

Scintillation Imaging with Coded Aperture Masks

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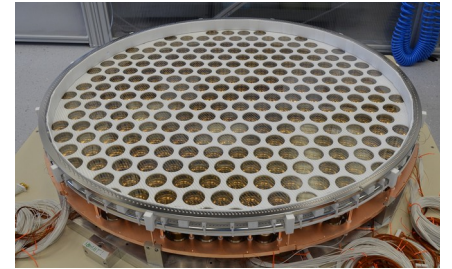
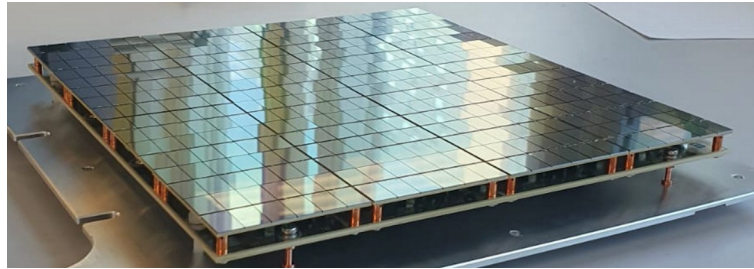
Ministero
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Italiadomani
PIANO NAZIONALE
DI RIPRESA E RESILIENZA



- In the vast majority of scintillation detectors, light is collected *without* optical elements capable of forming *images*

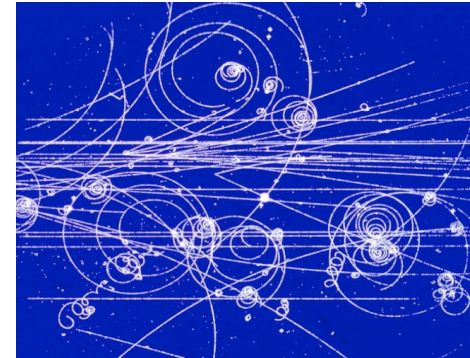


- Many experiments only use the amplitude and/or the timing of the scintillation signal. **None takes pictures of the tracks.***

*Several other prototypes and proposals exist, see for example [1], [2]. Also, one could argue Cherenkov detectors take pictures of tracks via the rings.

- **Why is scintillation imaging not used?**

- Taking pictures (of bubbles) is how we started...
- However, with scintillation *"light is not enough"*



- **Technological developments are now challenging this assumption:**

- More efficient, large area SiPMs detect more light on a broader spectrum
- More advanced ASICs enable higher channel densities
- More computing power allows for more complex reconstruction algorithms

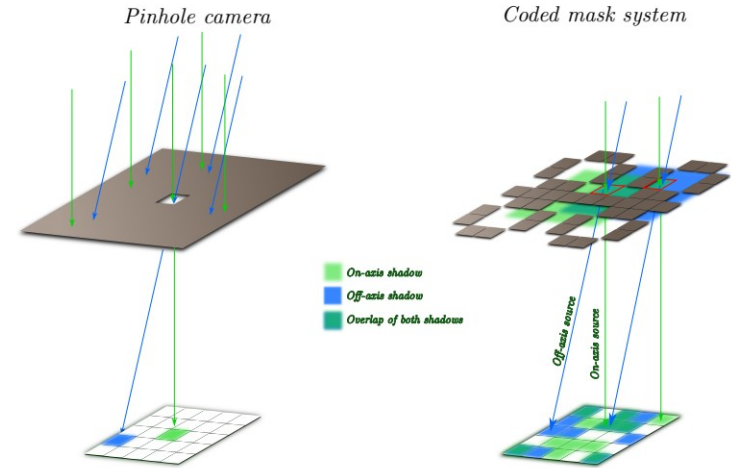
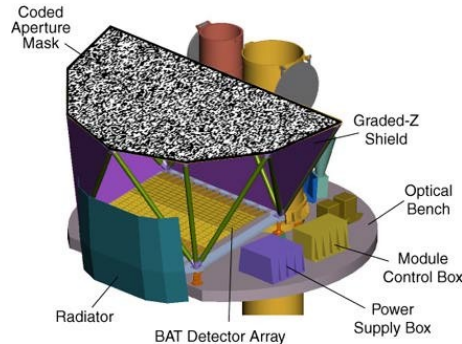
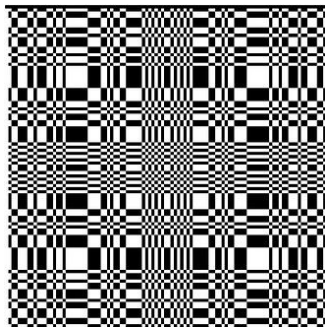
Scintillation imaging – the case for liquid Argon

- High photon yield per unit energy and especially per unit length
- Cryogenic operation is challenging, but negates SiPM noise
- **VUV spectrum is a problem for optics and sensors**
 - Coded aperture imaging ignores wavelength
 - Many R&D efforts focused on improved VUV sensitivity



What is coded aperture imaging?

- A technique developed for X and γ photons
 - These cannot be refracted or reflected
 - in astrophysics and in medical imaging



- An extension of the pinhole camera which captures more light

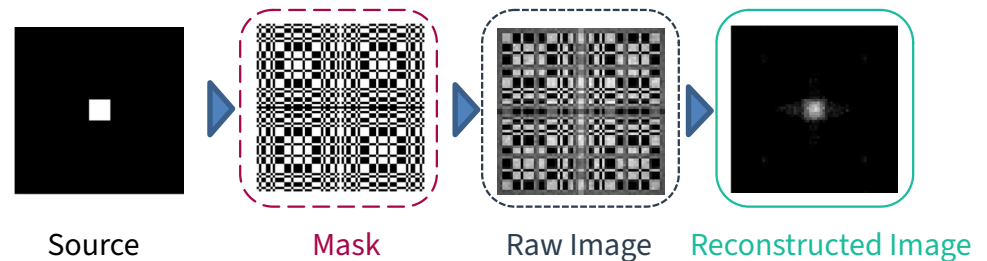
- For *far* field imaging (i.e. astrophysics):

- the original image can be obtained with a deconvolution process where the decoding matrix is derived from the mask pattern. The pattern matters here.

$$\text{Reconstructed Image } O'(x, y) = \text{Raw Image } I(x, y) \otimes H(x, y)$$

Where H is the “inverse” of M, such that

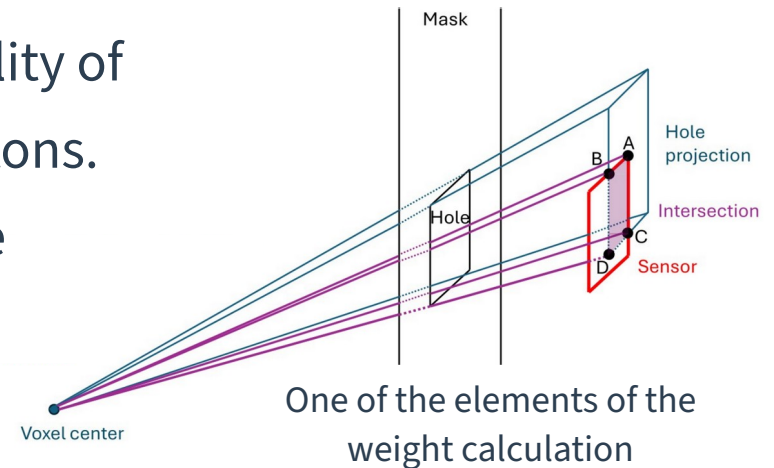
$$\text{Mask } M(x, y) \otimes H(x, y) = \delta(x, y)$$



- For *near* field imaging:

- more complex and computationally intensive algorithms can be implemented: Filtered Back Projection, Maximum Likelihood Expectation Maximization. Pattern can be random.

- Directly reconstructs in 3D the initial *photon source distribution* in a segmented volume (voxel array):
 - *measured photons* from all cameras are *propagated back* into the LAr volume with an appropriate weight, which is added to the voxel value
 - this weight represents the Bayesian probability of the voxel to be a source of the detected photons.
 - The *likelihood* of the resulting photon source distribution having produced the raw data is *maximized* through an iterative process.



- Photon counting is described by a Poissonian pdf:

$$f(H_s | [\lambda_s]) = e^{-[\lambda_s]} \frac{[\lambda_s]^{H_s}}{H_s!}$$

$$[\lambda_s] = \sum_j \lambda_j w(j, s)$$

H_s is the number of photons detected on sensor s (raw data)

λ_j is the (unknown) photon source value in voxel j

$[\lambda_s]$ is the expectation value of the detected photons

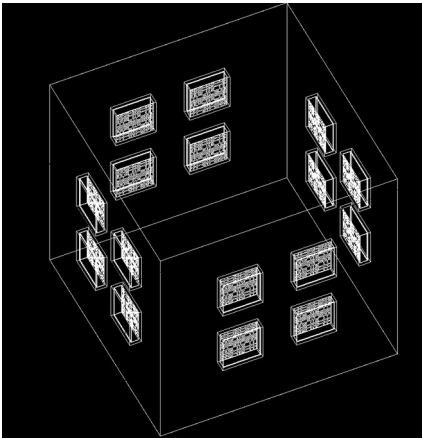
$w(j, s)$ is the weight (a very large precalculated matrix)

- The likelihood for all sensors must be maximized (iteratively) [3]

$$\prod_s e^{-[\lambda_s]} \frac{[\lambda_s]^{H_s}}{H_s!}$$

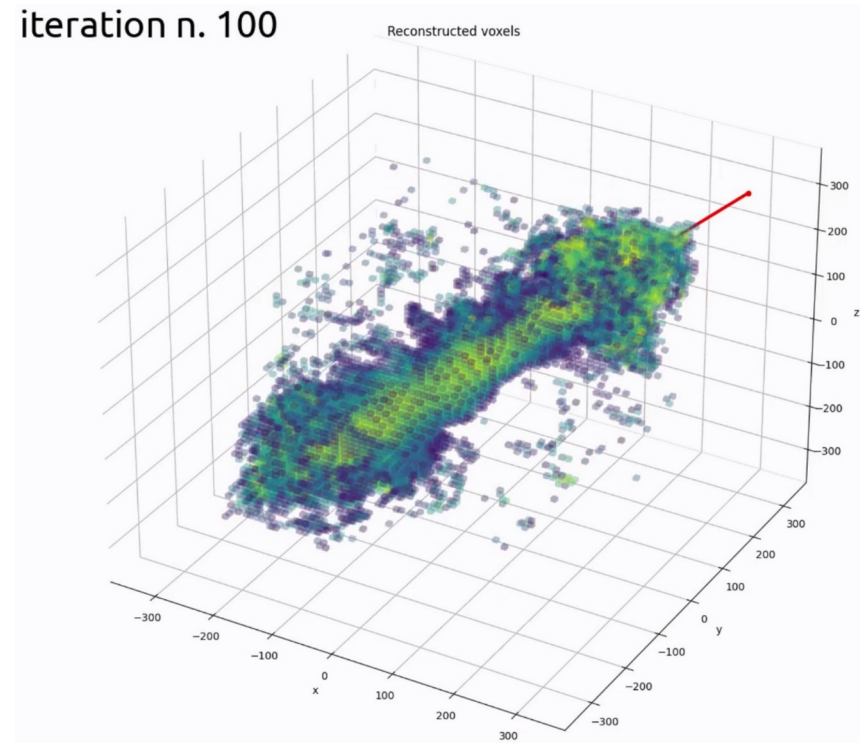
$$\lambda_j^{k+1} = \frac{\lambda_j^k}{\sum_s w(j, s)} \cdot \sum_s \frac{H_s \cdot w(j, s)}{\sum_j w(j, s) \cdot \lambda_j^k}$$

- Simulated 0.33 t LAr volume
 - 16 x 1024 channel cameras
 - 3x3 mm SiPM with TPB

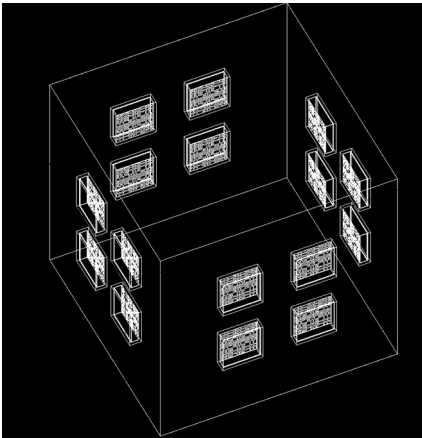


Images courtesy V. Cicero

iteration n. 100

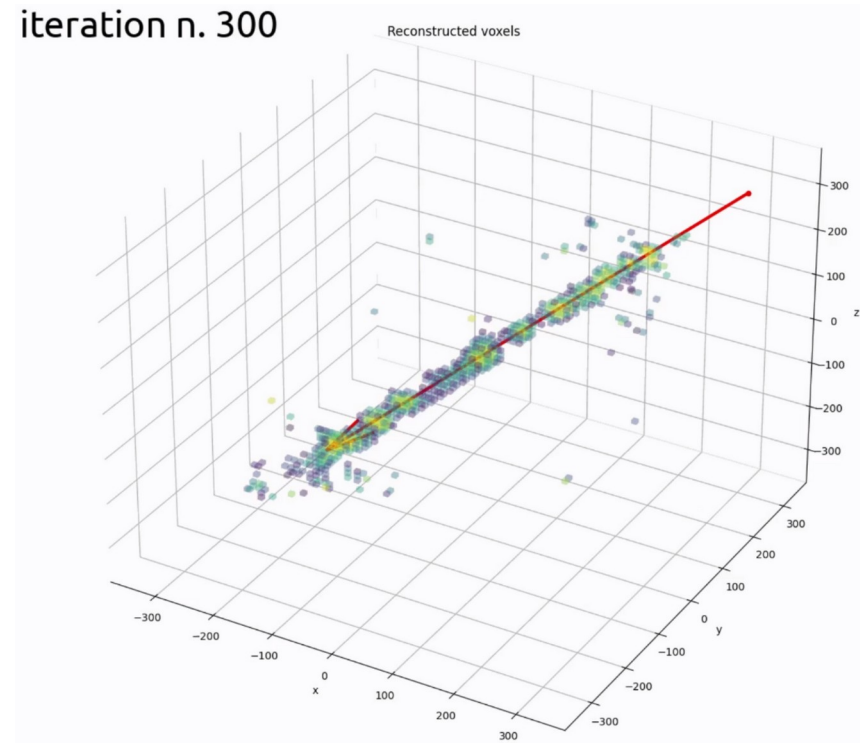


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Images courtesy V. Cicero

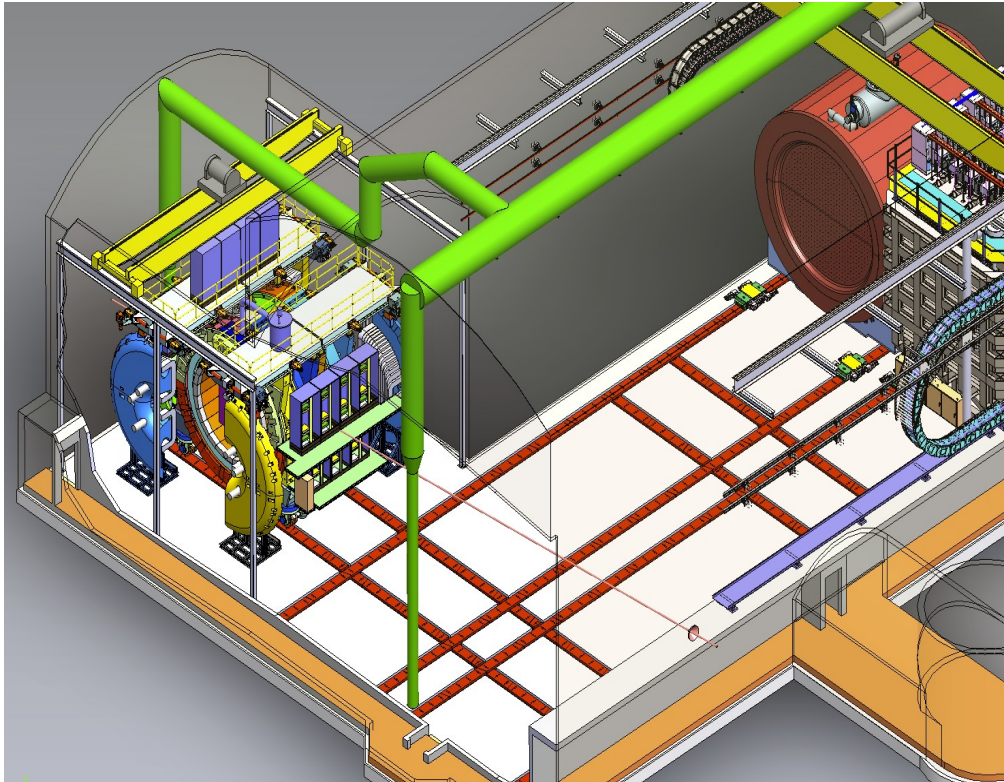
iteration n. 300



- **Achieving 3D tracking with 2D projections, from the periphery**
 - Good scaling to large volumes. Channel count only grows with size²
 - Only scintillator in the active volume, no passive material

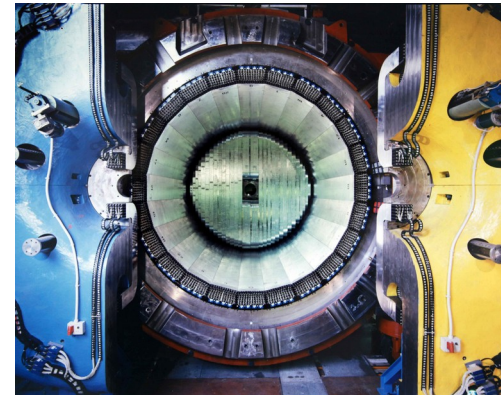
- **This could also be said of a TPC, where is the advantage then?**
 - $\sim\mu\text{s}$ vs $\sim\text{ms}$. A different compromise between rate and resolution
 - No HV, no field cage, potentially somewhat more robust in operation.

A real use case in the DUNE Experiment

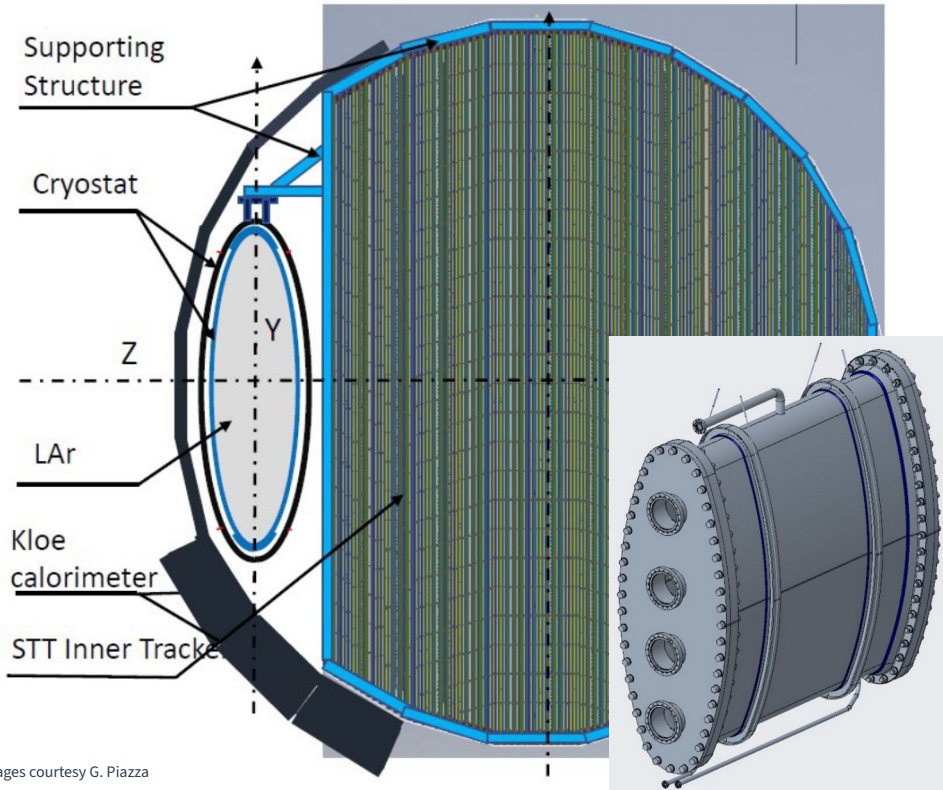


SAND is one of the three elements of the DUNE Near Detector complex [4]

- Re-use of the *KLOE* magnet and ECAL
- New gas Target Tracker and LAr “active target”



GRAIN, the Active Argon Target in SAND

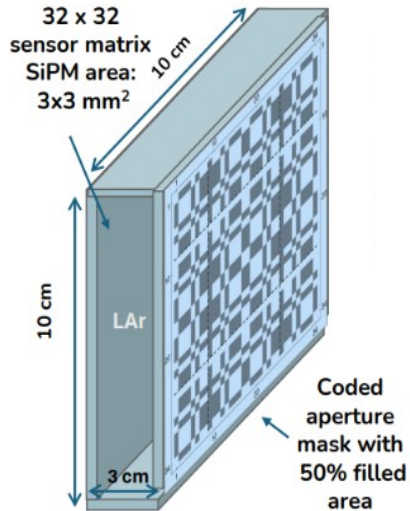


Images courtesy G. Piazza

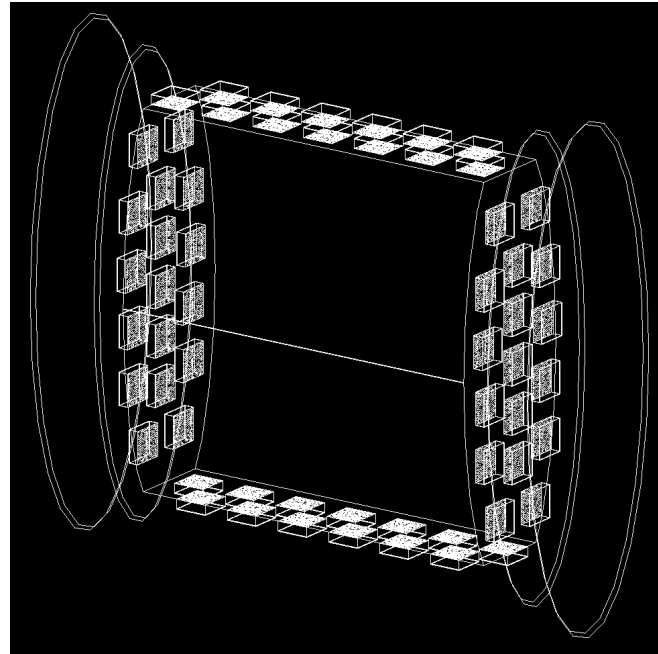
- A 1-ton target in a “thin” cryostat
 - Optical readout for rate
 - Several tracks/spill
- Main motivations:
 - constrain nuclear effects on Ar
 - have a complementary (to ND-LAr) target permanently located on-axis for cross-calibration

GRAIN read out with Coded Aperture cameras

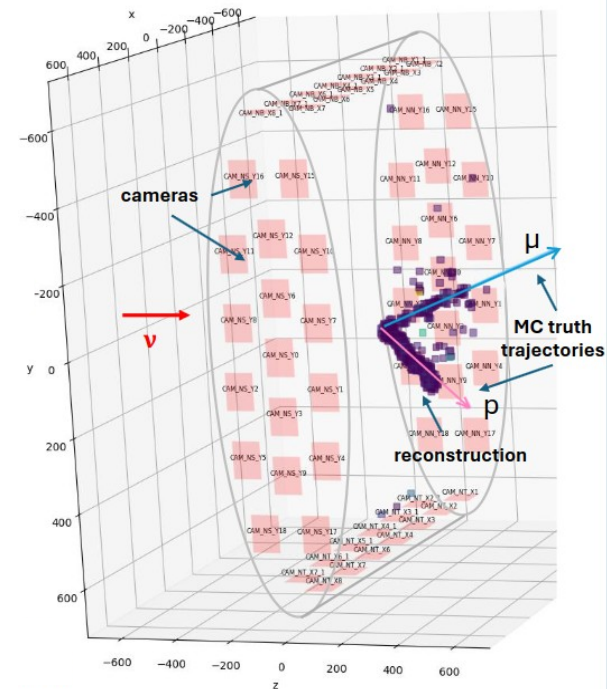
1024 pixel camera



60 cameras in GRAIN



MLEM reconstruction of ν_μ -CC event



Images courtesy V. Cicero

- **Coded aperture cameras can be used for near field imaging with compact (flat) detectors in fields other than neutrino physics**
 - Direct coupling to solid scintillators for high resolution calorimetry
 - Readout of LAr based “tracking” calorimeters, including proposed LAr PETs

- **Coded aperture cameras can still be coupled with charge readout**
 - Enhance the rate capability of TPCs, aid in event reconstruction

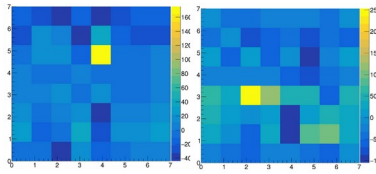
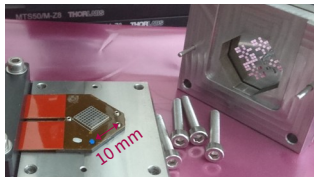
- Scintillation imaging with Coded Aperture cameras shows promise in simulation
- Reaching maturity for application in GRAIN, part of the Near Detector complex of the DUNE experiment, is presently our main goal
- Additional applications are undergoing early studies

- [1] J. Dalmasson et al. *Distributed imaging for liquid scintillation detectors* PHYSICAL REVIEW D 97, 052006 (2018)
- [2] A. Musumarra et al. *RIPTIDE: a novel recoil-proton track imaging detector for fast neutrons* JINST 16 C12013 (2021)
- [3] R. Willingale et al. *Advanced deconvolution techniques for coded aperture imaging* Nuclear Instruments and Methods in Physics Research 221.1 pp. 60–66 (1984)
- [4] A. Abed Abud et al. *DUNE Near Detector Conceptual Design Report* <https://arxiv.org/abs/2103.13910>

Backup

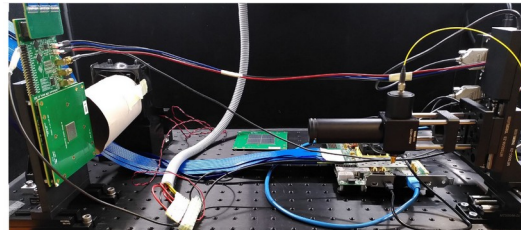
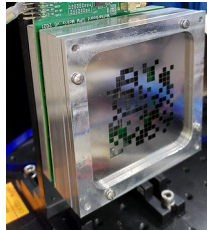
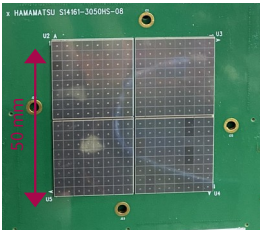
- Readout systems of increasing performance have been developed:

- Early warm demonstrator, 64 channels, 1x1 mm SiPMs, warm ASIC (TRIROC)



- could only reconstruct point sources

- Cold demonstrator: 256 channels, 3x3 mm SiPMs, 8 x prototype cryo ASIC



- laser calibration in progress

- Future development: a 1024 channel cryo ASIC for GRAIN

Coded aperture cameras with charge readout

- **Goal:**
 - Enhance the rate capability of TPCs
 - Aid in event reconstruction
- **The mask becomes active**
 - Filled squares become charge readout pads
 - Holes may also be made sensitive

