

# STEREO: Quick Review

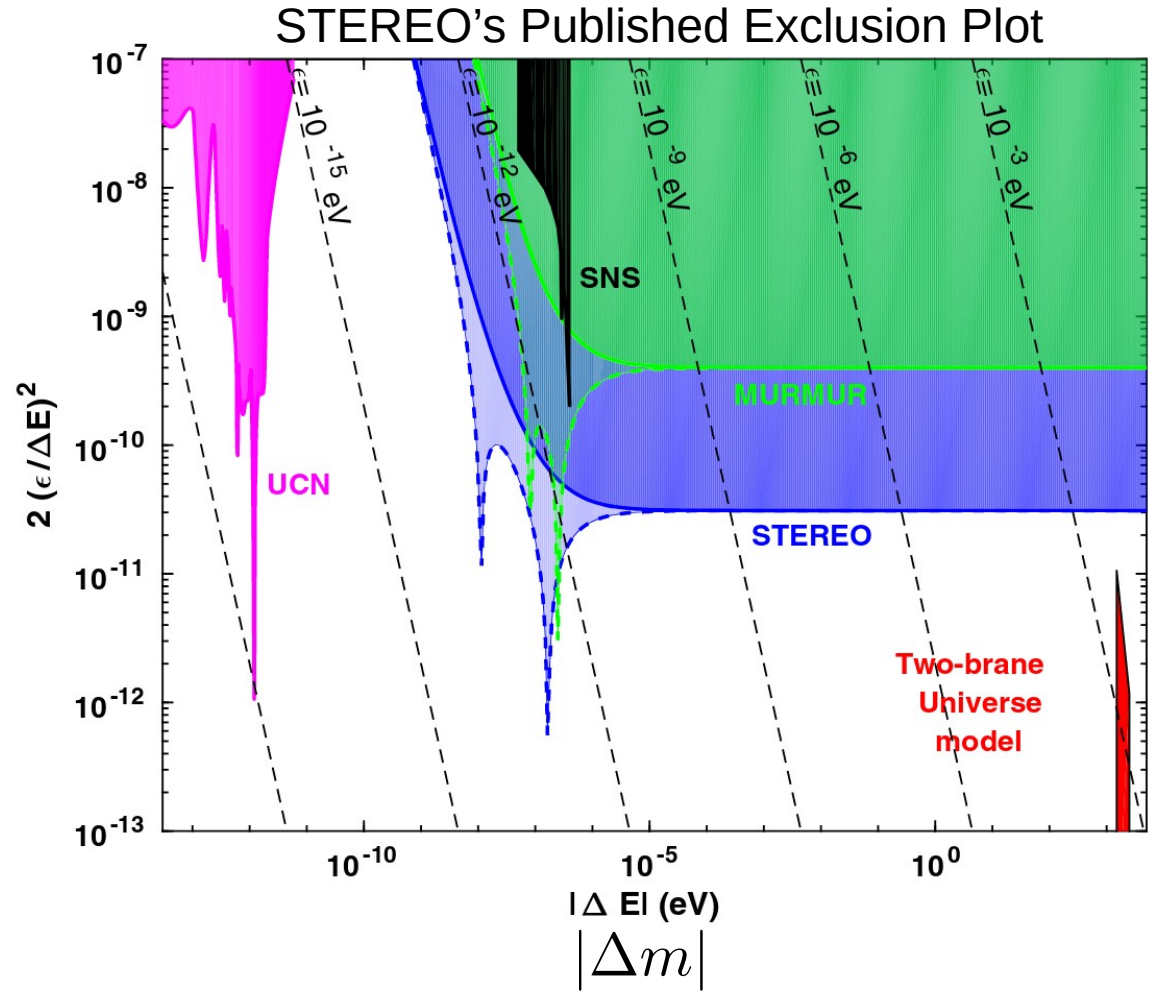
UT-ORNL n → n' Group Meeting

6/22/2023

Cary Rock

# STEREO: The Experiment and Claim

- Repurposed neutrino detector at the RHF of ILL, Grenoble
- Claim limit on  $n \rightarrow n'$  of  $p_{nn'} \sim 3.1 \times 10^{-11}$



# STEREO's Probability Calculation

STEREO's final calculation depends on:

1. computing the prob. per neutron,  $p$ ,

2. multiplying by the visible flux,  $\Phi_{vis}(\vec{r})$ , to give the volumic rate of projections,  $S_{hid}(\vec{r})$ ,

3. and then integrating over the reactor volume to get  $\Phi_{hid}(\vec{r}_d)$ .

$$\varepsilon = \Delta m \theta_0 \quad \Gamma = v \Sigma_{sc}$$

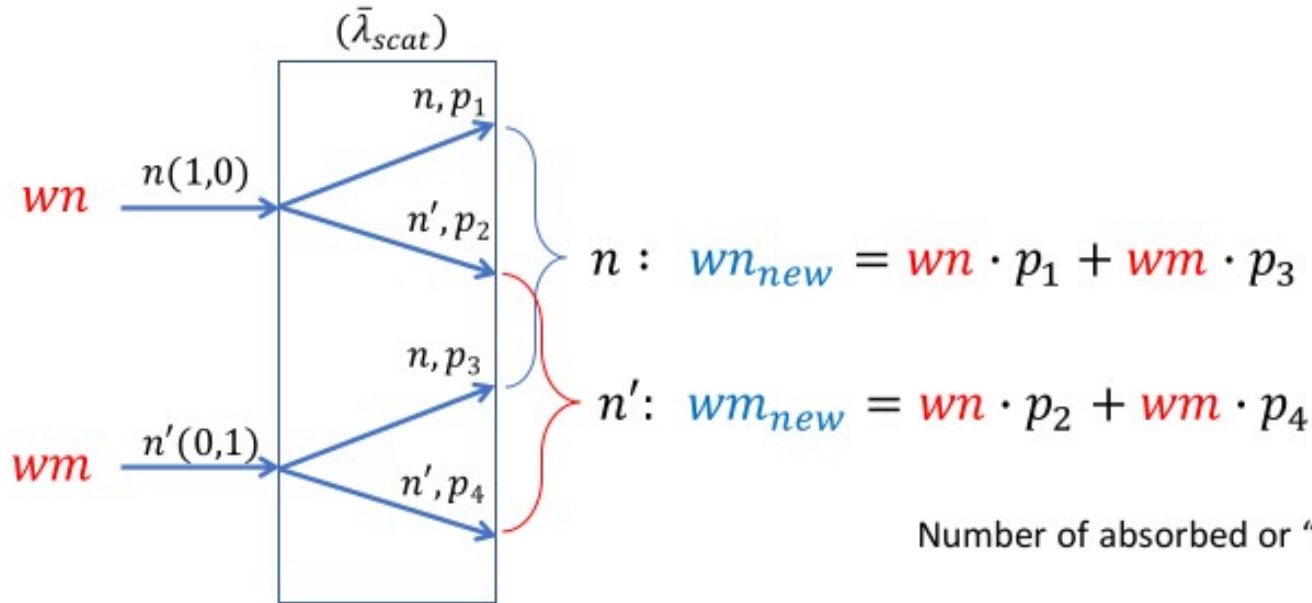
$$p = \frac{2\varepsilon^2}{(\pm\Delta m + V_F)^2 + 4\varepsilon^2 + \hbar^2\Gamma^2/4}$$

$$S_{hid} = p \frac{\Sigma_S}{2} \Phi_{vis}(\vec{r})$$

$$\Phi_{hid}(\vec{r}_d) = \frac{1}{4\pi} \int_{Reactor} \frac{S_{hid}(\vec{r})}{|\vec{r} - \vec{r}_d|^2} d^3r$$

# Method of Calculation

In the fast calculation method using Berezhiani's formulas (4)-(12) from the paper <https://arxiv.org/abs/2111.01791> the wave functions solution of Schrodinger equation is used instead of density matrix. This solution always starts with initial  $\begin{pmatrix} 1 \\ 0 \end{pmatrix}$  or  $\begin{pmatrix} 0 \\ 1 \end{pmatrix}$  for  $\begin{pmatrix} n(0) \\ n'(0) \end{pmatrix}$ , which allows to construct final density matrix from the solutions of S.E.  $\begin{pmatrix} n(t) \\ n'(t) \end{pmatrix}$ .



$wn$ : prob. of entering as  $n$

$wm$ : prob. of entering as  $n'$

$p_1$ : contribution of  $n$  to exit as  $n$

$p_2$ : contribution of  $n$  to exit as  $n'$

$p_3$ : contribution of  $n'$  to exit as  $n$

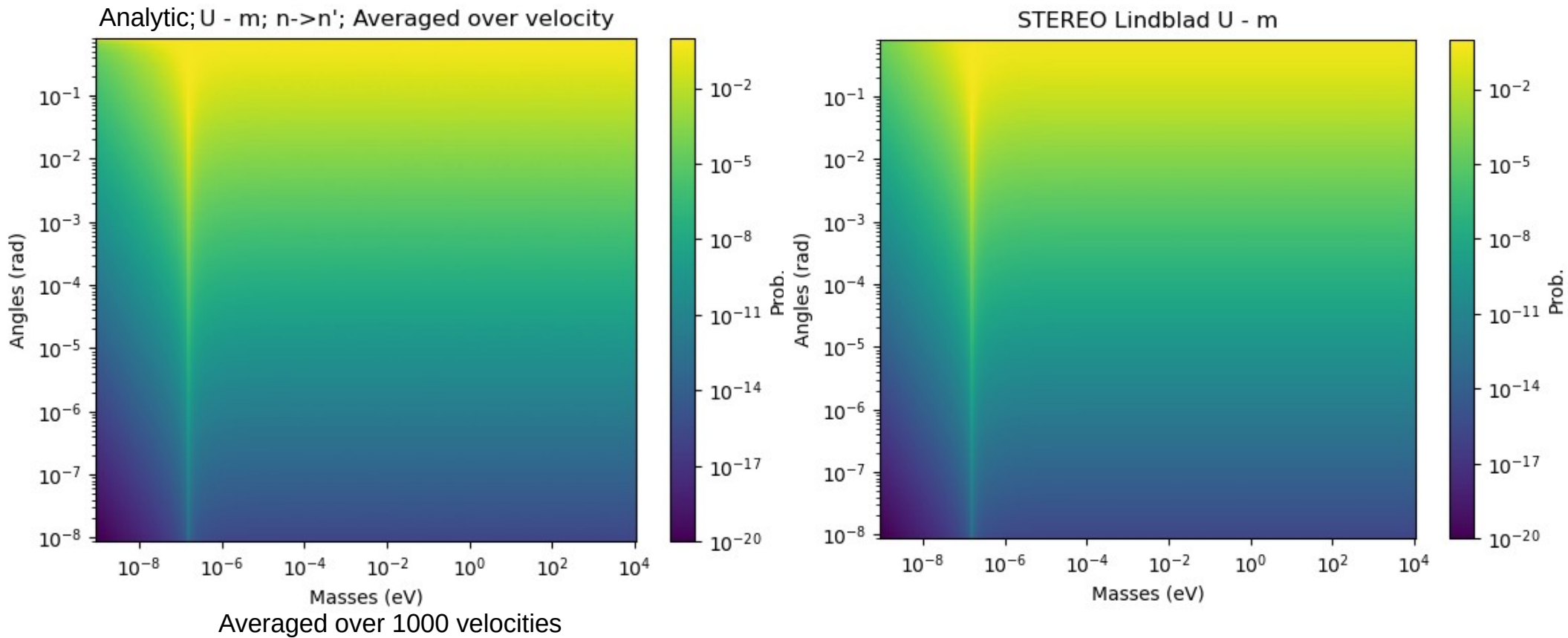
$p_4$ : contribution of  $n'$  to exit as  $n'$

Number of absorbed or "detected" in the layer:

$$\Delta = (wn + wm) - (wn_{new} + wm_{new})$$

If there is no "mirror absorber"

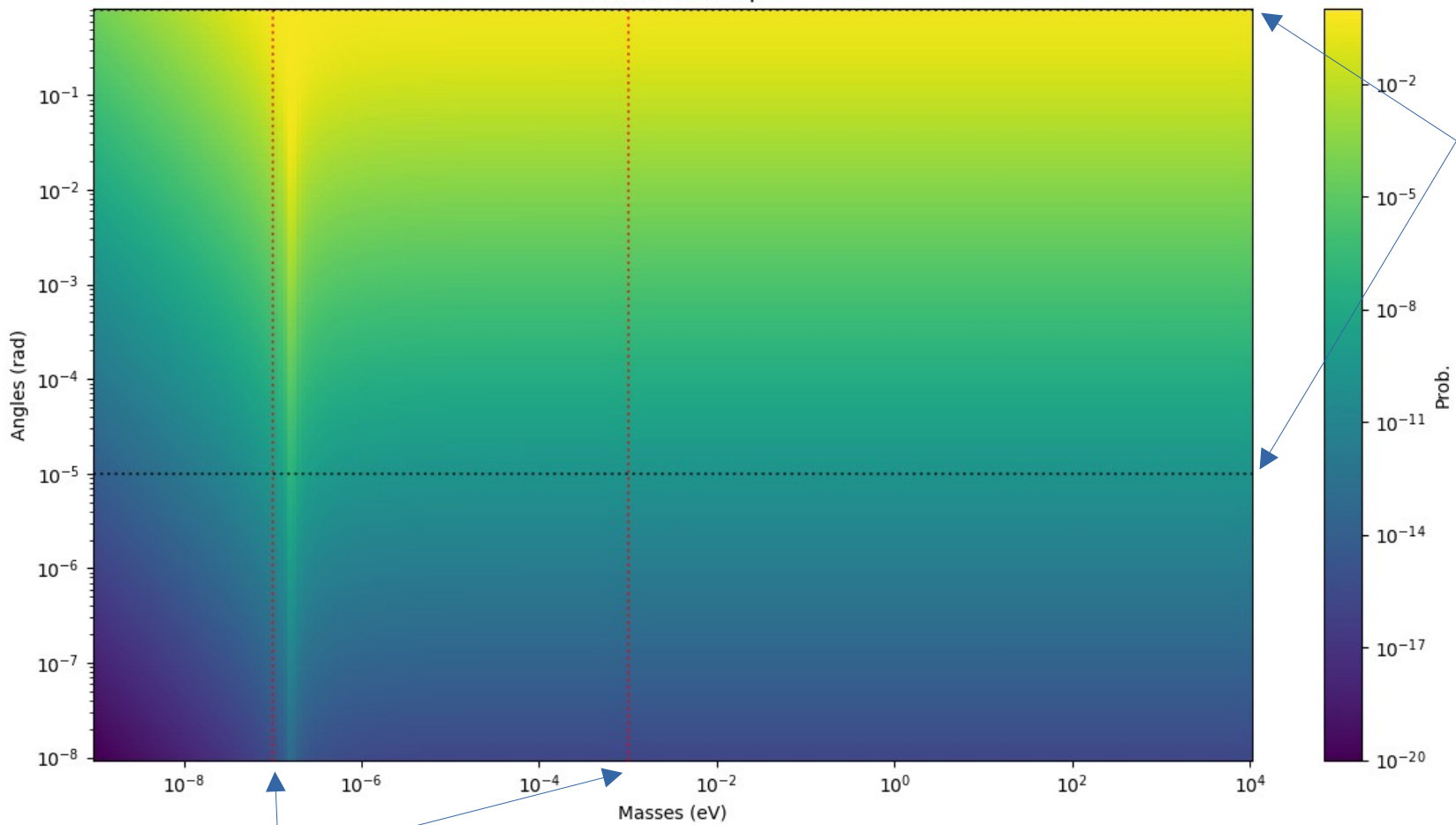
# Comparison in D<sub>2</sub>O



Any difference between these two is not very obvious at this level

# Slices for Better Comparison

Lindblad 2D Map

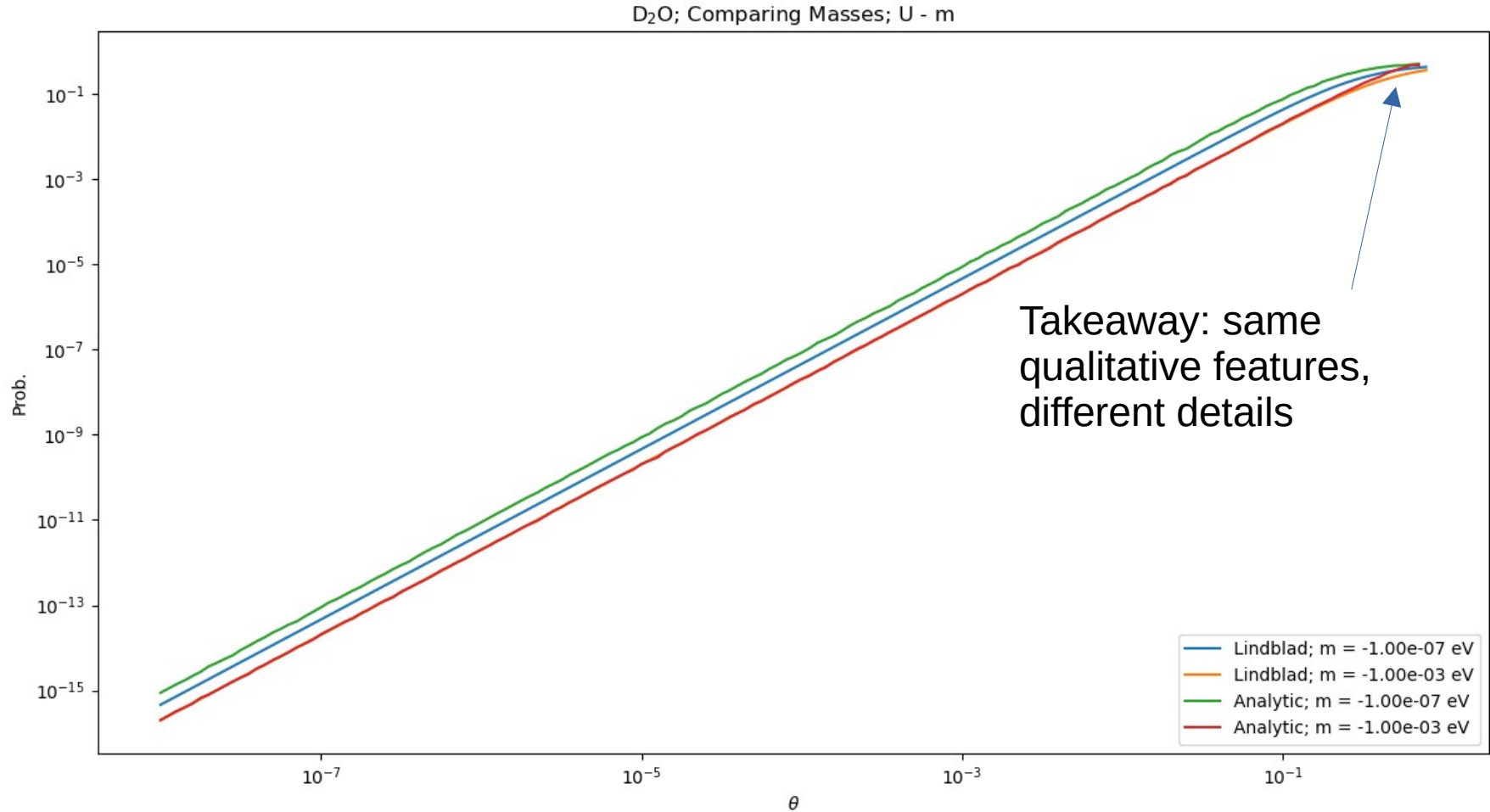


tan(45 deg) and  
10<sup>-5</sup> compared

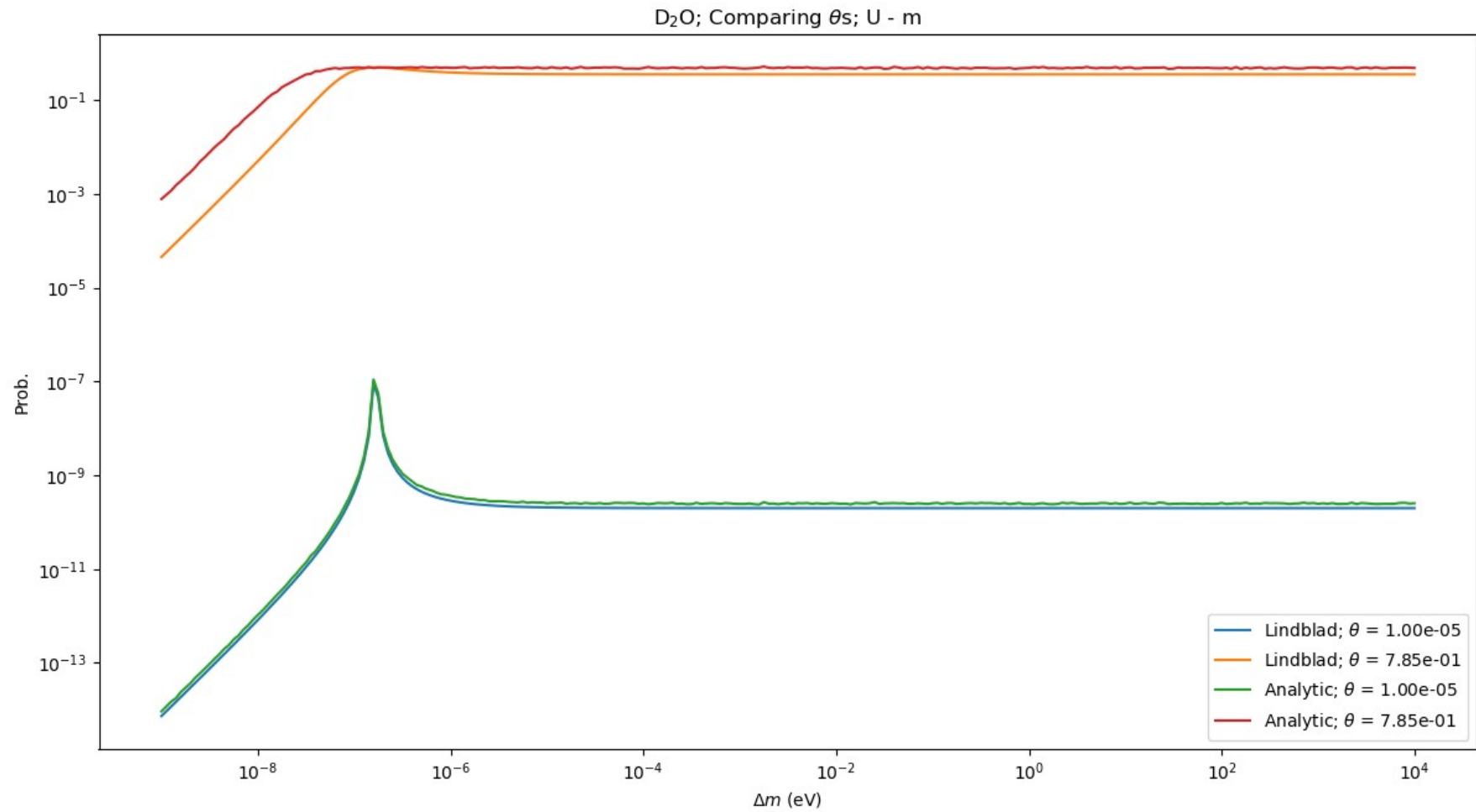
Varying “gamma”  
term in STEREO’s  
“p” was also  
investigated.

10<sup>-7</sup>, 10<sup>-3</sup> eV compared

# Comparing masses

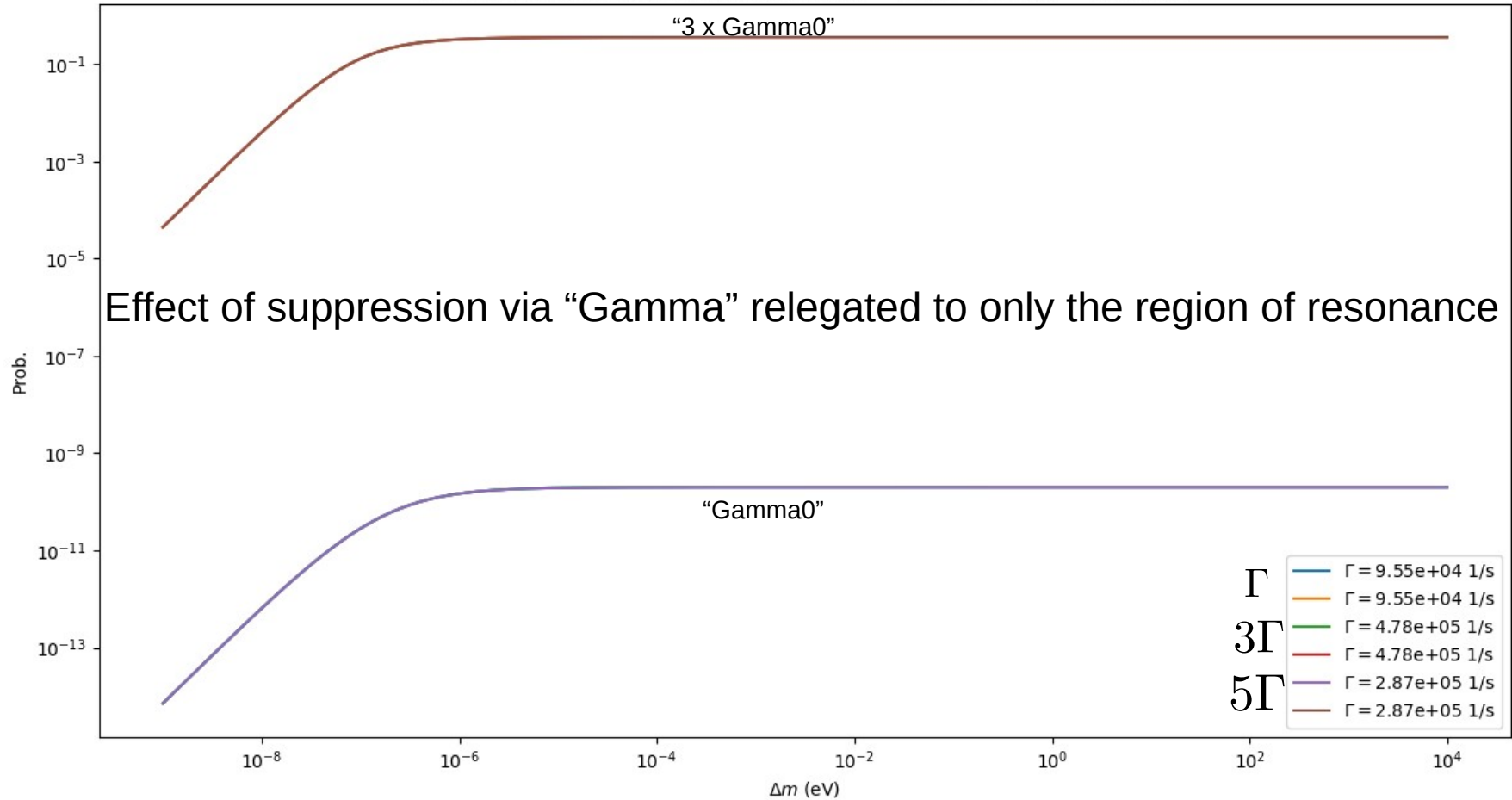


# Comparing angles



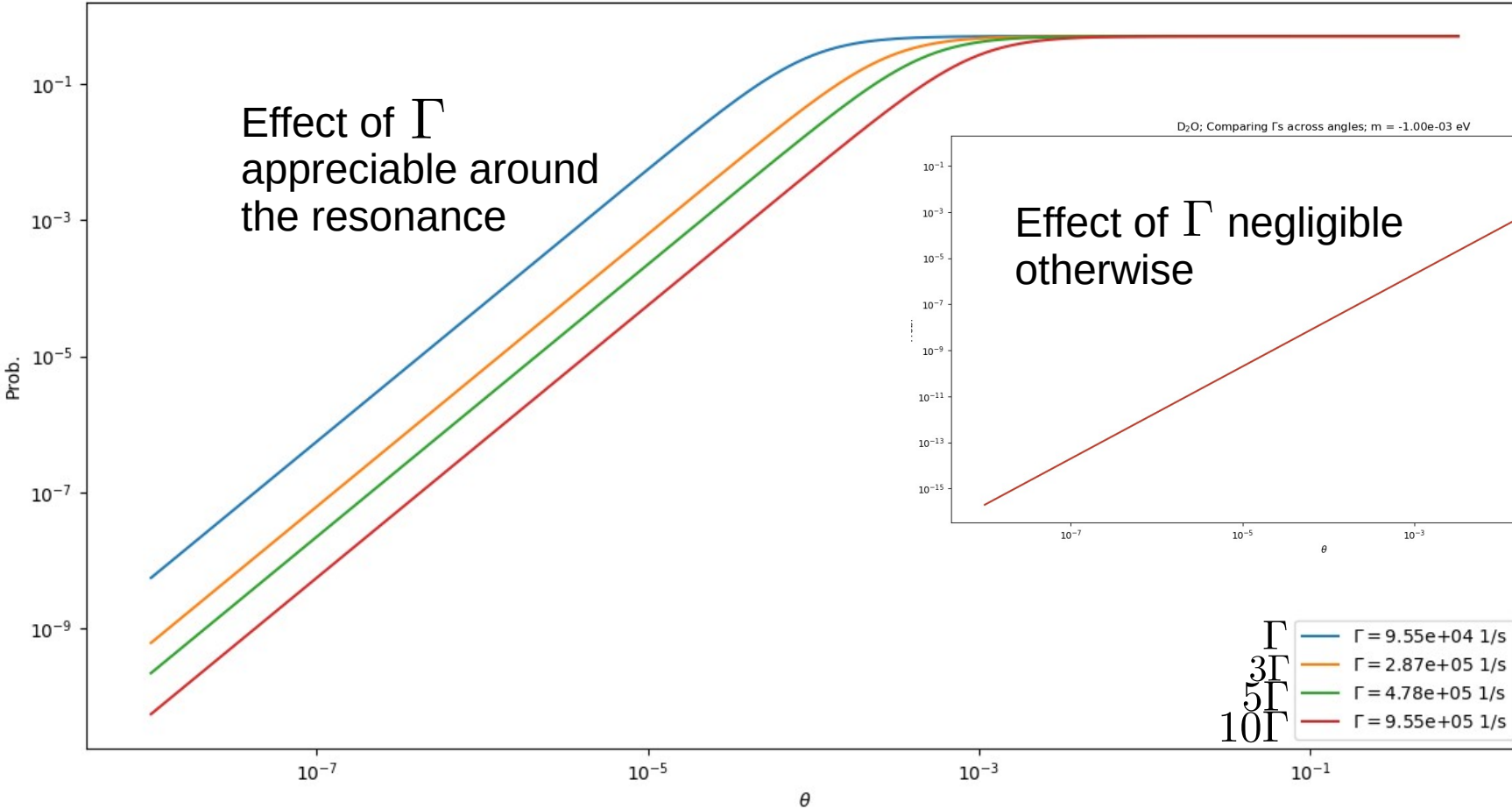
# $\Gamma$ across masses

D<sub>2</sub>O; Comparing  $\Gamma$ s across masses;  $\theta$  high = 7.85e-01,  $\theta$  low = 1.00e-05

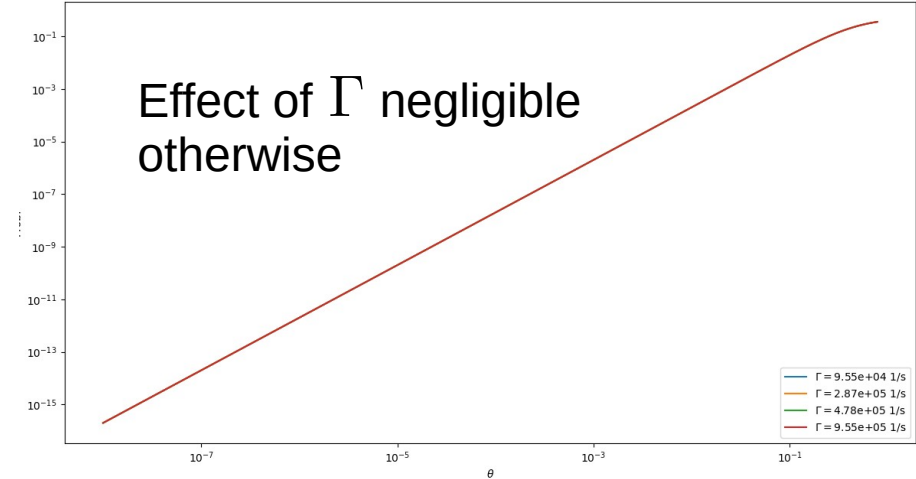


# $\Gamma$ across angles

D<sub>2</sub>O; Comparing  $\Gamma$ s across angles;  $m = -1.66e-07$  eV



D<sub>2</sub>O; Comparing  $\Gamma$ s across angles;  $m = -1.00e-03$  eV

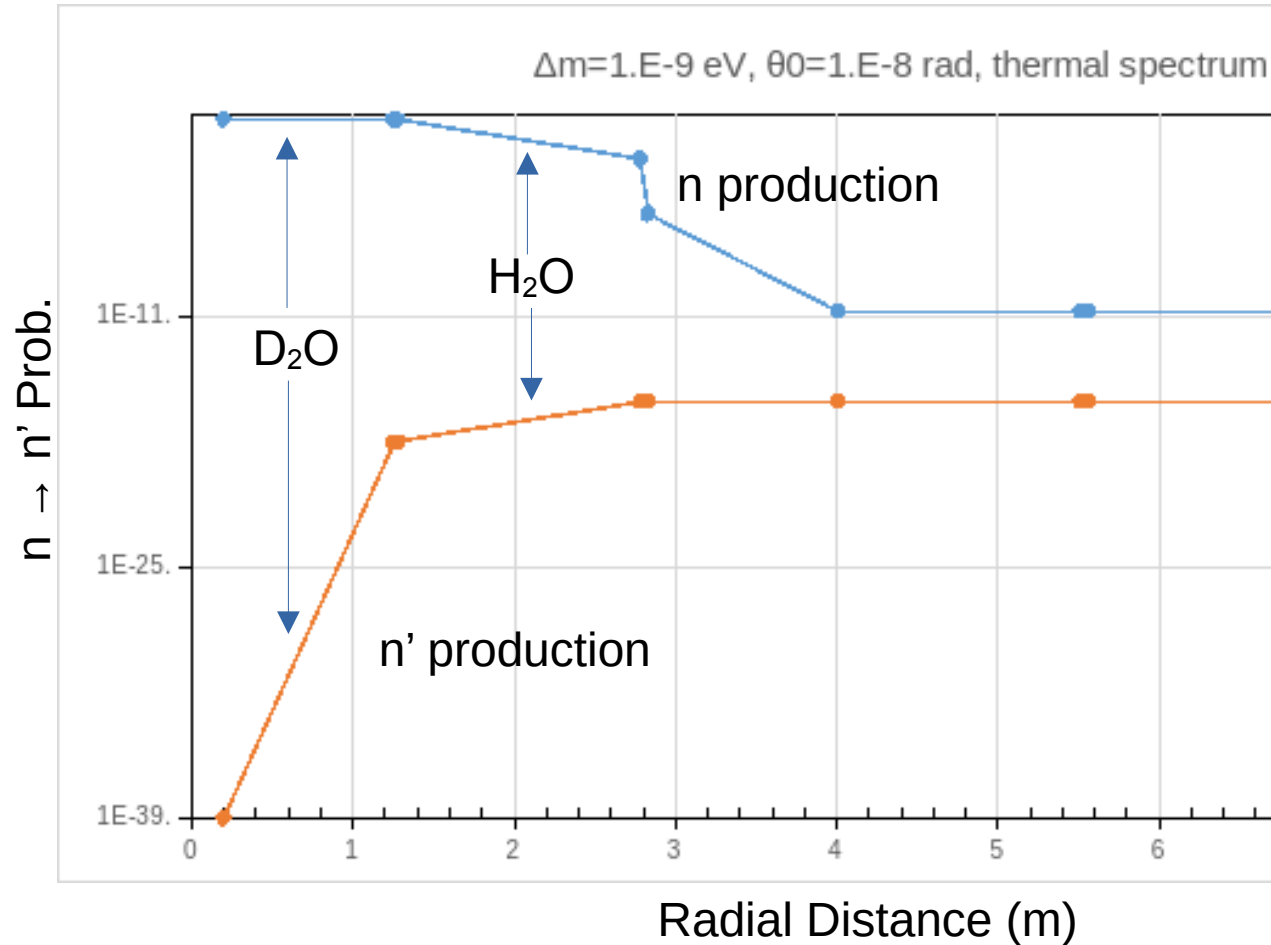


$\Gamma$   
 $3\Gamma$   
 $5\Gamma$   
 $10\Gamma$

$\Gamma$ (1/s)	Color
$9.55e+04$	Blue
$2.87e+05$	Orange
$4.78e+05$	Green
$9.55e+05$	Red

# Next Step: Computing $\Phi_h(\vec{r}_d)$ and comparing total numbers

- Code is currently being debugged for calculating values of  $n$  and  $n'$  flux as well as total  $n$  and  $n'$  from  $D_2O$
- $H_2O$  will also be computed
  - STEREO neglects effects of  $H_2O$ , our prior investigations suggest  $H_2O$ 's contribution is non-negligible.



End

# Flux Numbers - PRELIMINARY

```
=== STEREO HEAVY WATER ===
m = 1.0000000000000001E-009
a = 1.0000000000000000E-008
p = 7.3461885378738747E-021

Check that the results of the `compute_integral`
function are reasonable.

Phi_h(Rd) = 7.2031E+08
Total 4pi emit of mn = 9.6029E+15
Probability per init n = 3.6934E-03
rate of mirror neutrons= 4.6547E-07

=== ANALYTIC HEAVY WATER ===
m = 1.0000000000000001E-009
a = 1.0000000000000000E-008
p = 7.4193829999999996E-021

Check that the results of the `compute_integral`
function are reasonable.

Phi_h(Rd) = 7.8480E+08
Total 4pi emit of mn = 1.0463E+16
Probability per init n = 4.0241E-03
rate of mirror neutrons= 5.1220E-07
```