

Hidden Sector Experimental Challenges

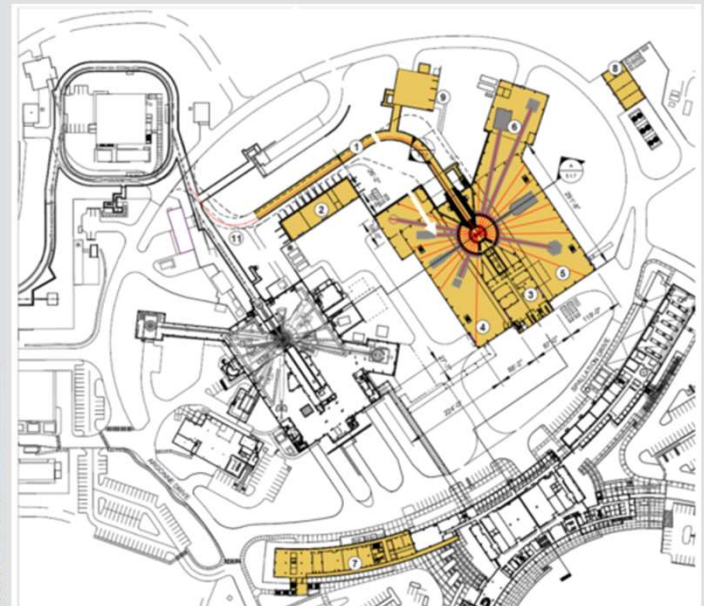
R. Tayloe, Indiana U.

Initial work on low-threshold, high-sensitivity detectors and facilities requirements for accelerator-DM searches and other physics at ORNL-SNS Second Target Station

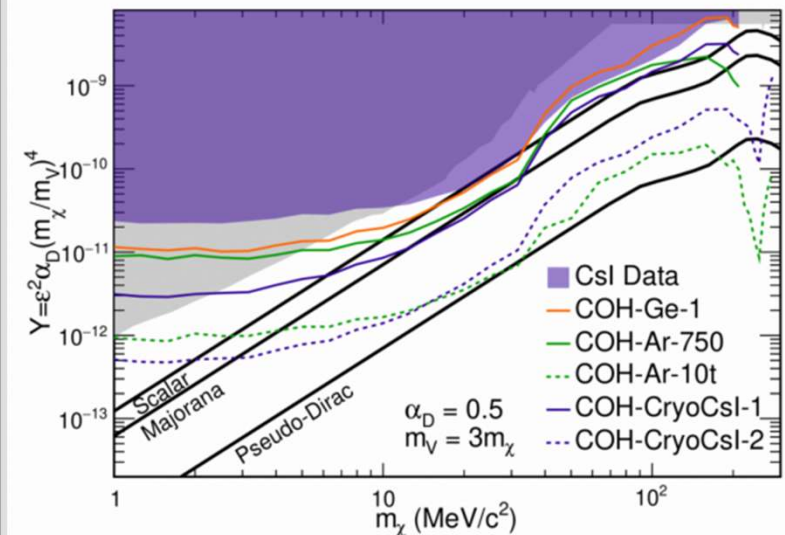
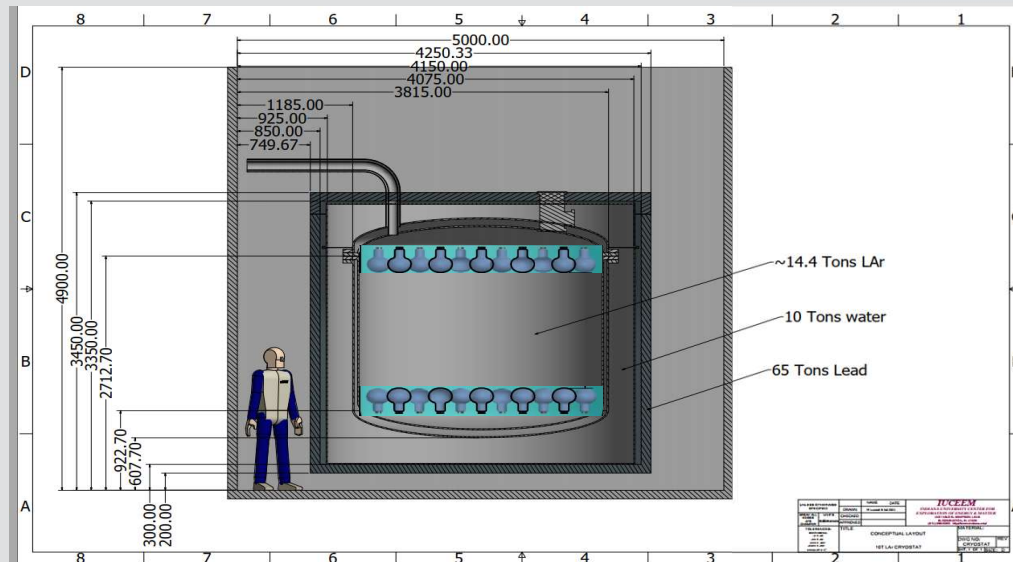
Neutrinos at ORNL Workshop

30 November 2021 to 2 December 2021

America/New_York timezone



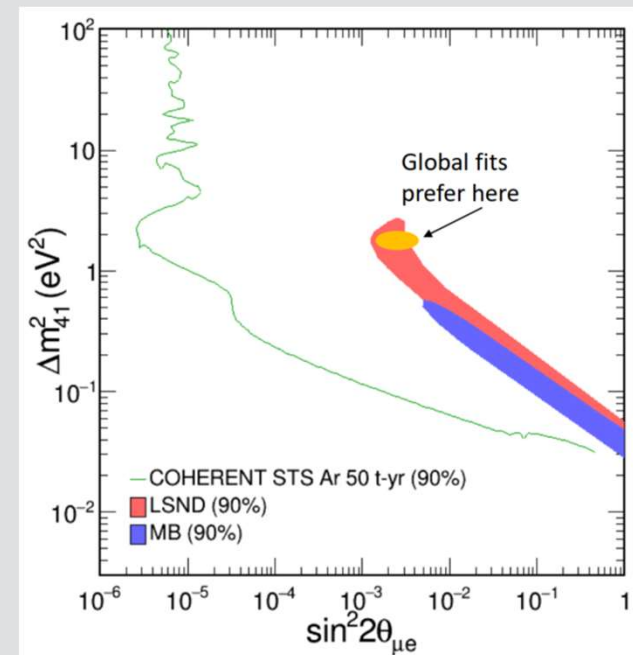
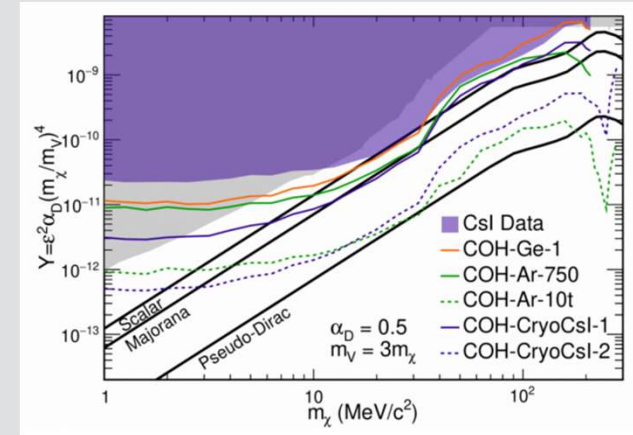
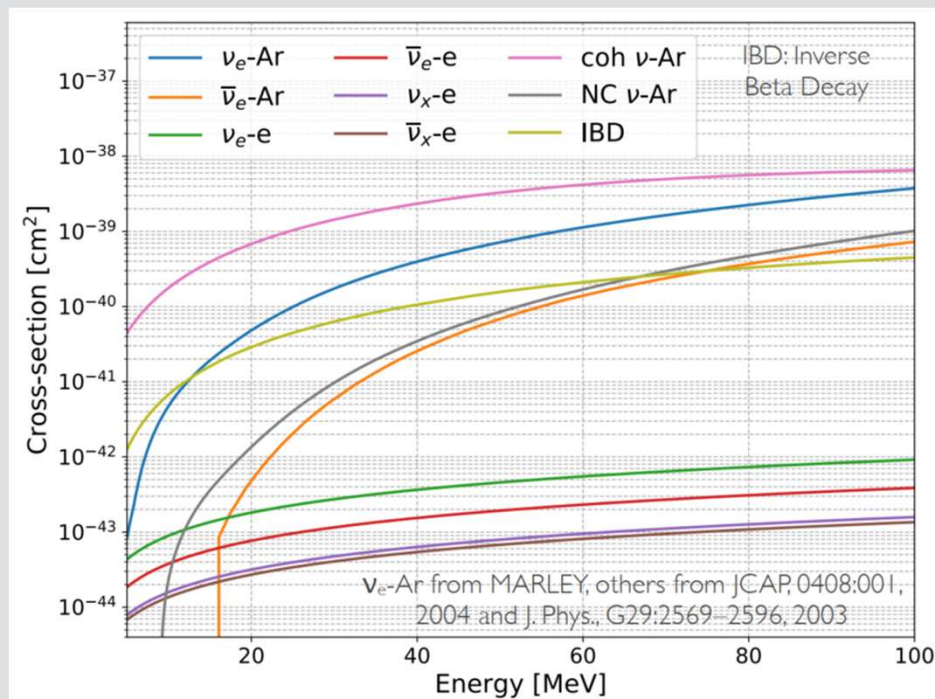
SNS Site with STS Facility



Motivation

Physics topics:

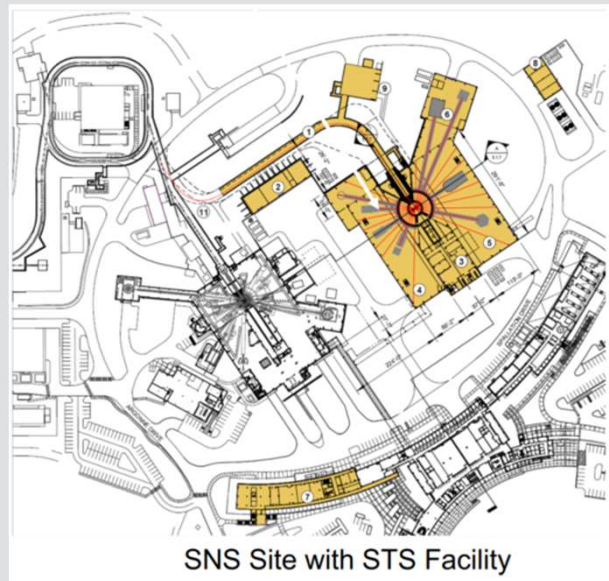
- accelerator-produced dark matter and...
- further (beyond Neutrino Alley) CEvNS studies of NSI, nuclear form factors, sterile ν searches
- higher-energy inelastic events for SN/DUNE



credits: Y-T Tsai, D. Persey, et al

Assumptions

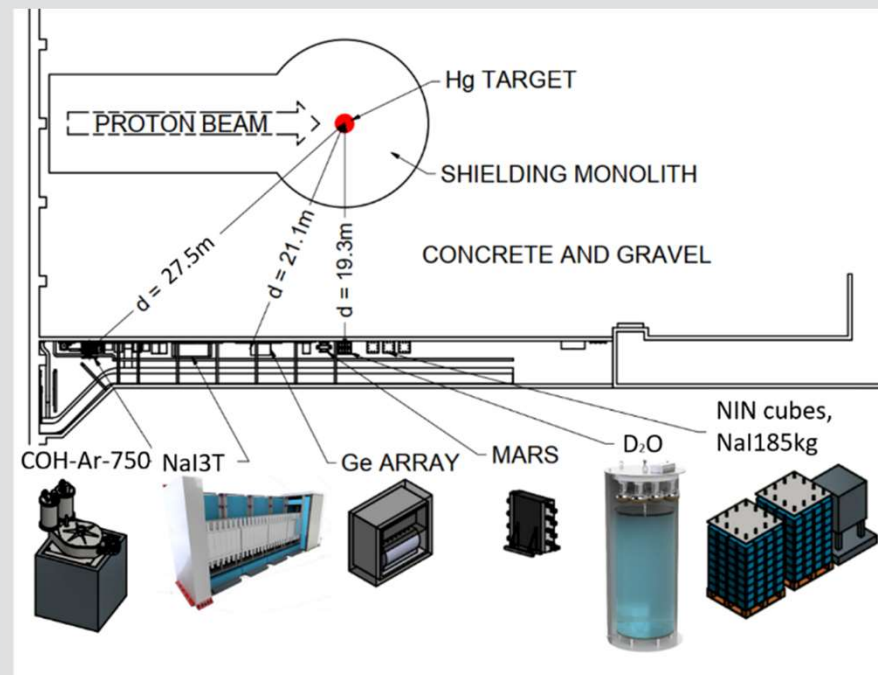
- Considering 10-ton scale detectors, building on ongoing work including ton-scale detectors from COHERENT in Neutrino Alley.
-
- 10-ton scale detectors require larger dedicated halls and therefore the SNS-STS with timescale ~ 5 yrs



SNS Site with STS Facility

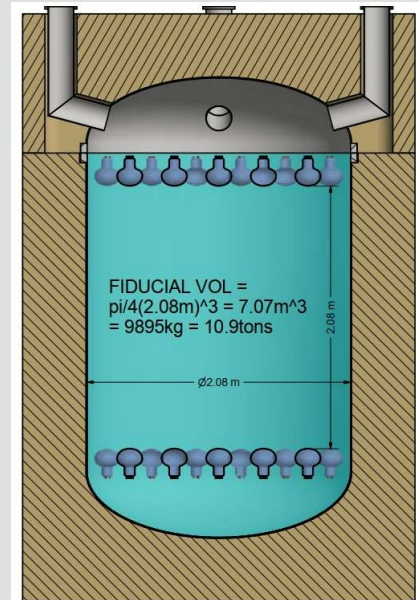
COHERENT CEvNS Detector Status and Farther Future

Nuclear Target	Technology	Mass (kg)	Distance from source (m)	Recoil threshold (keVr)	Data-taking start date	Future
CsI[Na]	Scintillating crystal	14.6	19.3	6.5	9/2015	Decommissioned
Ge	HPGe PPC	18	22	<few	2022	Funded by NSF MRI, in progress
LAr	Single-phase	24	27.5	20	12/2016, upgraded summer 2017	Expansion to 750 kg scale
Nal[Tl]	Scintillating crystal	185*/3388	25	13	2022, high-threshold deployment summer 2016	Expansion to 3.3 tonne , up to 9 tonnes

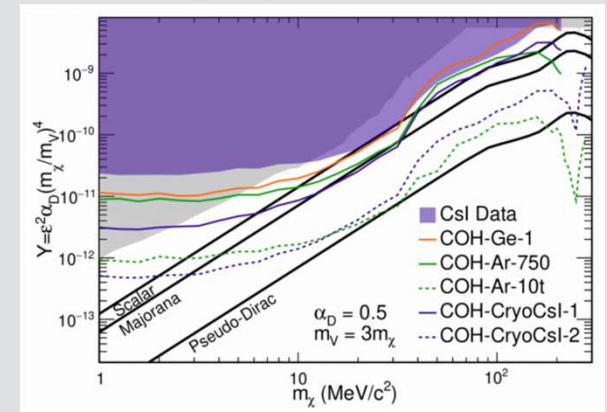


Detectors/Facilities for STS neutrino program

- Detector requirements
 - large mass for sufficient event rates
 - low-energy threshold for DM and CEvNS physics
 - high-spatial resolution for inelastic physics
 - well shielded
- Facility/building/room requirements
 - space
 - utilities for cryogenics, power needs
 - low backgrounds:
 - beam-related neutrons!
 - environmental (concrete etc)
 - cosmics
- Most work, so far, on design and facilities requirements for a ~10ton LAr detectors, since:
 - scales to 10ton naturally
 - overlap with DUNE etal, physics and design experience
 - drives the facility requirements for size and utilities (eg: cryogenics)
- Other detectors possible in these facilities (eg Cryo-CsI)



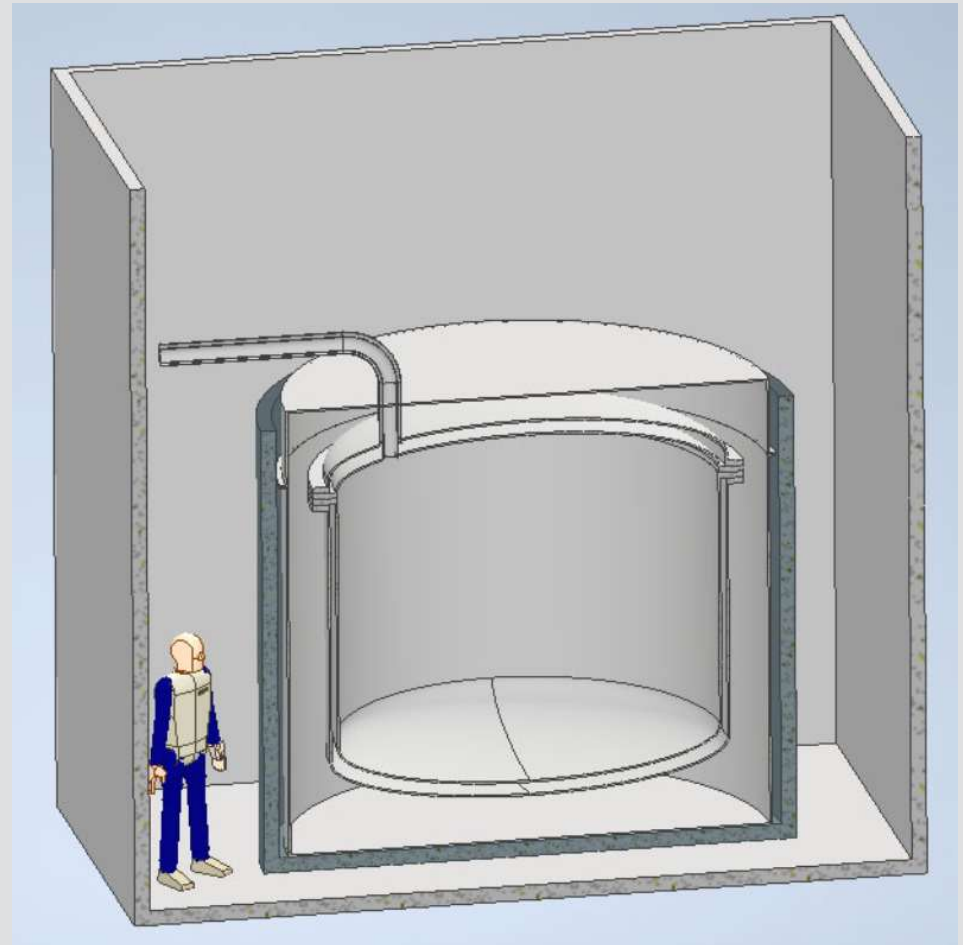
early concept
for 10ton LAr



Concept for ~10ton LAr detector for scintillation or TPC detector

Cryostat/room specs:

- 4.9m x 5m x 5m room in SNS STS basement
- 2.5m dia x 2m high cryostat, Design from cryofab example from SLAC.
- detector shielding: outside of cryostat, 10" water, 4"Pb.
- For cryogenic plumbing and related, require adjacent 5mx5m room, likely can reduce that.
- Similiar cryostat works for scintillator or TPC detector.
- For best room shielding, suggesting a "shielded cave" approach. Inner detector assembled in high bay, forked into place. Modular H2O tank, Pb assembled in place (eg: no high-bay crane access here)



Shielding

From Neutrino Alley work, we have developed shielding requirements (for low-energy measurements, reduction of beam-related neutrons most crucial)

General Shielding Requirements

- 20 meters of line of sight shielding between target and neutrino instrument (45-50 interaction lengths, HD concrete reduces distance)
- 3.5 meters of overburden for cosmic ray shielding (could include neutron instrument shielding above)
- 3 meters hermetic shielding around instrument in directions open to beam (learning from FTS BL1 instrument trenches)

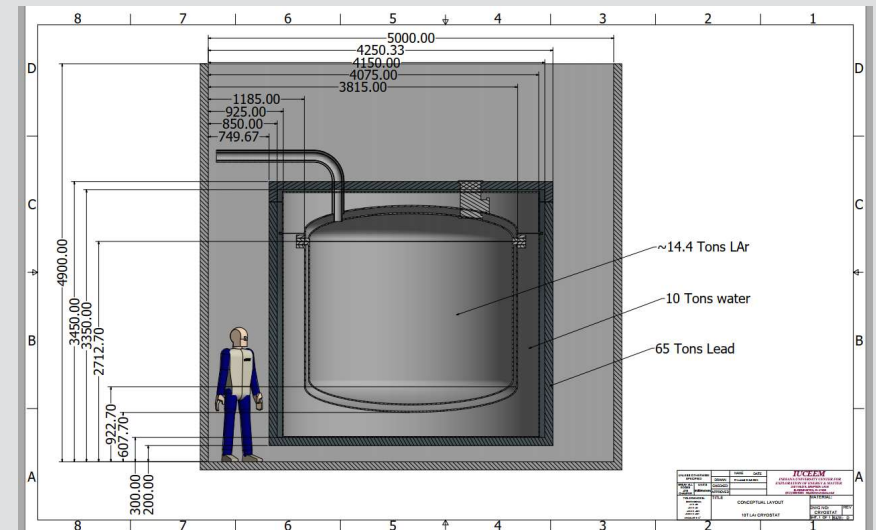
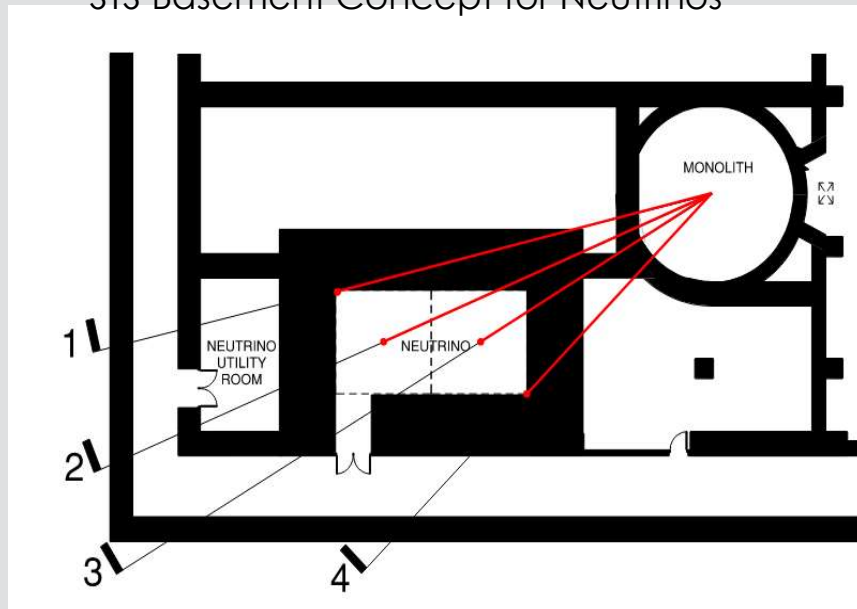
This shielding can be comprised of a combination of structural materials, target monolith, etc. and dedicated shielding for neutrinos instruments.

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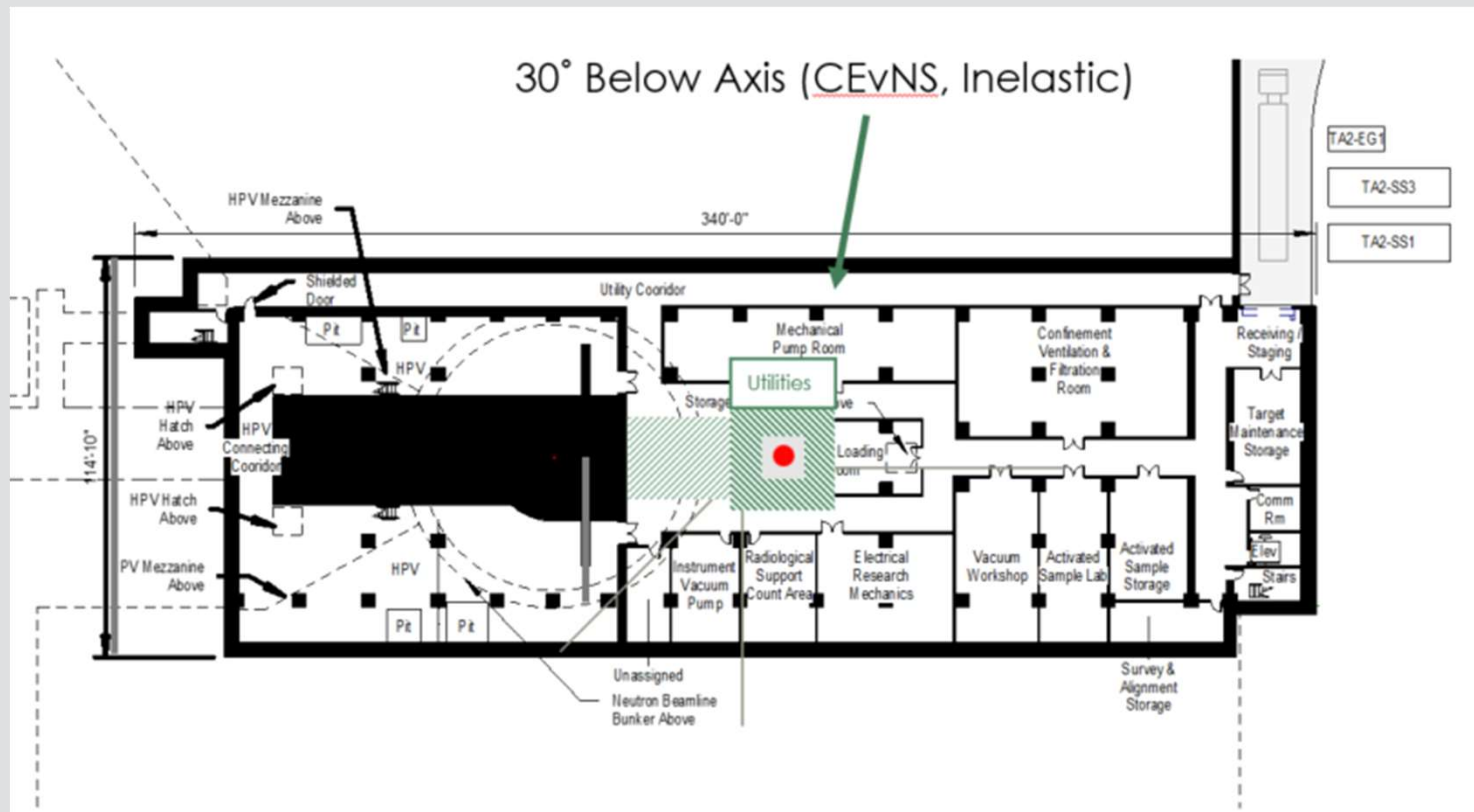
Recent design concept

STS Basement Concept for Neutrinos



STS Basement concept:

Recall that forward production of heavier DM suggests a small-angle hall.
A more challenging specification.... work in progress.



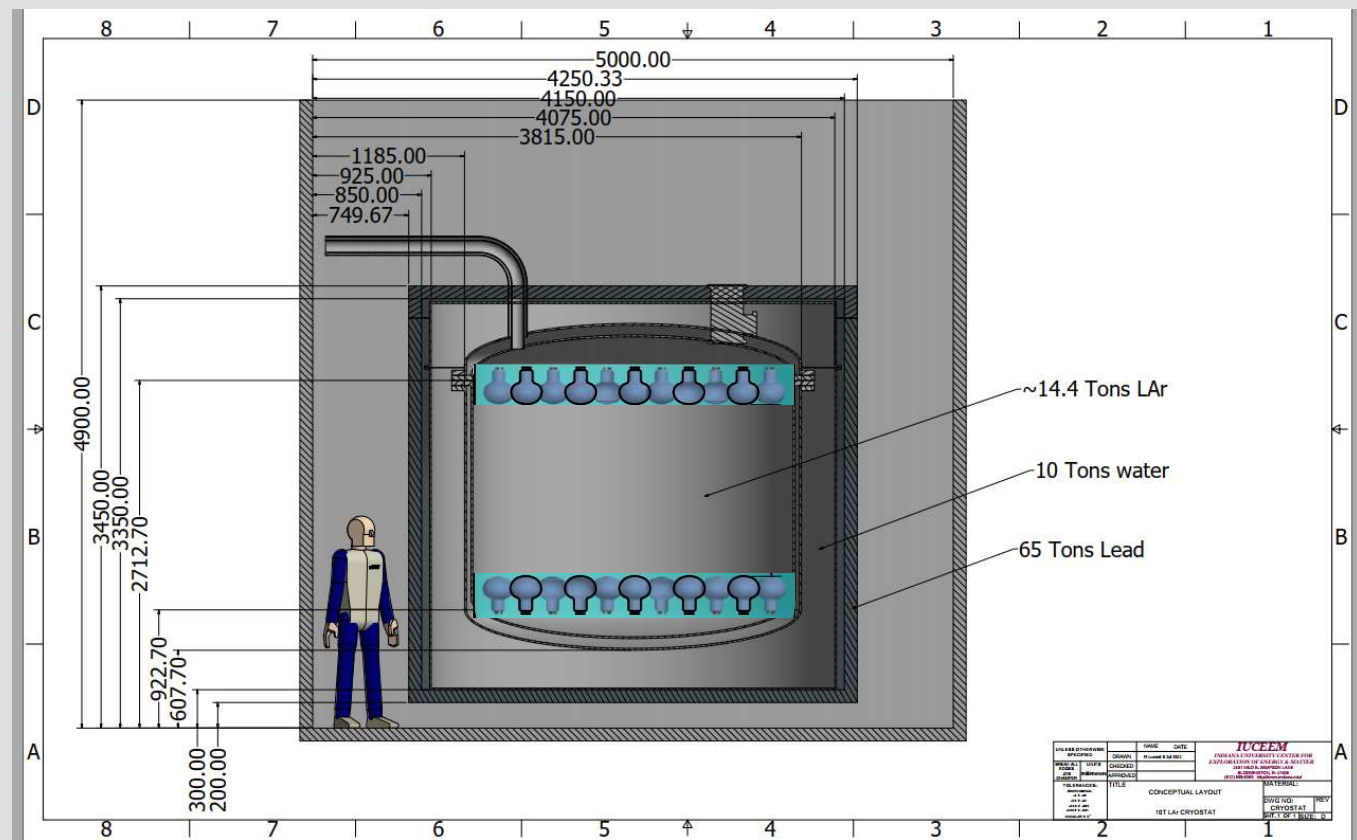
Detector details for STS neutrino program

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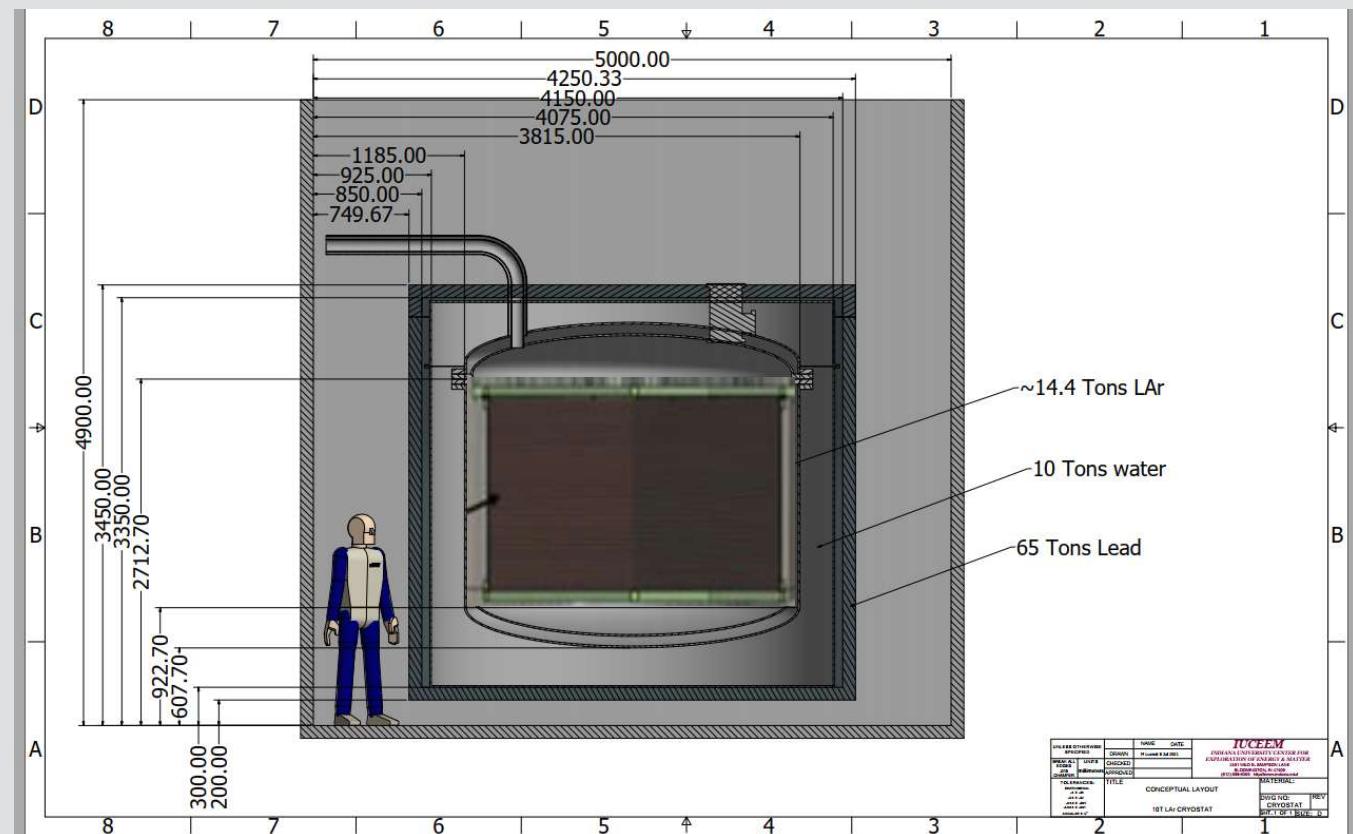
Concept for a ~10ton LAr **scintillation** detector



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Concept for a ~10ton LAr TPC detector



more on TPC design
from Y-T Tsai talk

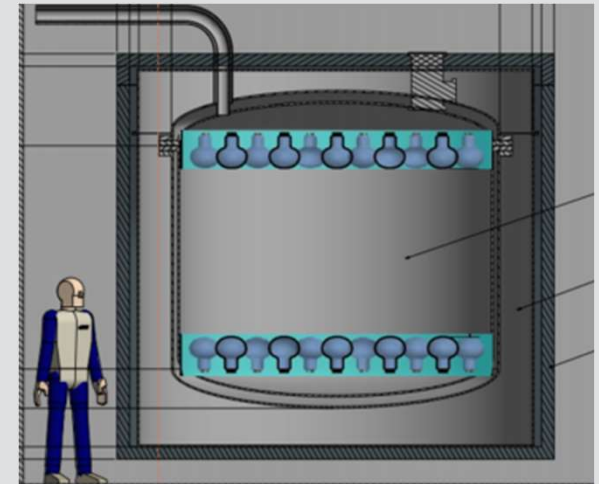
Detector details for STS neutrino program

Detector technology, possible R&D, and further work

- Underground argon: All cryostats wait for UAr from Urania(Kinder-Morgan) in Colorado. Other avenues? ([2018 Workshop writeup arXiv:1901.10108](#))
- Xe-doping: May eliminate need WLS-coatings while maintaining ER/NR PID (eg: [inspirehep.net/literature/1737812](#))
- SiPM implementation: optimize channel count/segmentation for position reconstruction in CEvNS and hi-energy detectors, heat load a concern in cryogenic detectors, onboard digitization.
- In-situ sources such as ^{83m}Kr for calibration (eg: <https://inspirehep.net/literature/1824411>) .
- In-situ nuclear recoil calibration (eg: <https://inspirehep.net/literature/1811232>)
- cryogenic CsI (eg: <https://doi.org/10.1140/epjc/s10052-020-8111-7>)
- dual-phase and/or other noble gases should be considered (see J. Daughetee talk tomorrow)

Summary

- Opportunities for neutrino and DM physics at SNS-STS exist
- program will yield high-impact physics for modest investment
- Facilities and detector planning underway
- More participation and ideas welcomed



Cryogenic CsI detector

