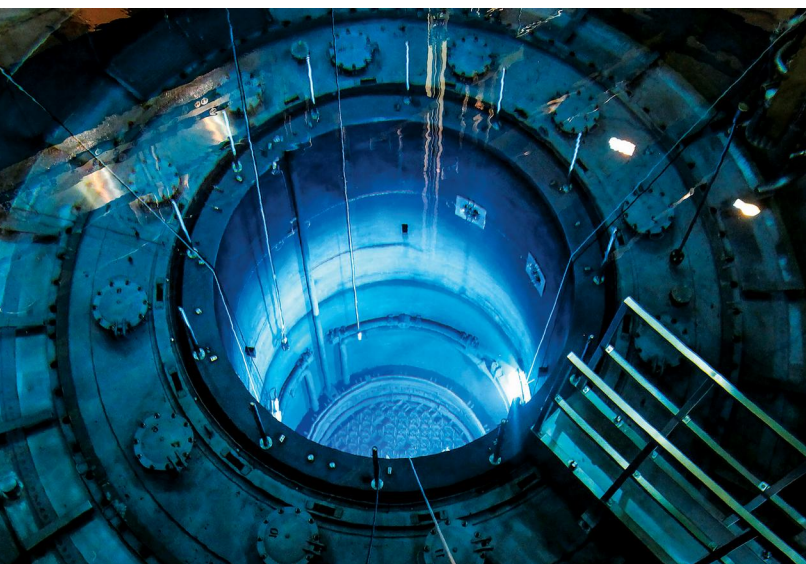
²³⁹ Cm	²⁴⁰ Cm	²⁴¹ Cm	²⁴² Cm	²⁴³ Cm	²⁴⁴ Cm	²⁴⁵ Cm	²⁴⁶ Cm
					n, γ	n, γ	
²³⁸ Am	²³⁹ Am	²⁴⁰ Am	²⁴¹ Am	²⁴² Am	²⁴³ Am	²⁴⁴ Am	²⁴⁵ Am
					n, γ	n, γ	
²³⁷ Pu	²³⁸ Pu	²³⁹ Pu	²⁴⁰ Pu	²⁴¹ Pu	²⁴² Pu	²⁴³ Pu	²⁴⁴ Pu
			n, γ	n, γ	n, γ		
²³⁶ Np	²³⁷ Np	²³⁸ Np	²³⁹ Np	²⁴⁰ Np	²⁴¹ Np	²⁴² Np	²⁴³ Np
			n, γ				
²³⁵ U	²³⁶ U	²³⁷ U	²³⁸ U	²³⁹ U	²⁴⁰ U		²⁴² U
			n, γ				



Heavy Element Production

■ Fermium Gap

- Fm decay rate too fast for reactors
- Still possible in much high neutron fluxes, i.e in stellar explosion



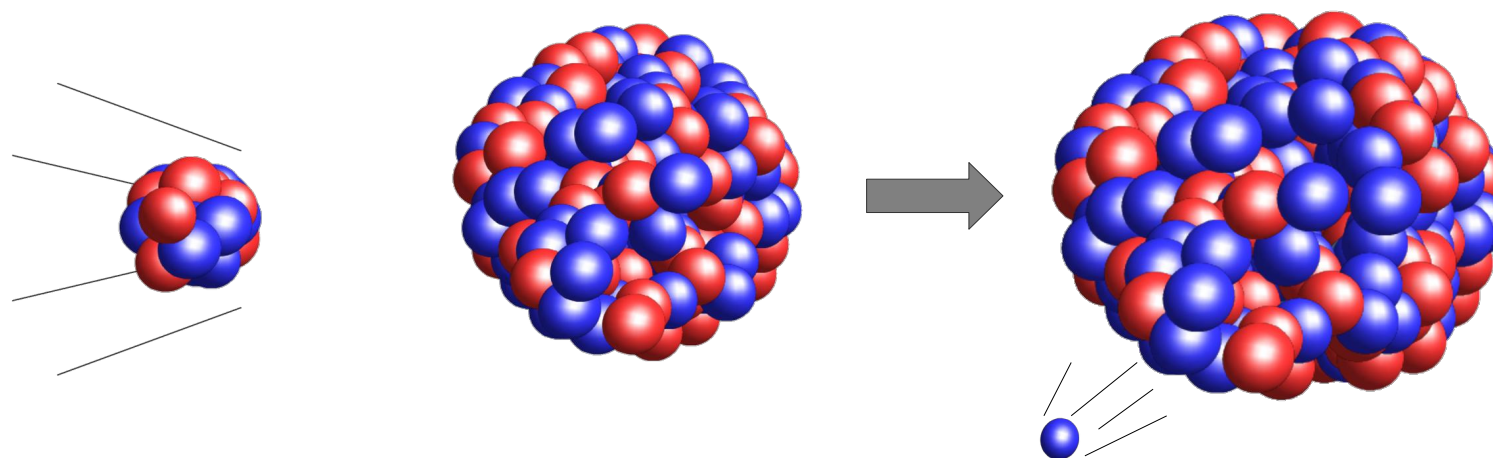
255No	256No	257No	258No	259No	260No	262No
254Md 28 m	255Md 27 m	256Md 77 m	257Md 5.5 h	258Md 51 d	259Md 95 m	260Md 32 d
253Fm 25.4 h	254Fm 3.0 d	255Fm 3.24 h	256Fm 20.1 h	257Fm 100 d	258Fm 0.4 ms	259Fm 1.5 s
252Es	253Es	254Es	255Es	256Es	257Es	
251Cf	252Cf	253Cf	254Cf	255Cf	256Cf	

Decay paths are indicated by arrows: red for α , blue for β , and green for SF. Labels n, γ indicate neutron capture and gamma emission.



Compound Nuclear Reactions

- Fusion Evaporation
 - Light(er) ion beam on heavy(ier) target
 - Accelerated above Coloumb barrier
 - Post-fusion excitation dissipated via particle emission





Compound Nuclear Reactions

- Reaction Theory:
 - Three step process, depends on E , Z , A

Production cross section $\rightarrow \sigma_{EVR}(E_{c.m.}) = \sum_J \sigma_{cap}(E_{c.m.}, J) P_{CN}(E_{c.m.}, J) W_{sur}(E_{c.m.}, J)$

Capture cross section Probability of CN formation Survival

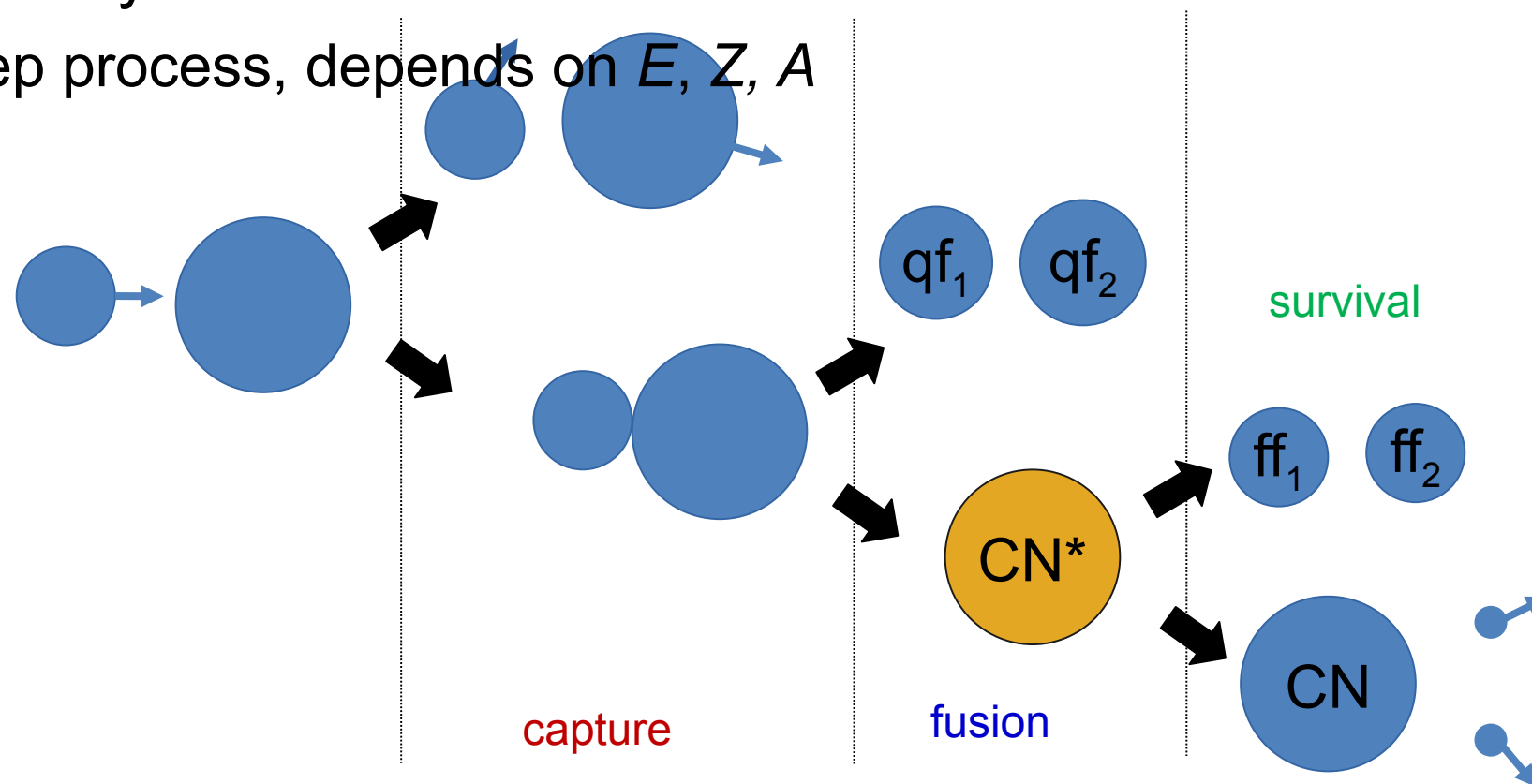
$$\text{Fusion Evaporation} = (\text{sticking}) \times (\text{fusion}) \times (\text{survival})$$

- Survival: dissipating excitation energy without breaking apart



Compound Nuclear Reactions

- Reaction Theory:
 - Three step process, depends on E , Z , A

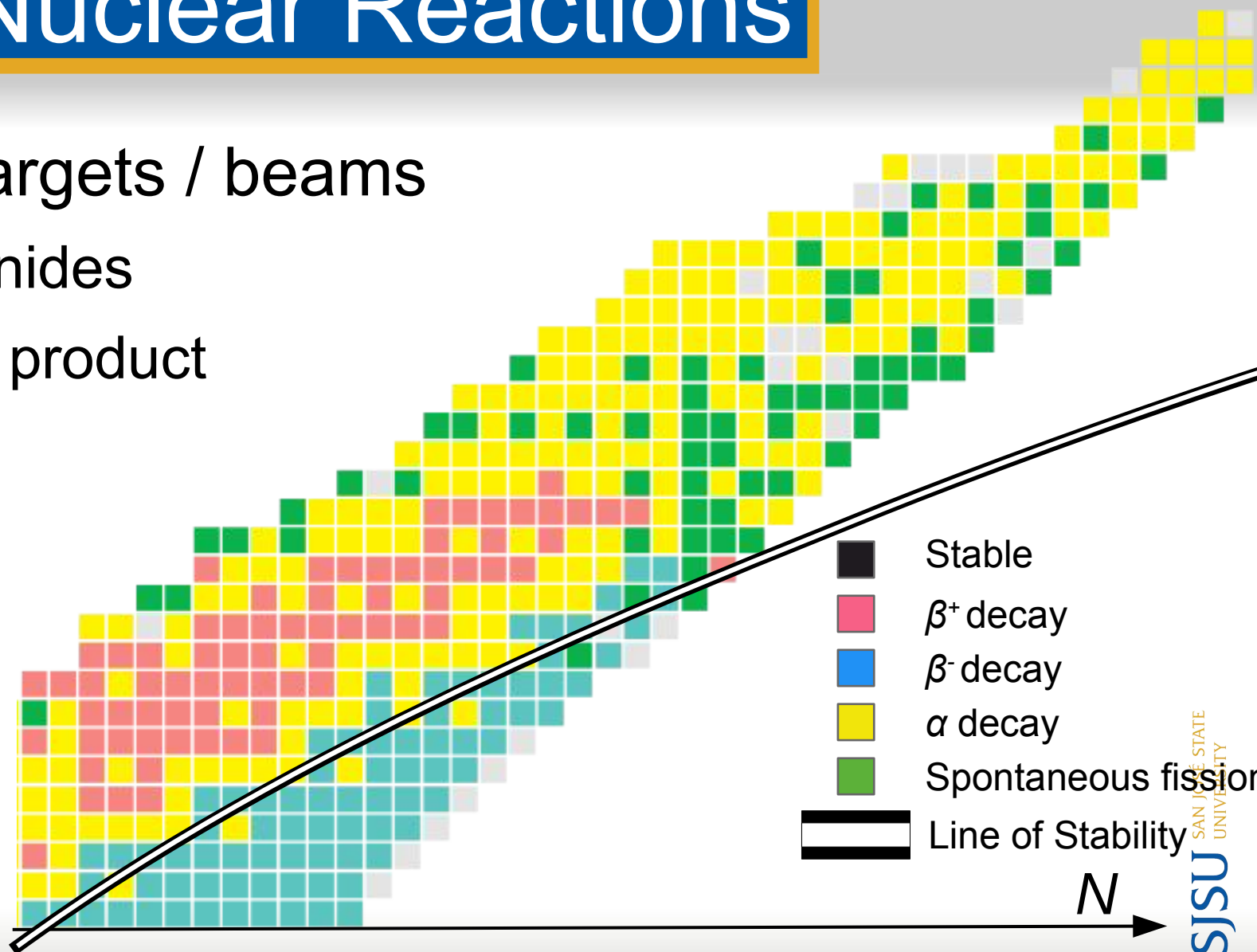


- Survival: dissipating excitation energy without breaking apart



Compound Nuclear Reactions

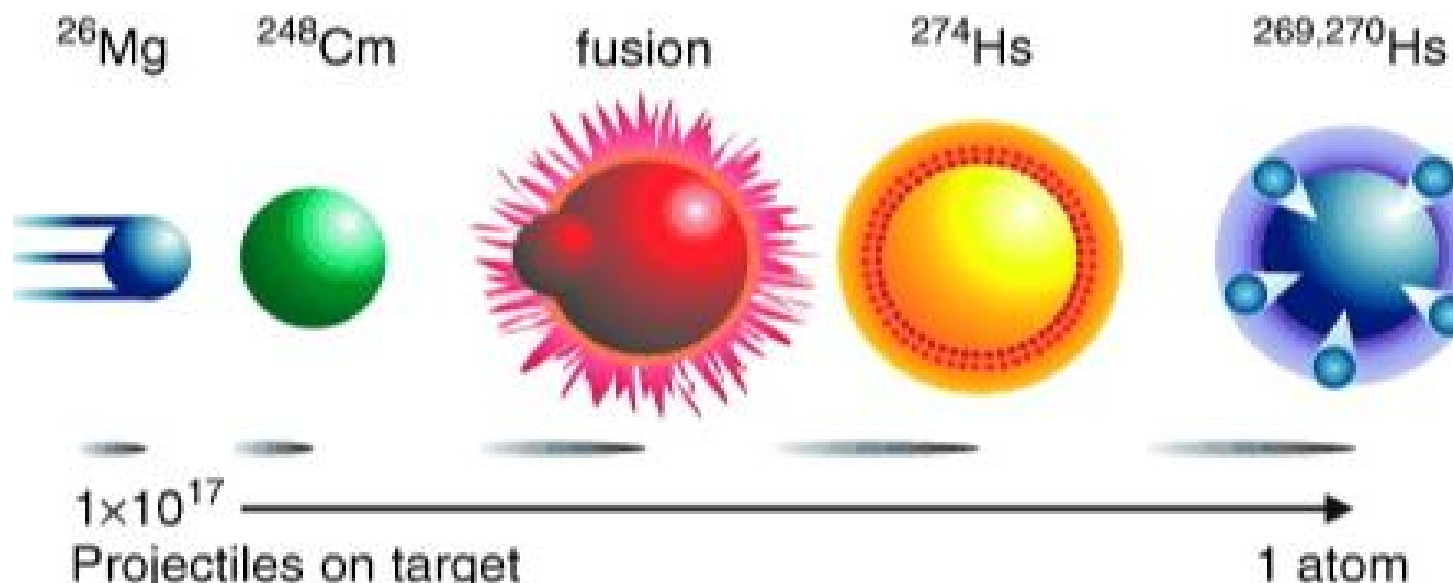
- **Need** neutron-rich targets / beams
 - Targets: Pb / Bi, actinides
 - Beams: based upon product





Compound Nuclear Reactions

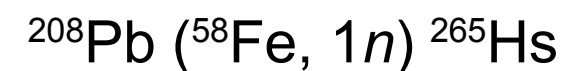
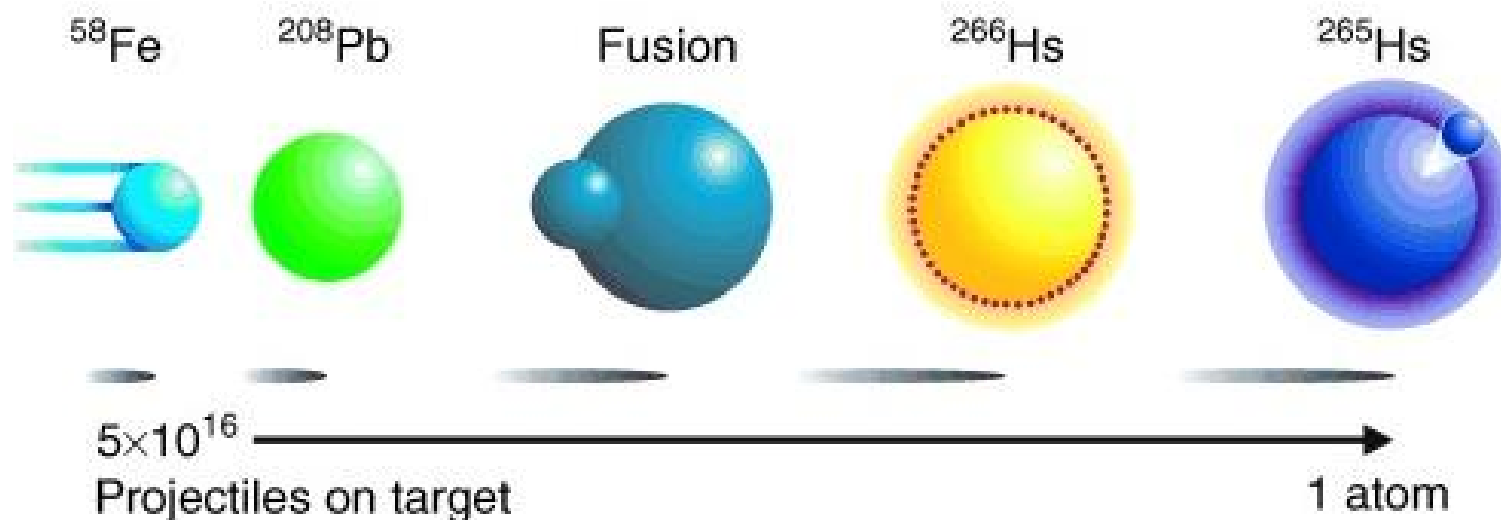
- “Hot” Fusion vs. “Cold” Fusion
 - Hot Fusion: actinides + light ion beam, higher CN excitation ($E^* \sim 40$ MeV)
 - Take Advantage of greater n in actinides





Compound Nuclear Reactions

- “Hot” Fusion vs. “Cold” Fusion
 - Hot Fusion: actinides + light ion beam, higher CN excitation ($E^* \sim 40$ MeV)
 - Cold Fusion: Pb / Bi target + beam, lower CN excitation ($E^* \sim 10$ MeV)
 - Take advantage of shell closure in target



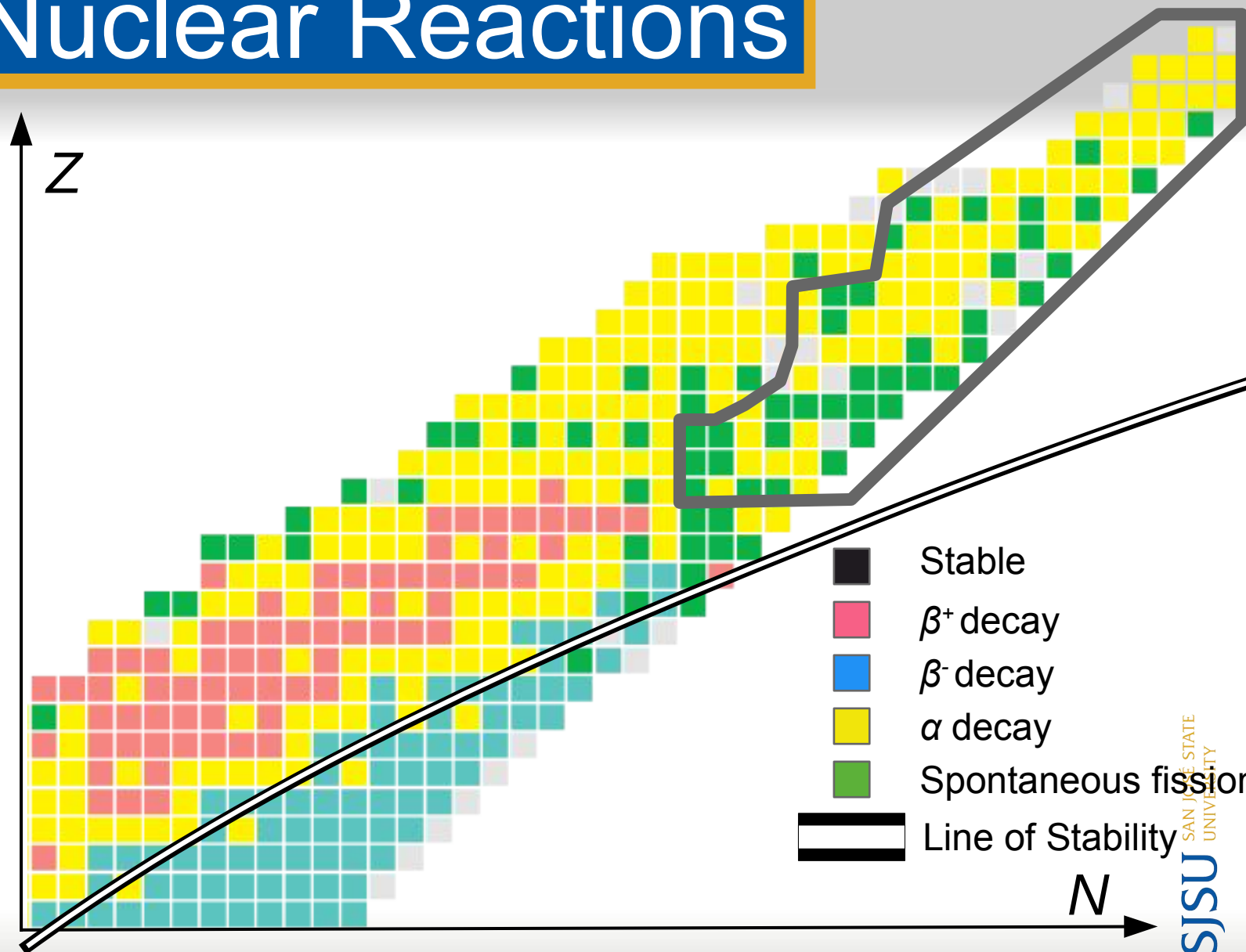
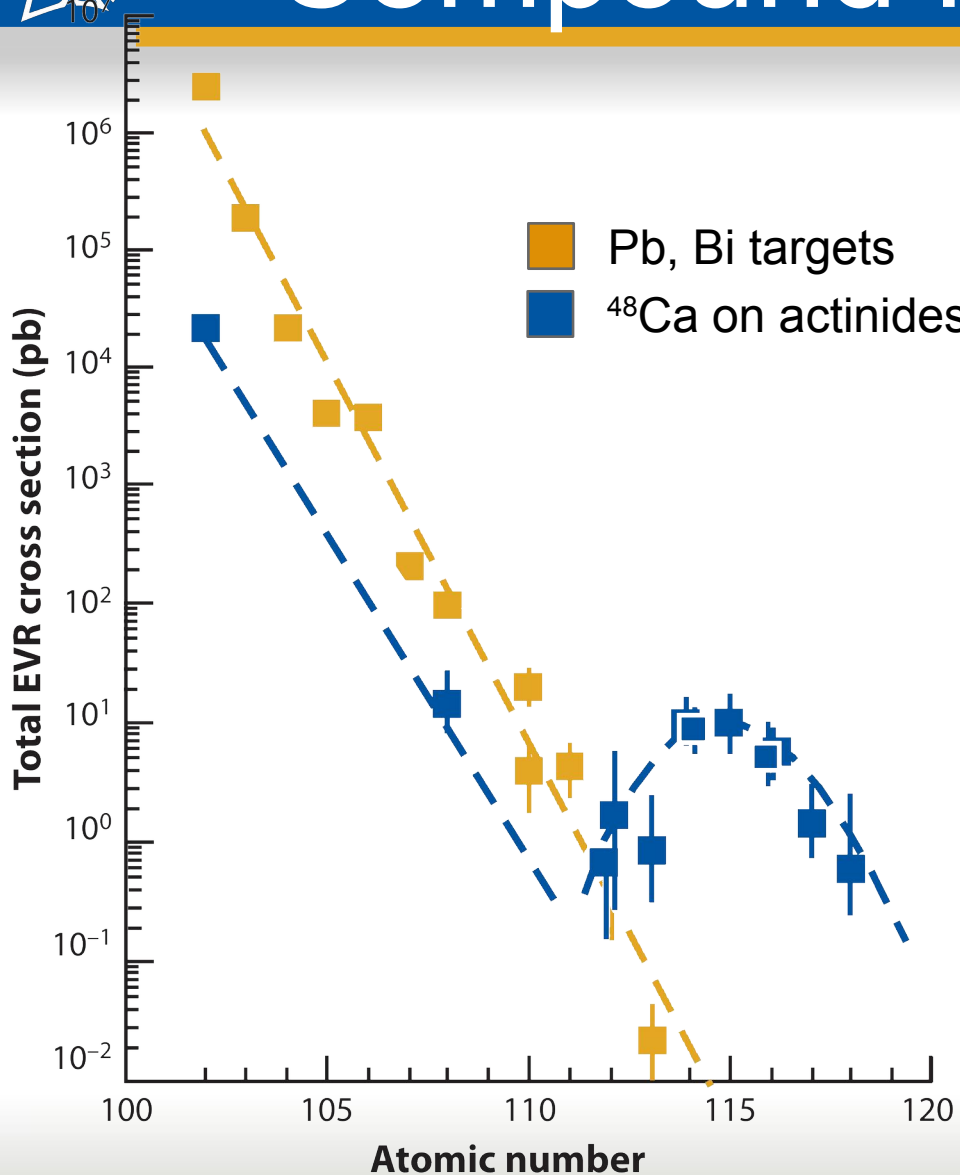


Compound Nuclear Reactions

- “Hot” Fusion vs. “Cold” Fusion
 - Hot Fusion: actinides + light ion beam, higher CN excitation ($E^* \sim 40$ MeV)
 - Cold Fusion: Pb / Bi target + beam, lower CN excitation ($E^* \sim 10$ MeV)
- “Lukewarm” Fusion
 - ^{48}Ca + actinides, CN excitation ($E^* \sim 20 - 40$ MeV)
 - Take advantage of highest neutron economy from beam + target,
 - Take advantage of doubly magic ^{48}Ca

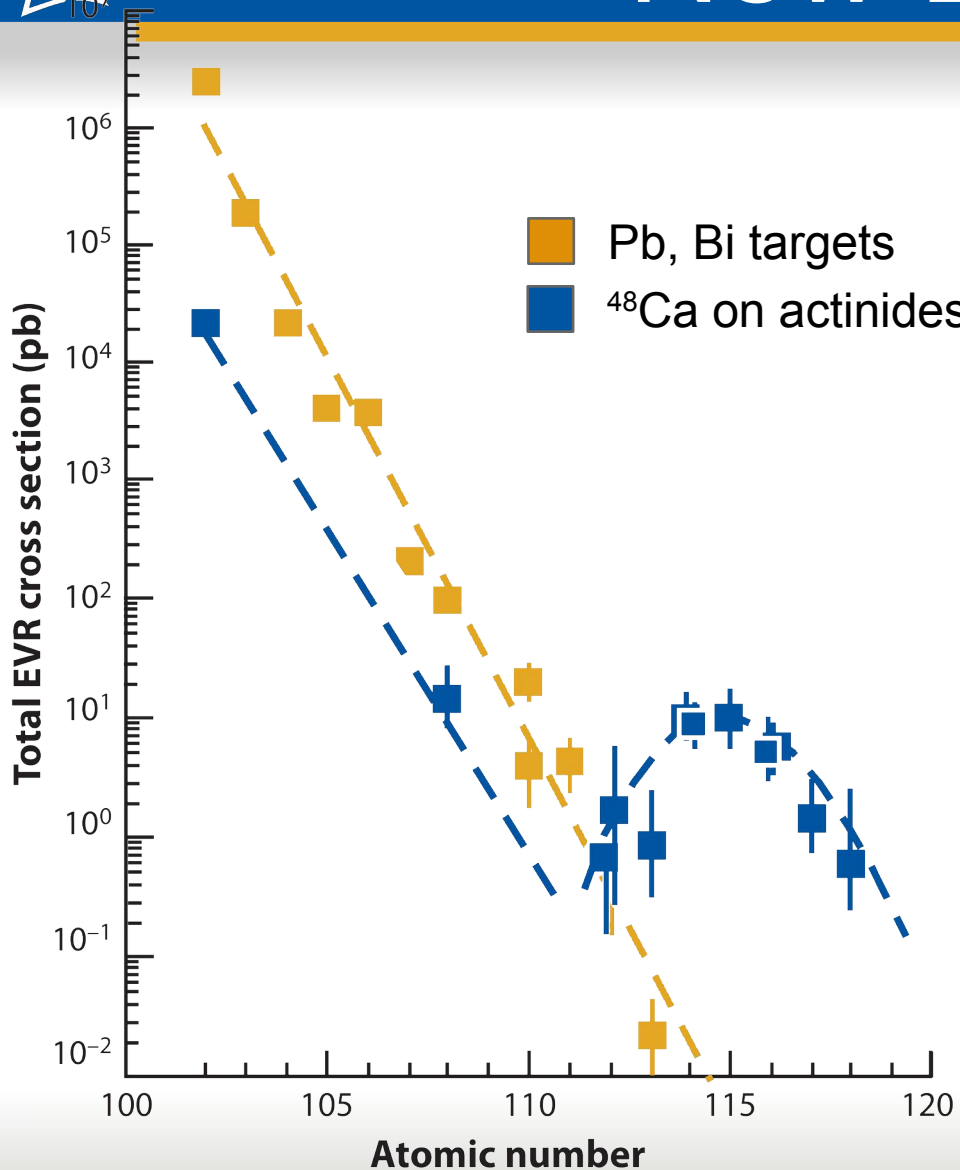


Compound Nuclear Reactions





New Element Synthesis



- Nuclei at the very bounds of stability
 - Low production cross sections
 - Short lived isotopes

102	103	104	105	109	115
^{254}No	^{255}Lr	^{257}Rf	^{258}Db	^{276}Mt	^{288}Mc
55 s	22 s	4.7 s	4 s	0.54 s	171 ms
1/s	10/min	1/min	1/hr	1/wk	1/day

- Og(Z=118) is heaviest synthesized element
 - ^{249}Cf (Z=98) + ^{48}Ca
 - Cannot get enough Es (Z = 99)



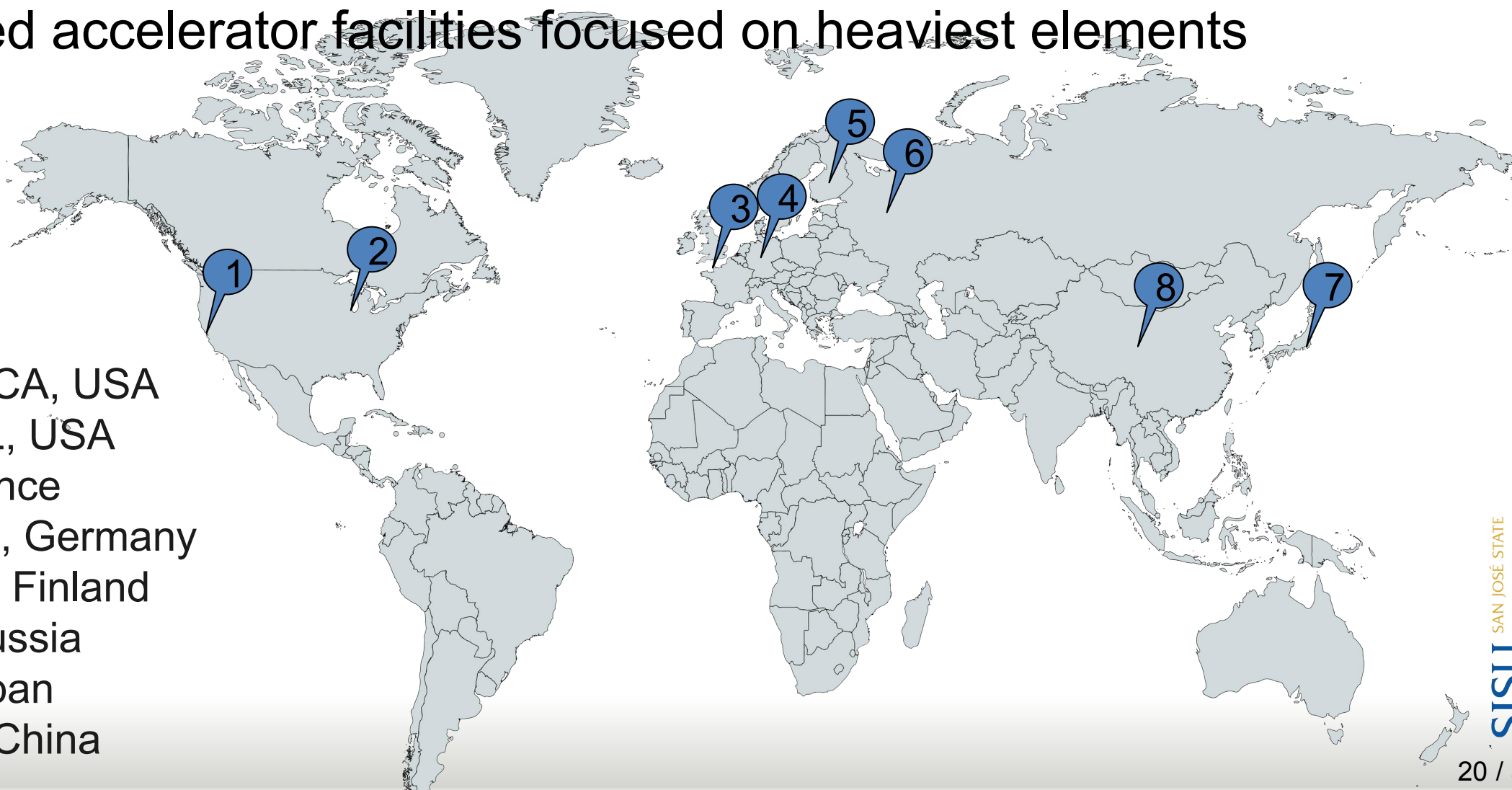
Experimental Approach

- Low cross sections means high flux beams
 - Limited accelerator facilities focused on heaviest elements
- Dedicated recoil separators
 - Extremely low statistics, incredible background suppression needed
- Need right target, right beam & beam flux, and luck



Experimental Approach

- Limited accelerator facilities focused on heaviest elements

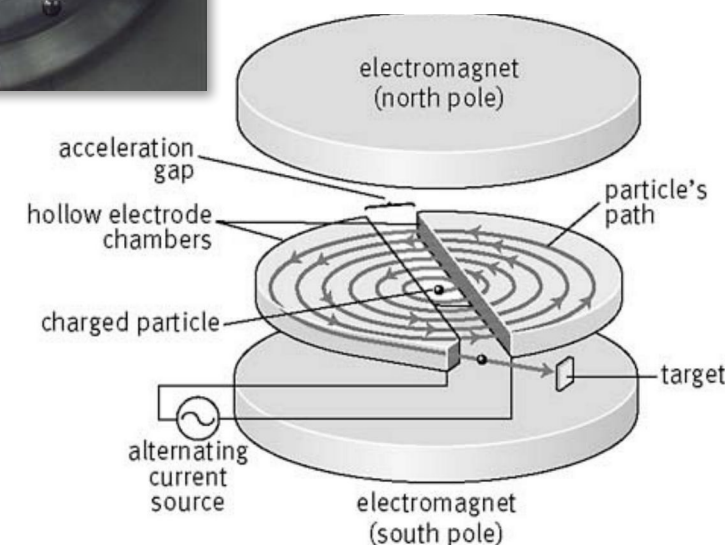
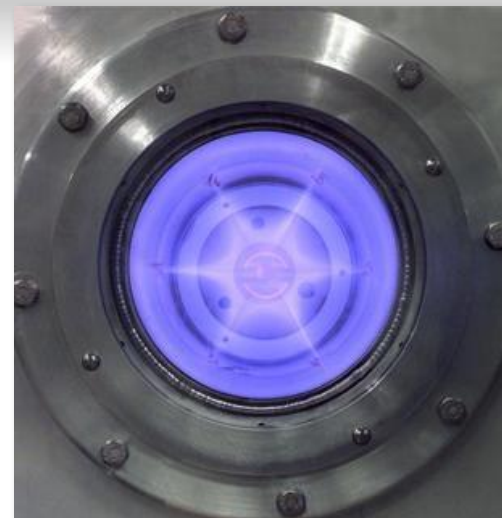


1. Berkeley, CA, USA
2. Lemont, IL, USA
3. Caen, France
4. Darmstadt, Germany
5. Jyväskylä, Finland
6. Dubna, Russia
7. Tokyo, Japan
8. Lanzhou, China



Ion sources & Accelerators

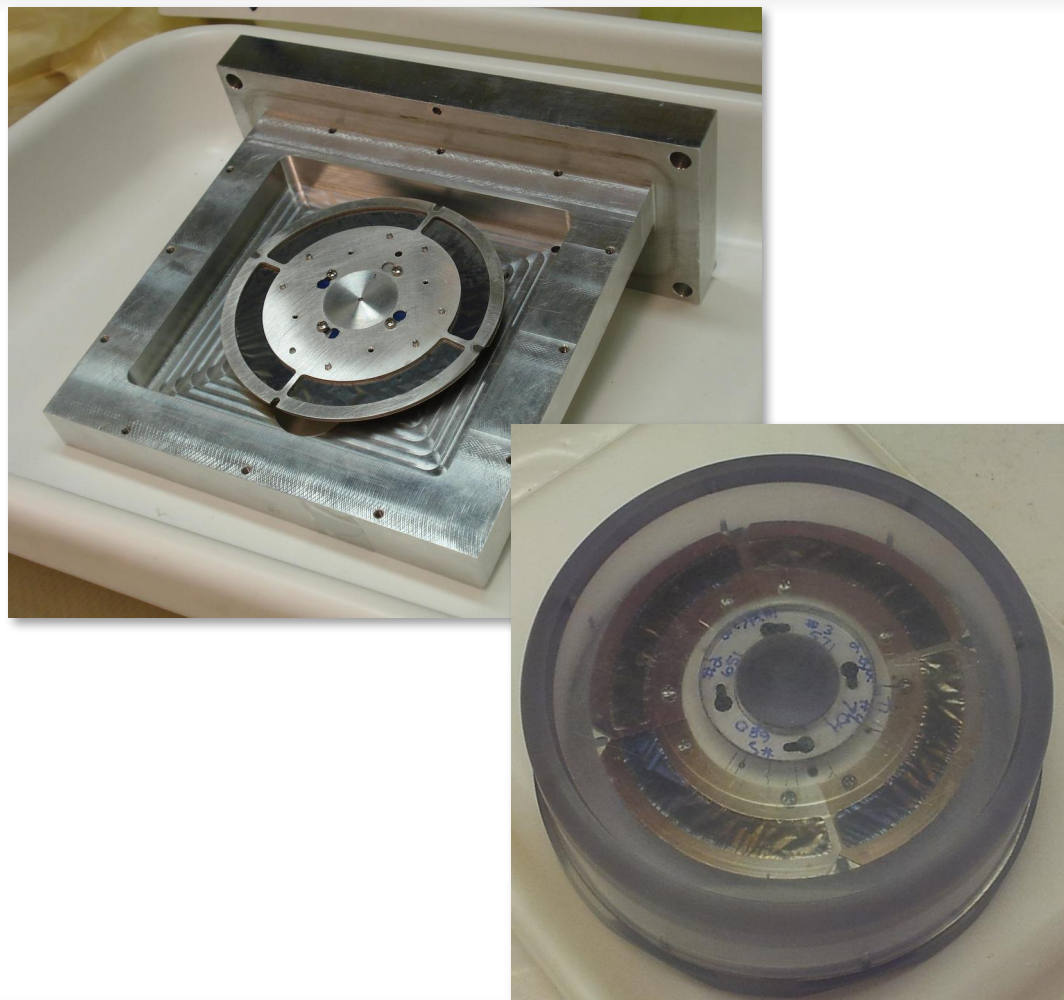
- Ion sources:
 - Deliver highly charged plasmas to accelerators
- Accelerators:
 - Transport ion beam to experimental facility
 - SHE production: linac, cyclotron



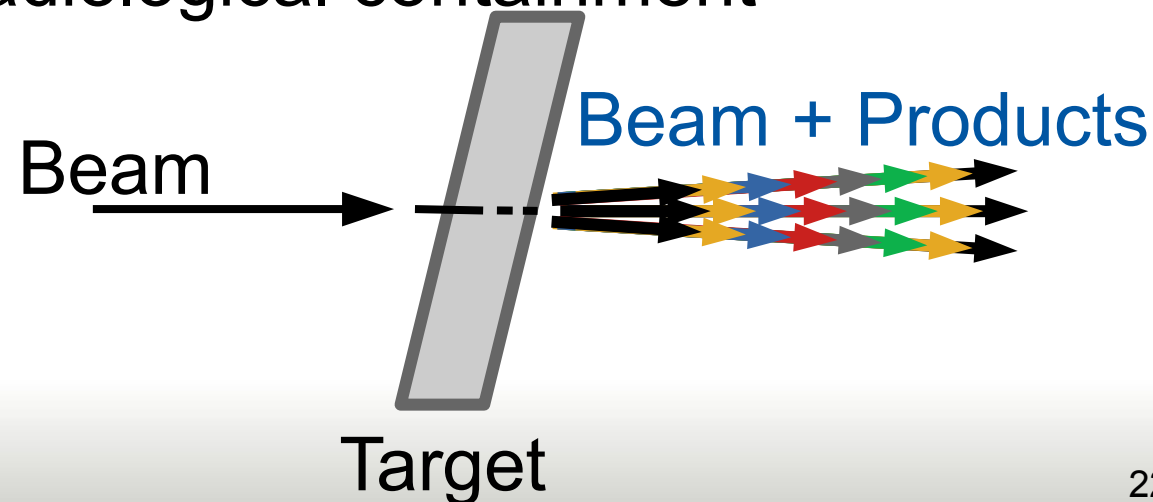
Precision Graphics
<http://images.vourdictionary.com/cyclotron>



Targetry and Reaction Products



- Thickness $\sim 500 \mu\text{g}/\text{cm}^2$
 - Optimize energy loss
 - ~ 2000 nuclei for beam interactions
- High beam flux = rotating targets
 - Heat dissipation
- Actinide targets = additional radiological containment

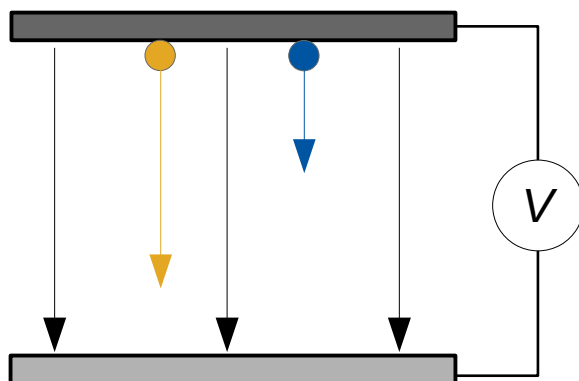




Manipulating ions

▪ Electrostatics

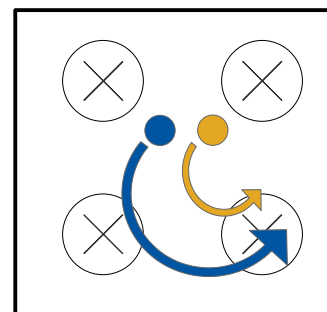
$$\vec{F}_{\vec{E}} = q \cdot \vec{\mathcal{E}}$$



▪ Magnetics

$$\vec{F}_{\vec{B}} = q \vec{v} \times \vec{\mathcal{B}}$$

$$m_{\bullet} = 2 m_{\circ}$$



$$\begin{array}{l} \bullet v_i = \text{<very small>} \\ q = +1 \end{array}$$

$$\rho_{\bullet} = \rho_{\circ}$$

$$\begin{array}{l} \bullet v_i = \text{<very small>} \\ q = +1 \end{array}$$

$$\rho_{\bullet} = 2 \rho_{\circ}$$



Recoil Separators

- In-flight separation of reaction products

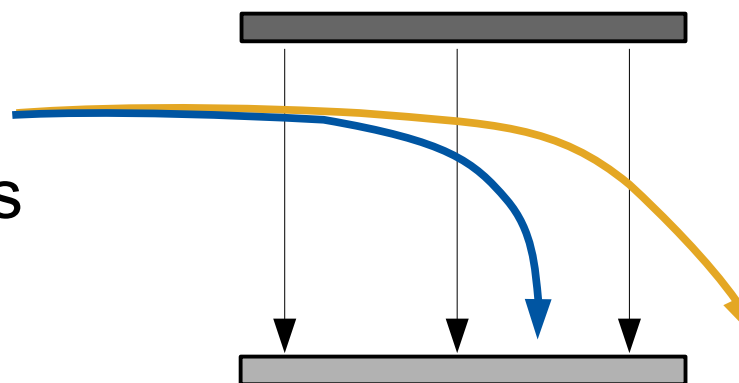
- Purpose:

- Product selection
- Separates from beam, unwanted reaction products

- Electrostatic

$$\vec{F}_{\vec{E}} = q \vec{\mathcal{E}}$$

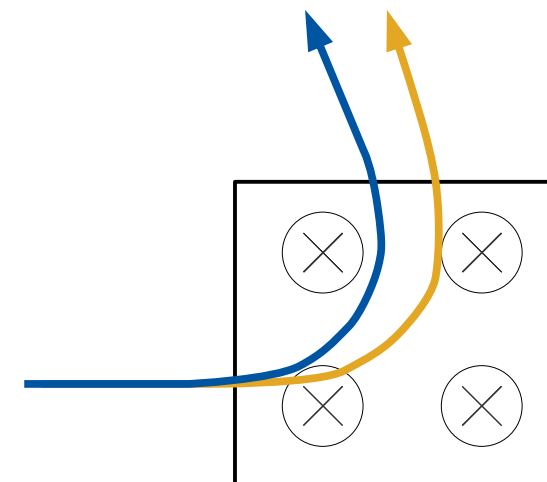
$$\mathcal{E} \rho = \frac{m v^2}{q} = \frac{2E}{q}$$



- Magnetic

$$\vec{F}_{\vec{B}} = q \vec{v} \times \vec{\mathcal{B}}$$

$$\mathcal{B} \rho = \frac{m v}{q} = \frac{p}{q}$$



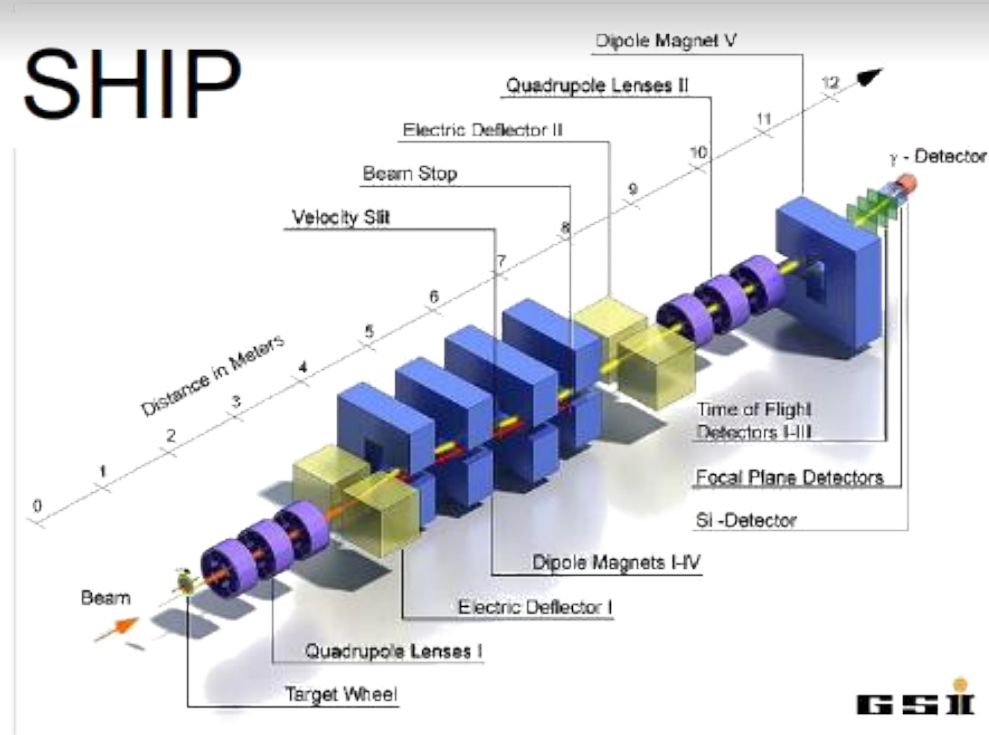
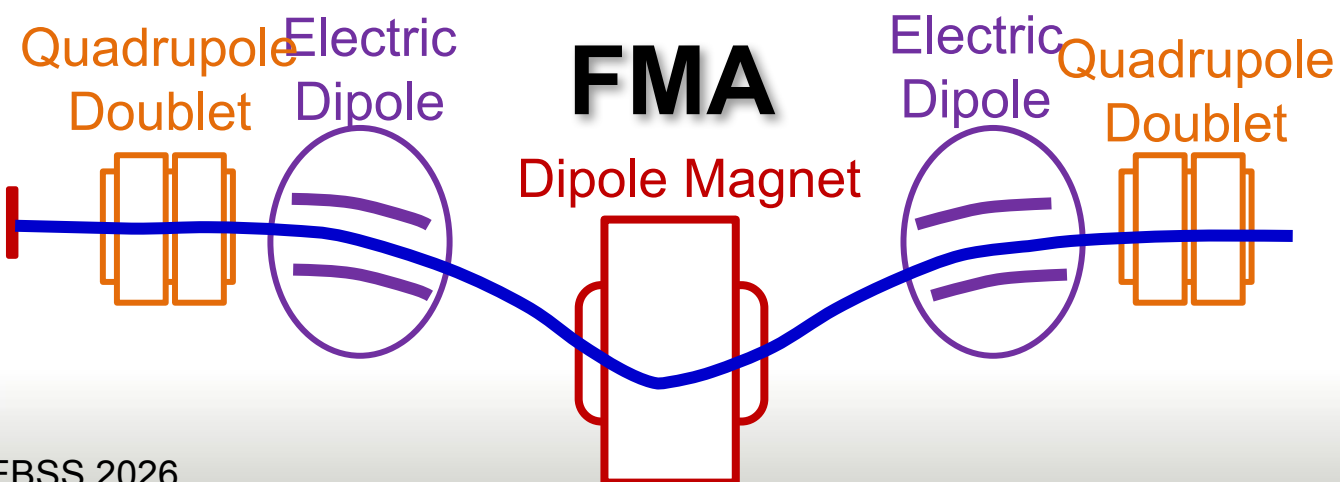


Vacuum Mode Separator

Standard Configuration:

- Quadrupole doublet (or triplet)
- Electric Dipole
- Magnetic Dipole
- Electric Dipole
- Quadrupole doublet (or triplet)

Examples: FMA, SHIP, SHELLS



Velocity Filters:

- Velocity of Beam and SHE are very different

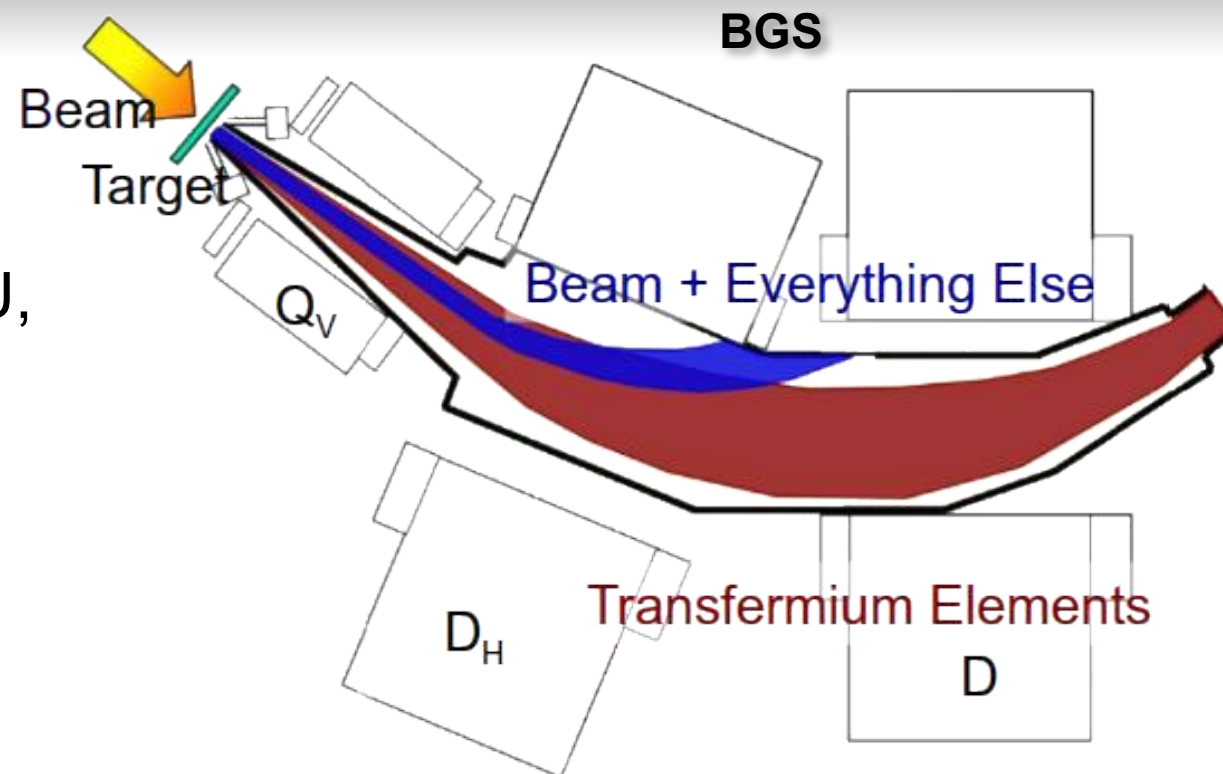
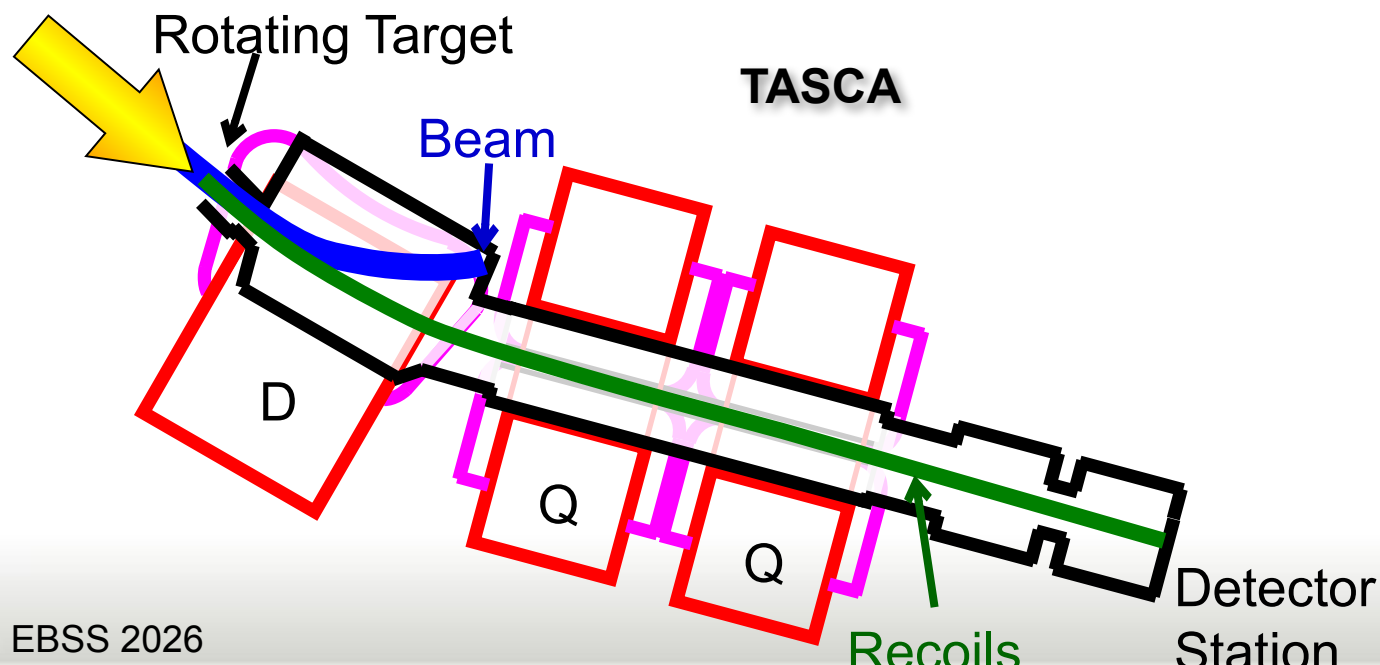


Gas-filled Separators

Standard Configuration:

- Dipole
- Quadrupole doublet

Examples: DGFRS, GARIS(I-III), RITU, SASSY, TASCA



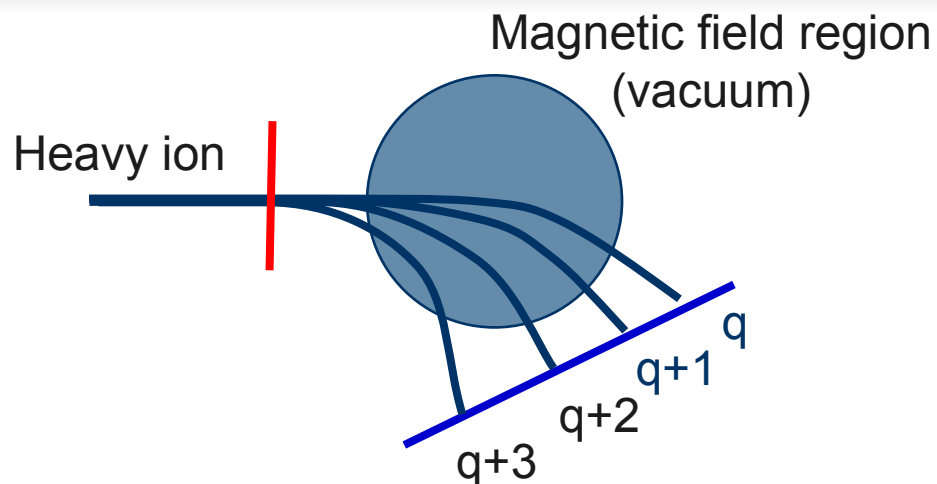
Other Configuration:

- Quadrupole
- Focusing dipole
- Dipole

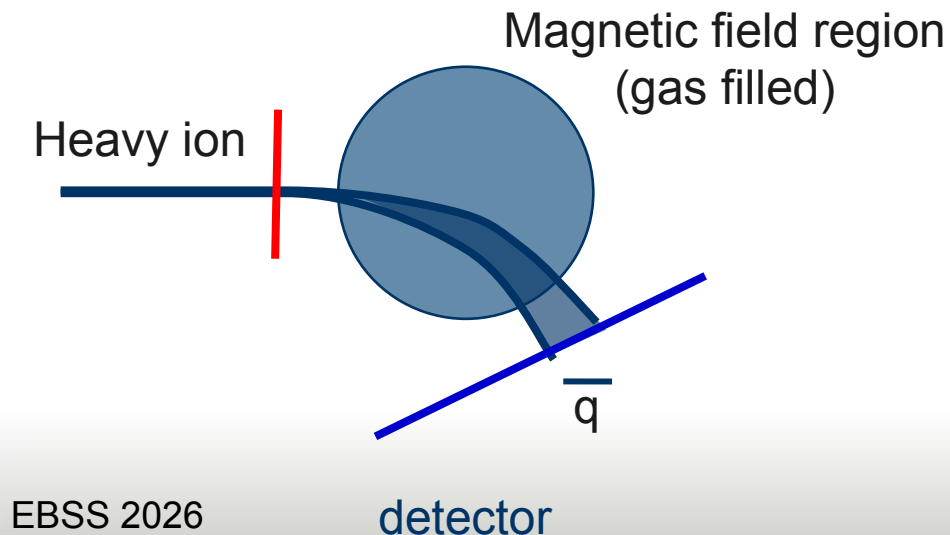
Examples: BGS



Vacuum vs Gas-filled Separators



- Vacuum mode separators:
 - Higher mass resolution
- Gas-filled Separators:
 - Higher transmission efficiency





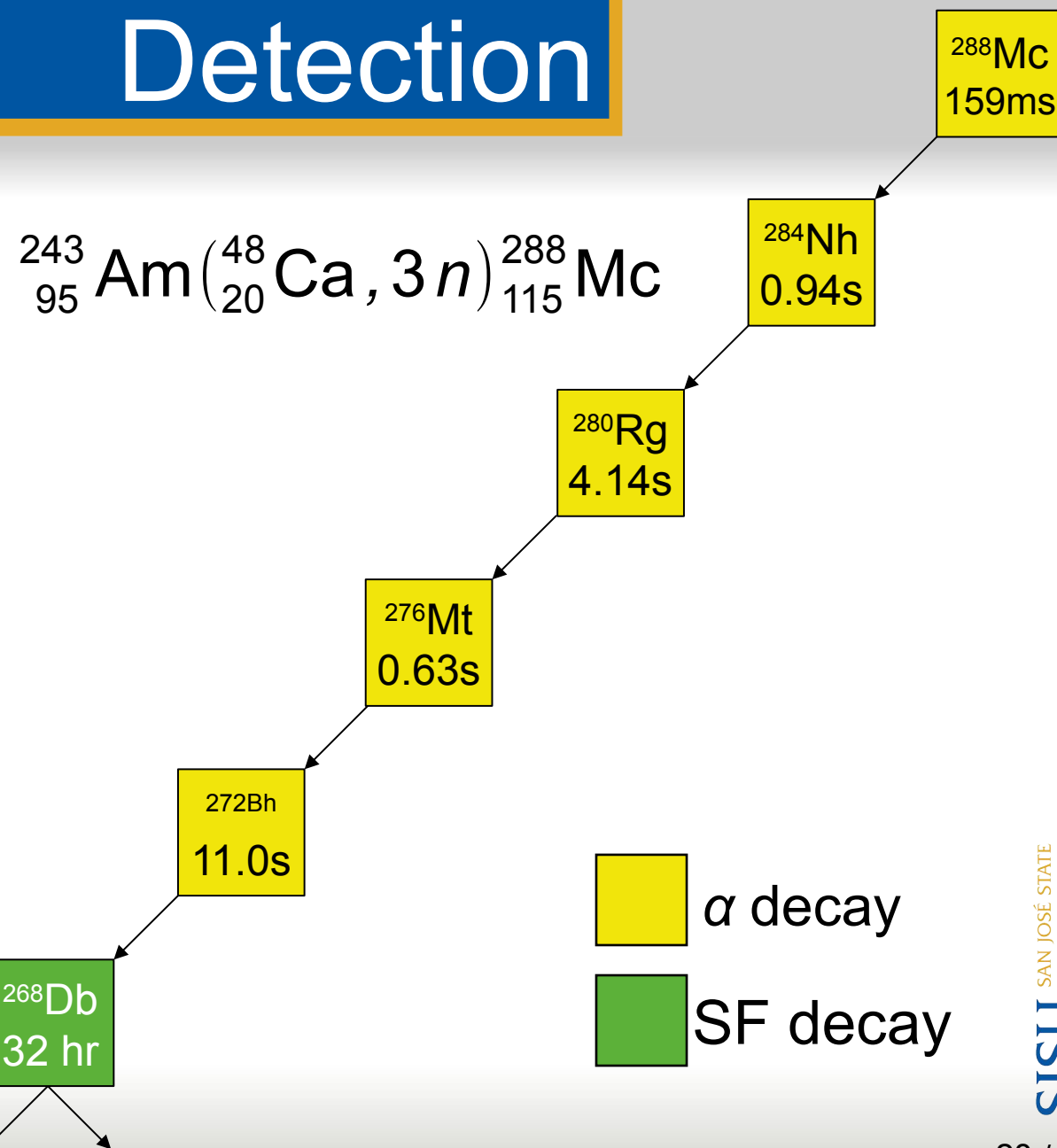
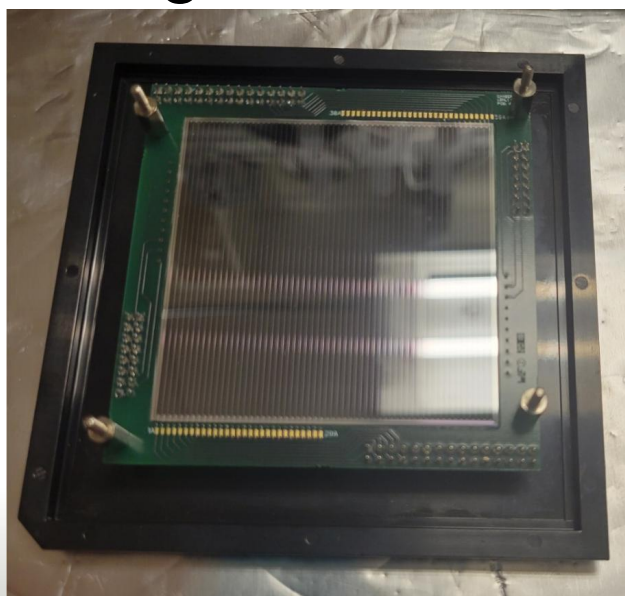
Vacuum vs Gas-filled Separators

	Vacuum Separators	Gas-filled Separators
Examples	FMA, SHIP, VASSILISSA, etc	BGS, DGFRS, GARIS, RITU, TASCA, etc
Angular acceptance	Moderate (up to 16 msr)	Large (up to 45 msr)
Velocity acceptance	Moderate (up to $\pm 25\%$)	Large (up to 100%)
Charge acceptance	Moderate (up to $\pm 7\%$)	Large (up to 100%)
Efficiency: $^{48}\text{Ca} + ^{208}\text{Pb}$	Up to 30%	Up to 70%
Mass resolution	Good ($M/\Delta M > 200$)	Poor ($M/\Delta M < 30$)



Detection

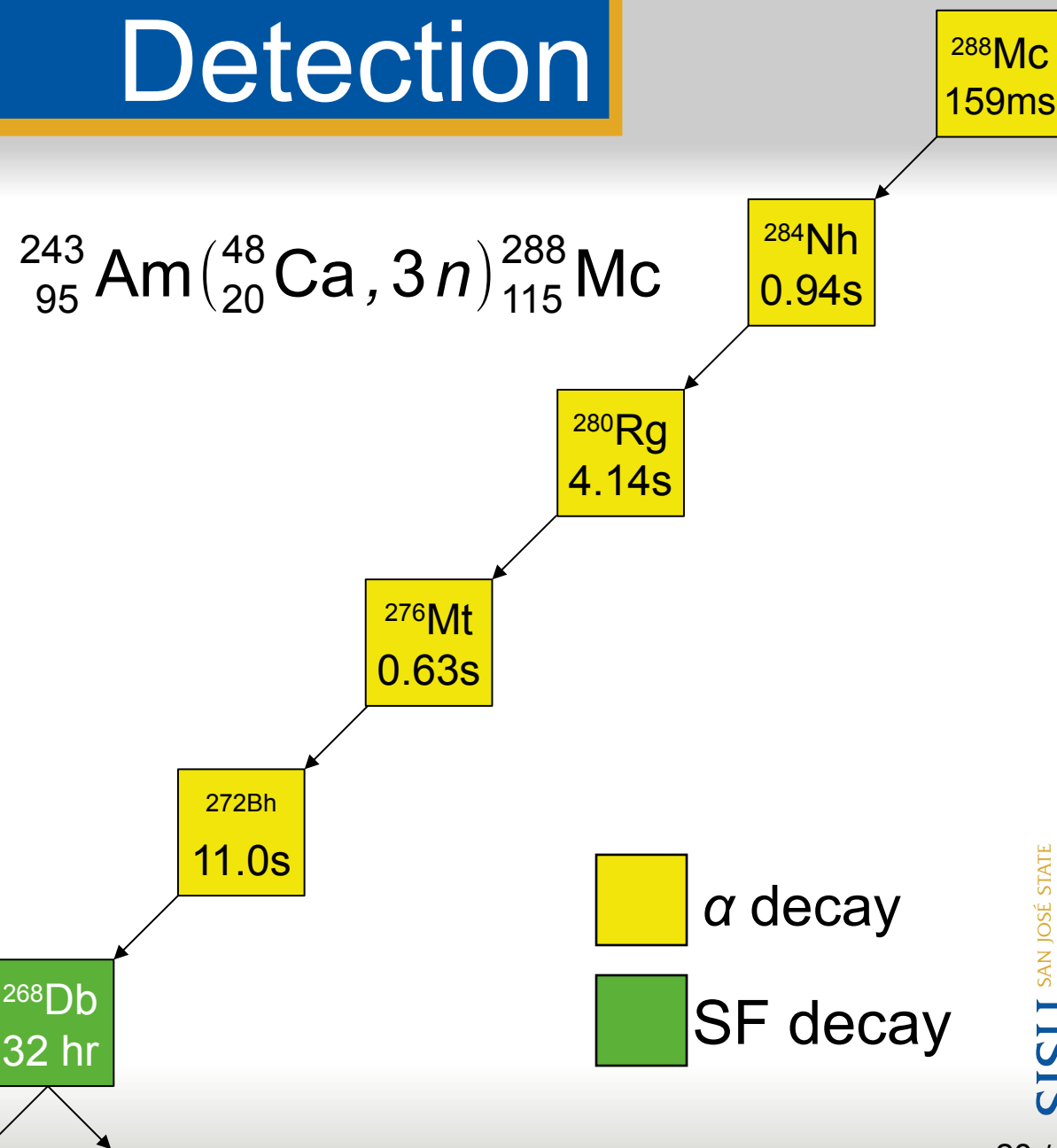
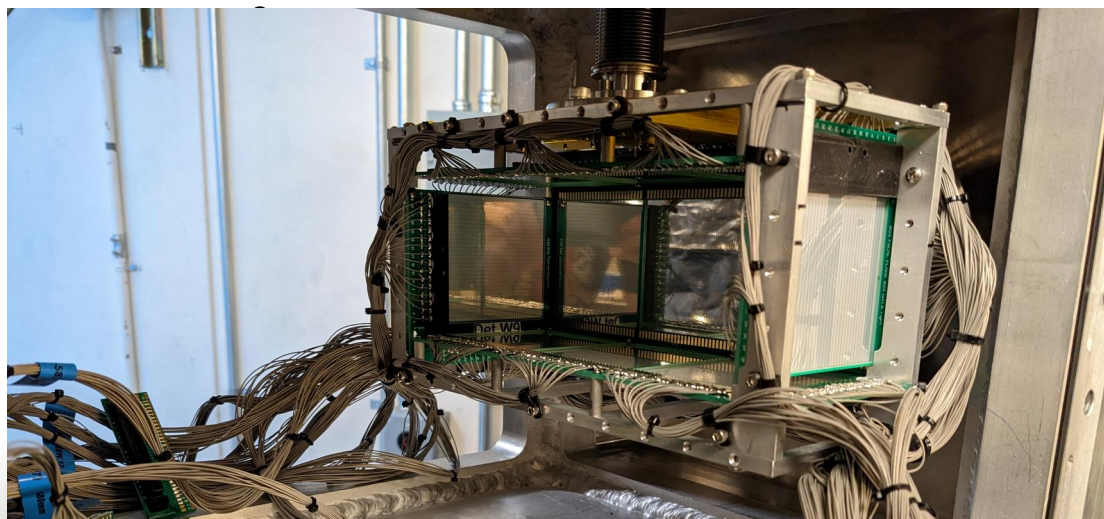
- Focal Plane Detection
 - Highly pixellated Si charged particle detectors
 - “Genetic Fingerprint” to distinguish from 1-2 kHz





Detection

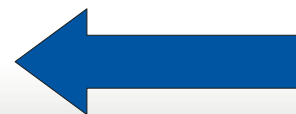
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 - Highly pixellated Si charged particle detectors
 - “Genetic Fingerprint” to distinguish from 1-2 kHz





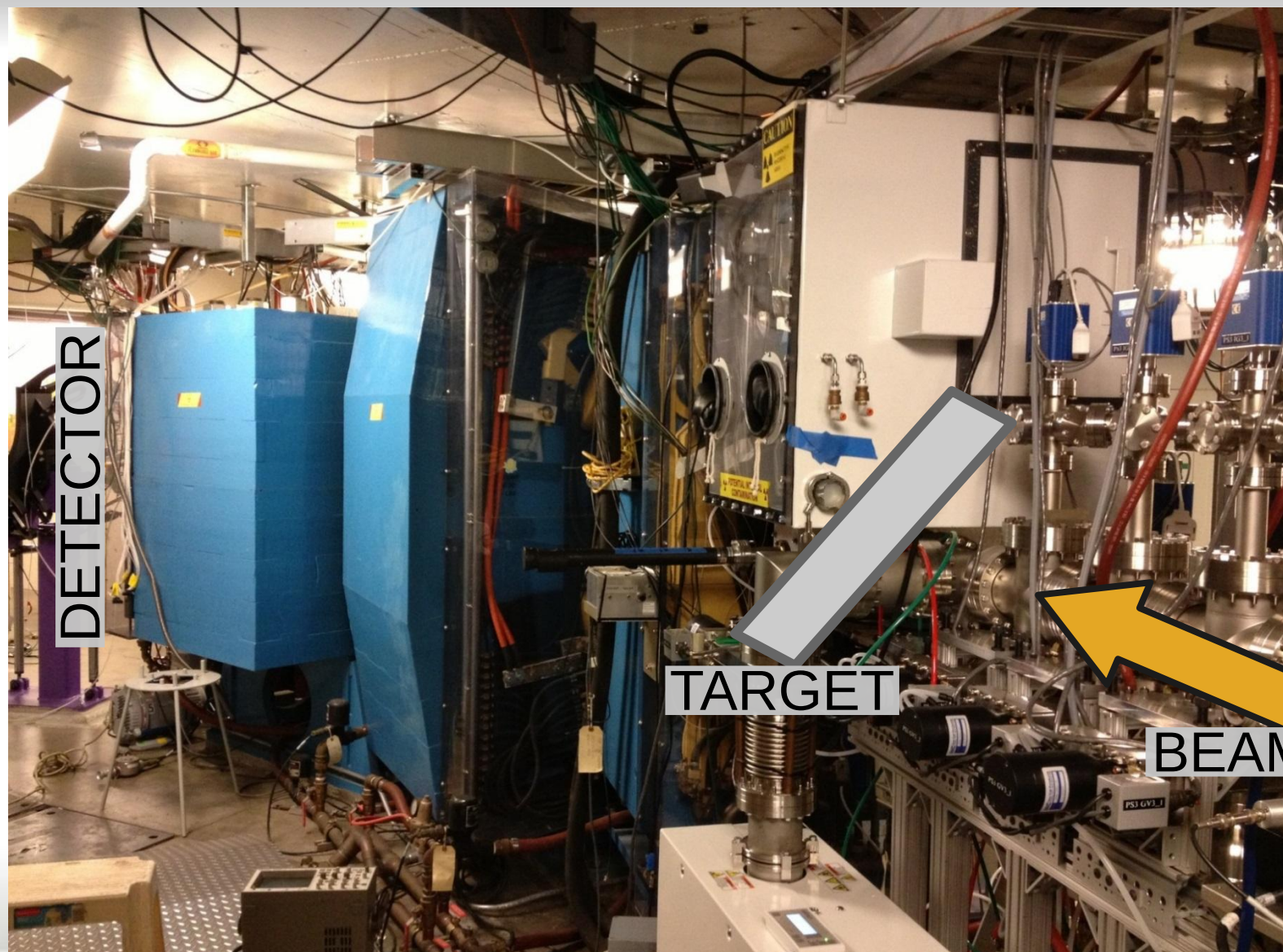
Current Efforts

- Production / Cross Section Measurements
- Chemistry of SHEs
 - Gas phase chemistry
 - Liquid phase chemistry
 - Thermal gradients
- Nuclear Structure
 - In-beam spectroscopy
 - Recoil – decay tagging on high spin isomers
- Mass Spectroscopy
 - High precision





Berkeley Gas-filled Separator

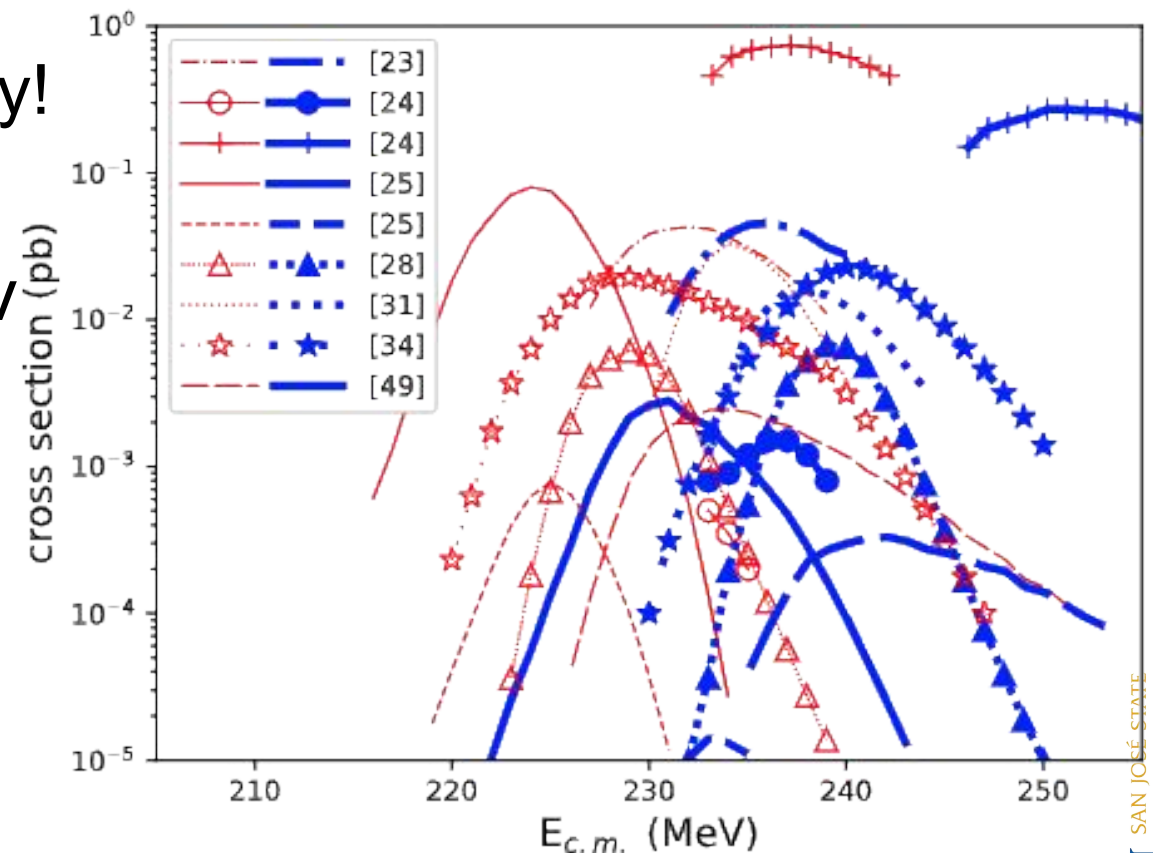


- BGS
 - Differential pumping
 - Glovebox for radioactive targets
 - QDD arrangement
 - BGS Focal Plane Detection area



Going beyond ^{48}Ca

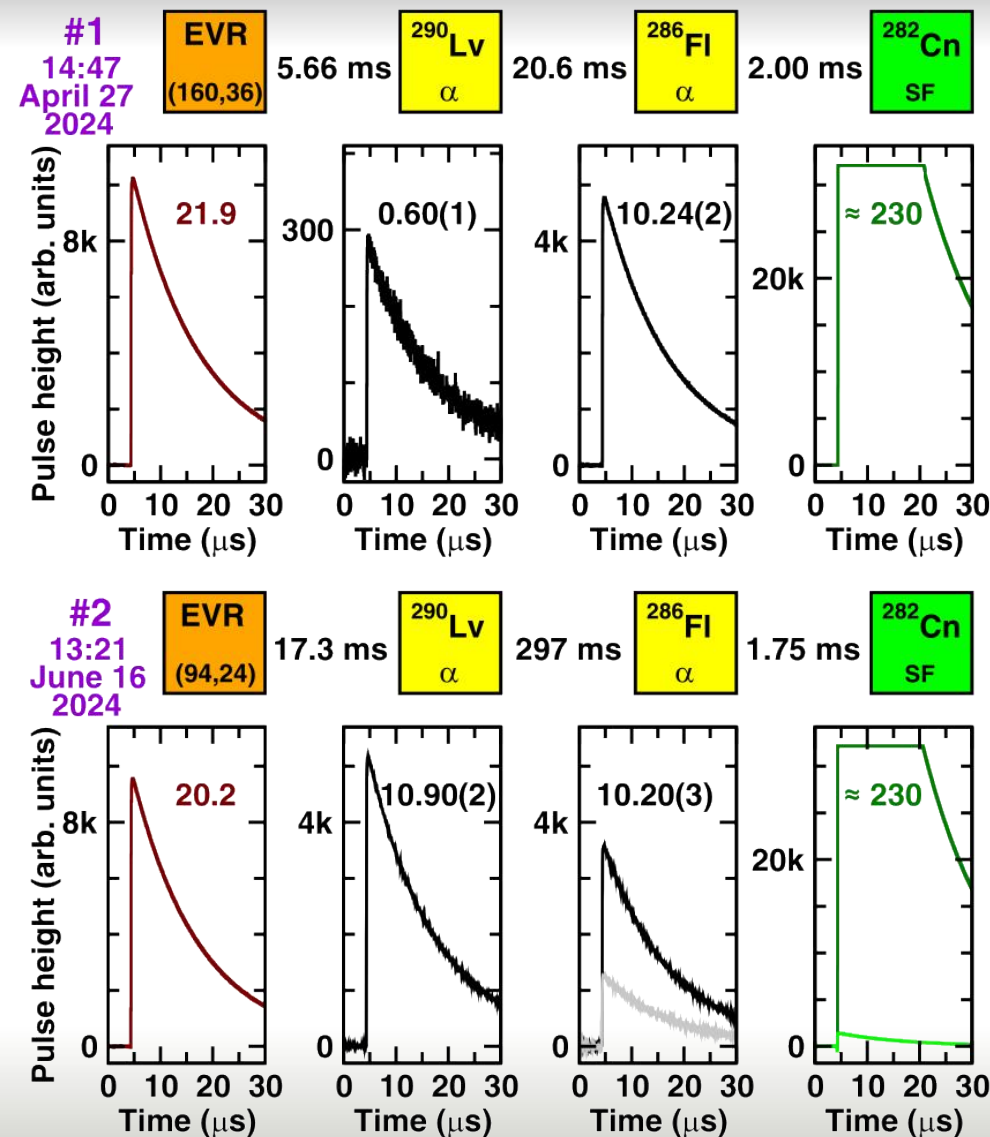
- Beam of ^{50}Ti is a promising next step
 - Been attempted before, need to verify!
- Experimental Plan: $^{244}\text{Pu} (^{50}\text{Ti}, 4n)^{290}\text{Lv}$
 - ^{50}Ti : 6×10^{12} ions / sec
 - ^{244}Pu : $\approx 430 \mu\text{g}/\text{cm}^2$





Going beyond ^{48}Ca

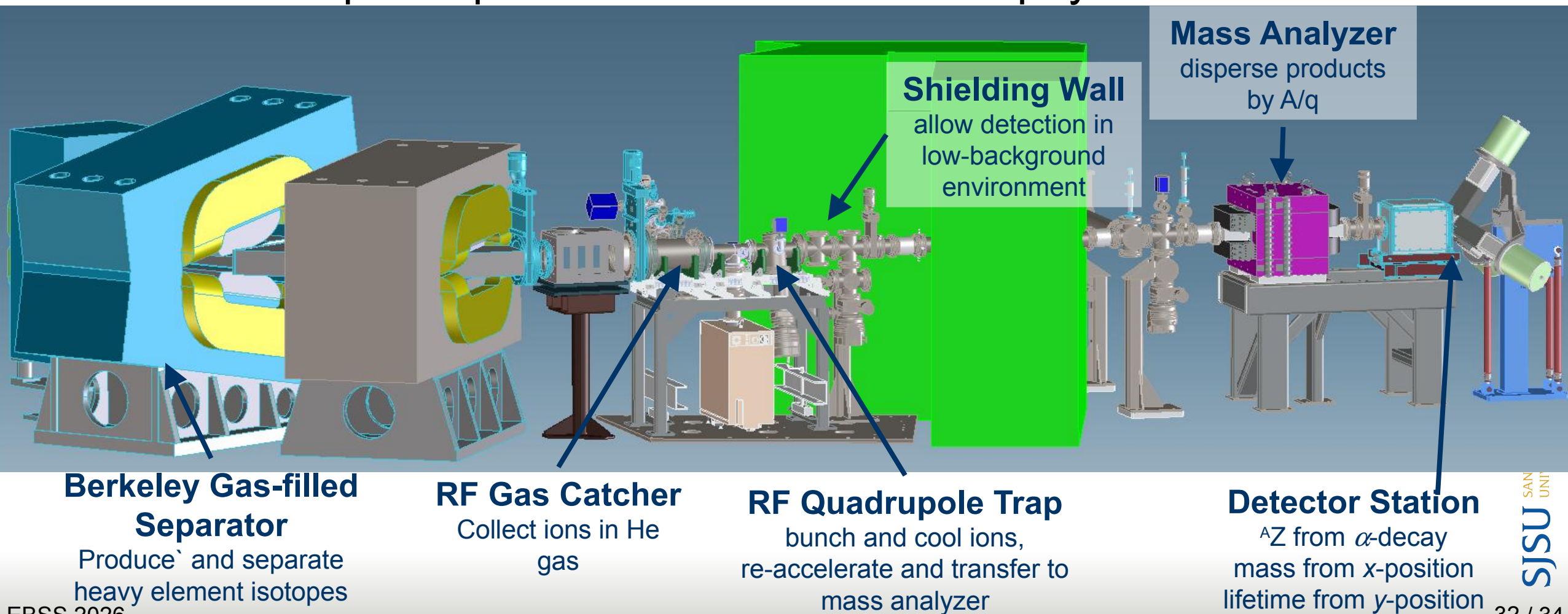
- Beam of ^{50}Ti is a promising next step
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 - ^{50}Ti : 6×10^{12} ions / sec
 - ^{244}Pu : $\approx 430 \mu\text{g}/\text{cm}^2$
- Results:
 - 22.1 days for 2 events!
 - E120 might be possible?!





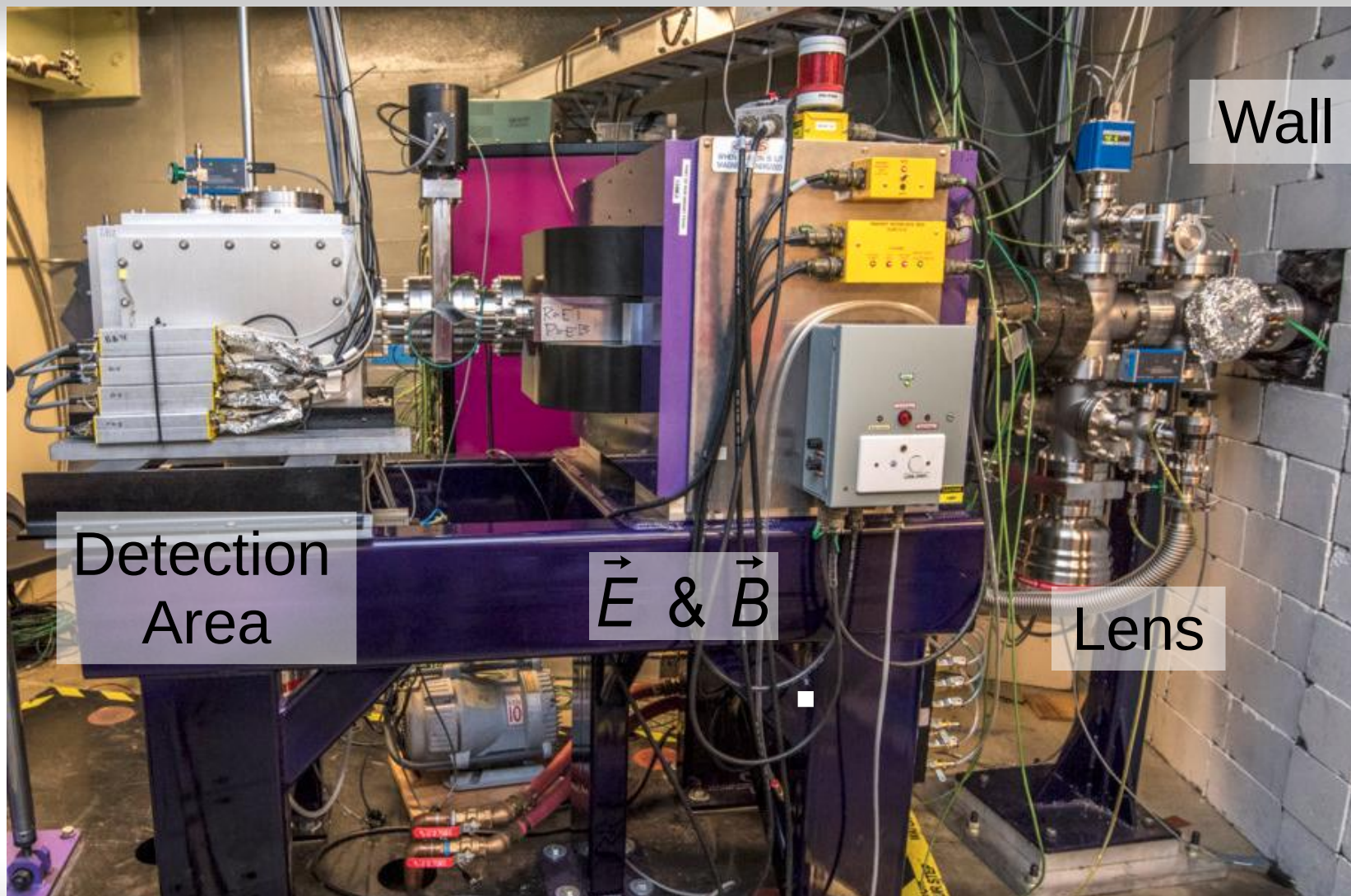
BGS + FIONA

- BGS as a pre-separator for a whole host of physical measurements!





BGS + FIONA



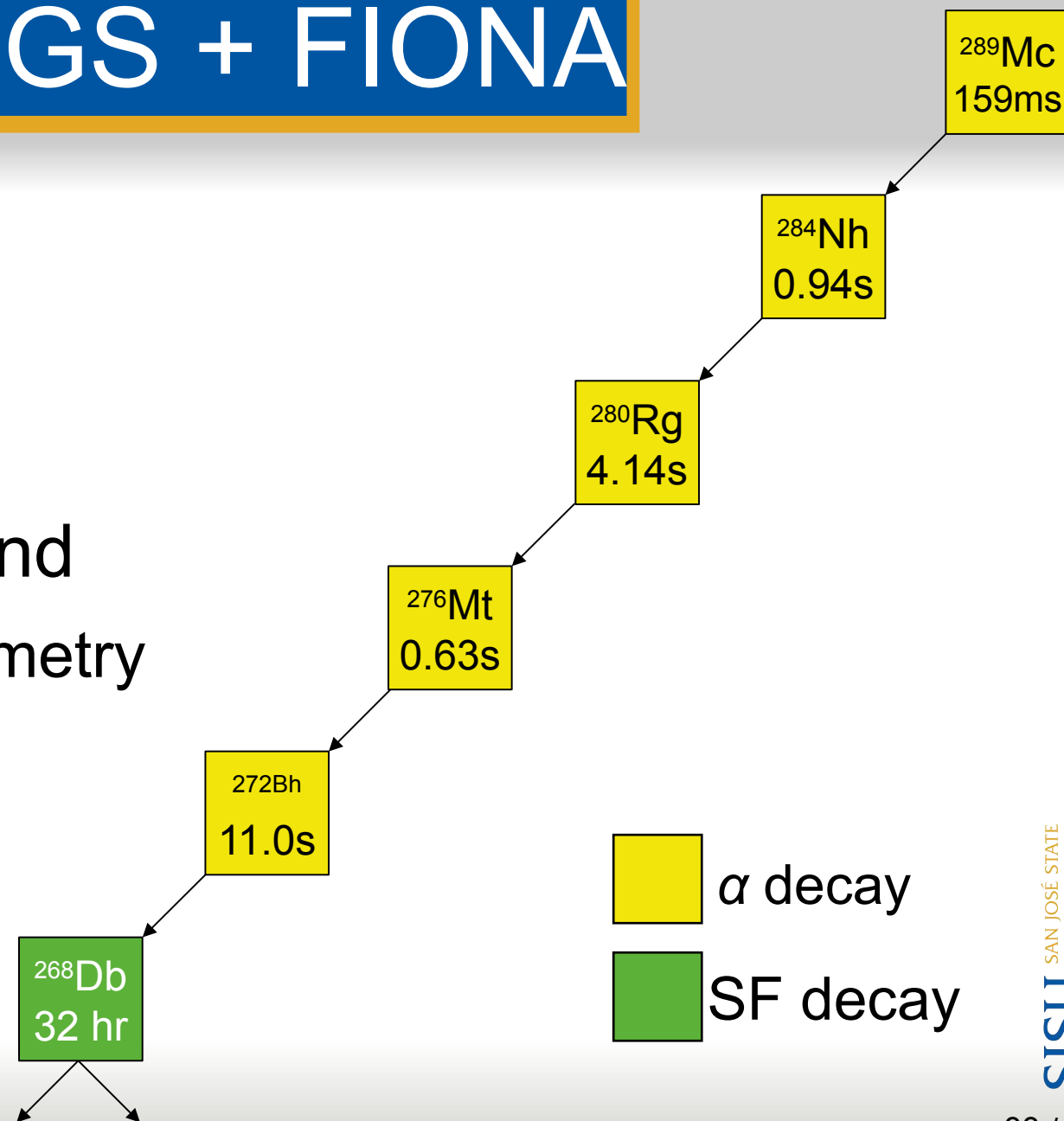
Experimental Setup:

- BGS:
 - 70% transmission
 - $\leq 1 \mu\text{s}$ transport time
- Thermalization:
 - 30% transmission
 - $\sim 30 \text{ ms}$ transport time
- FIONA
 - 90% transmission
 - $1 \mu\text{s}$ transport time



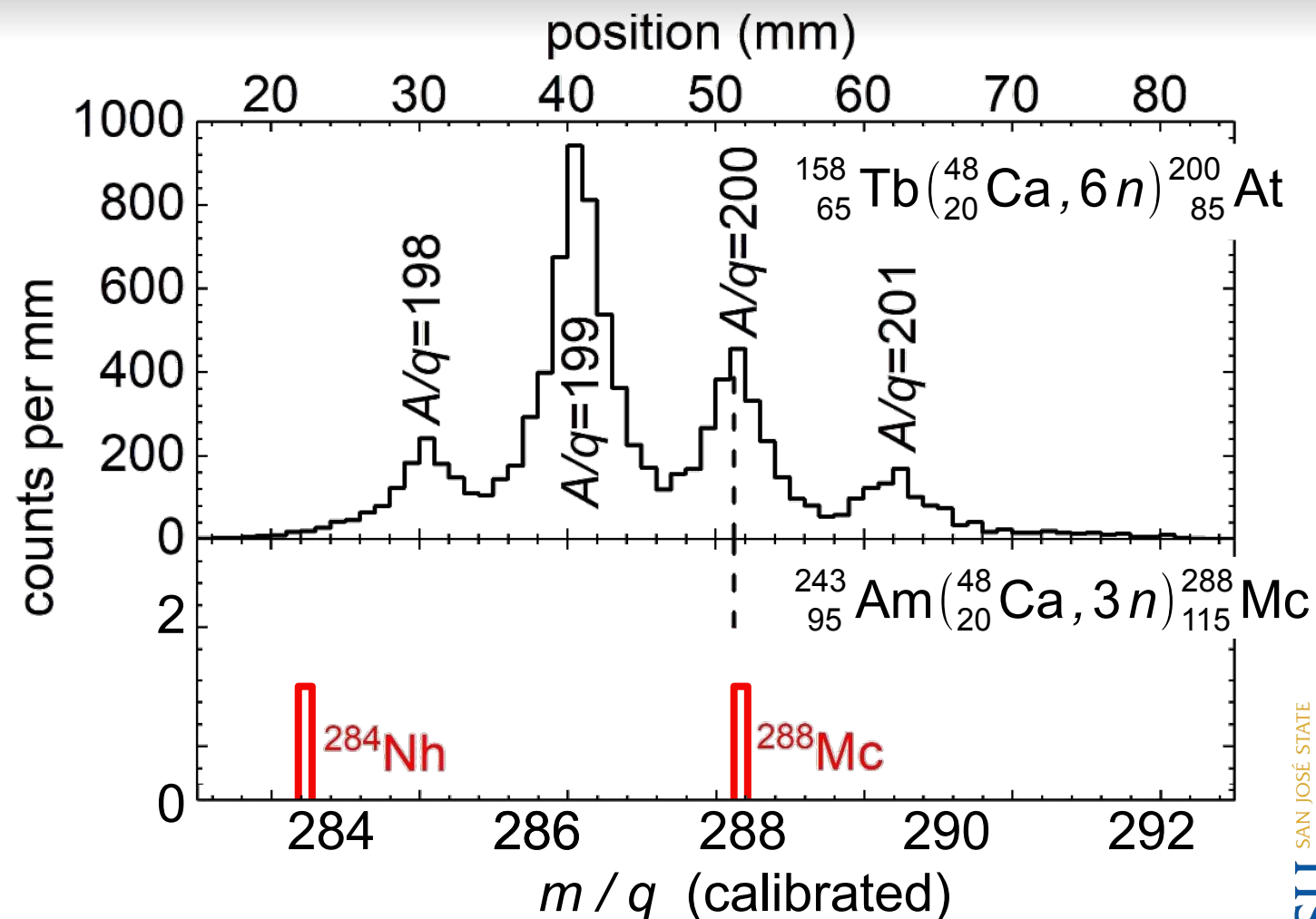
BGS + FIONA

- Headline experiment:
 - ${}_{95}^{243}\text{Am}({}_{20}^{48}\text{Ca}, 3n){}_{115}^{288}\text{Mc}$
 - Detection via decay chain
- Huge reduction in background
 - Atom at a time mass spectrometry
 - Nuclear structure studies
 - Gamma, X ray spectroscopy
 - Fission studies





- First direction mass determination of a superheavy element!
- Implications:
 - Mass model confirmed
 - Anchor 50+ isotopes





In Conclusion

- Why study the superheavies?
- How are they produced and studied?
 - Recoil Separators and production studies
- Heavy element science at 88" cyclotron
 - Verified synthesis of ^{290}Lv ($Z=116$) using ^{50}Ti
 - Mass Identification of SHEs with FIONA

Thank you!

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