

Agenda of nn' Collaboration meeting September 18, 2026 at ORNL

- | | |
|---|--|
| 1. Yuri Kamyshev | Overview of finished measurements and data for nTMM-2 experiment |
| 2. Alina Moore | Analysis of nTMM-1 2025 data |
| 3. Carolyn Haviland | Analysis of nTMM-1 2025 data |
| 4. Mubi Khan | Magnetic calculations for mu-prototype and more |
| 5. Linus Persson | Magnetic field reconstruction in nTMM-2 |
| 6. Evan Michael | Pressure gauge image recognition |
| 7. Yuri Kamyshev | Plans and projects for 2026 data analysis |
| 8. Discussion on first PRL nTMM publication preparation | |
| 9. Discussion of NIM publication for mu-prototype studies | |
| 10. Yuri Kamyshev 25' | Measurements with mu prototype |

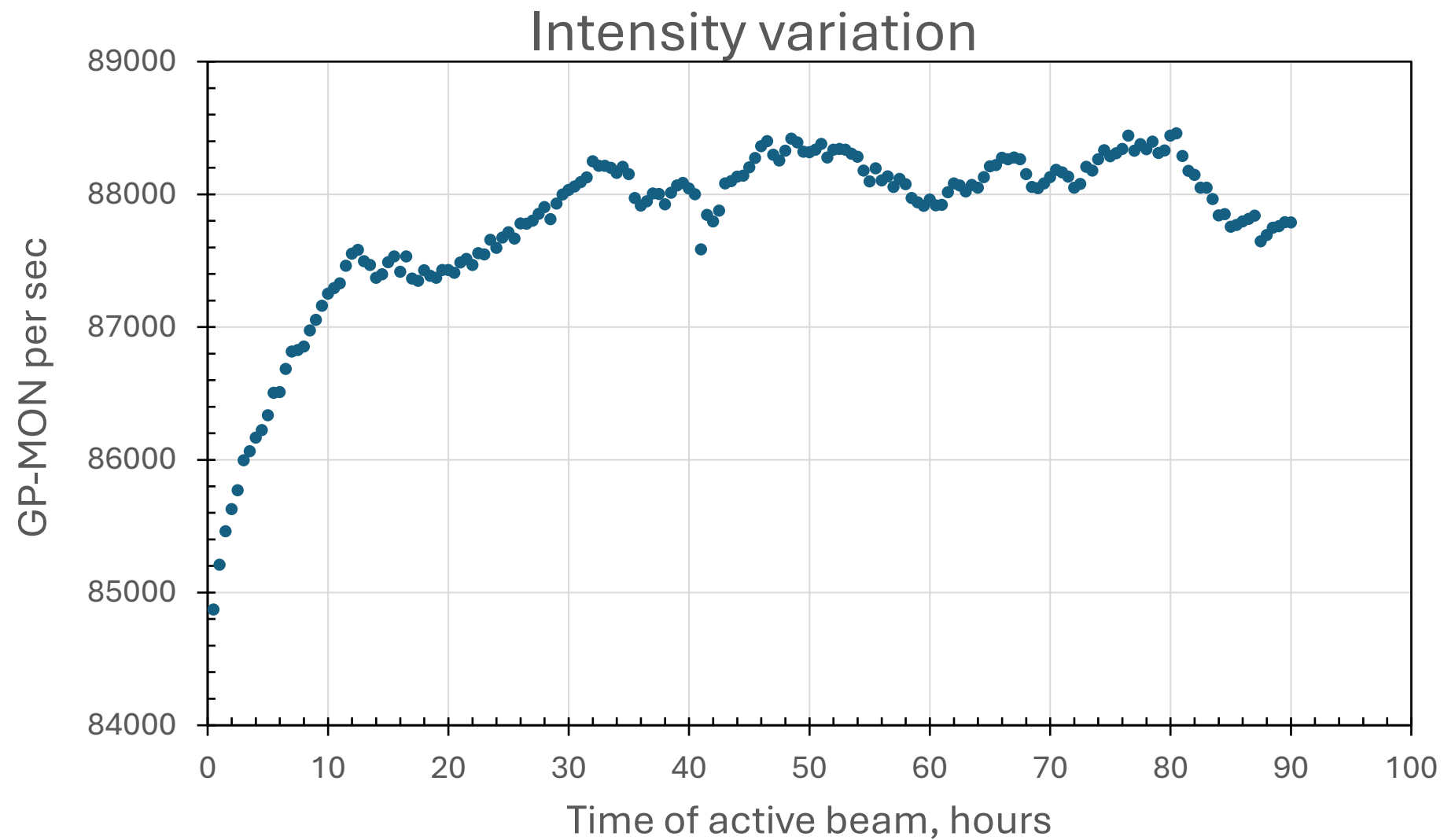
All speakers: please send your slides to kamyshev@utk.edu for INDICO posting

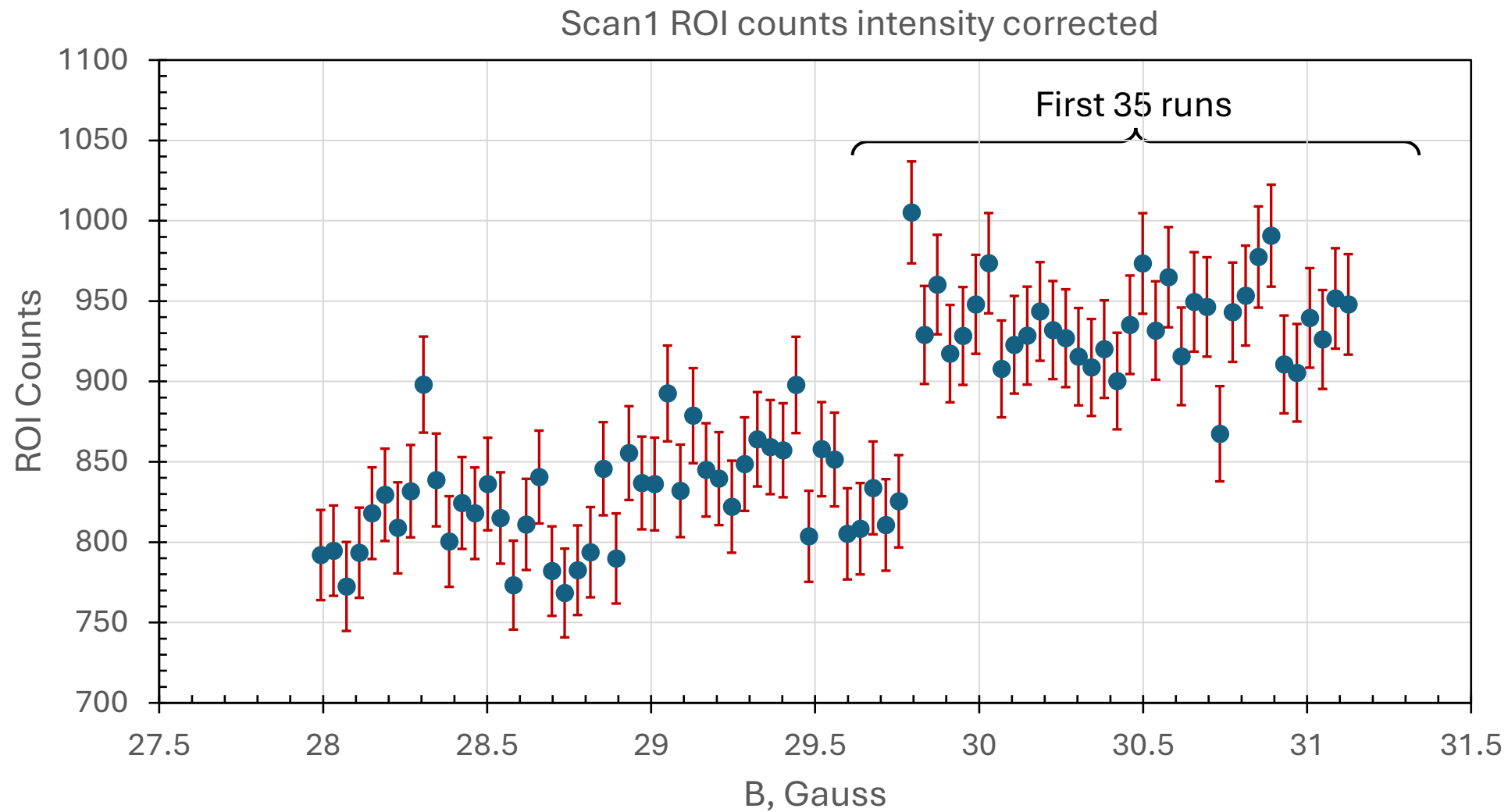
Overview of finished measurements and data for nTMM-2 experiment

Yuri Kamyshev / UTK

OVERVIEW

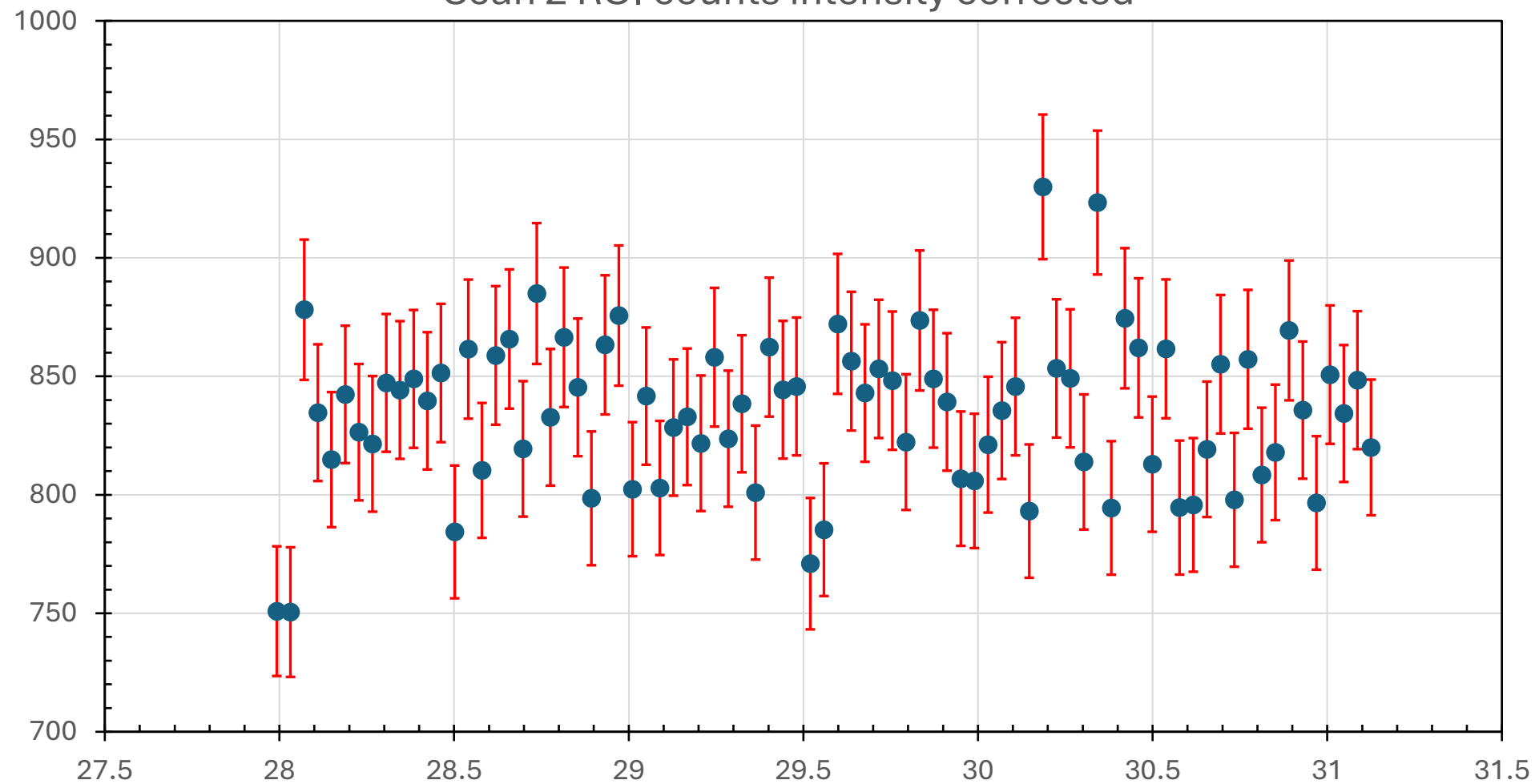
- All data runs are 30 min
- 1-st run started on Thursday Sept. 10 at 16:56
- Last run ended on Monday Sept. 14 at 12:33
- First sequential scan # 170127-170207
- Second random scan # 170212-170292
- Third central scan # 170293-170310
- We measured for 90 hours out of 96 allocated
- There are also few runs with “shutter closed”
- ICE magnet of CG-4 instrument was OFF



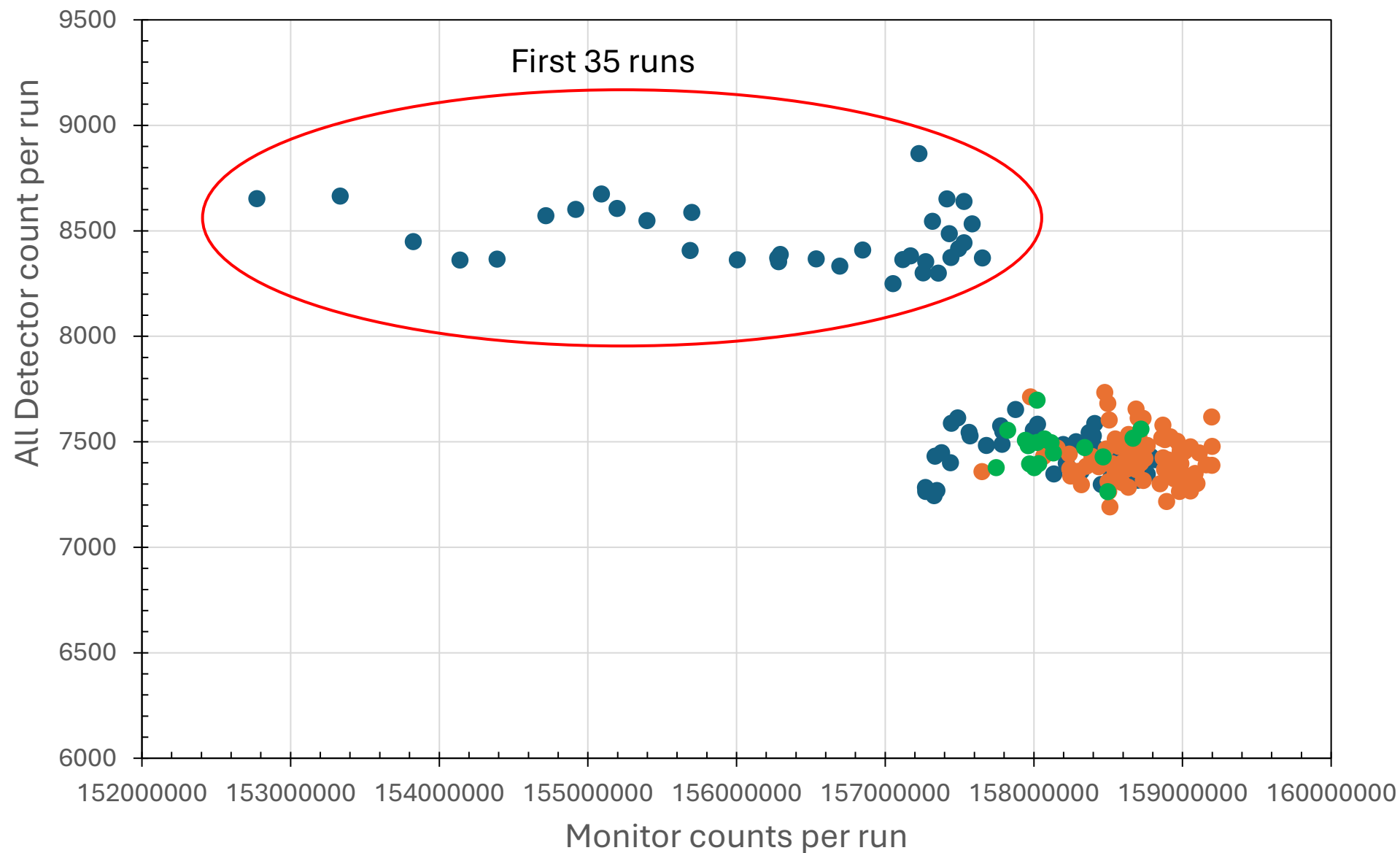


Detector count $\lesssim 900 / 1800 \text{ s} = 0.5 / \text{s}$; error is $\sim \pm 30$

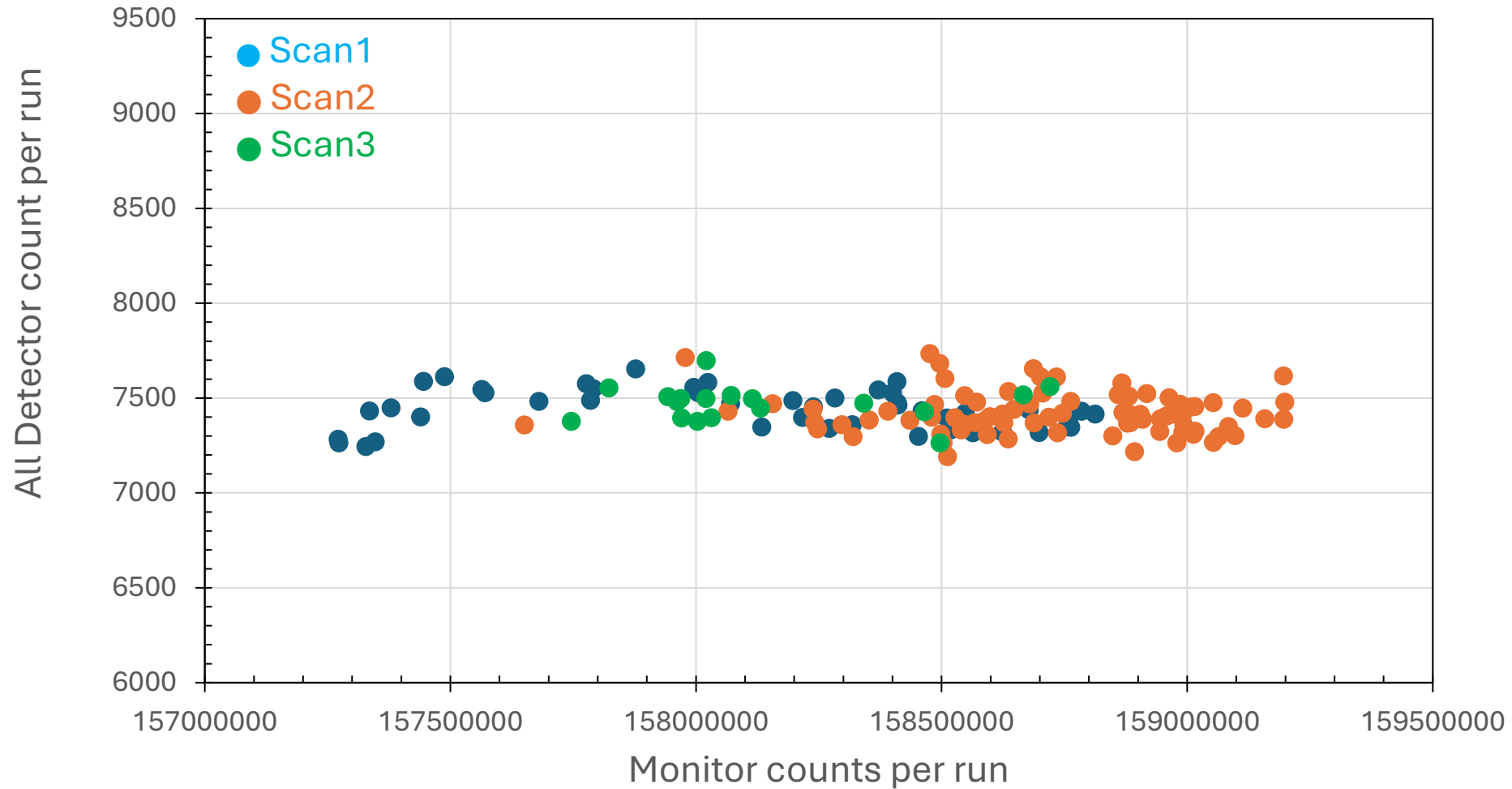
Scan 2 ROI counts intensity corrected



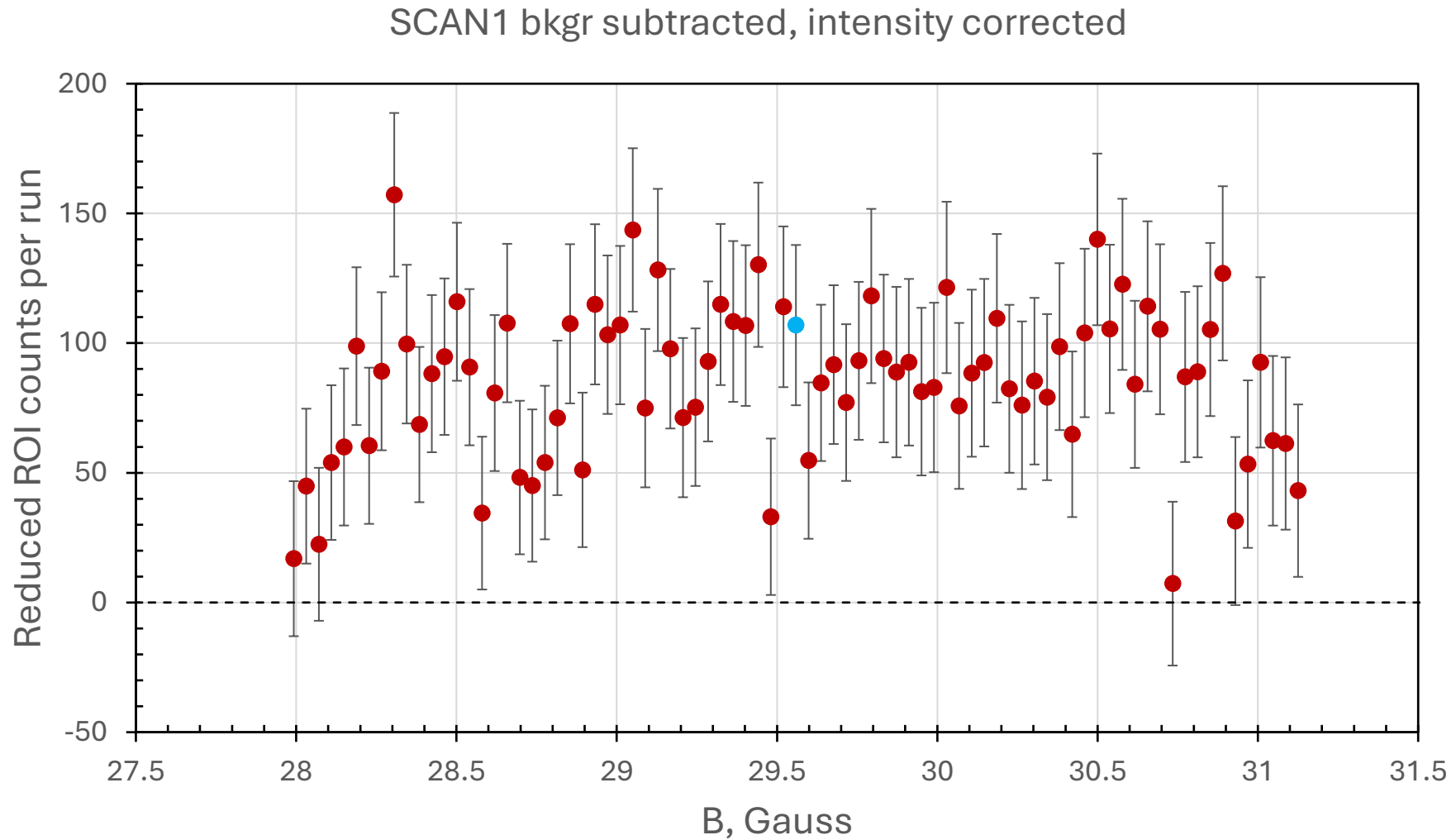
Detector counts vs Monitor counts



Without first 35 runs. Is detector counts proportional to monitor?

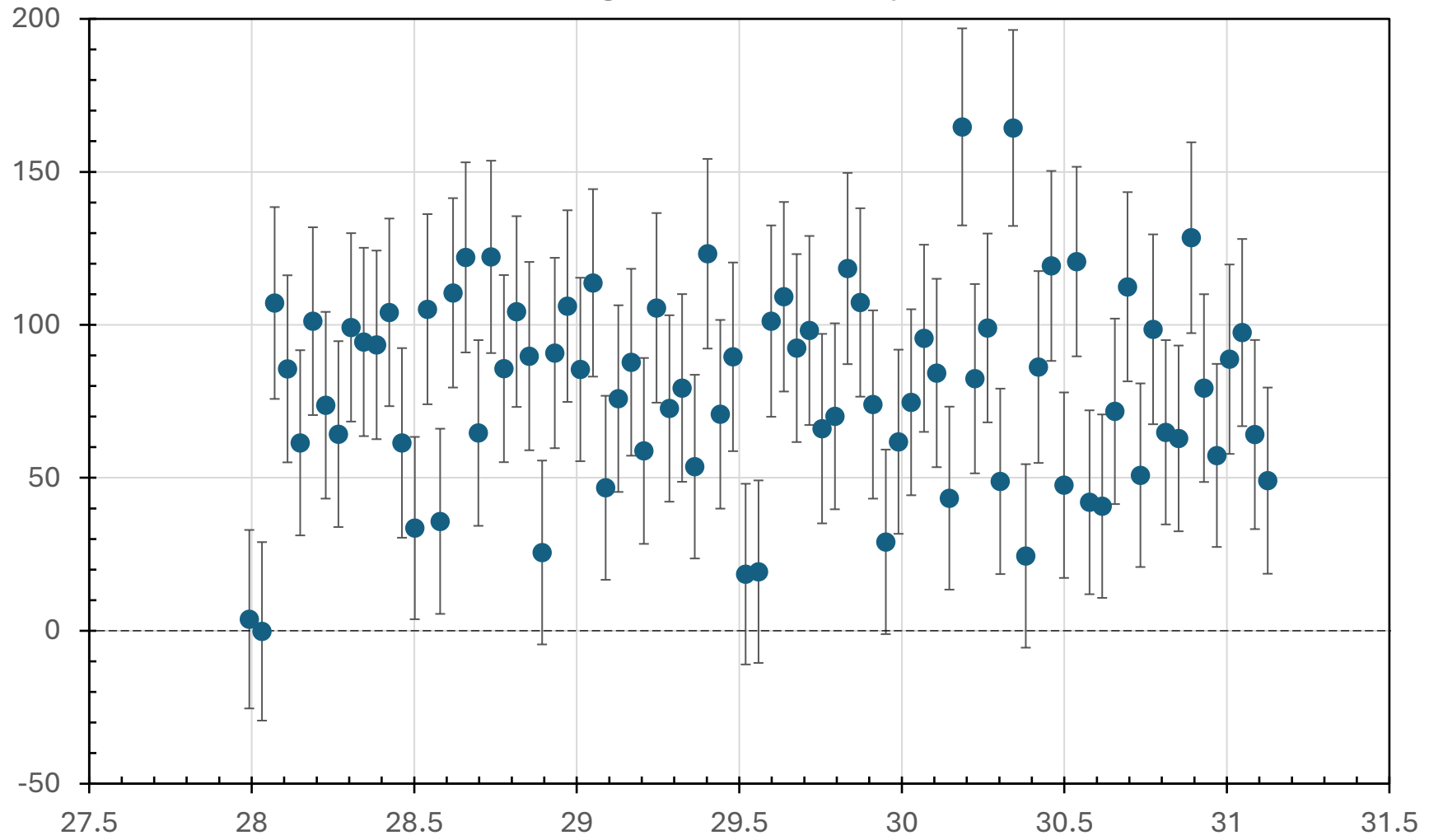


Detector counts are independent on Monitor: **detector background is not a GP-SANS beam background.**
With B-Poly and Cd we absorb GP-SANS beam but we have no control of the background in the hall.

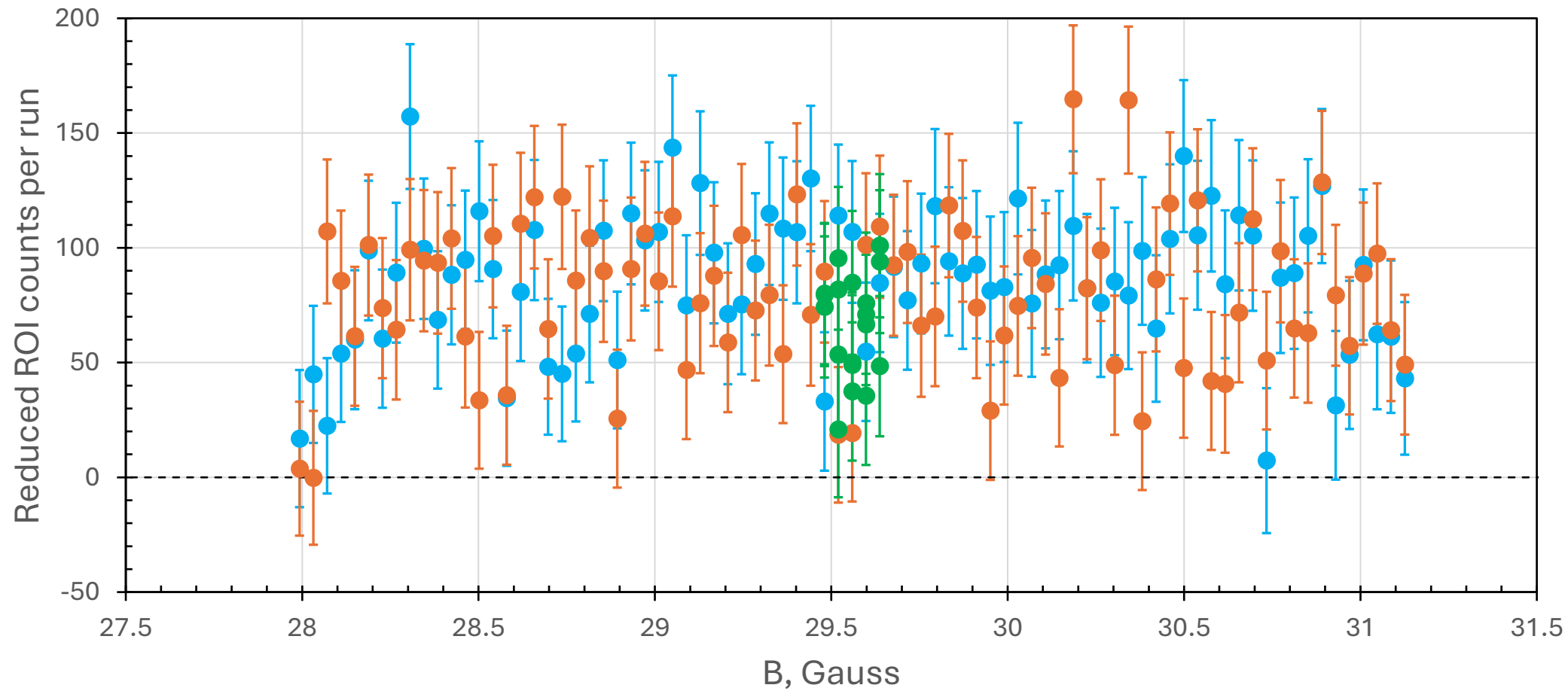


First 35 points were due to background increase uniformly in the whole detector.
Blue point is position of expected nTMM peak is $V_F = \mu B$ ($\Delta m = 0, B' = 0$)

SCAN 2 Bkgr subtracted, intensity corrected

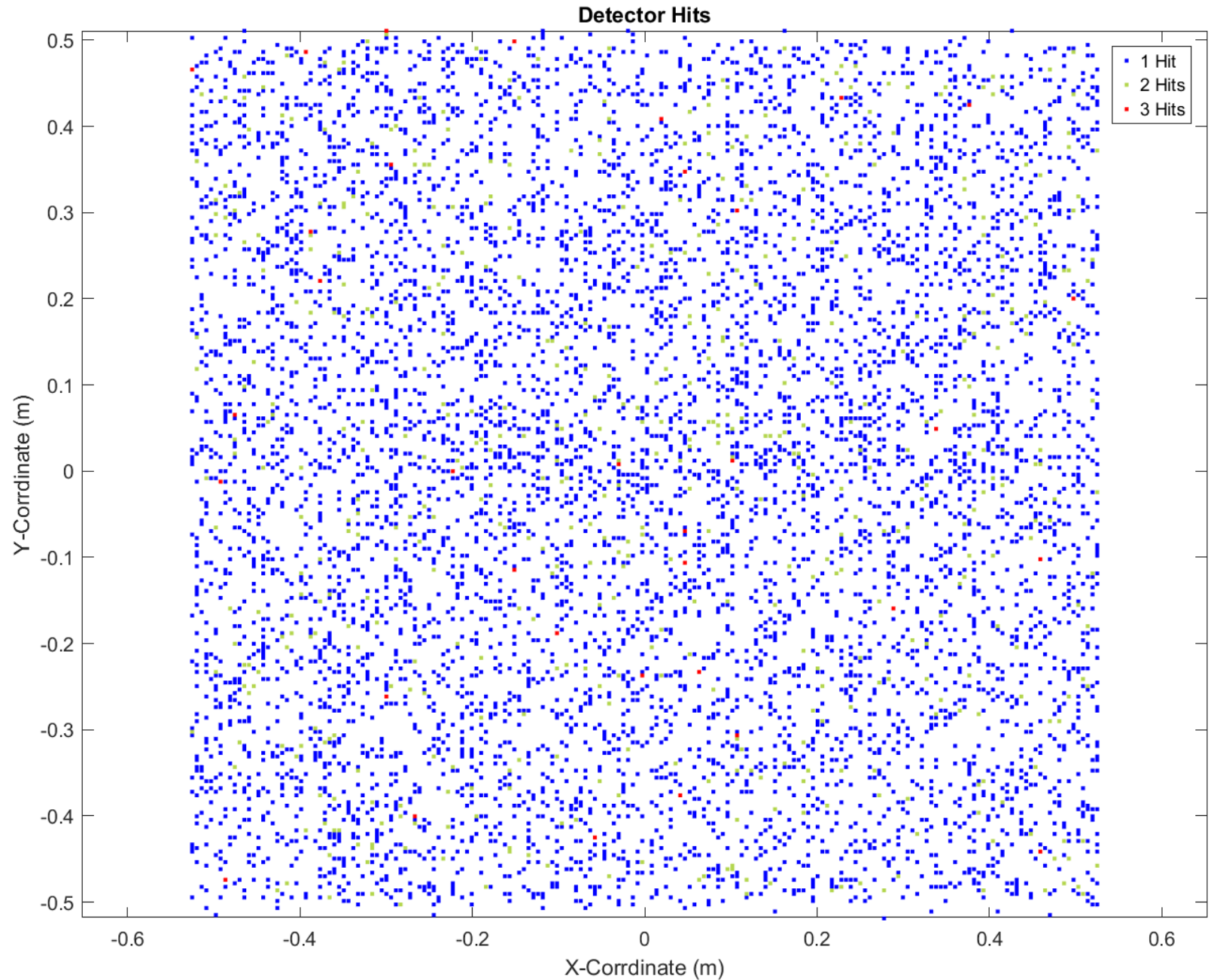


SCAN1,SCAN2,SCAN3 bkgr subtracted, intensity corrected



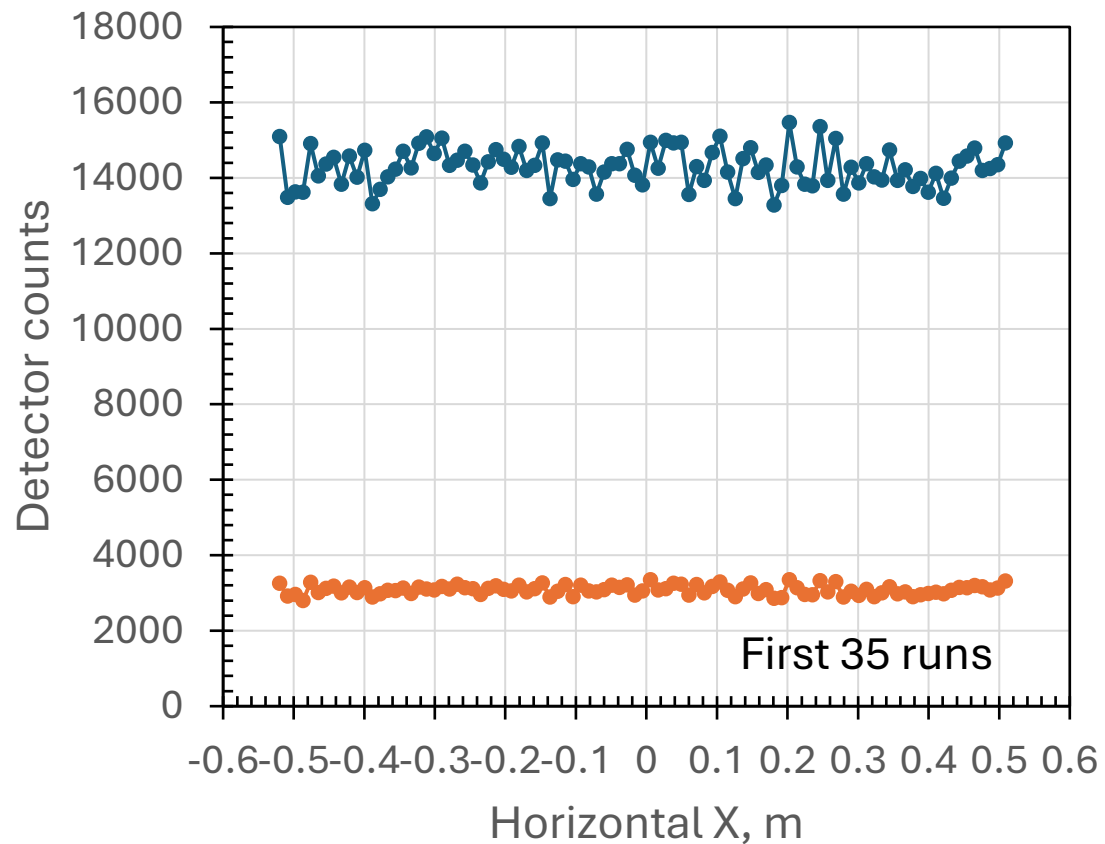
No central peak
in the detector

You need to have a sharp eye
of Micheal Lumsdaine to find
a horizontal band here

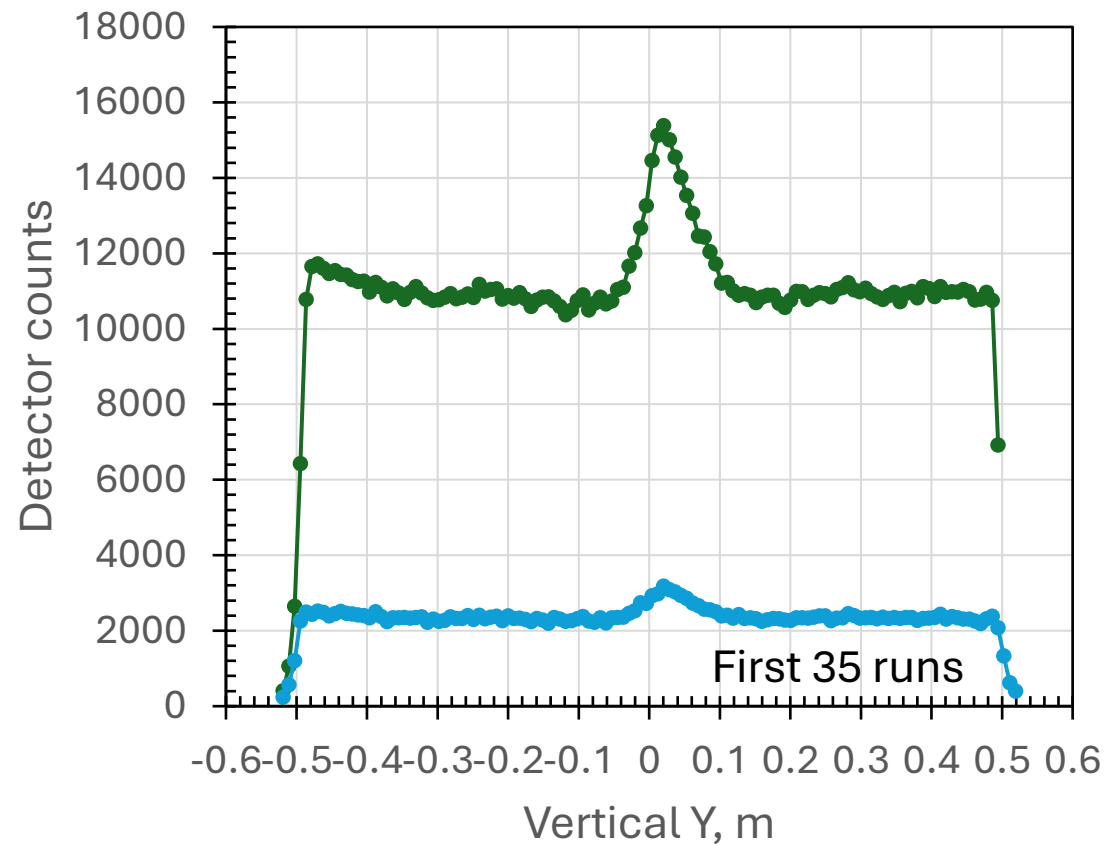


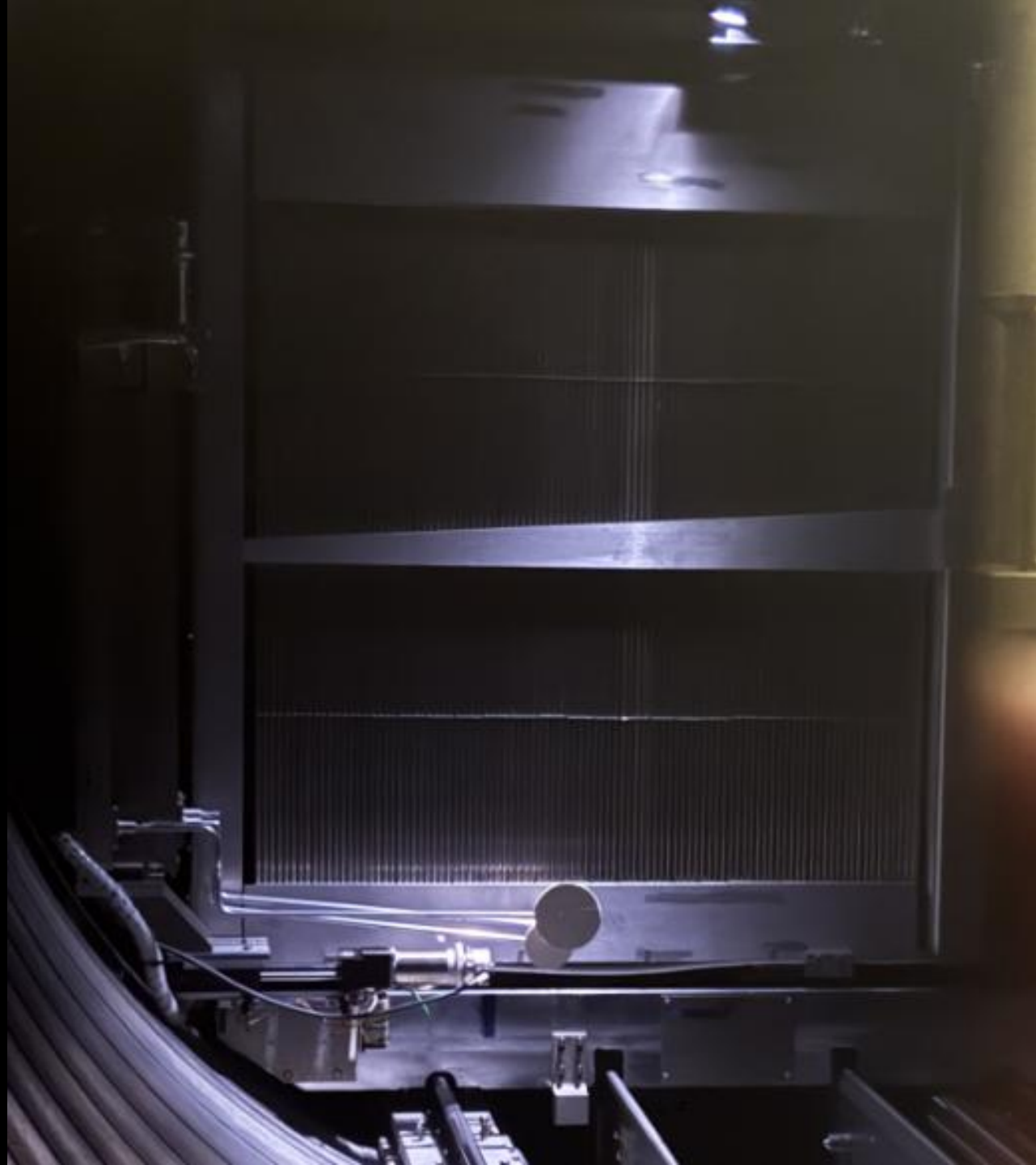
X & Y detector distributions for all 180 runs

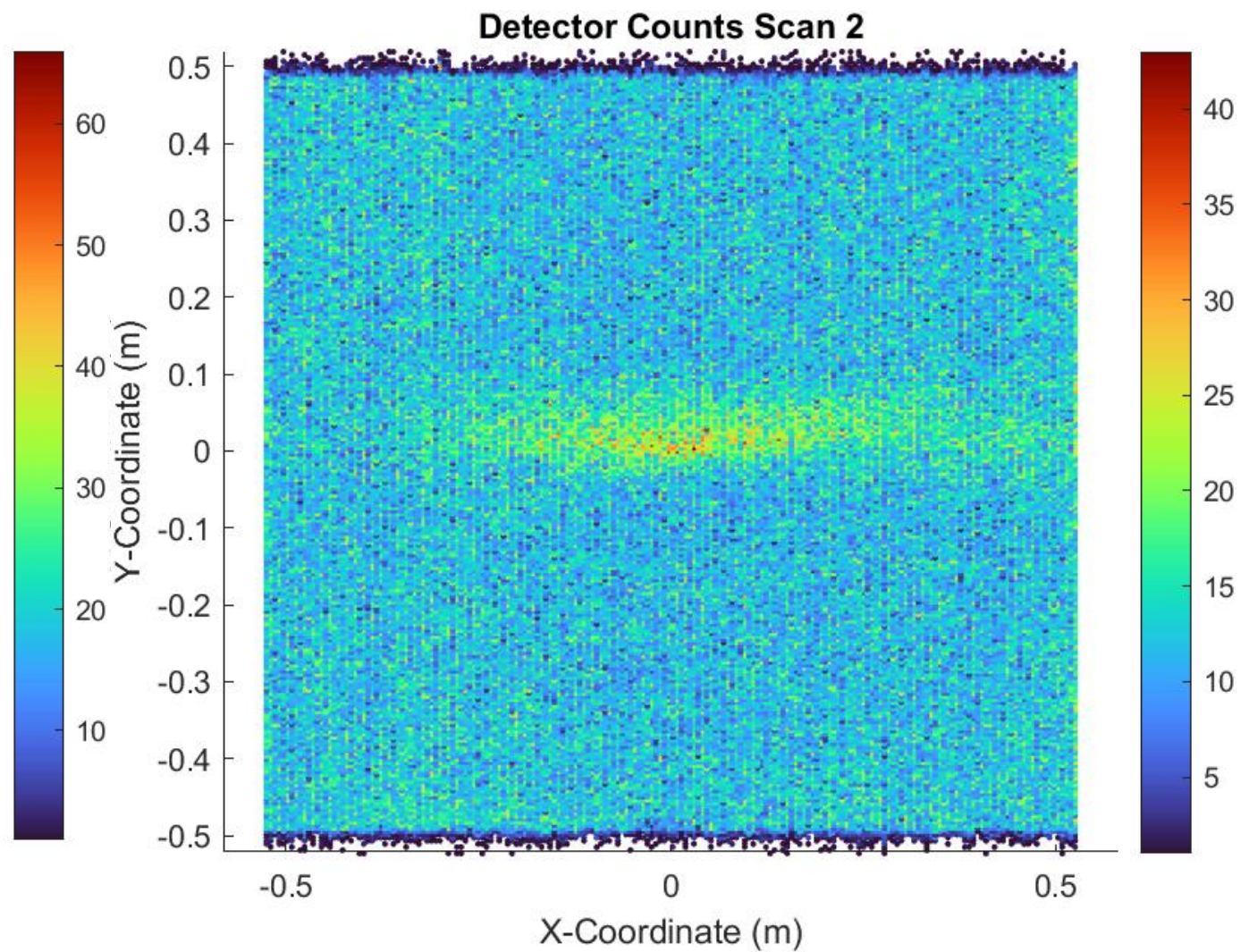
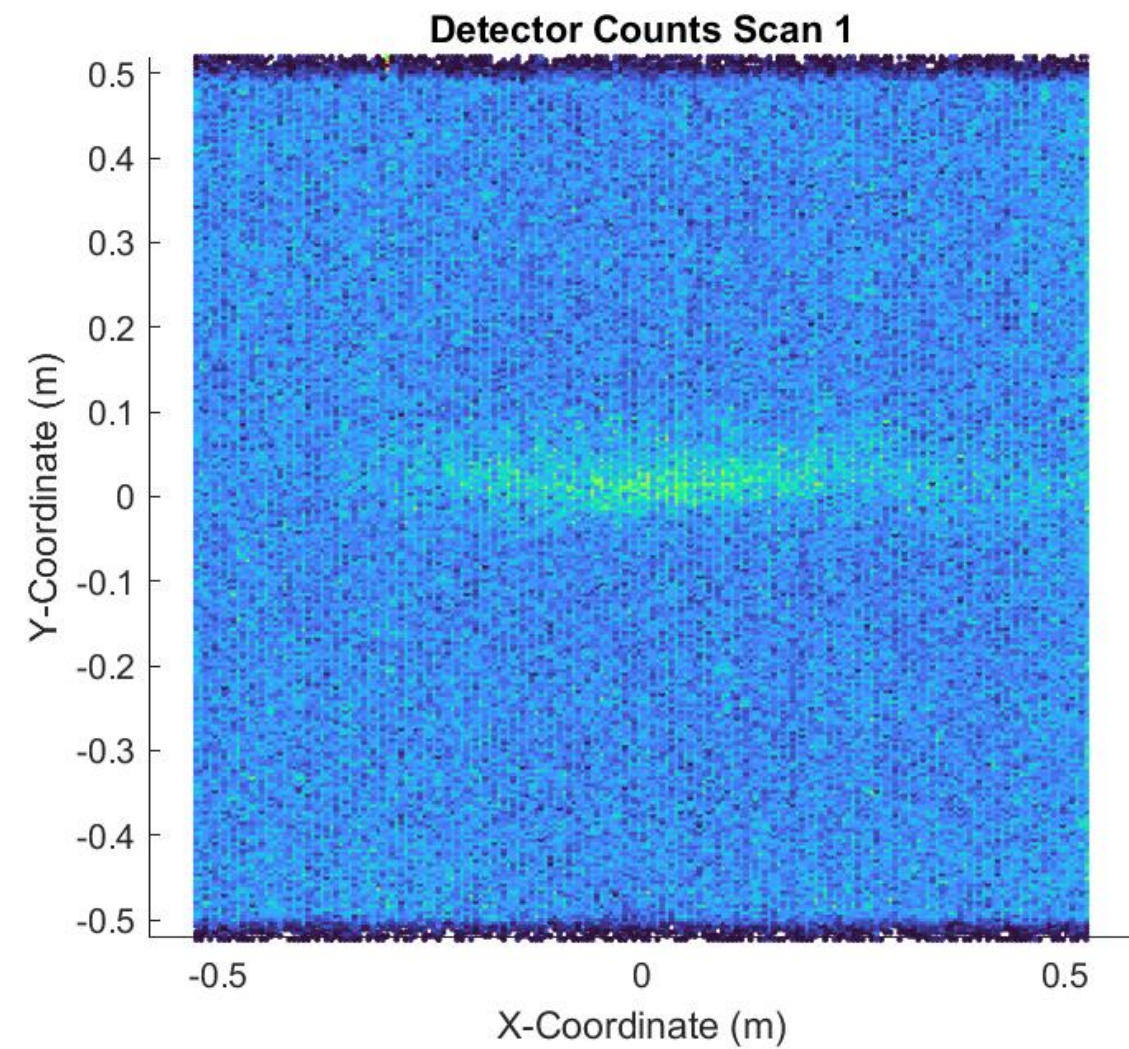
Detector X-projection for All Runs



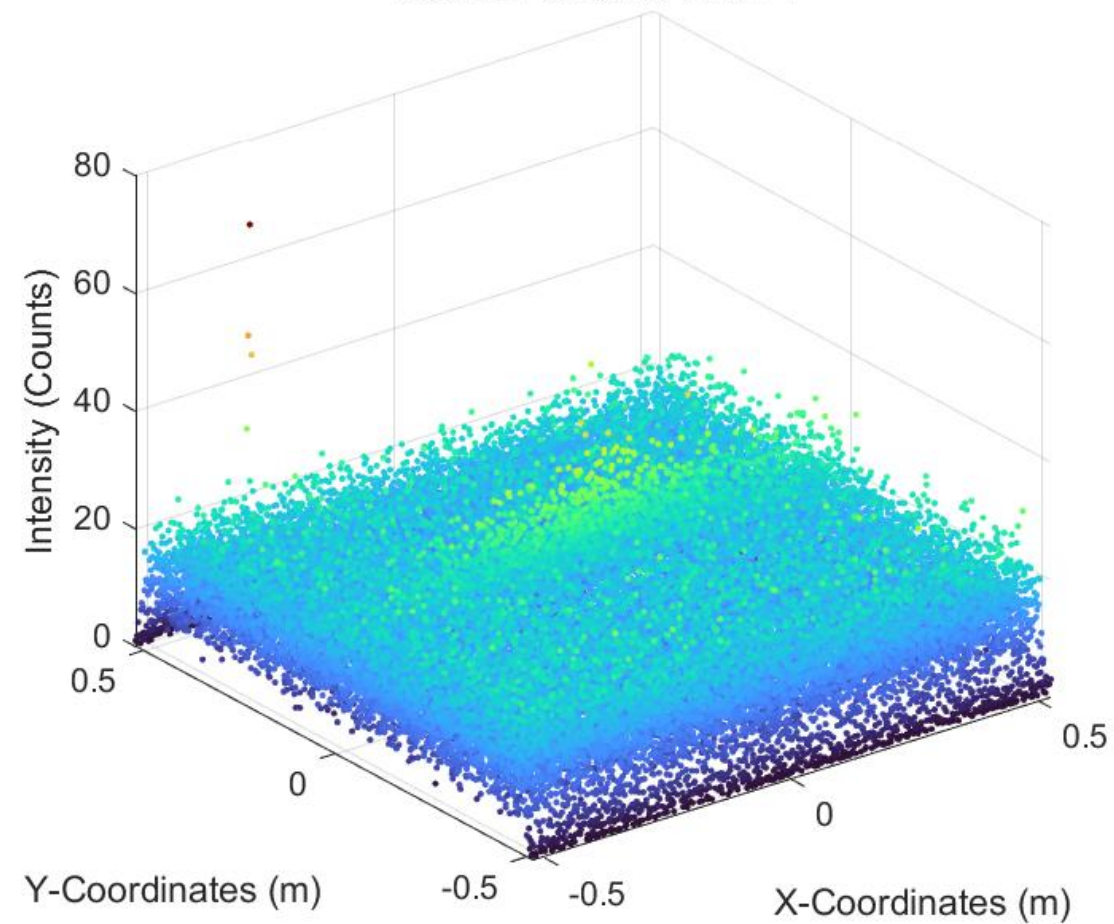
Detector Y-projection for All Runs



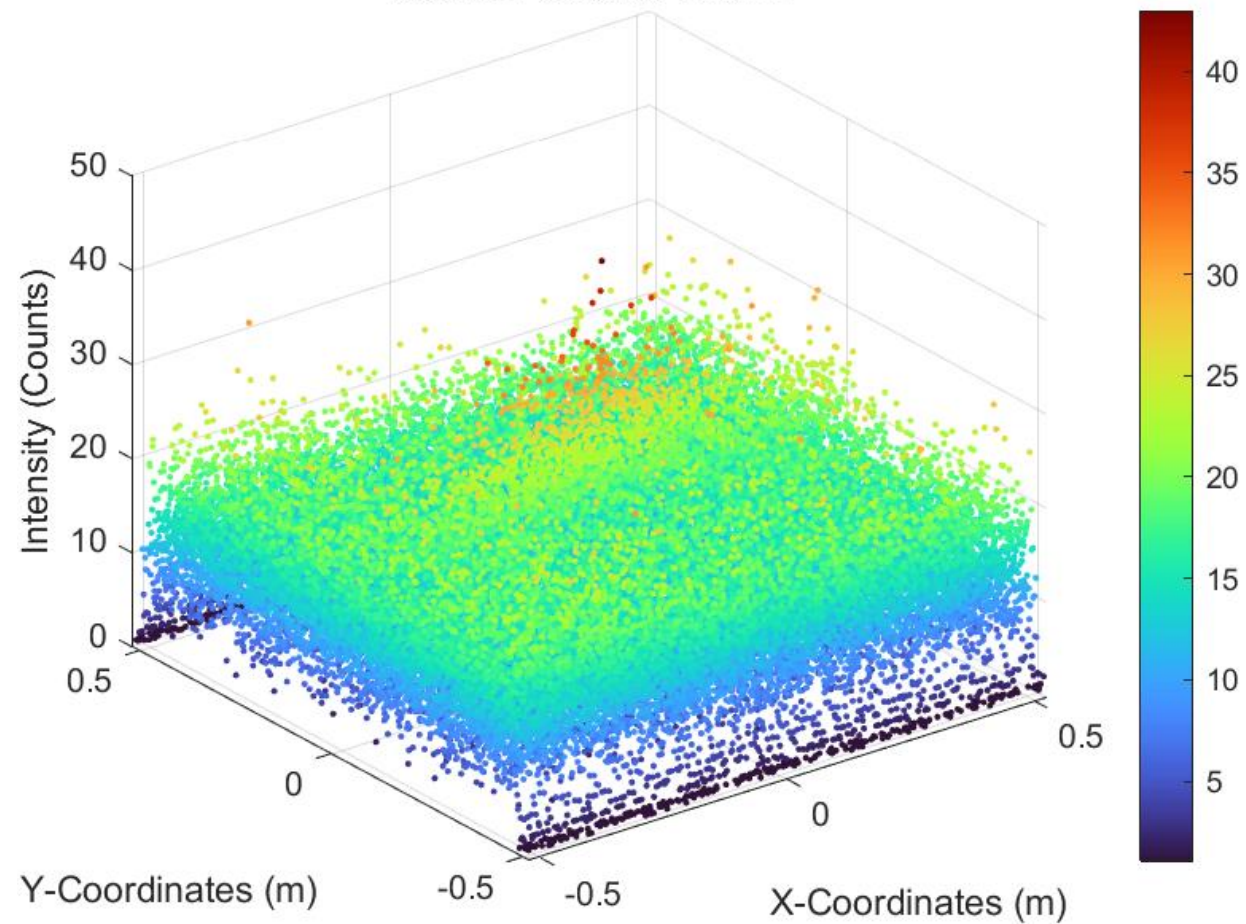




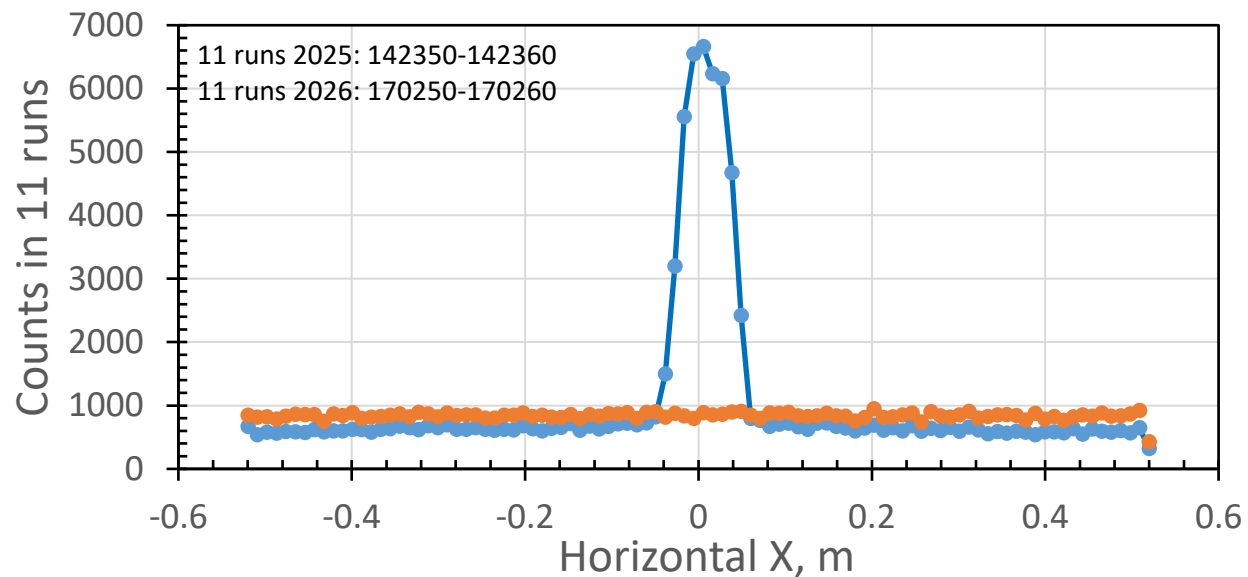
Detector Counts Scan 1



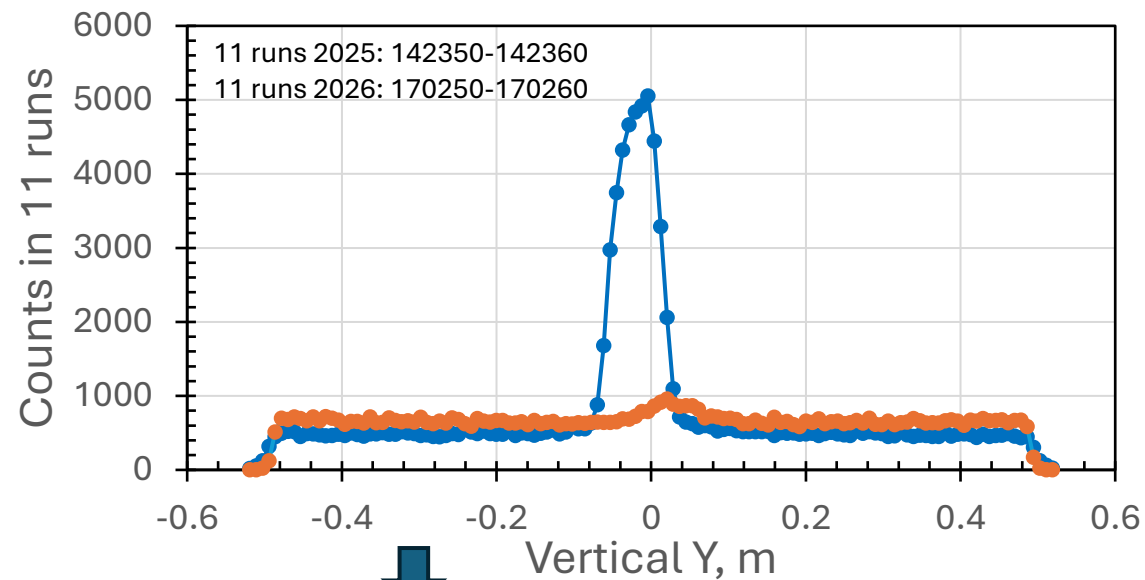
Detector Counts Scan 2



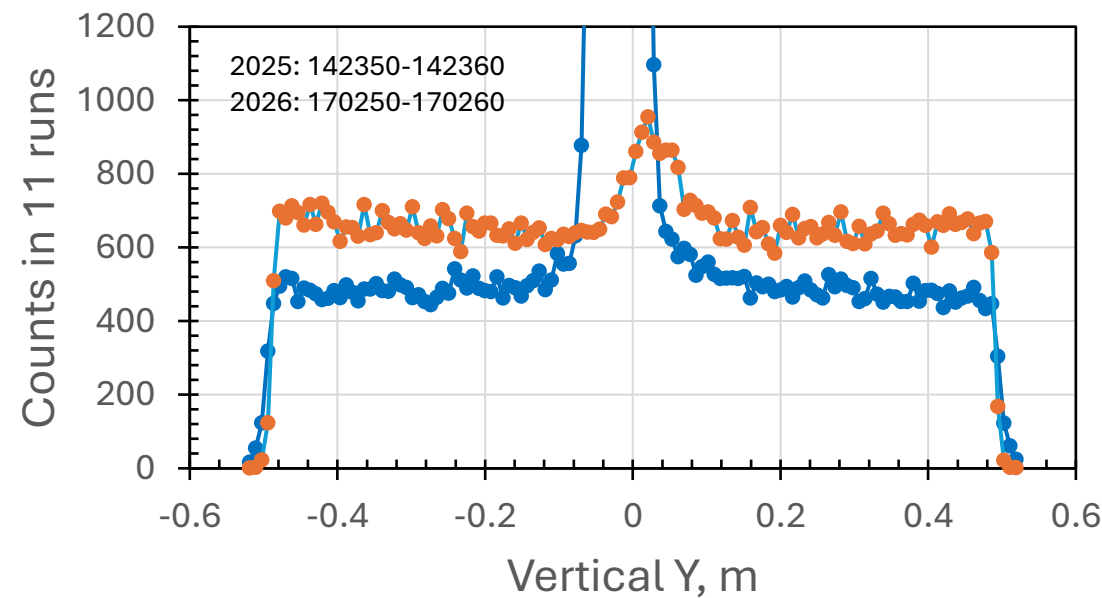
Detector X-distribution from 11 runs in **2025** and **2026**



Detector Y-distribution from 11 runs in **2025** and **2026**

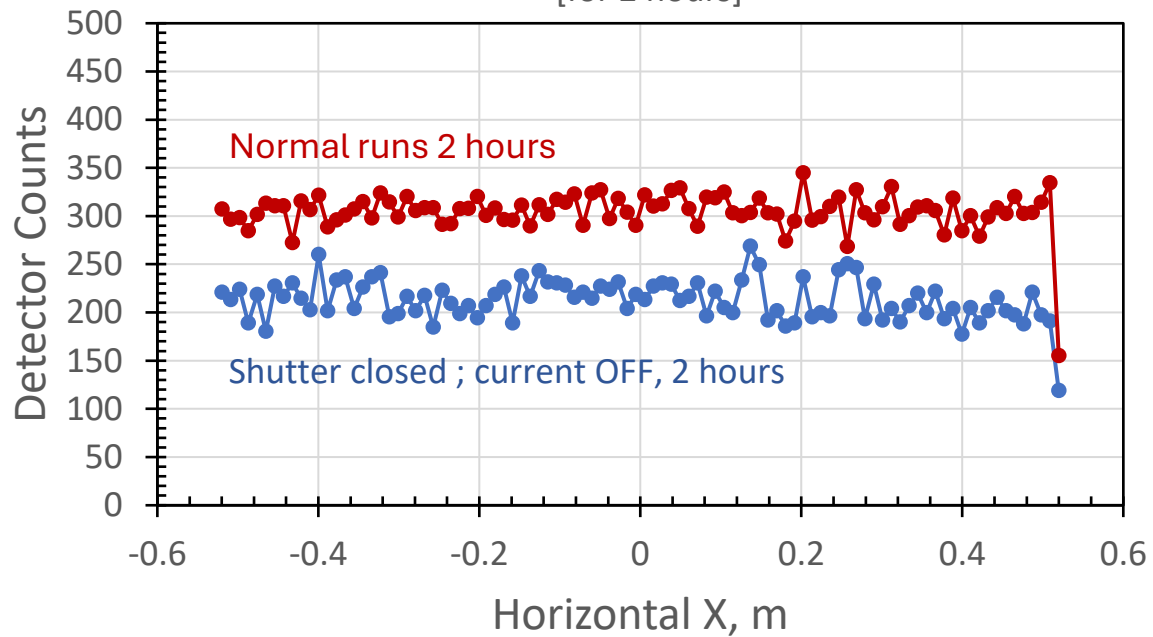


Detector Y-distribution from 11 runs in **2025** and **2026**

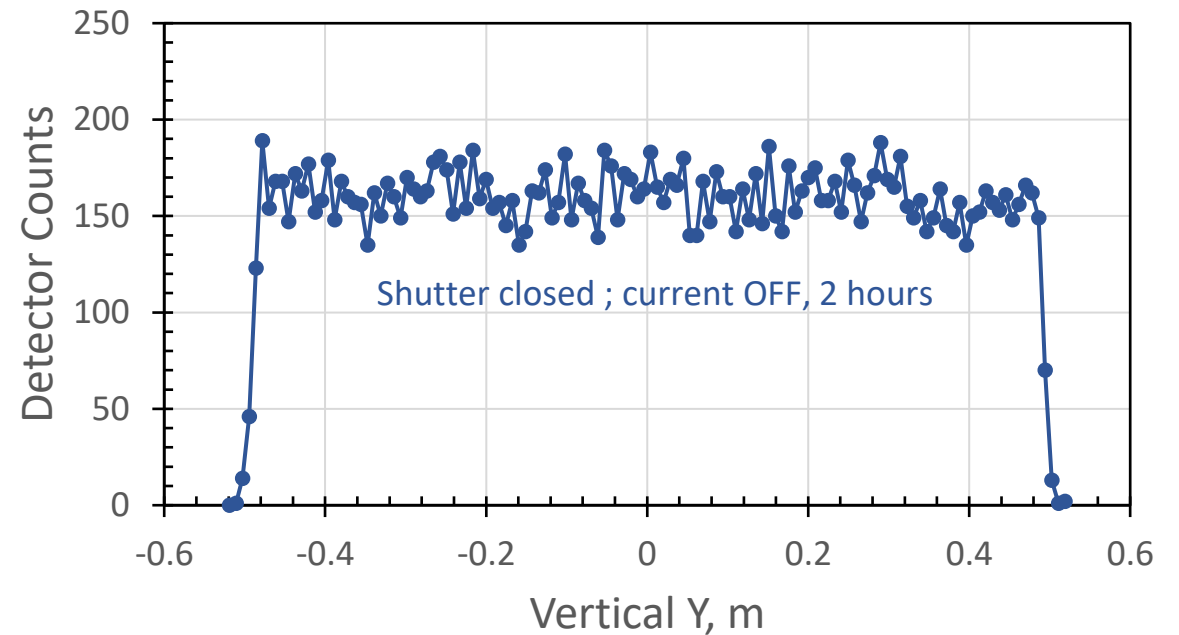


Shutter closed, Current OFF 2-hours at the end of measurements

All Detector X-distribution from **Runs 170311+170312**
[for 2 hours]

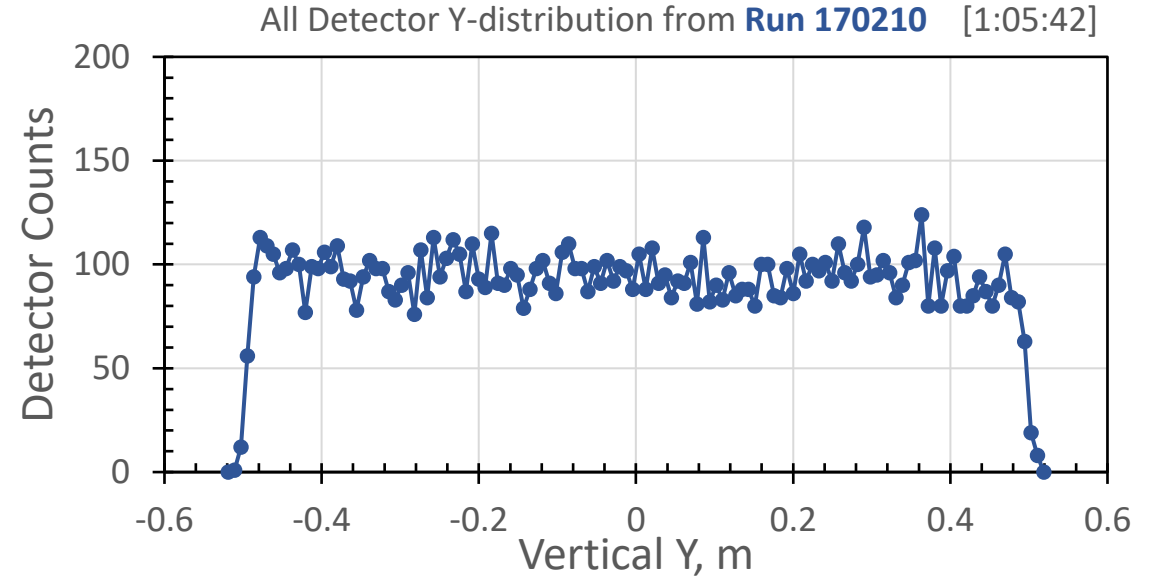
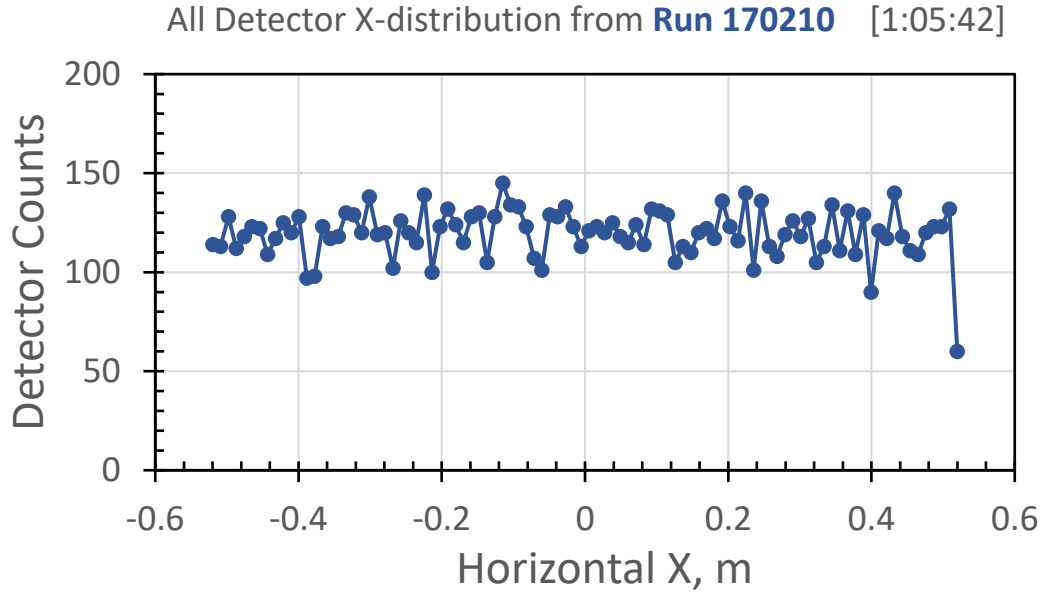


All Detector Y-distribution from **Runs 170311+170312**
[for 2 hours]

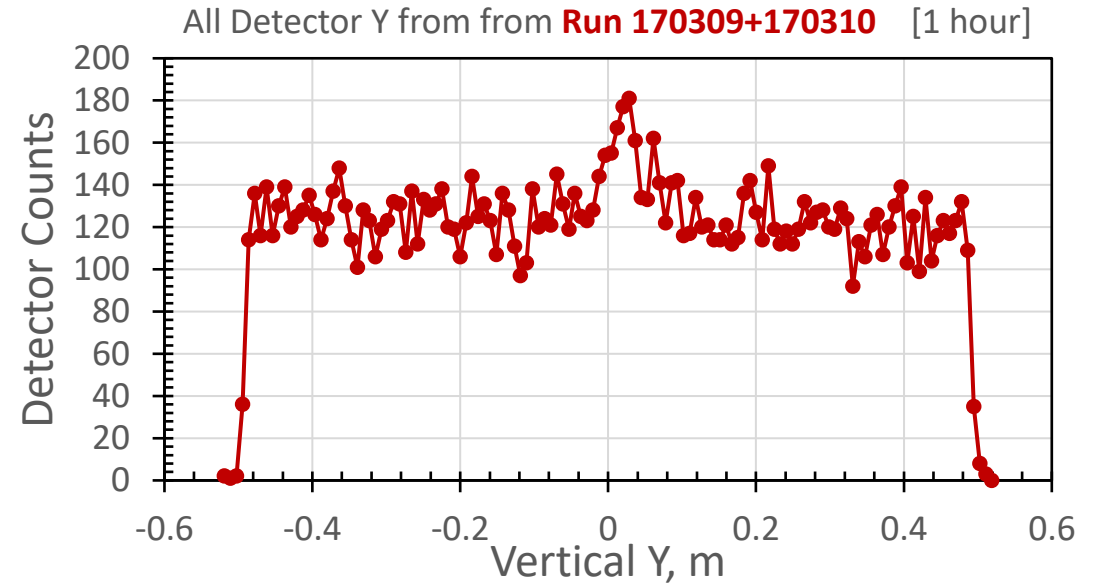
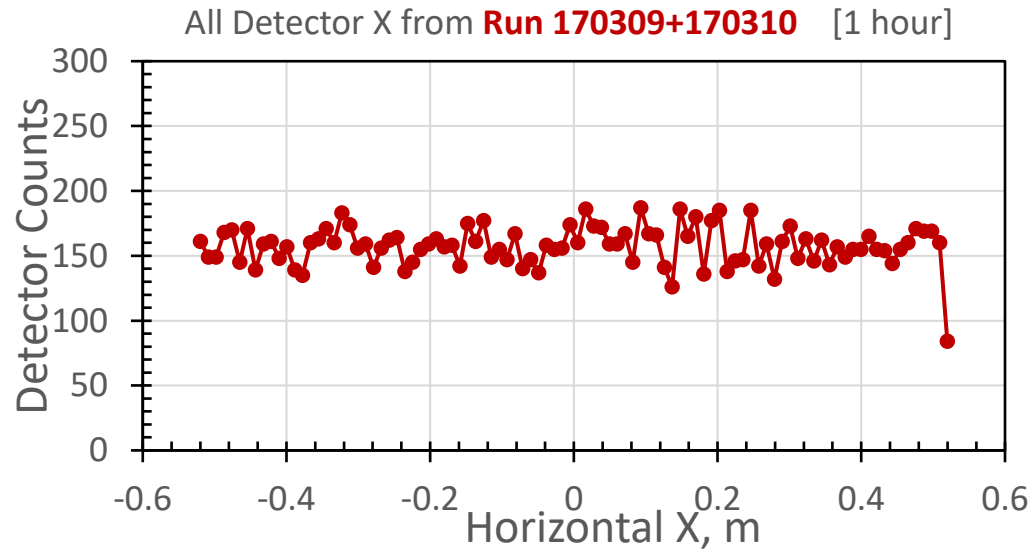


~ 2/3 of detector background are not coming from GP-SANS beam stopped by B-Poly and Cd

Shutter closed, Current ON 1-hour 6 min



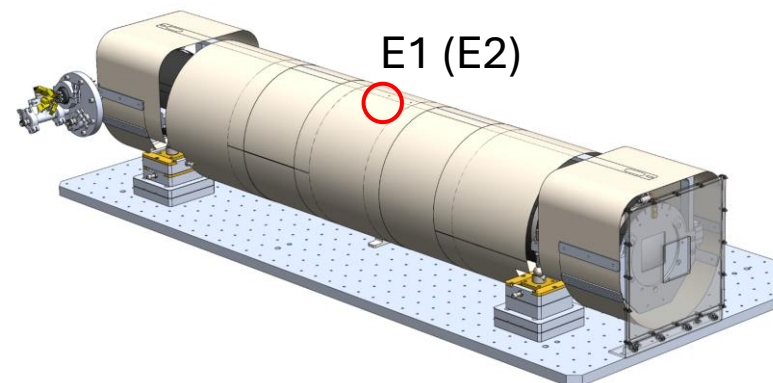
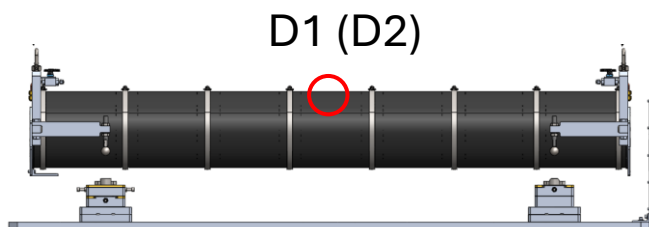
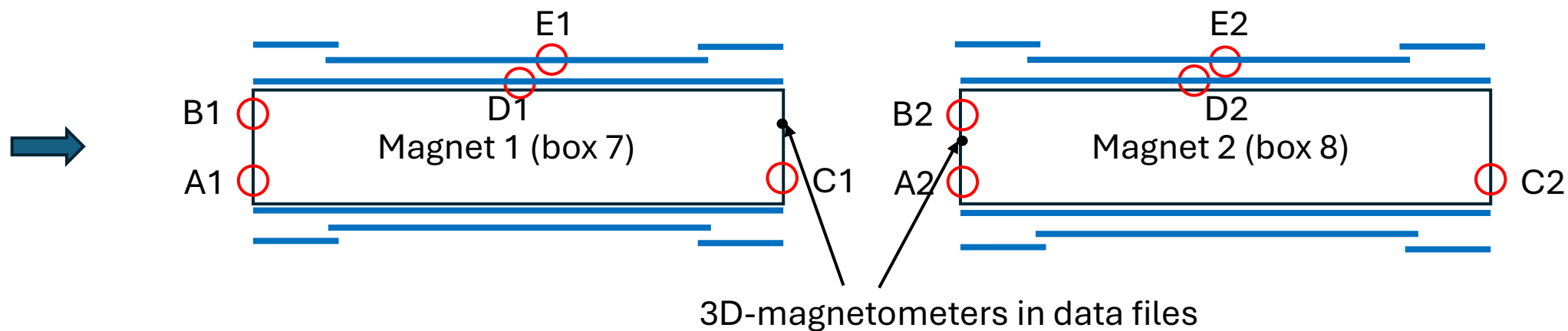
Shutter open, Current ON 1-hour



No measurement with Shutter open, Current OFF

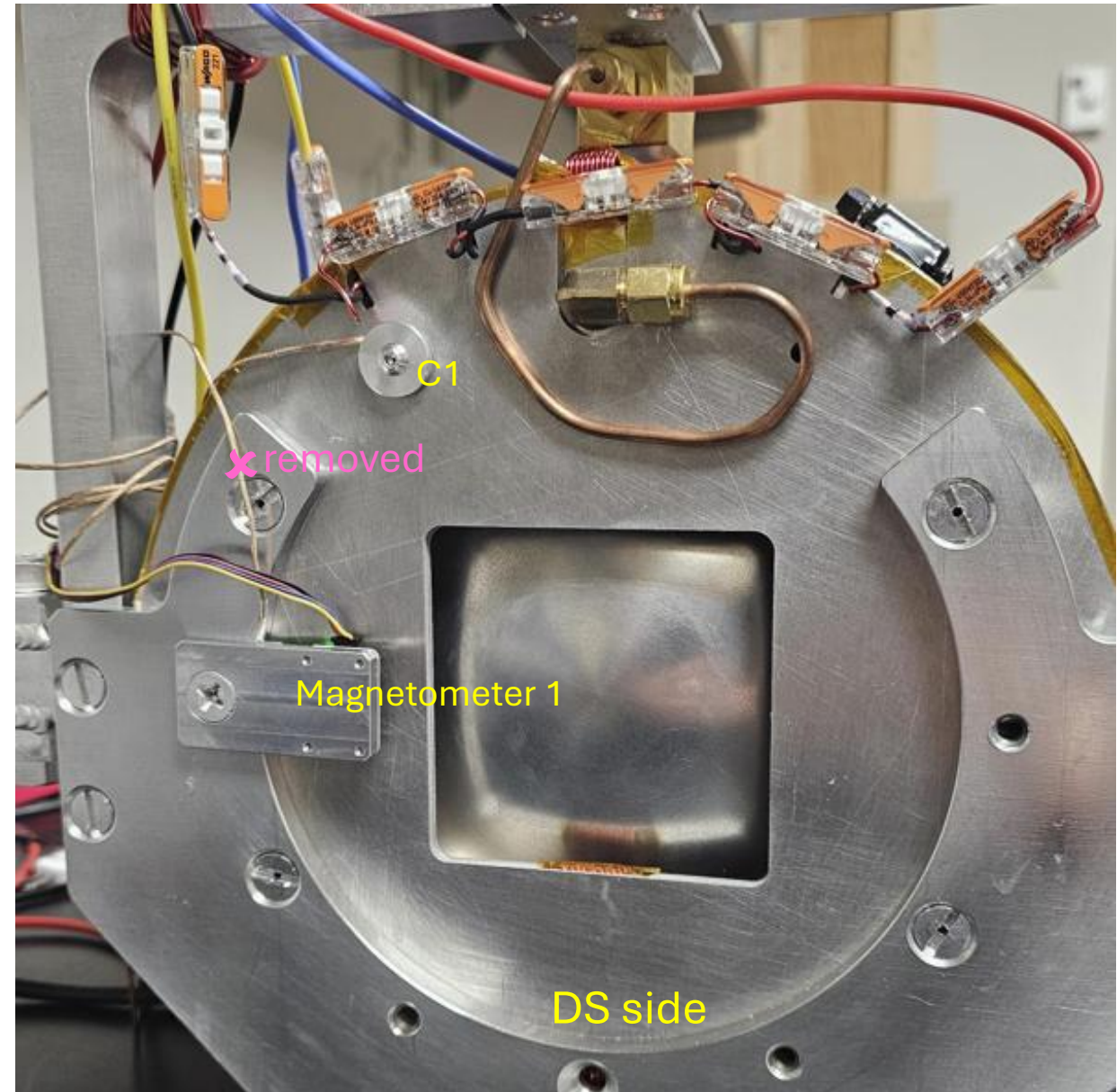
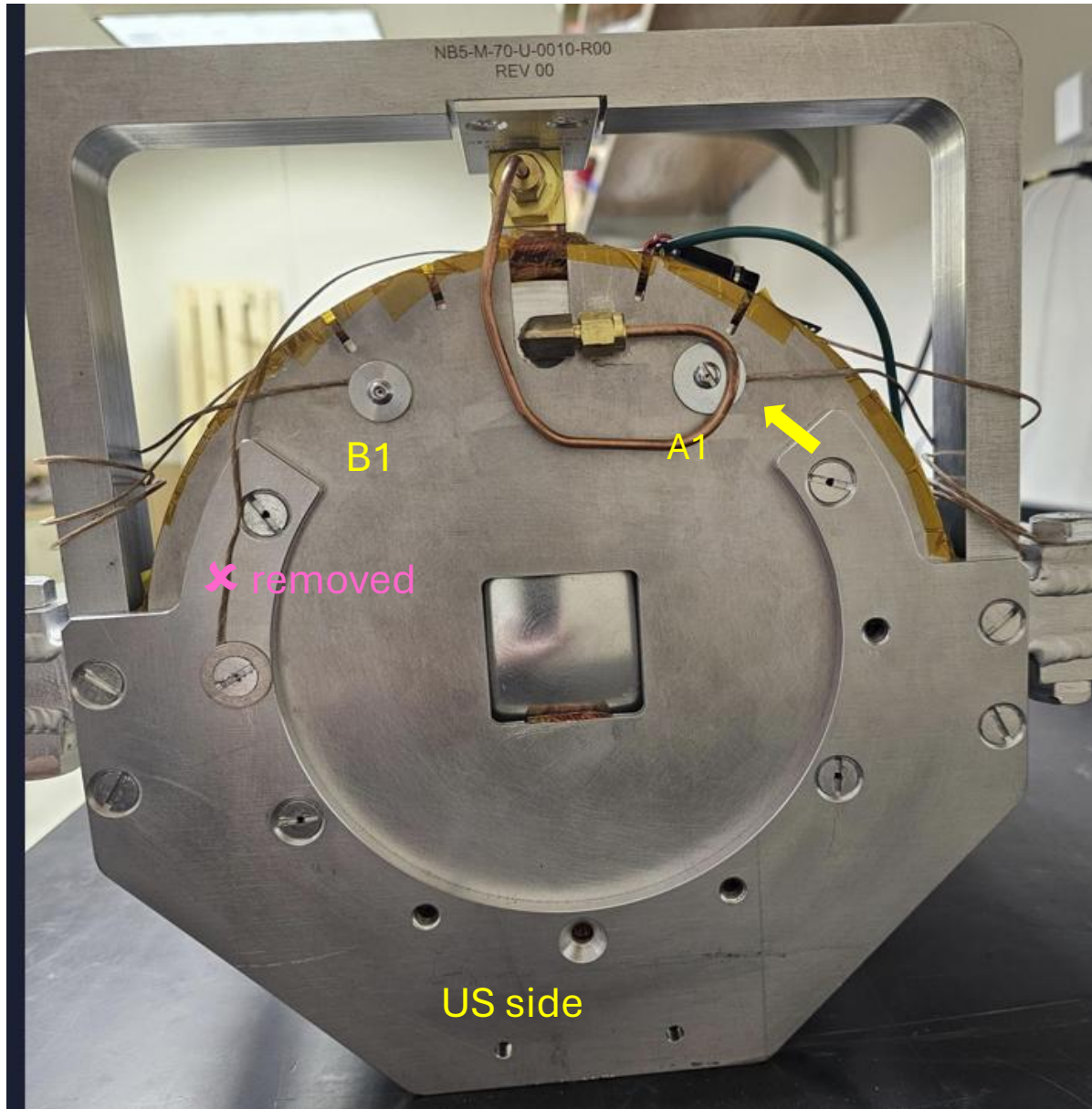
Origin of X-stripe in detector background (peak in Y) might be related to the presence of FAST neutrons in the beam with energies ~ 1 MeV. Neutrons scattered from the mirror of the “optical filter” and then scattered in Borated Poly that can not sufficiently moderate them. Possibly Charlie Snow can easily simulate that in GEANT-4

nTMM-2 Thermocouple Location Schematic 2026

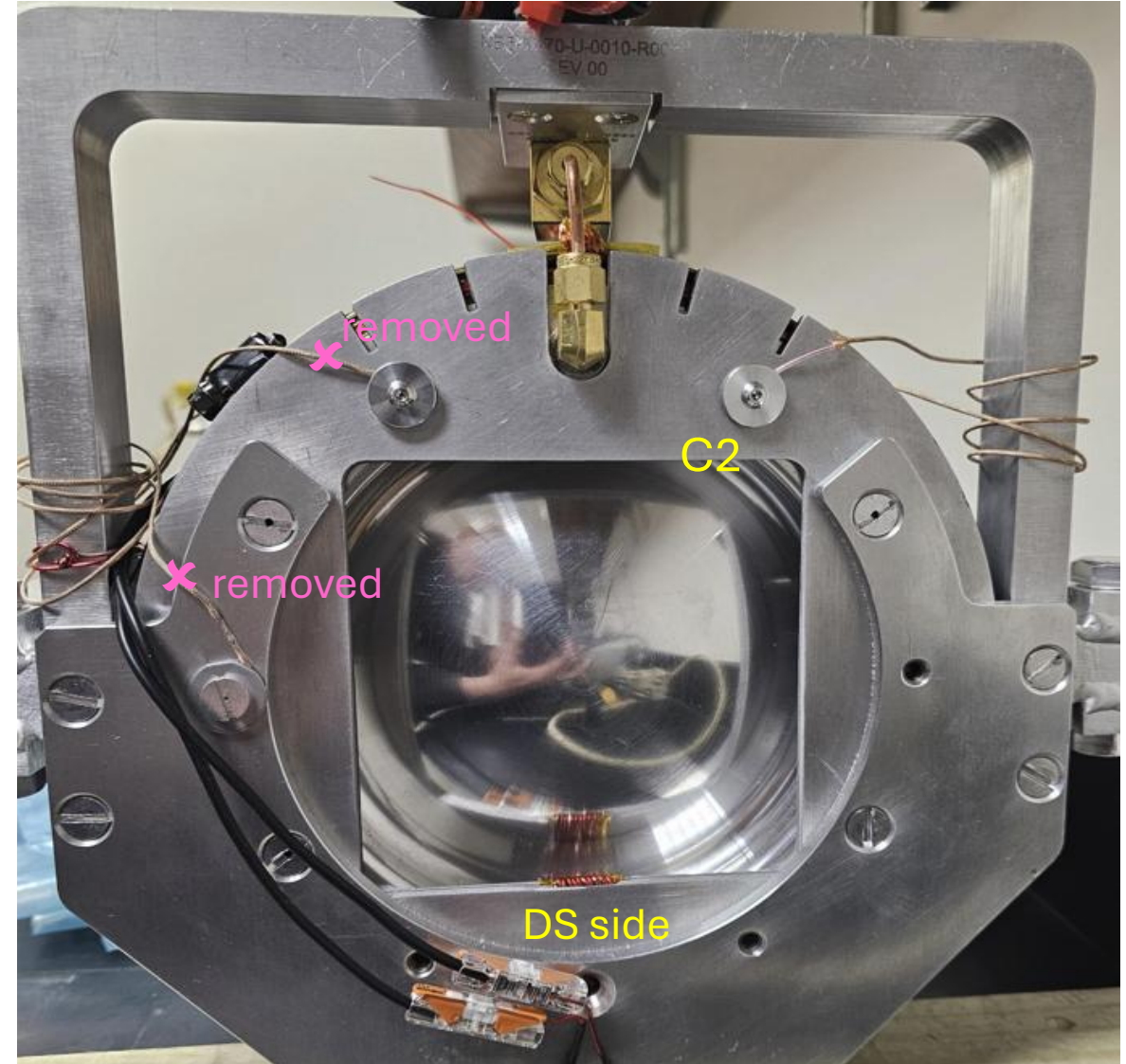
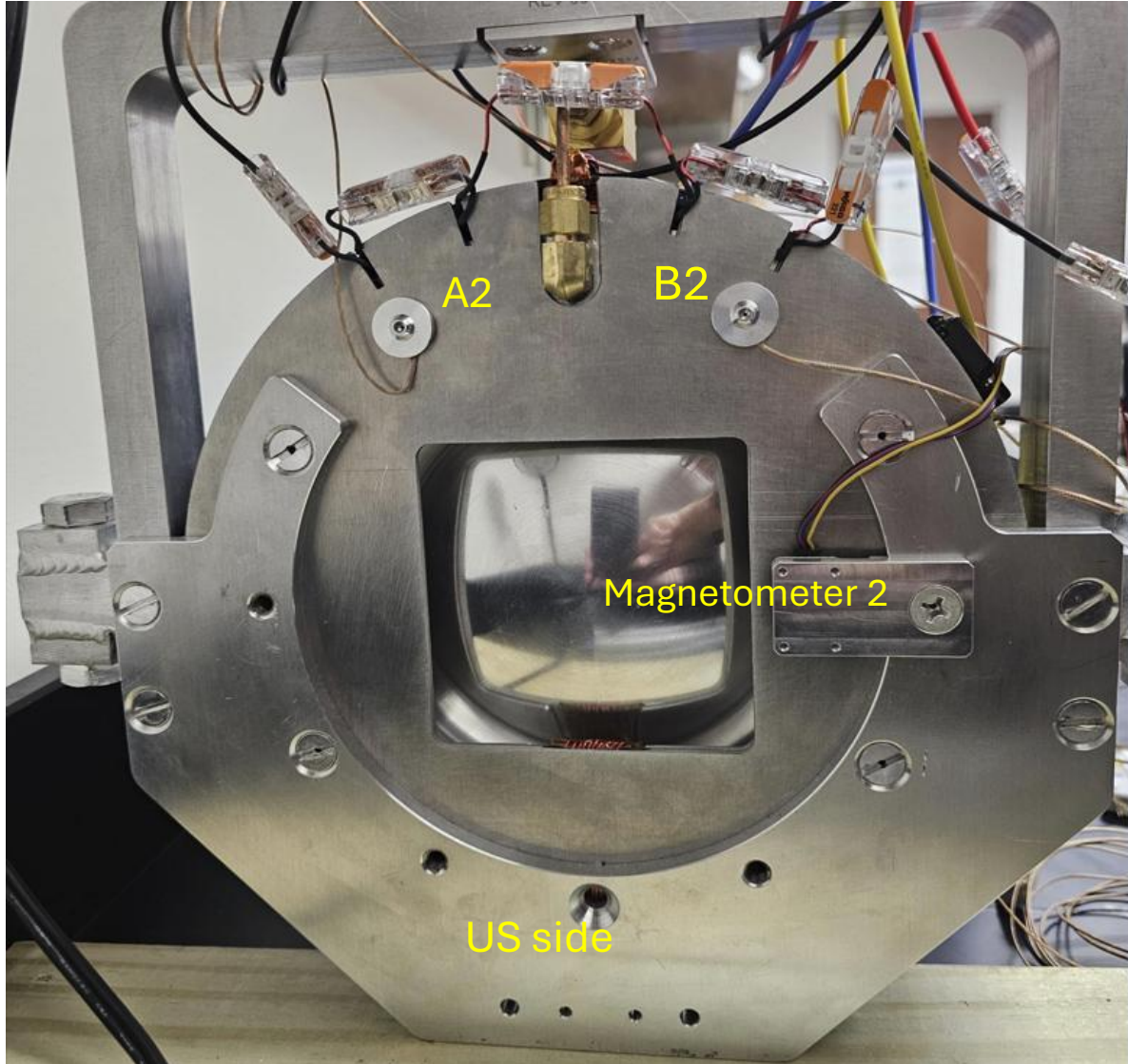


A1, A2 , D1, D2 thermocouples are in data files

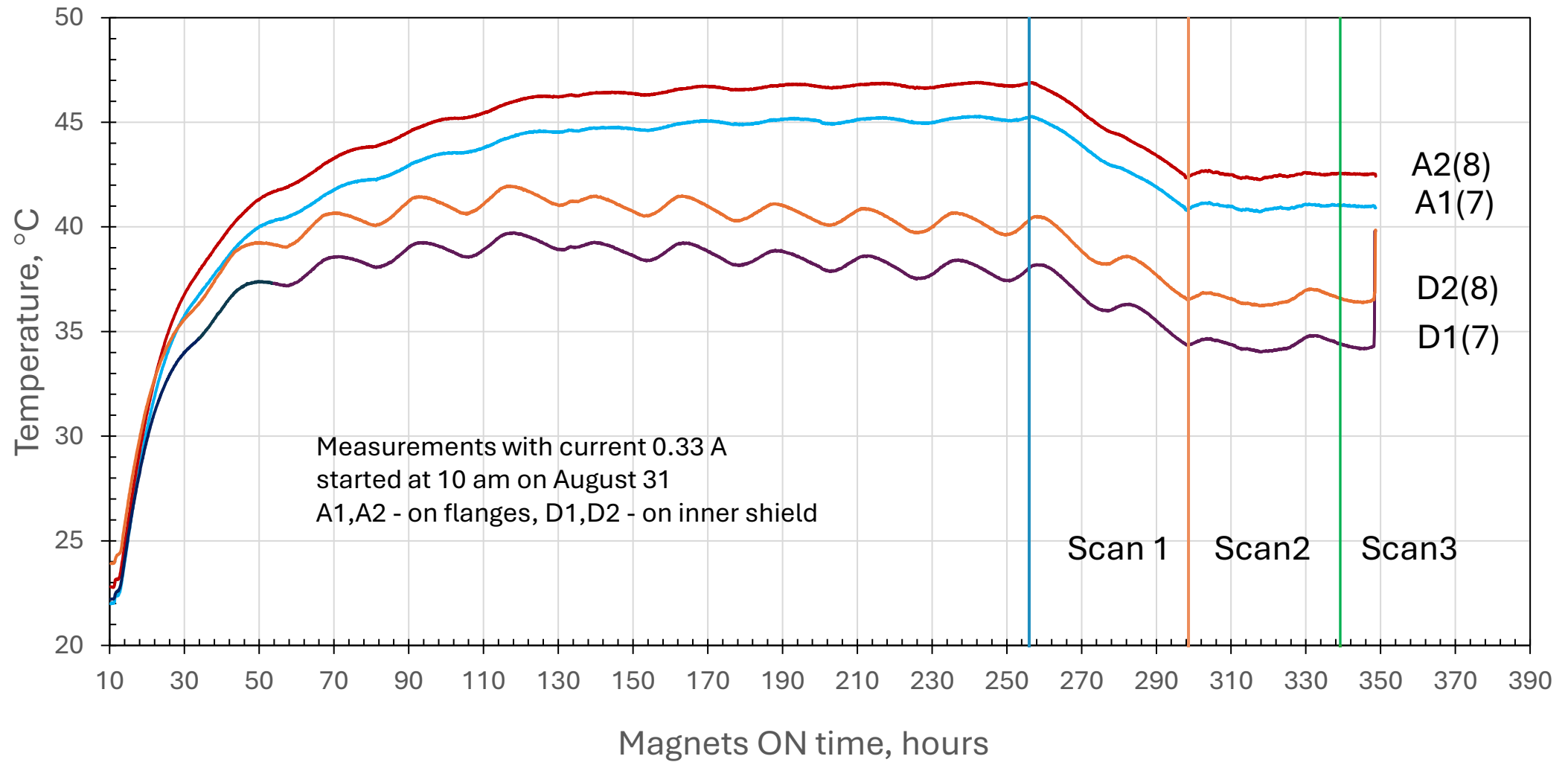
2026 Thermocouple's location on Magnet 1 (in box 7)



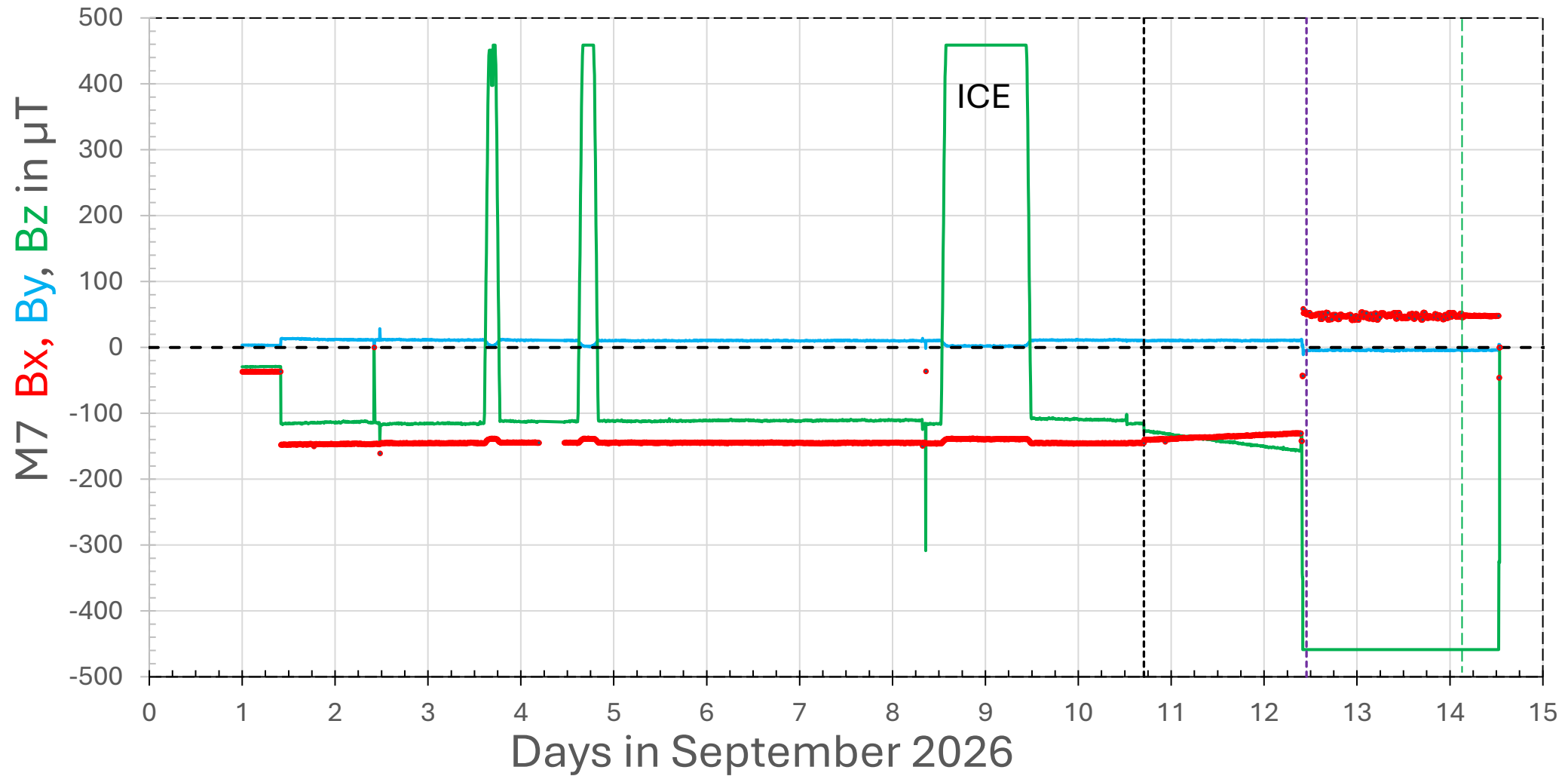
2026 Thermocouple's location on Magnet 2 (in box 8)



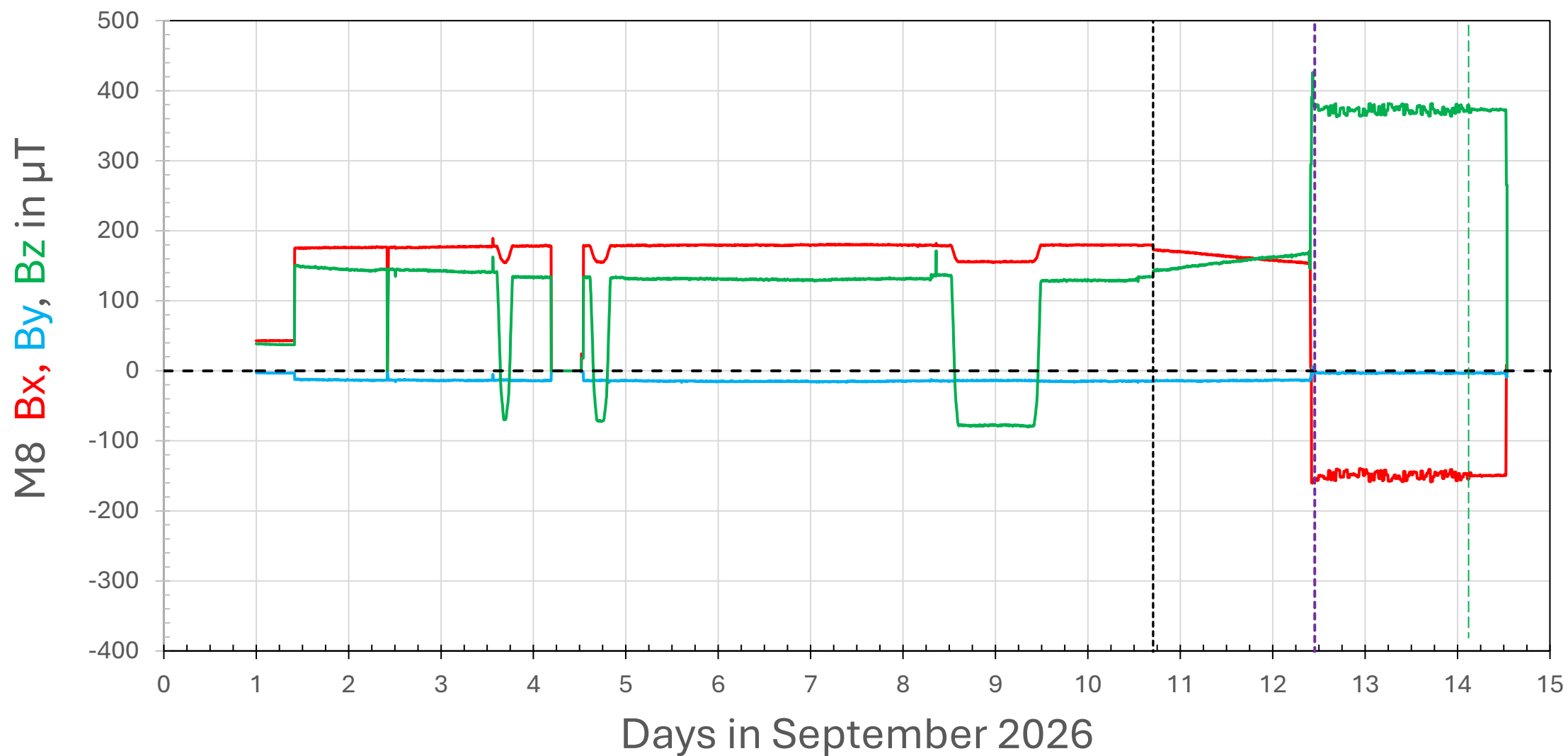
Temperature of magnets 7 and 8



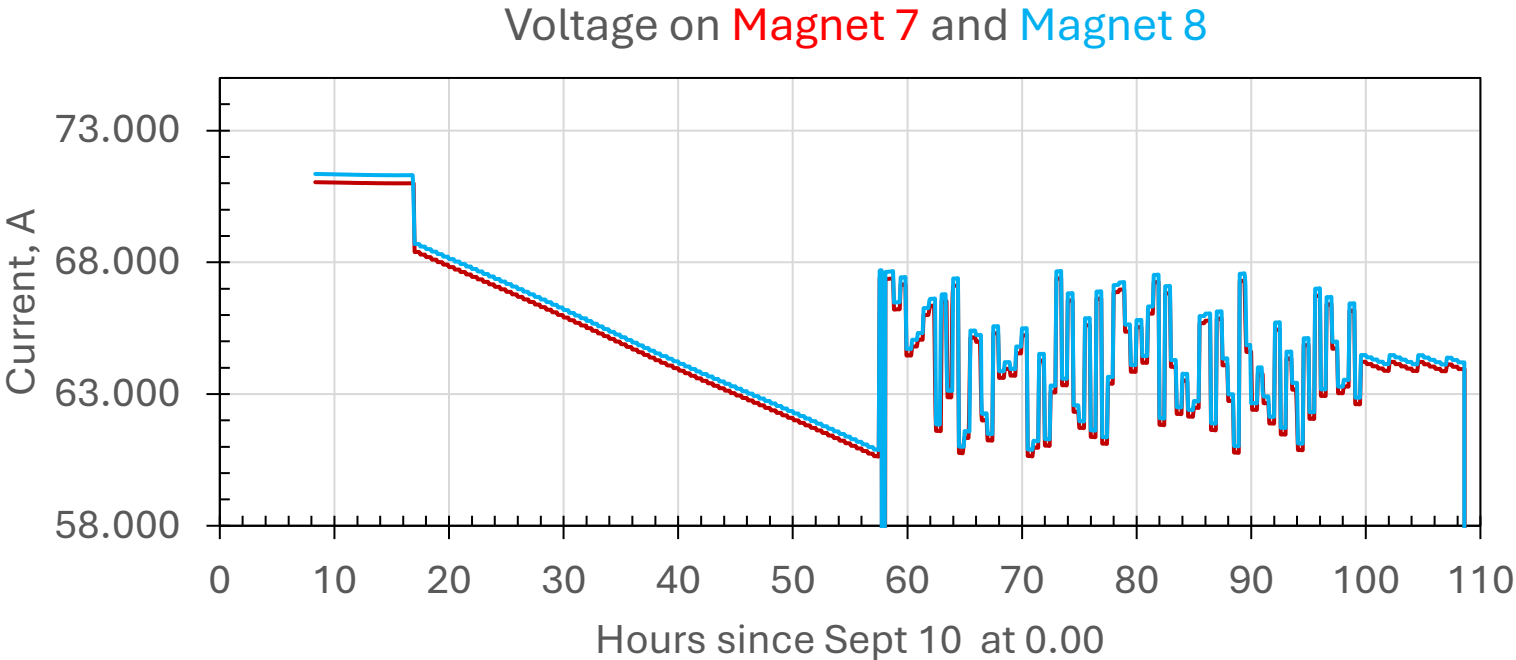
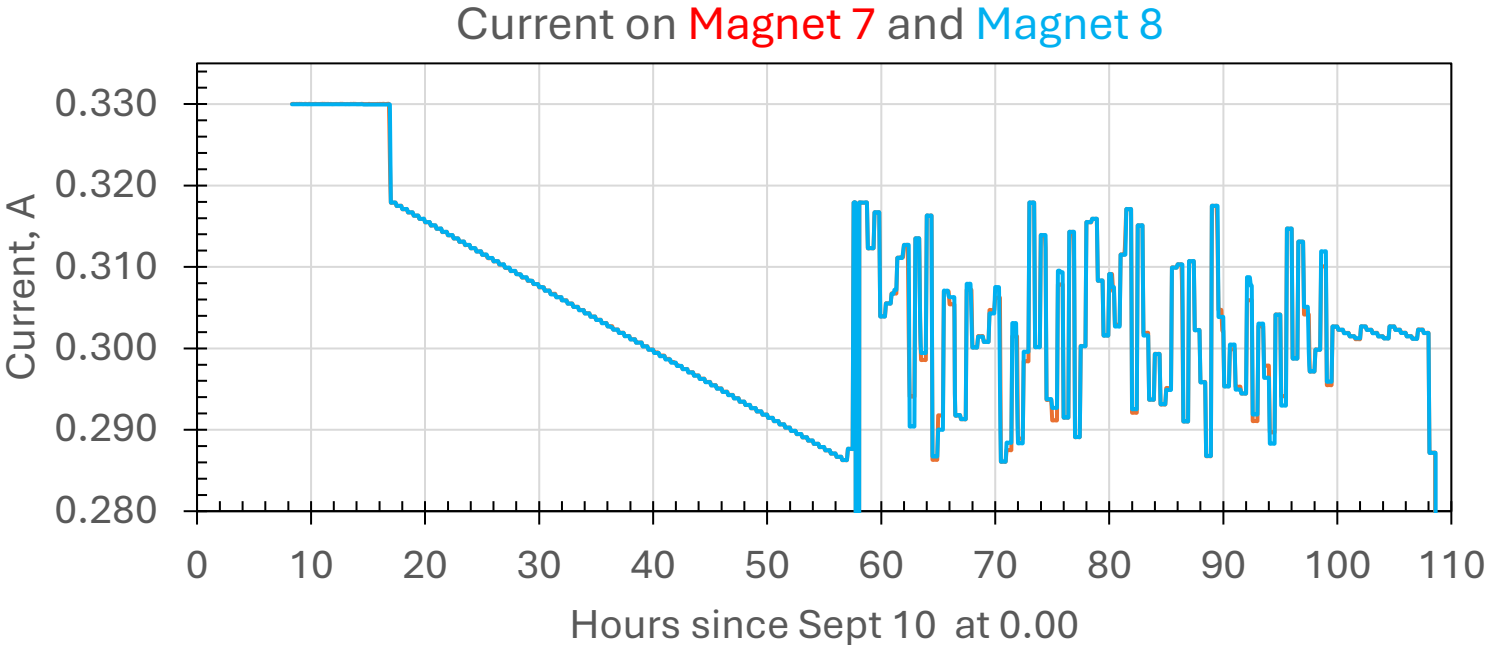
3D CHIP magnetometer on box 7



3D CHIP magnetometer on box 8



Current and Voltage
on Magnet 7 & 8



ROI background comparison 2026 vs 2025

Run #	142244 (2025)	Scan 2 (2026)
GPM	93,600 / s	~88,000 / s
All Detector	13,011 / 30'	~ 8,200 / 30'
ROI	6,102 / 30'	768 / 30'
Comment	Typical run	Typical run

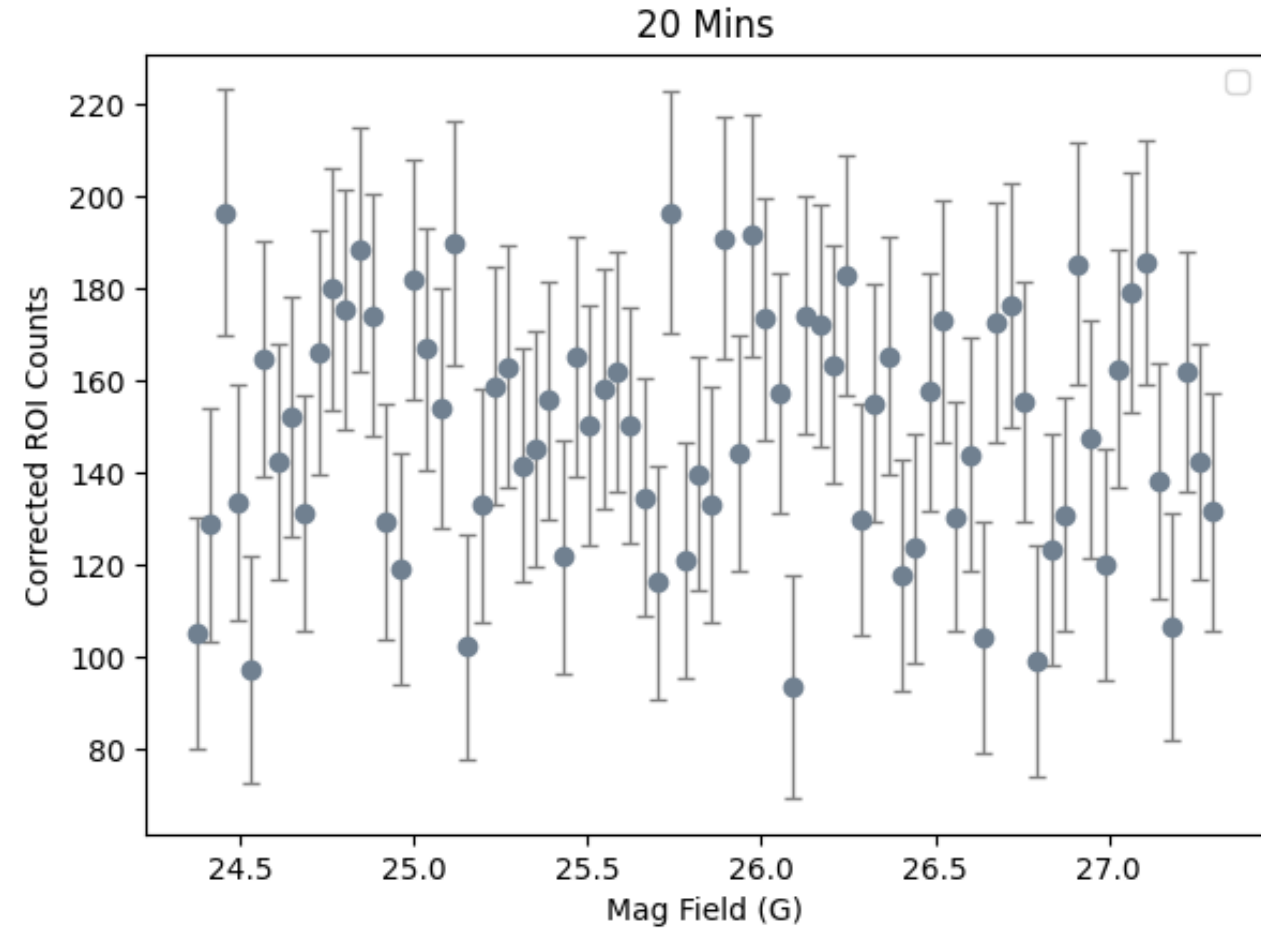
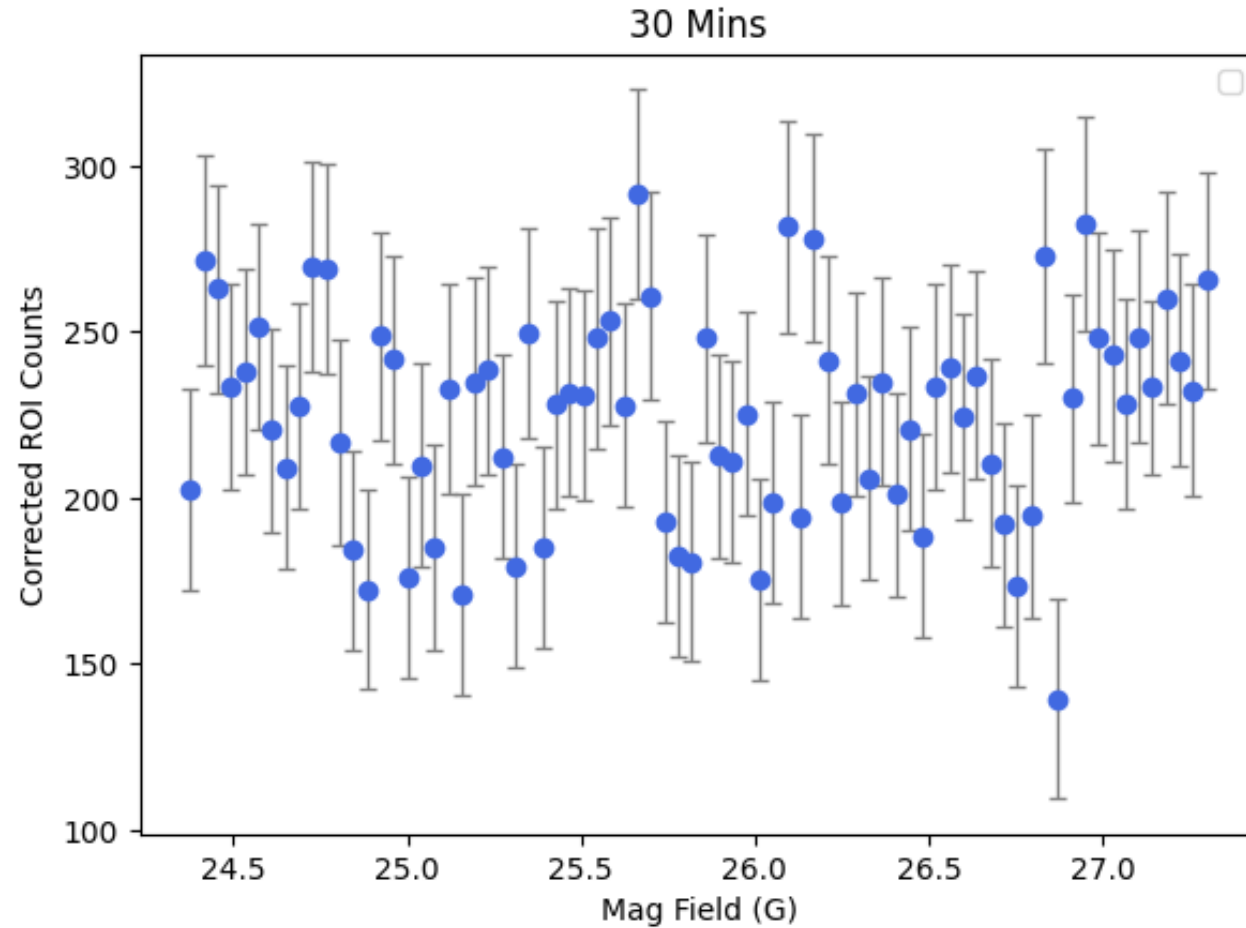
For ROI (30 cm x 36 cm) background is better by ~ 8 times

Summer Work and Results

Alina Moore – ORNL Meeting 9/18/2026

2025 Experimental data

Observed Ci



* Our expected peak position is 25.896 G

What goes into experimental data?

- Correction for intensity variation (found last year) applied due to slight change over time

$$\text{Corrected Counts} = \frac{C_{x,i}}{\left(\frac{GPMCounts_i}{ReferenceGPMCounts_x} \right)}$$

- Rescaled background by area ratio to subtract it from the ROI counts
 - See reference slides!!!

GPM relative intensity correction with ref. run **253** for 30-min series

GPM relative intensity correction with ref. run **346** for 20-min series

ROI 30 x 36 cm² with fast n peak removed area 13.4 x 13 cm²

ROI area 905.8 cm² ; ROI efficiency 90.15%

ROB 94 x 96 cm² - dead rim 5 cm around ROI 40 x 46 cm²

ROB area 7376 cm², with $A_{ROI}/A_{ROB} = 0.1228$

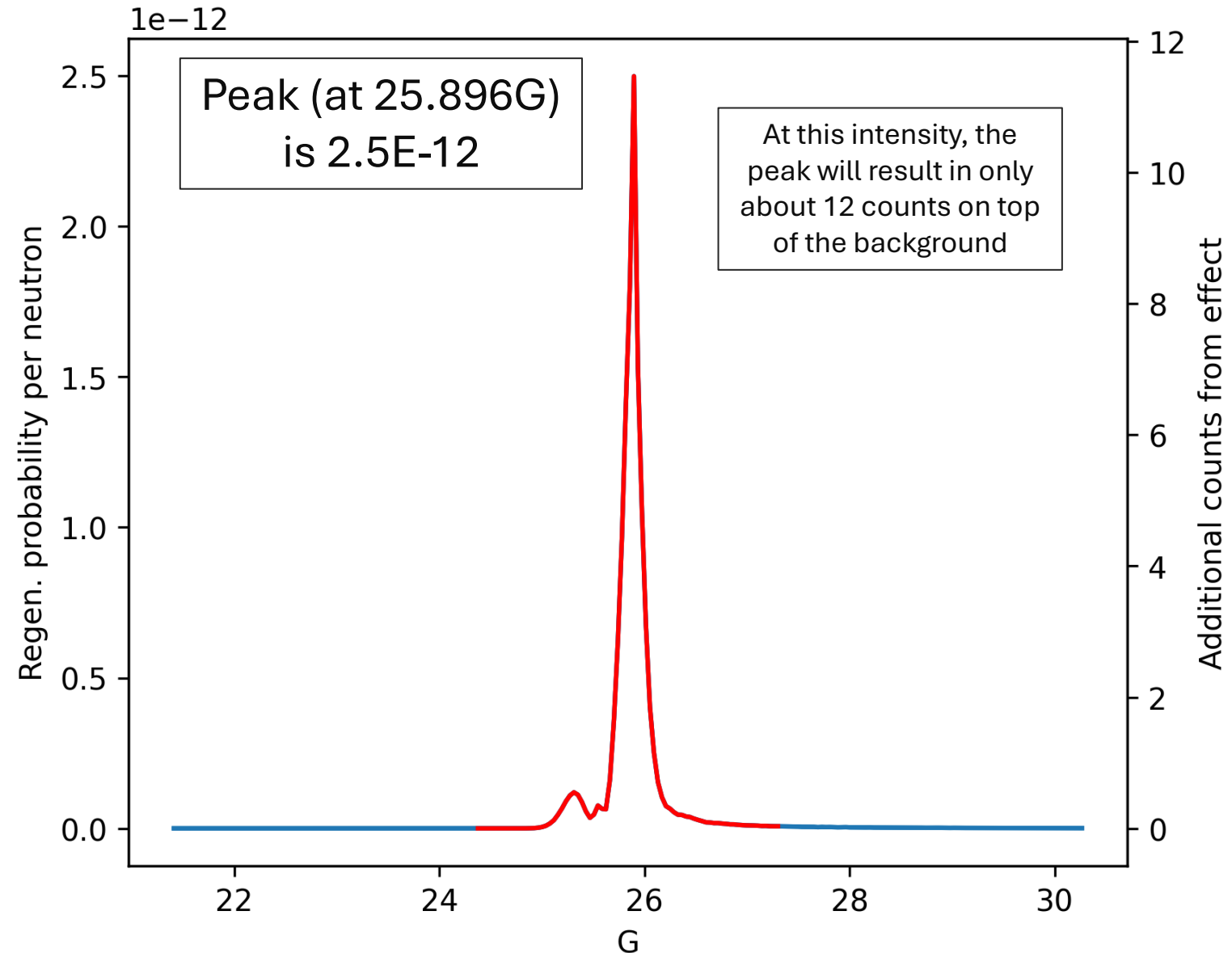
Corrected counts in ROI :

$$C_{corr} = C_{ROI} - C_{ROB} \cdot \frac{A_{ROI}}{A_{ROB}}$$

Fit functions + explanation

- The fit function used accounts for lower sensitivity from the pressure difference during the experiment
- Corresponding to $k = 5\text{E-}6$
- δ_m (mass difference between n and n') is not included, will have to be additional parameter
- Given in regeneration prob. per neutron, and multiplication by intensity gives the expected counts from effect
 - This is not larger than error of each run's counts

Fit function accounting for pressure difference

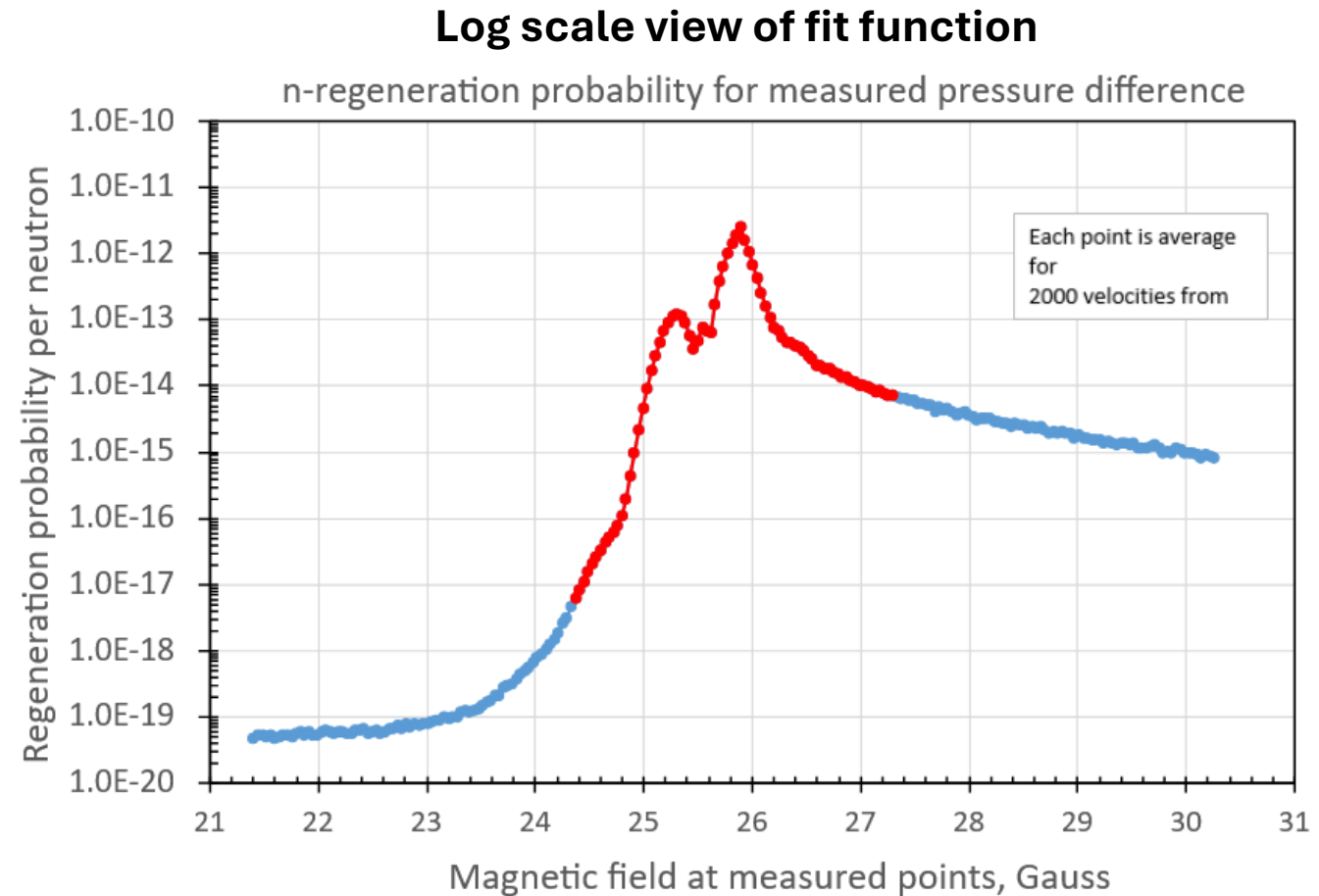


Fit functions + explanation continued

For all fits, the model is

$$aP(N) + B$$

- a is the amplitude of the effect and is proportional to κ^4 .
 - $P(N)$ is the function of regeneration
 - When multiplied by intensity, it's # of additional counts from effect
 - Probability for each magnetic field value
 - B is the constant background
-
- Function tails are extended so we can try shifting the peak to be at each point in our magnetic field range

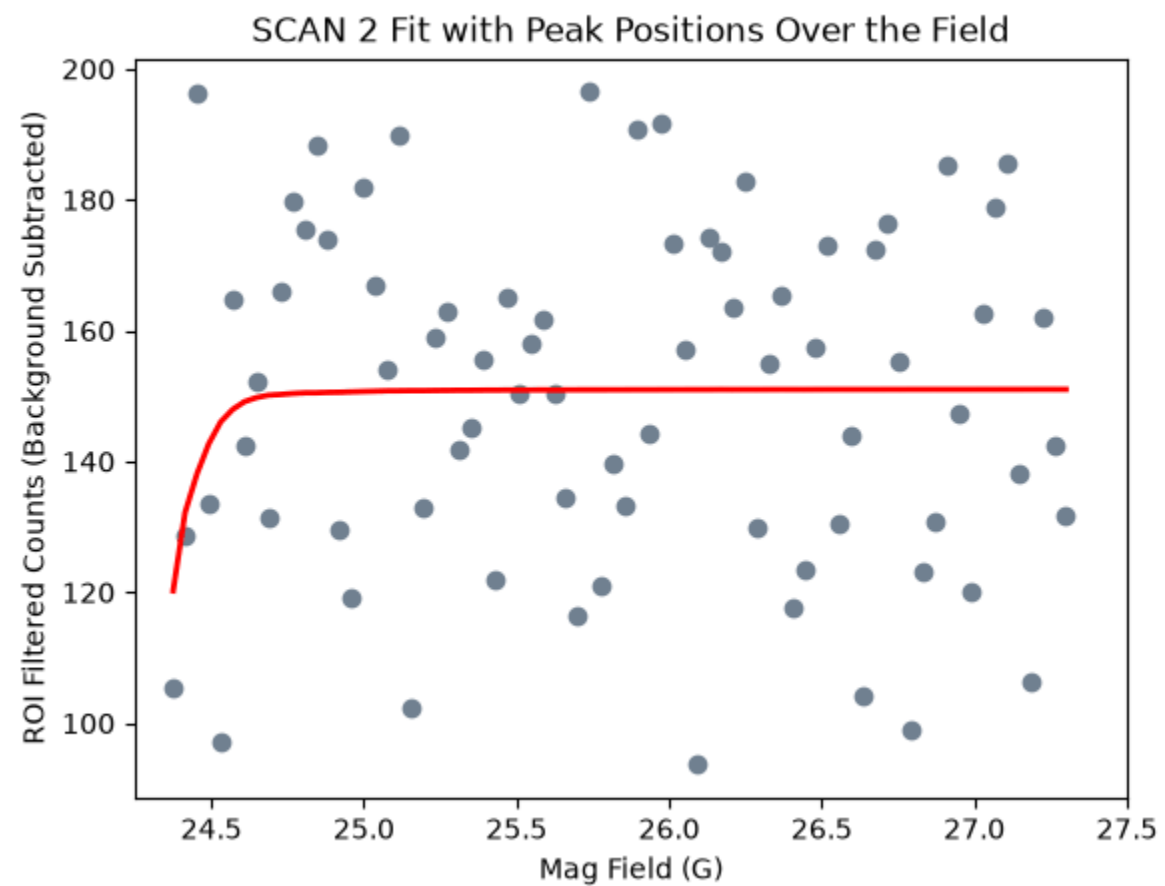
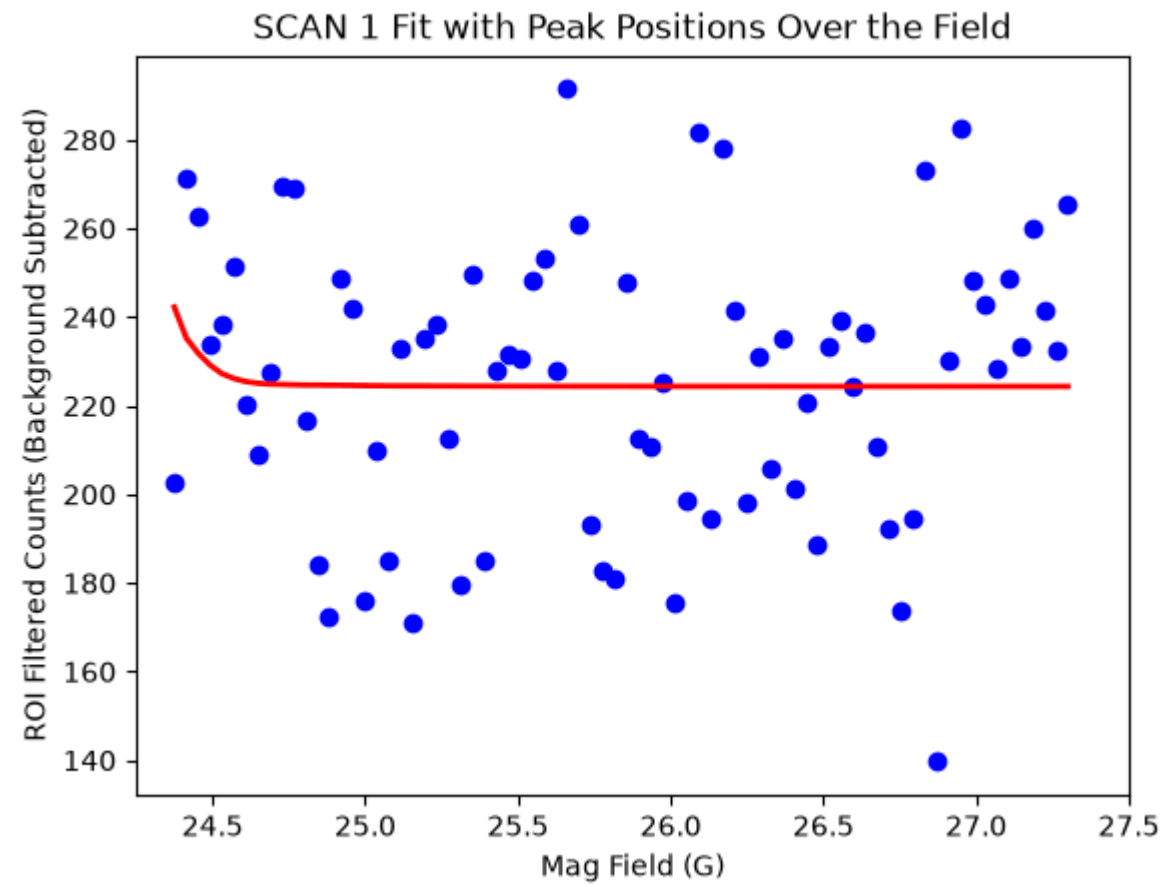


Context for Summer Work

Calculated intensity: $4.59\text{E}12$

- Includes correction factors from detector and its hardware
- There have been 3 main fitting tasks
 - Curve fit to experimental data
 - Try it for the peak fixed at each position in the magnetic field
 - Curve fit to pure background (generated using the B parameter estimate from part 1)
 - Curve fit to generated distributions with effect sitting on top of background
 - For calibration only
 - Compare with real data
- 95% Feldman-Cousins confidence intervals for the first two tasks

Animation to show fixing the peak through our field range



The fit can estimate a negative effect even though we assume a is not negative in real life

Shifting the fit function to fix the peak at each point and doing curve fitting

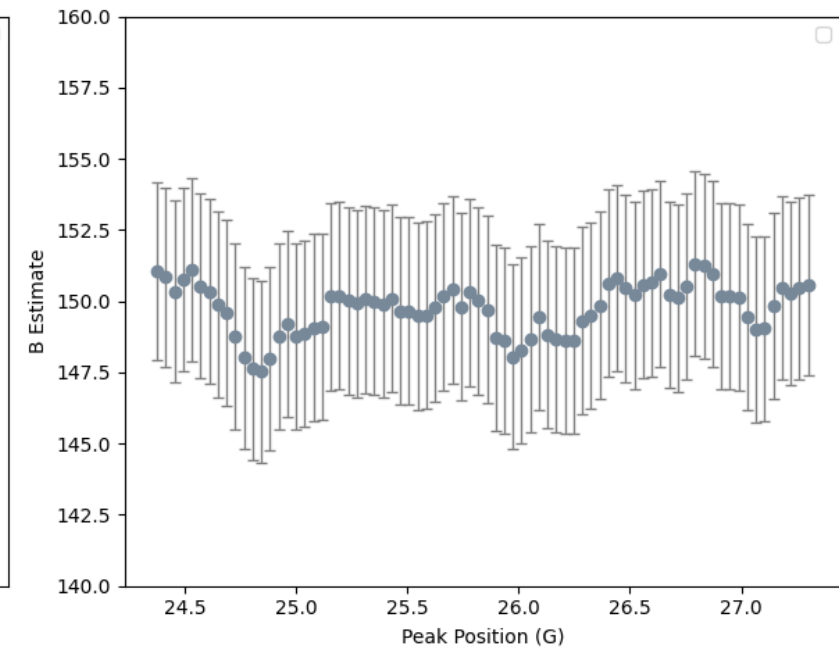
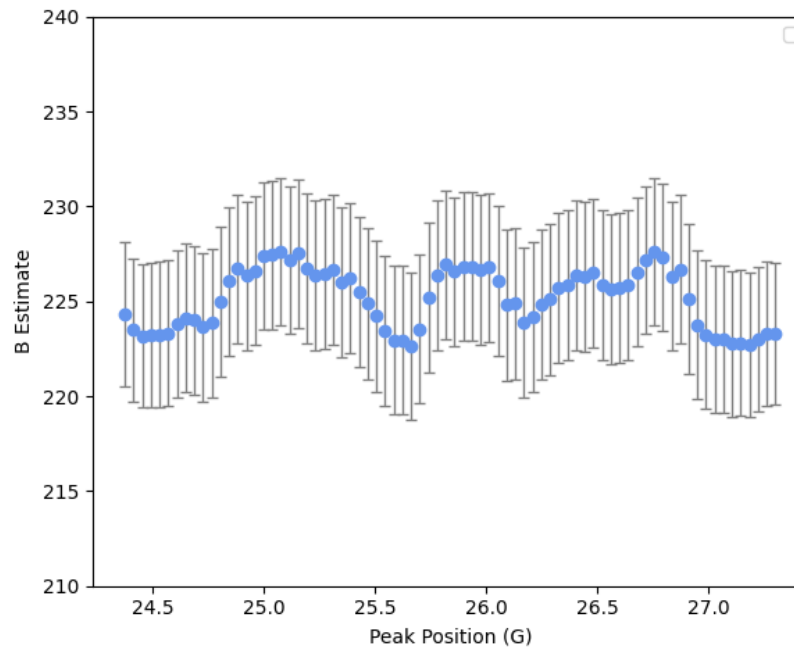
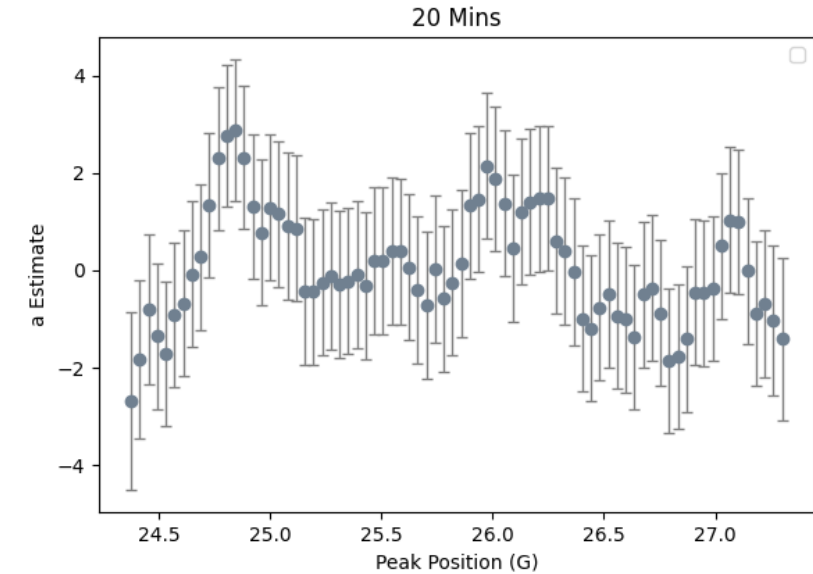
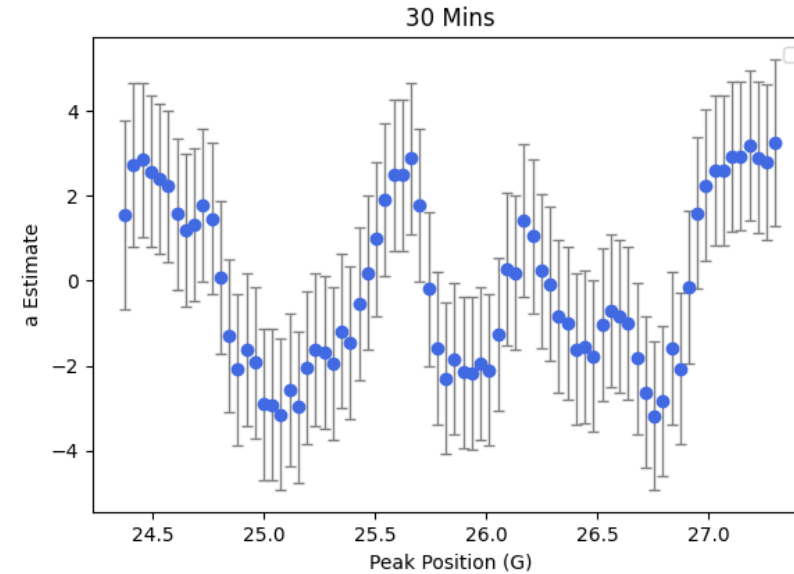
- ‘what if the peak was here?’ x76

B estimates vary within about 5 and are realistic for our measured points

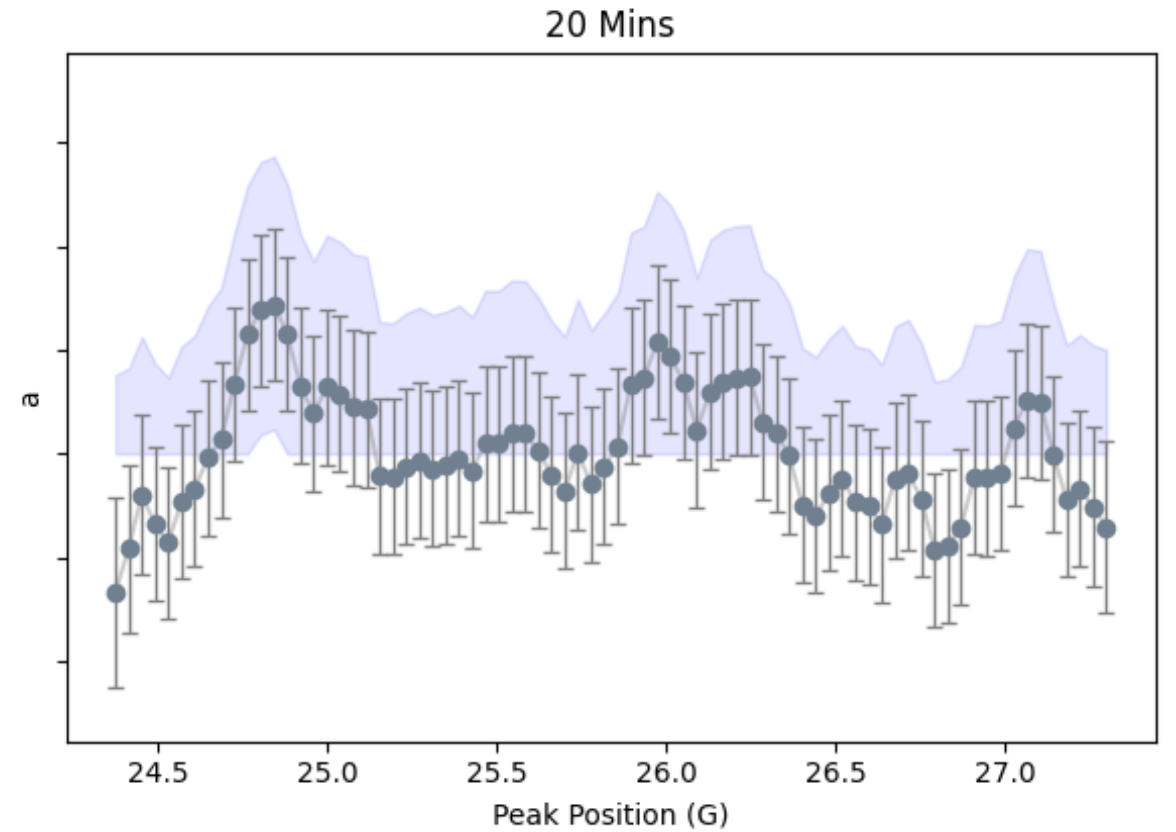
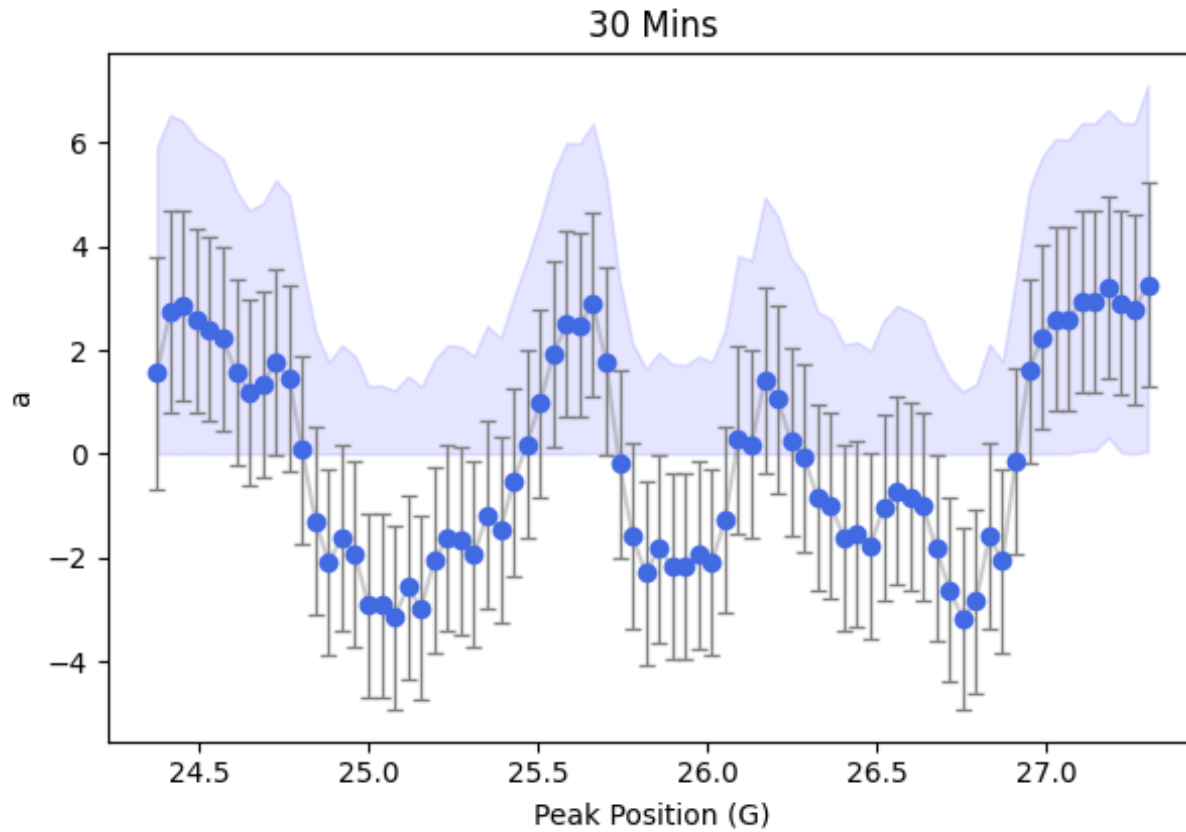
A and B estimates are inversely related to each other

- In real life BG should be independent
- Errors come from fit's correlation matrix

a and B Estimates with Peak Fixed at each Point in Magnetic Field



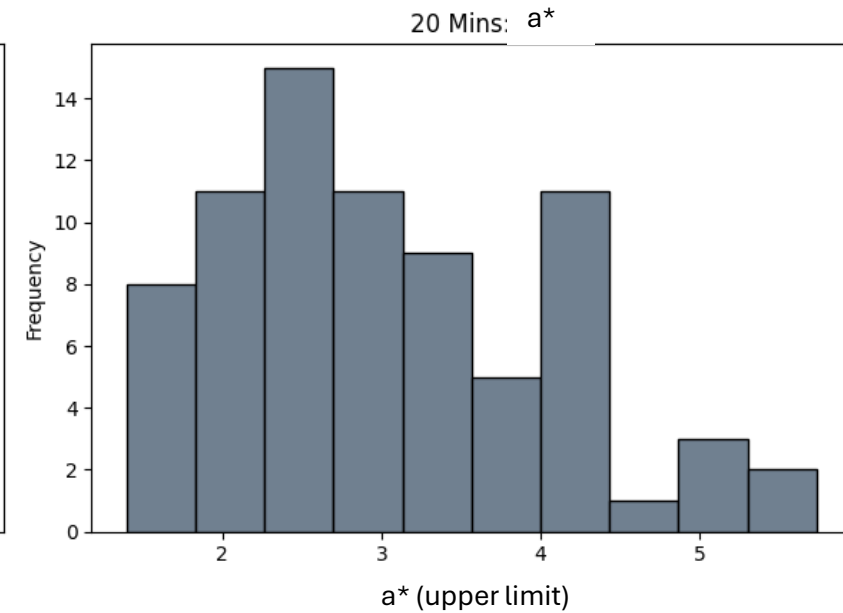
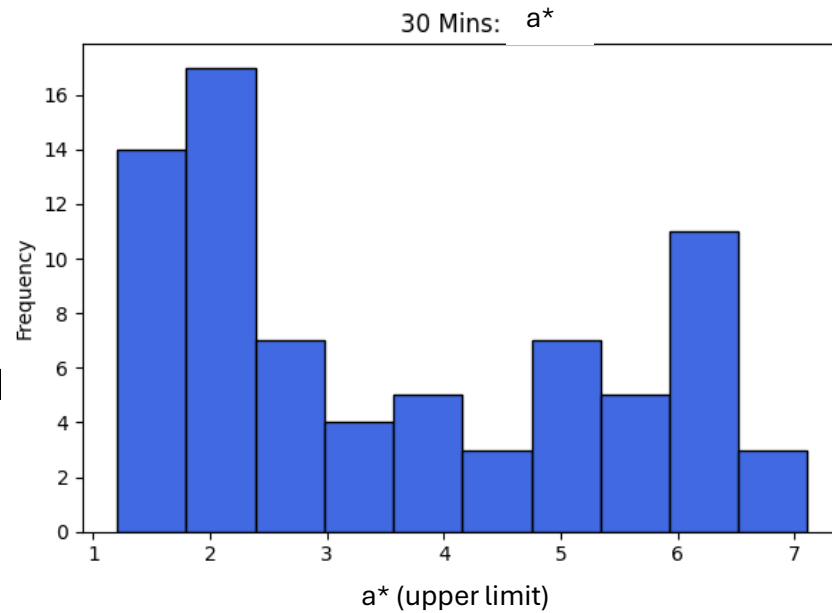
95% FC Intervals (a)



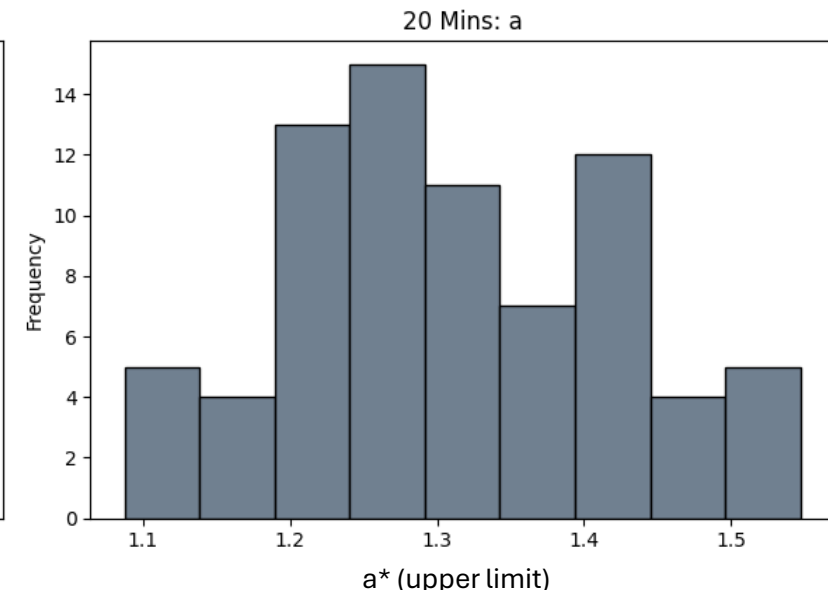
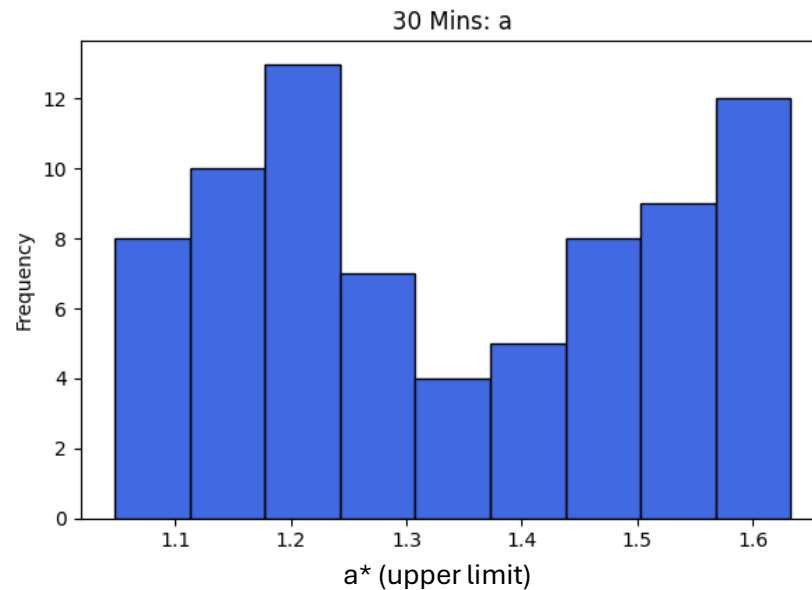
- Using FC limits for the gaussian bounded at 0 case (table 10 in the paper)
- $x_0 = \frac{aEstimate}{aEstimateError}$ where aEstimateError is given in the correlation matrix from the fit
- Rescale by aEstimateError to get the interval limits back in terms of a (shown in plot)

- Comes from the real data, one experiment for each length of time
- Histograms built from 76 a^* estimates (result from each fit)
- Using the confidence interval upper limits removes values for a negative effect that we assume is impossible

Distributions of a parameter upper limit (from fits to experimental data)



Distributions of upper limit 4th root



Simulating pure background

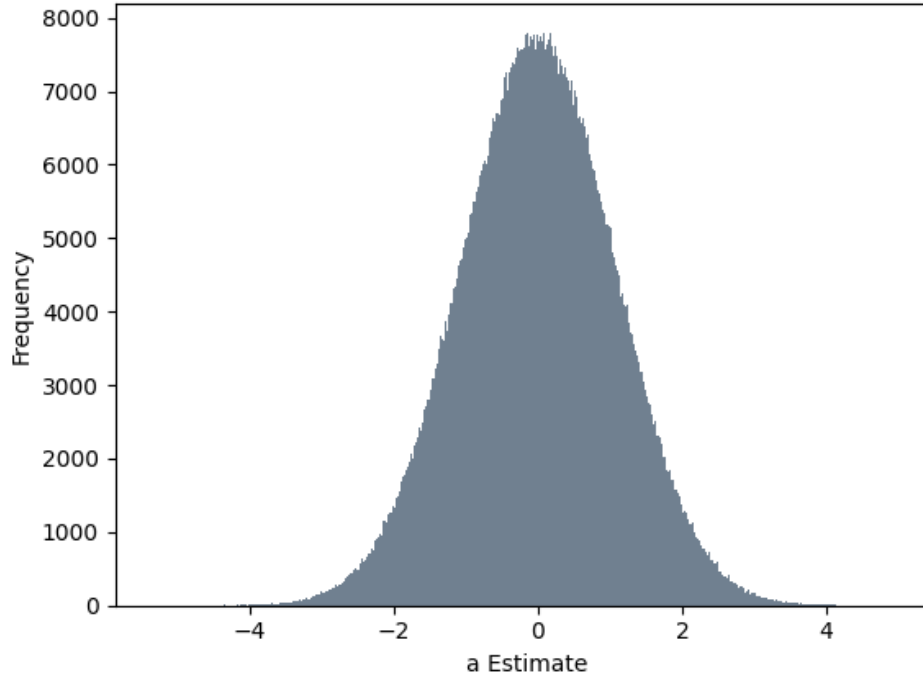
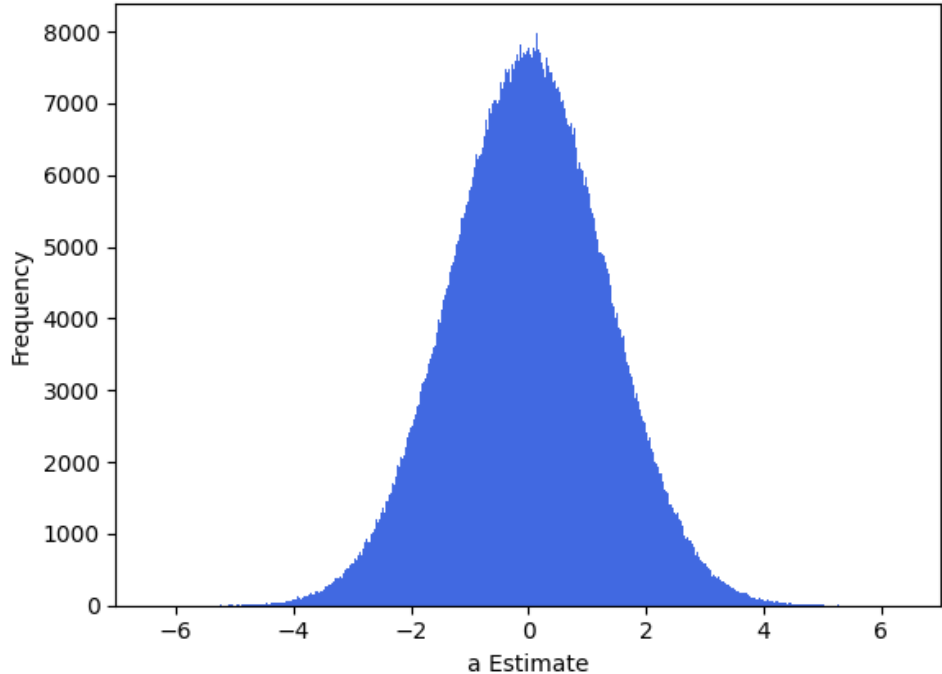
- Generate 1 mil experiments (of 76 runs, same as the real one) with only poisson (approximately normal) distributed background
 - Use $\bar{B}$ for μ
 - $\sqrt{\bar{B}}$ for σ
- No effect added
- Repeat the curve fitting procedure on each experiment
 - These should be sets of 76 randomly generated values from the background distribution
- Use same method for finding each estimate's 95% confidence intervals

1,000,000 iterations

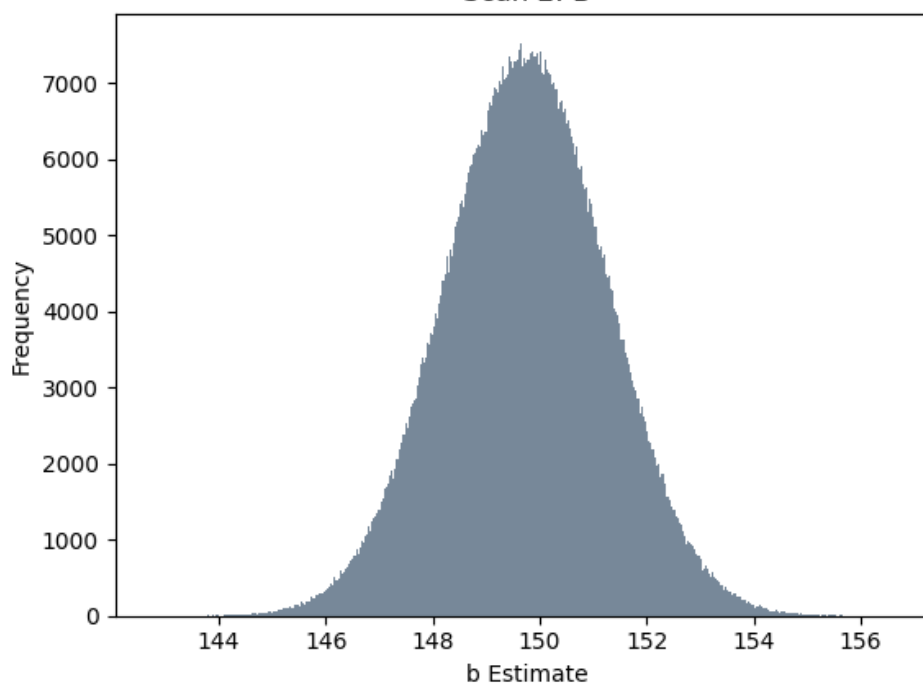
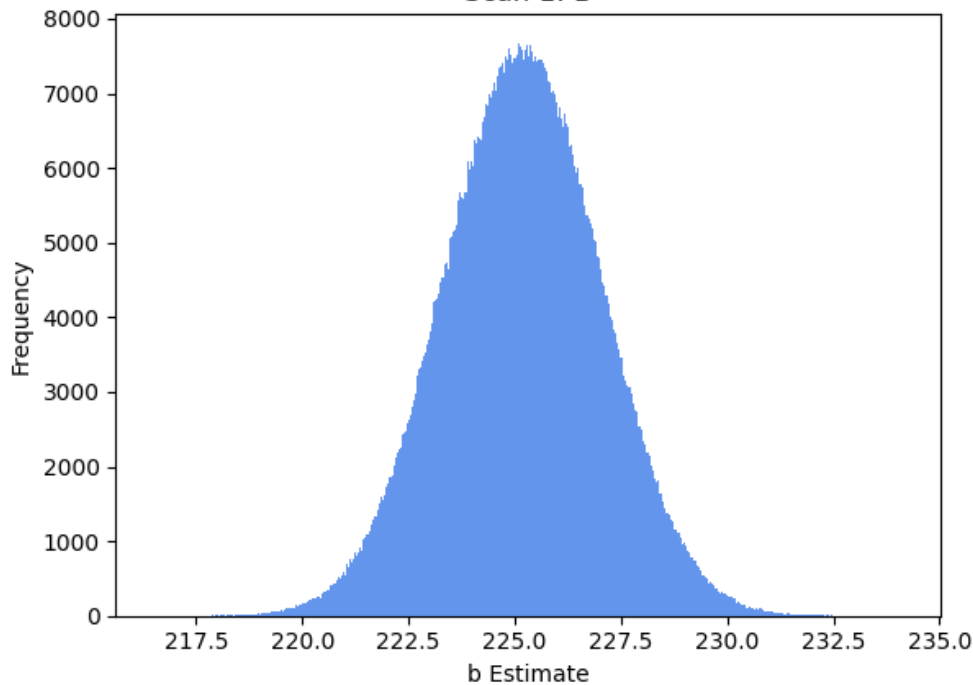
Results from
simulation: what
you'd expect

a centered
around 0 and B
around $\bar{B}$

a



B



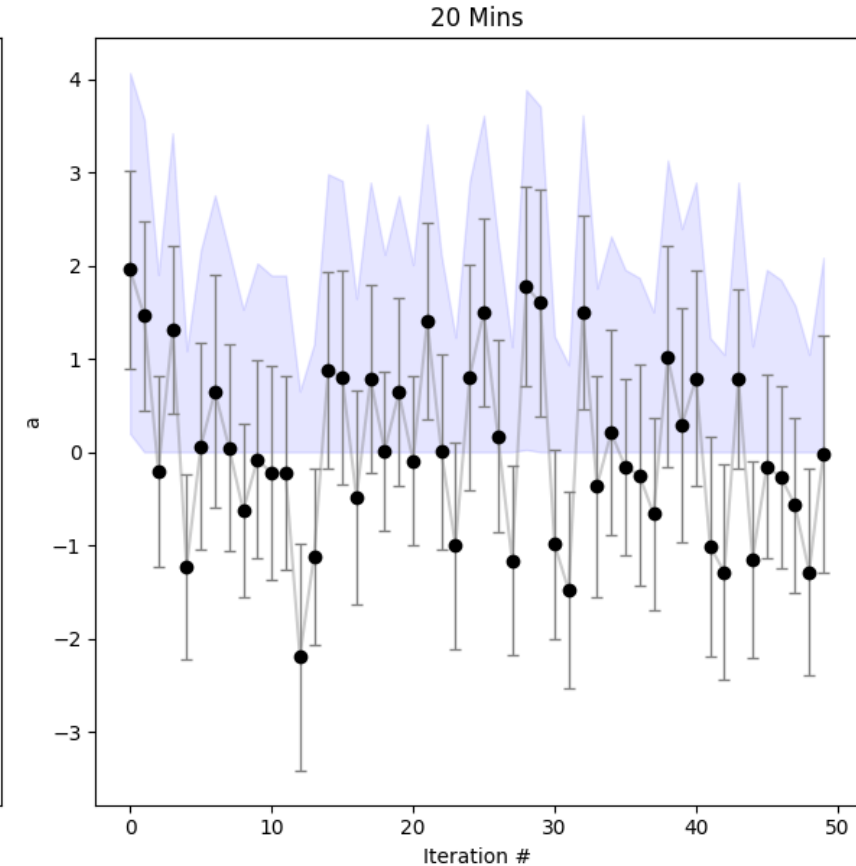
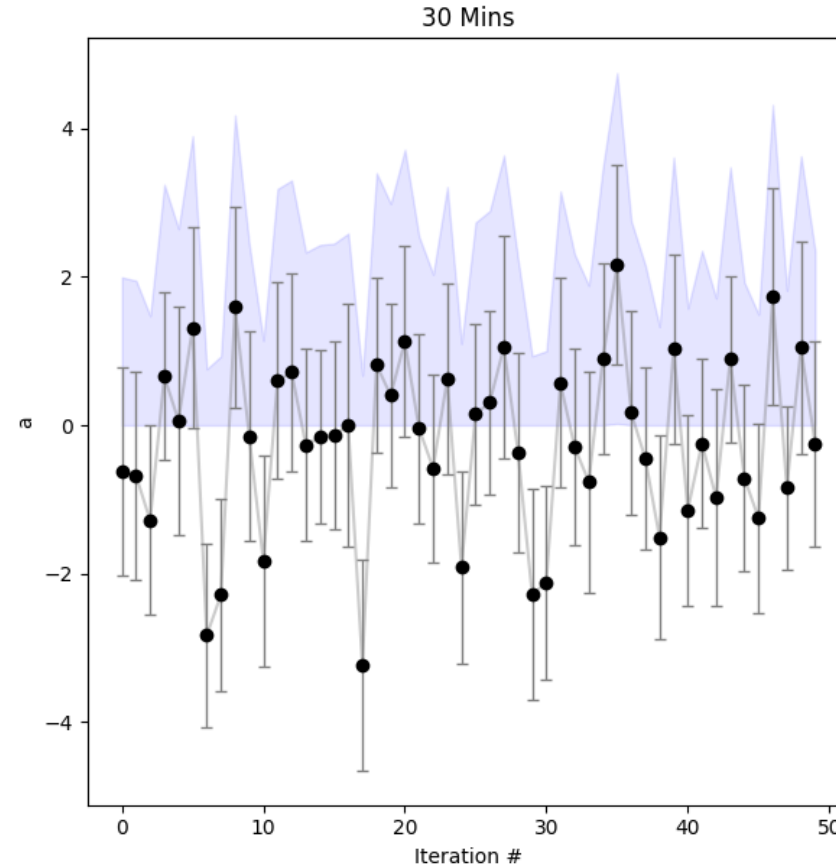
95% FC Intervals for 50 Randomly Sampled Background Fits (a)

Example estimations:

- Plotting 1mil points would not be useful
- Randomly picked 50 to show in the plots with FC intervals
- Here, the distribution the estimates come from is created, so

$$x_0 = \frac{aEstimate}{\sigma}$$

Where σ is the standard deviation of the distribution of 1mil a estimates

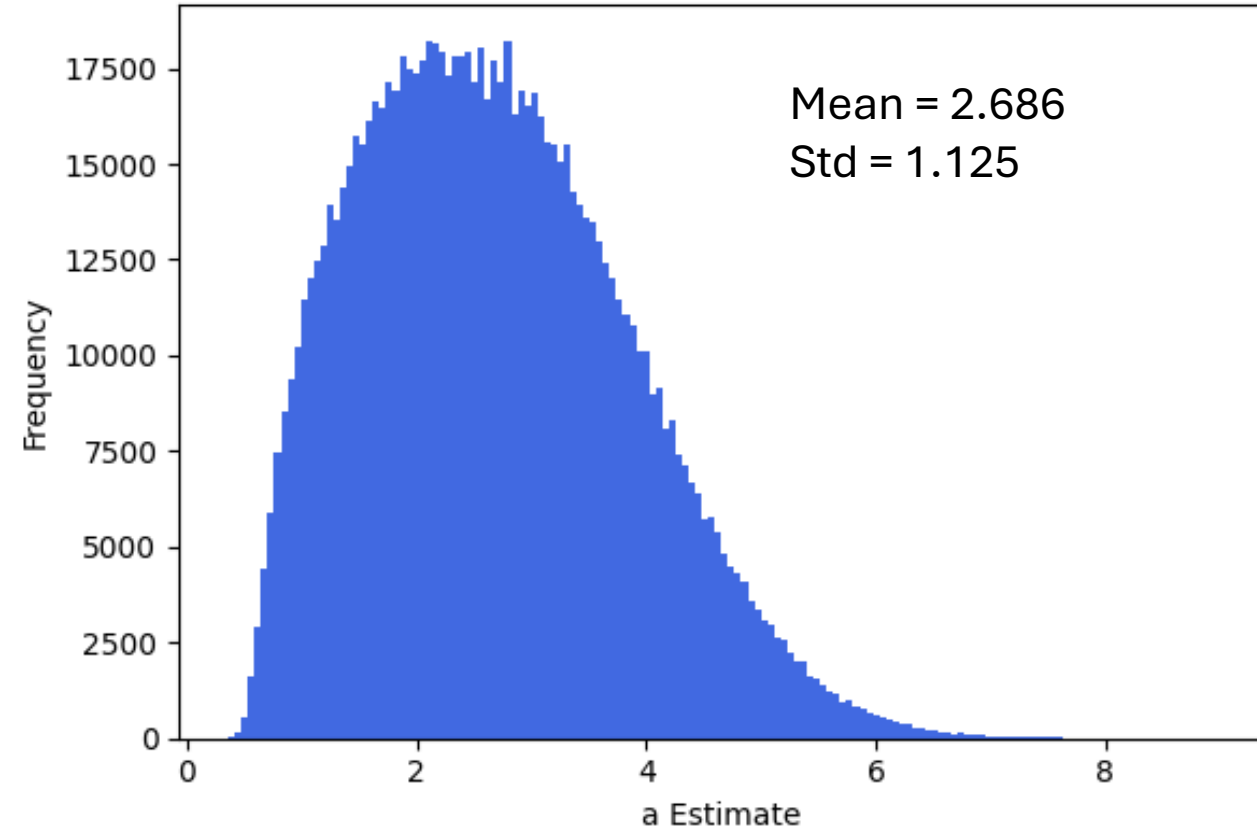


*multiply the interval limits by σ to get them in terms of a

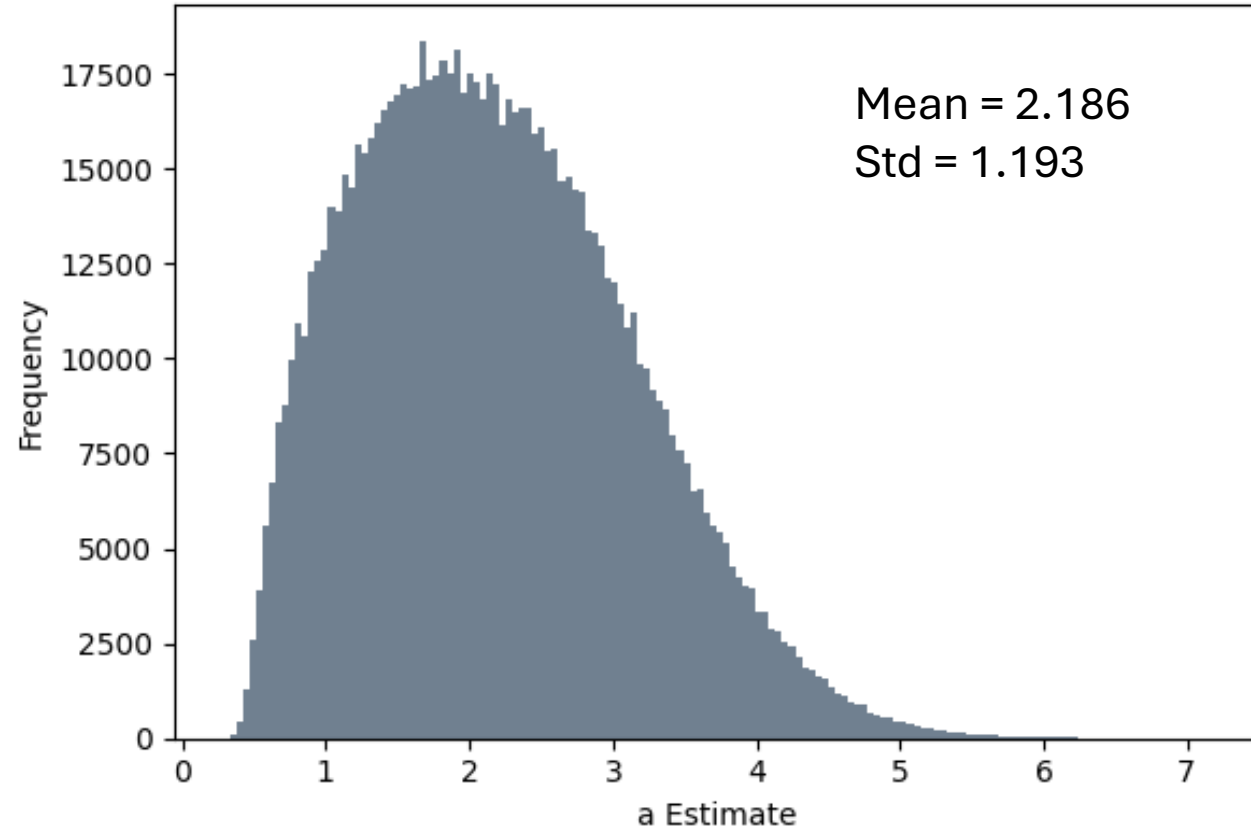
Distributions of upper FC limits (a^*) for background

Distributions of a parameter upper limit (Generated pure background)

30 Mins: a



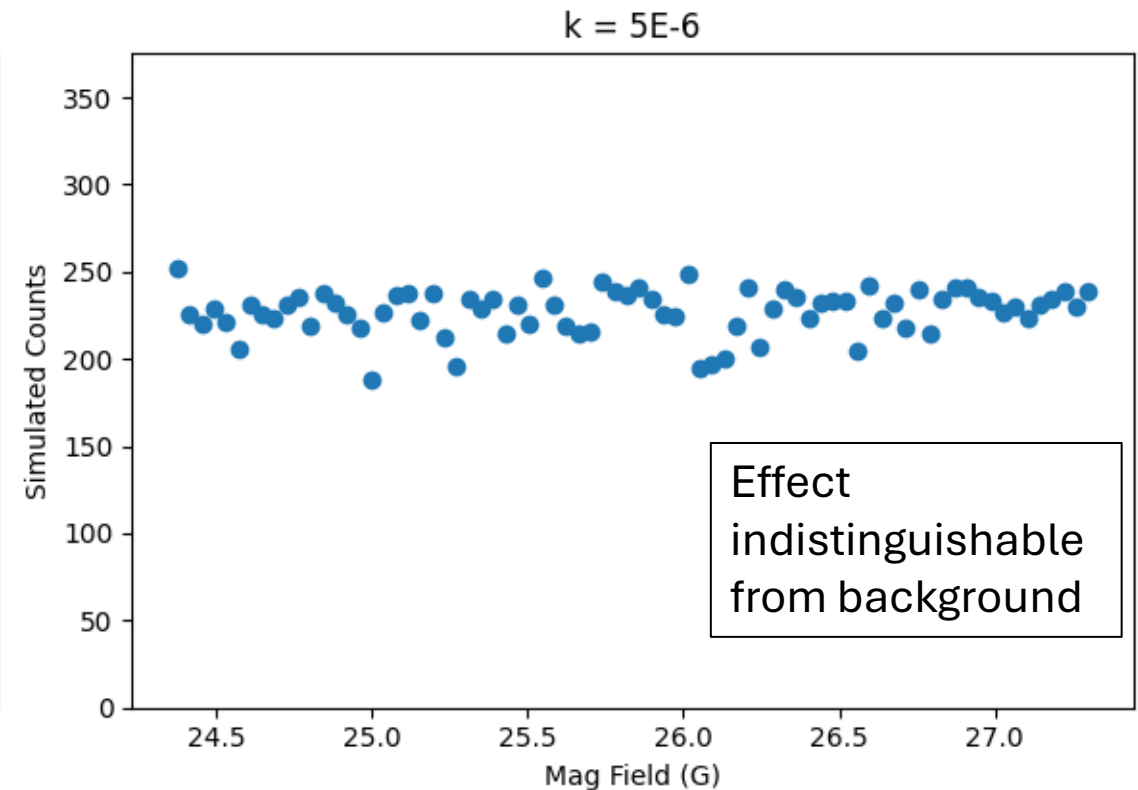
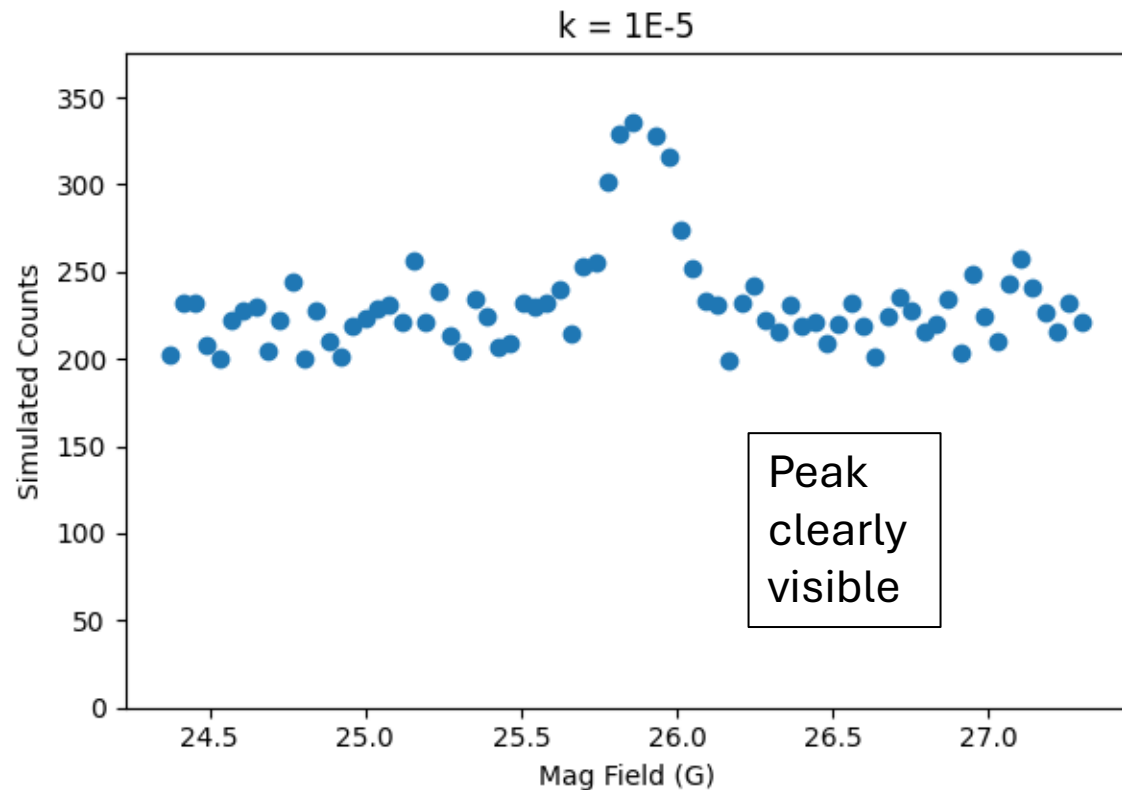
20 Mins: a



Generated distribution with effect

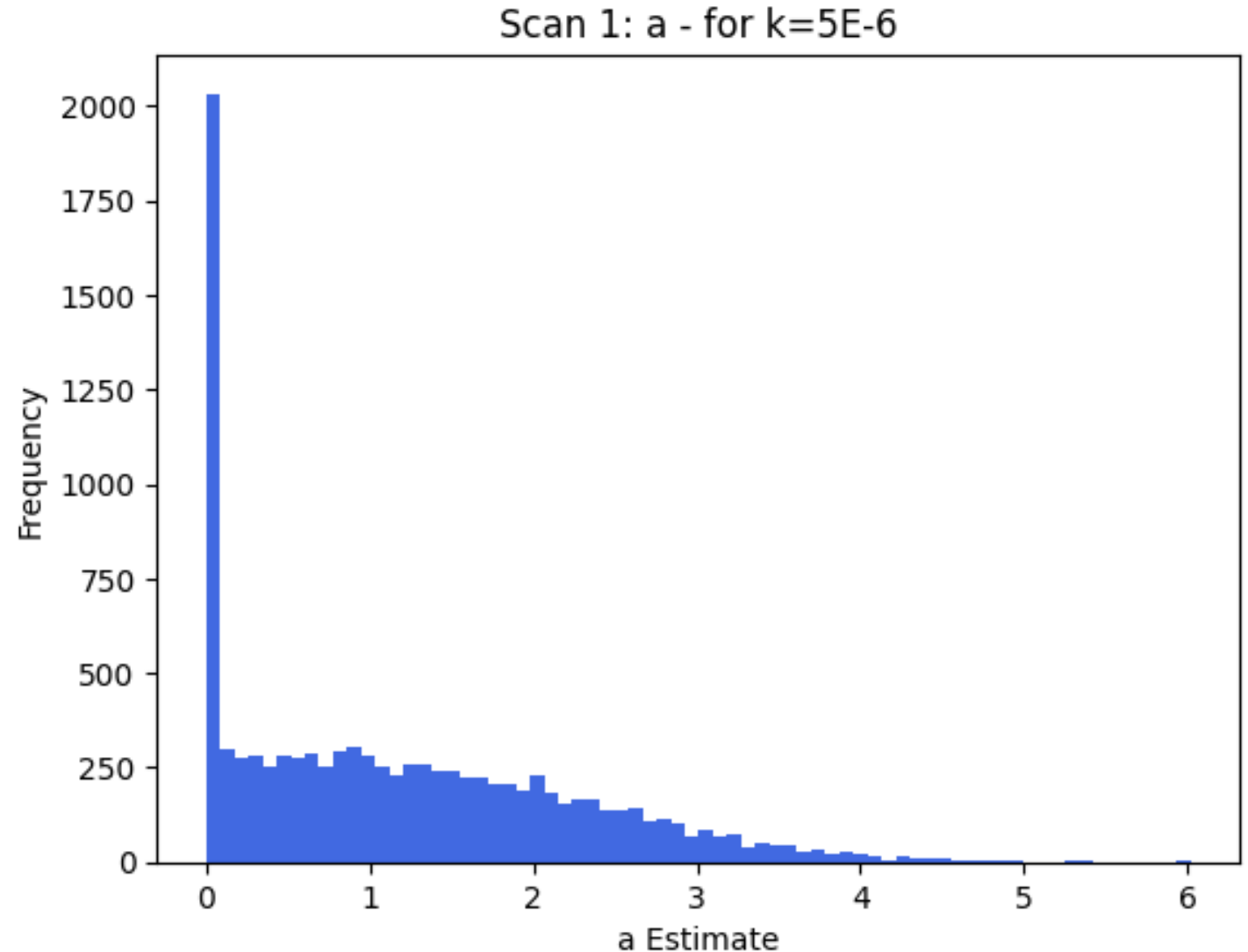
- Meant for calibration of a parameter (how does it relate to k ?)
- Even though we assume a $5E-6$ effect, use $1E-5$ for this case only for clarity.
- Effect is added on top of the background like usual

Example Generated Distributions w/ Effect - 30 Mins

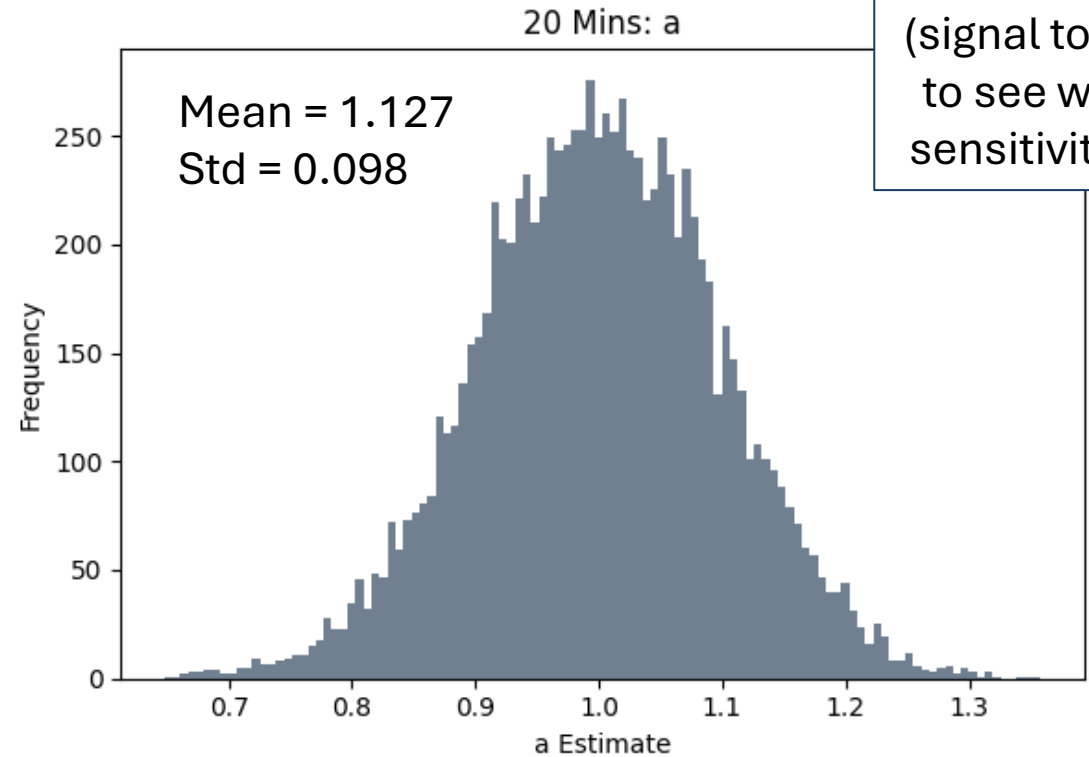
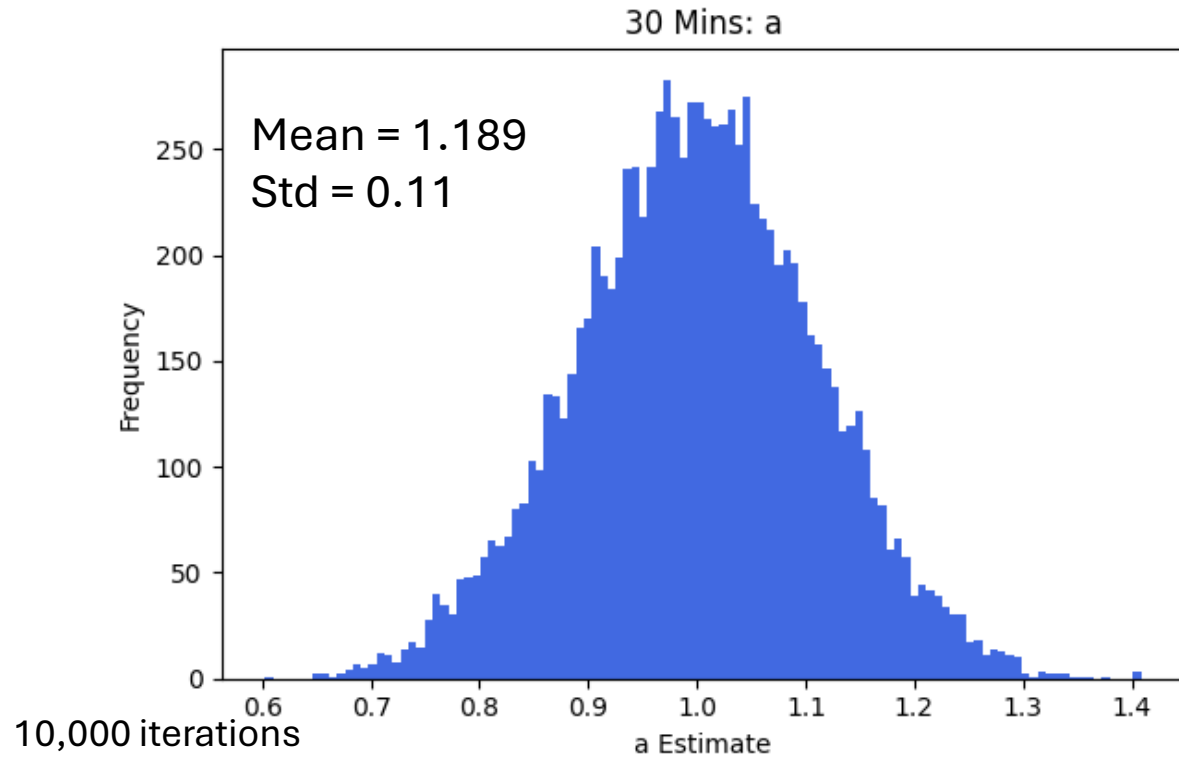


Result for fit with function for $k=5E-6$ (why we won't use it)

- Points less than 0 are forced to be 0 due to the boundary set for the fit
- No confidence intervals because the error can't be estimated (consequence of fitting at a single data point)
- We fit only with the peak point for calibration because A is related to it



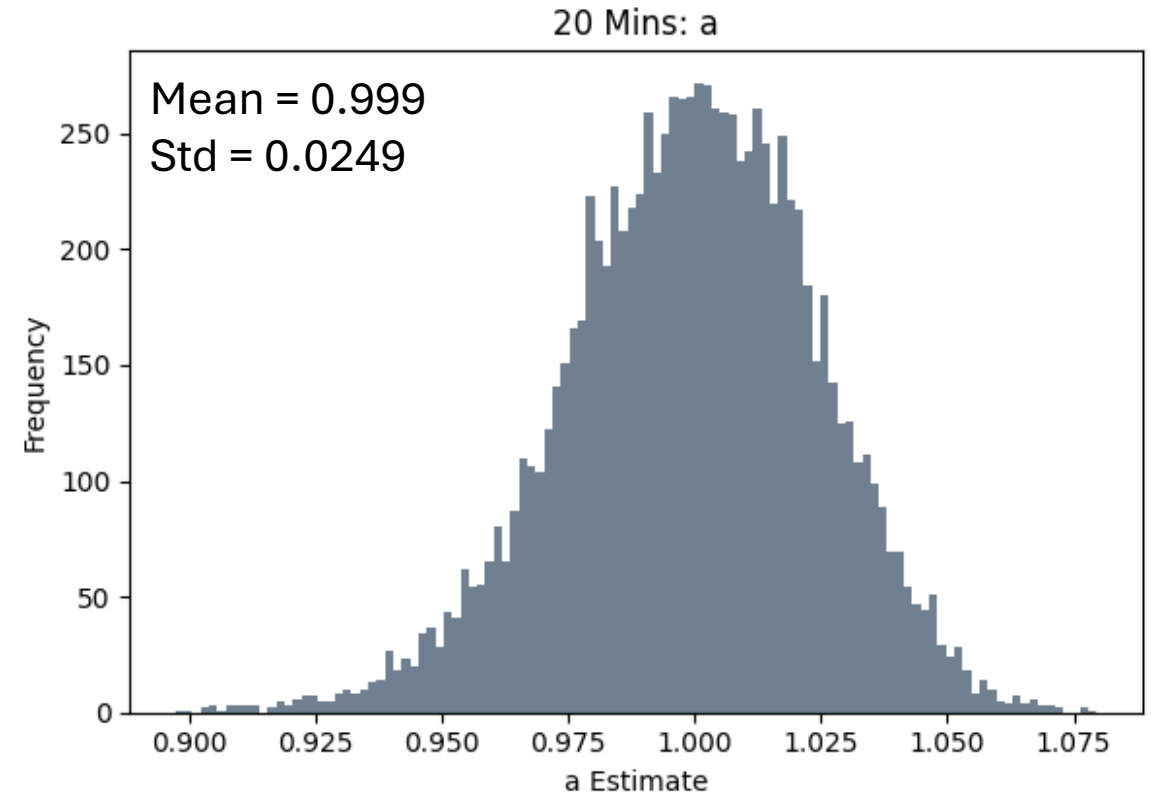
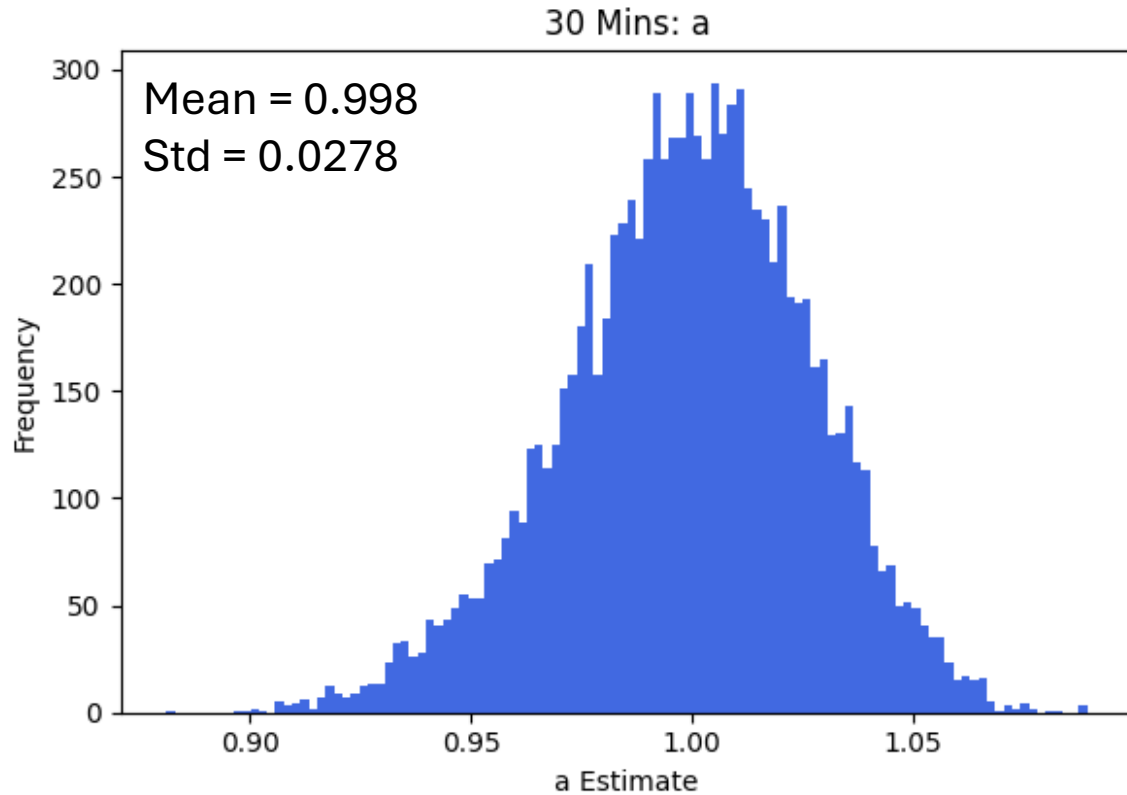
Distributions of a parameter for Generated Distributions with Peak



For if $k = 1\text{E-}5$
instead of $5\text{E-}6$
(signal too small
to see with our
sensitivity level)

- Had to force $a > 0$ for this to run because it is fit only with the peak point
 - Single point curve fits have INF correlation matrices
- a is proportional to κ^4 , so $\kappa = c\sqrt[4]{a}$ where c is some constant.
- These effects are all positive due to the higher sensitivity. No FC limit needed for cutting negatives

Distributions of a parameter for Generated Distributions with Peak (1/4 root)



Histogram of the 4th root of the values from the distributions on previous slide

Finding a_c (a for calibration)

Finding upper limit of k from our data sets & calculations

Notation:

- a : estimate of a parameter
- δ_a : error of a estimate from correlation matrix from fits to experimental data
- a^* : upper limit of Feldman-Cousins confidence interval for a given a estimate
- $\kappa = 5E - 6 * \sqrt[4]{a^*}$ in all cases but for calibration

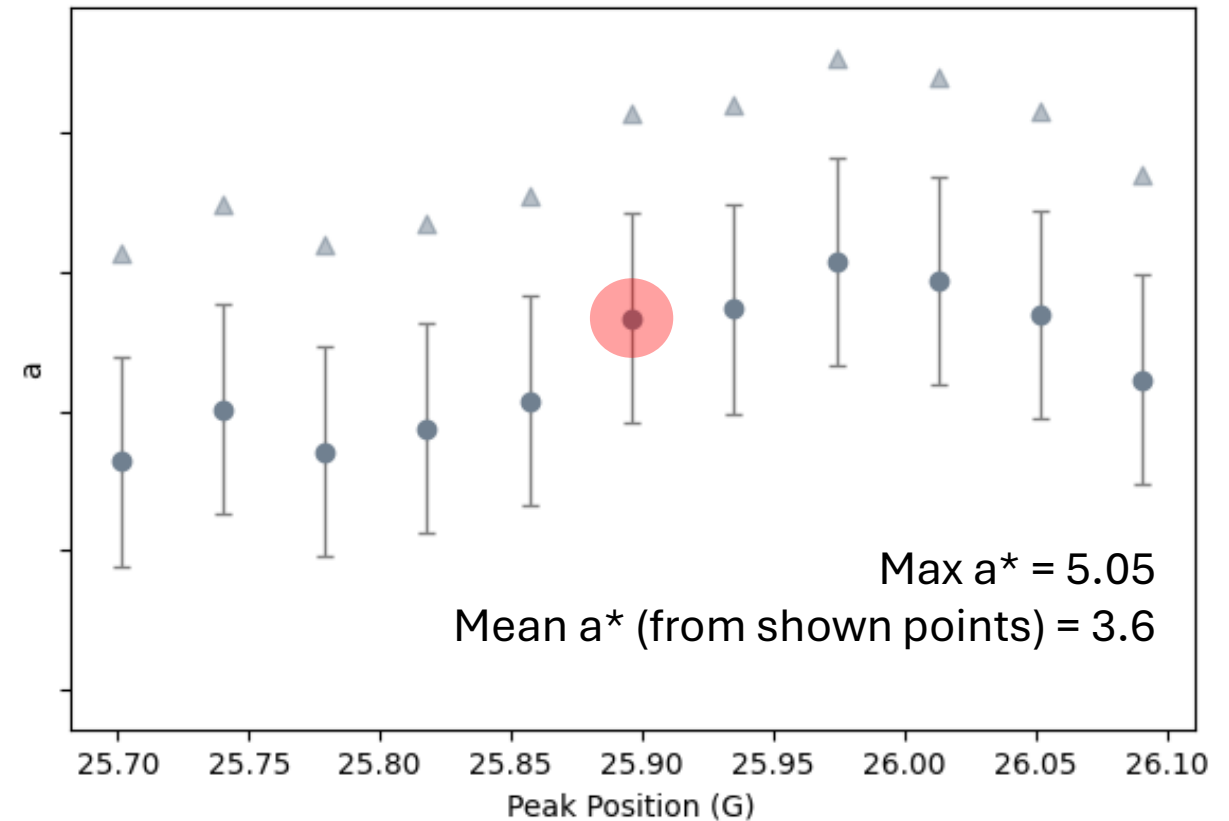
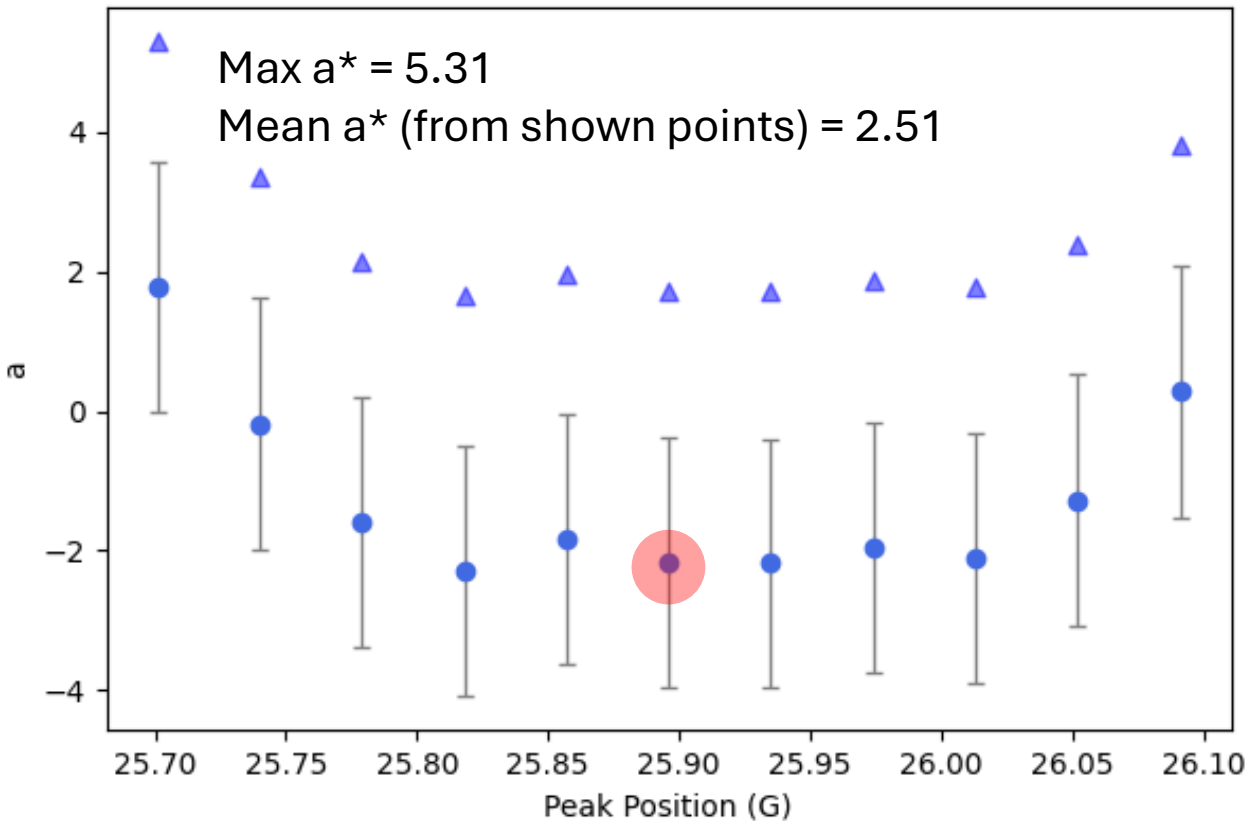
a and a* around the expected peak position for fits to experimental data

▲ a*
● a

25.896G

30 Mins

20 Mins



$$\sqrt[4]{\max(a^*)} = \sqrt[4]{5.31} = 1.518 * 5 \times 10^{-6} = \mathbf{7.59 \times 10^{-6}}$$

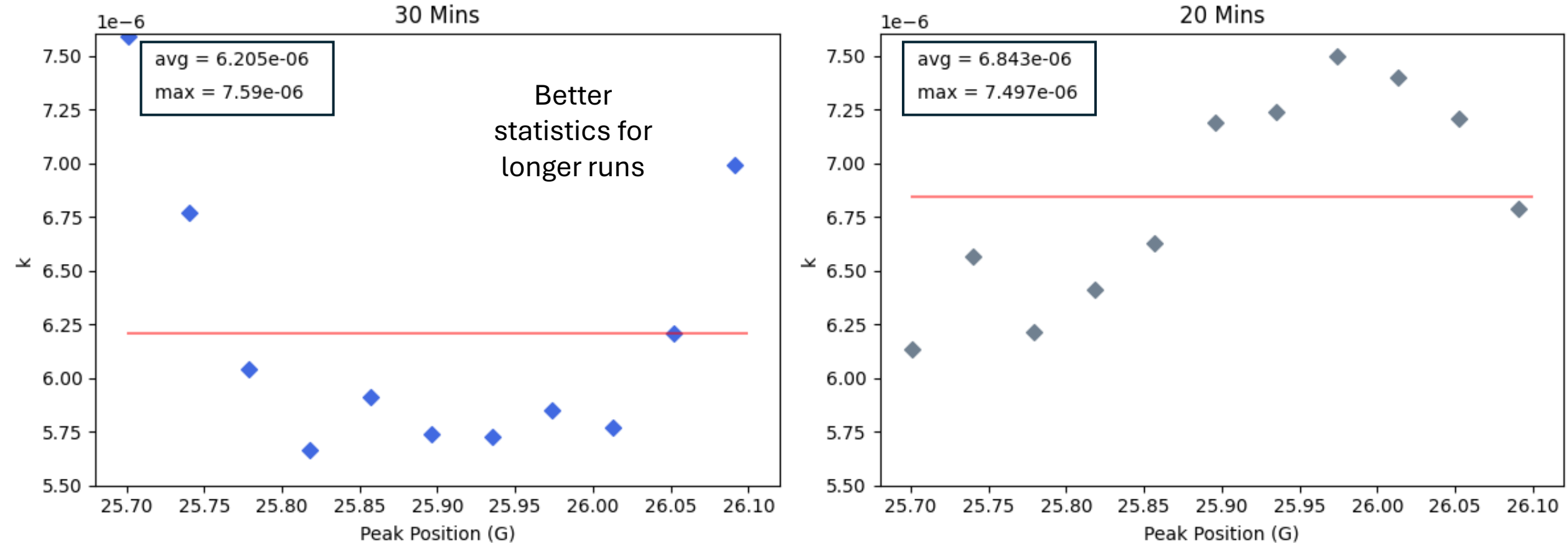
$$\sqrt[4]{\text{mean}(a^*)} = \sqrt[4]{2.51} = 1.259 * 5 \times 10^{-6} = \mathbf{6.295 \times 10^{-6}}$$

$$\sqrt[4]{\max(a^*)} = \sqrt[4]{5.05} = 1.499 * 5 \times 10^{-6} = 7.495 \times 10^{-6}$$

$$\sqrt[4]{\text{mean}(a^*)} = \sqrt[4]{3.6} = 1.383 * 5 \times 10^{-6} = \mathbf{6.915 \times 10^{-6}}$$

Upper FC Limits in terms of κ

From experimental data around 25.896 G. $\kappa = 5E-6(\sqrt[4]{a^*})$



- In ITPS proposal, our experiment's goal is $5E-6$ was the level of κ to reach. Came close to original goal but pressure difference during the run lowered sensitivity
- In 2026 run, sensitivity will be improved by 81x with constant pressure and better shielding. Factor ~ 3 improvement in κ limit. Ideally the limit will go from 6 to about 2

Final Result From Experimental Data

30 mins

$$\kappa = 6.205\text{E-}6$$

20 mins

$$\kappa = 6.84\text{E-}6$$

Final Results From Simulations

Pure Generated Background (for $k=5E-6$)

From simulated 30 min runs

$$\sqrt[4]{\text{mean}(a^*)} = \sqrt[4]{2.686} = 1.28 * 5 \times 10^{-6} = \mathbf{6.4 \times 10^{-6}}$$

From simulated 20 min runs

$$\sqrt[4]{\text{mean}(a^*)} = \sqrt[4]{2.186} = 1.21 * 5 \times 10^{-6} = \mathbf{6.08 \times 10^{-6}}$$

Generated distributions with effect (for $k=1E-5$)

From simulated 30 min runs

$$\sqrt[4]{\text{mean}(a_c)} = \sqrt[4]{1.189} = 1.044 * 1 \times 10^{-5} = \mathbf{1.044 \times 10^{-5}}$$

From simulated 20 min runs

$$\sqrt[4]{\text{mean}(a_c)} = \sqrt[4]{1.127} = 1.03 * 1 \times 10^{-5} = \mathbf{1.03 \times 10^{-5}}$$

Final κ value table

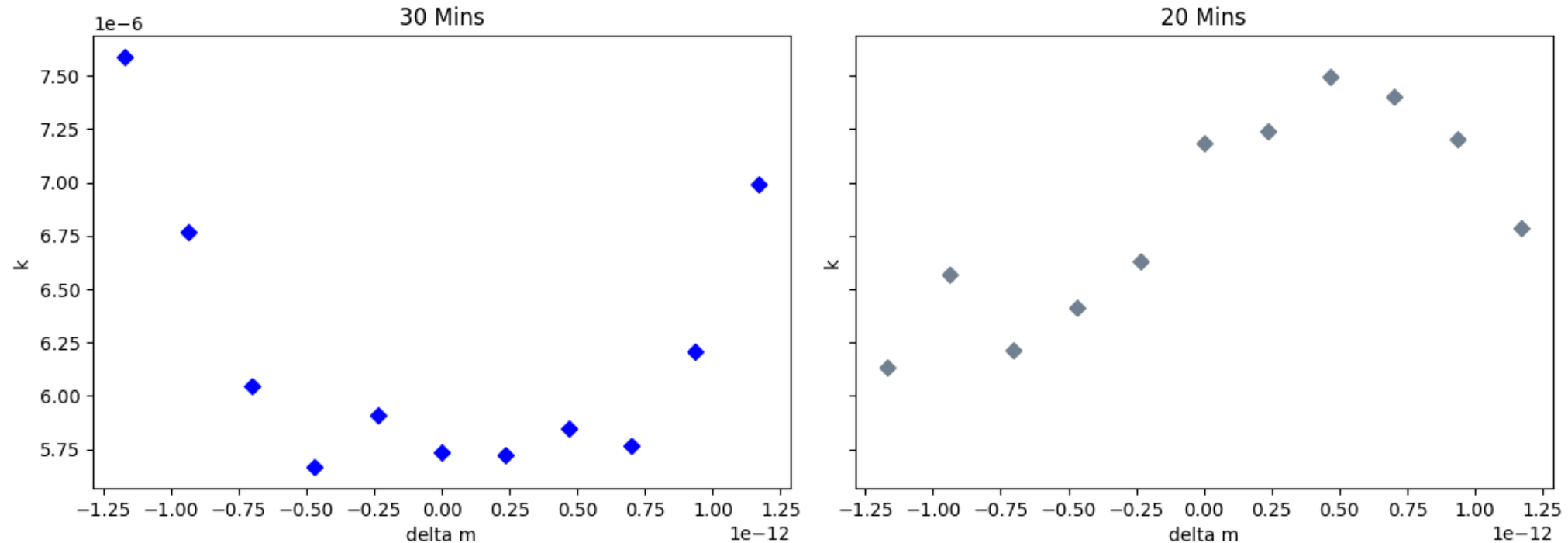
Source	30 mins	20 mins
Experimental	6.205E-6	6.84E-6
Simulated background	6.4E-6	6.08E-6
Simulated distribution with background (for calibration)	1.044E-5	1.03E-5

Additional

In terms of δ_m (a different x axis)

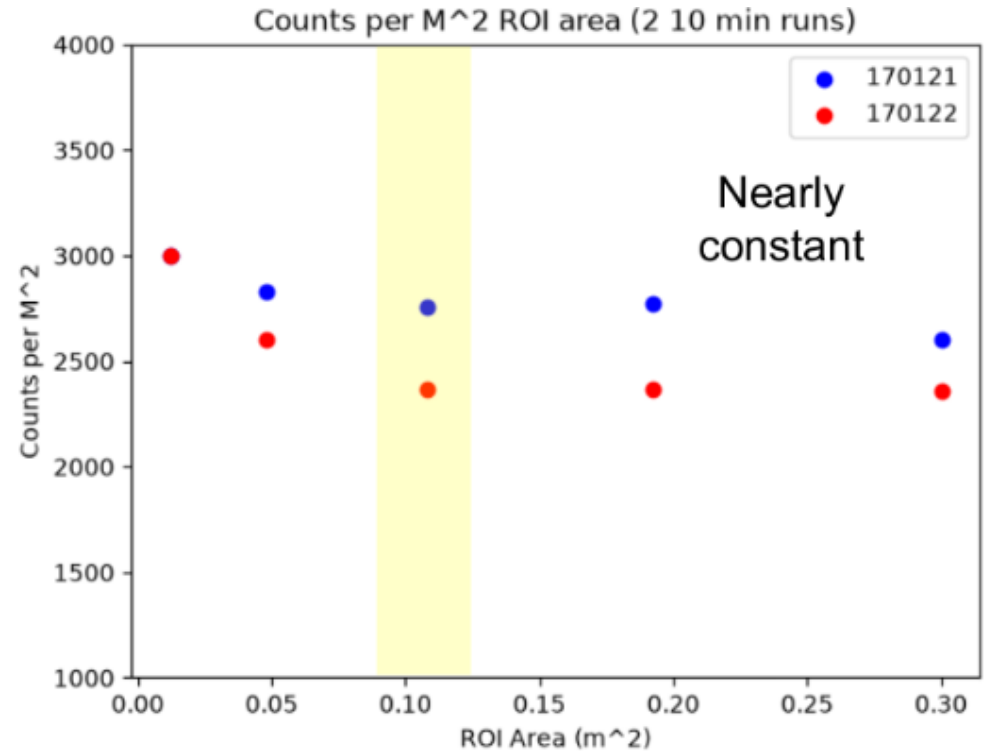
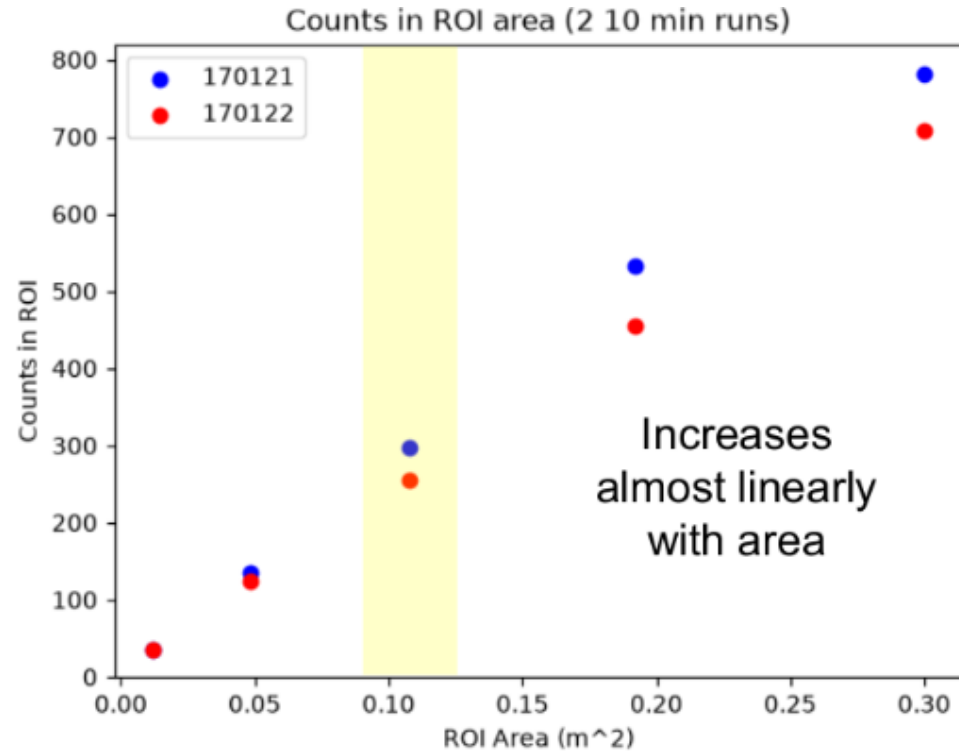
- Using the experimental data, take distance from expected peak (B in G) for each position tried. This is positive and negative around 0
- Multiply the distances by 6E-12 to get it in δ_m scale
- For a model where δ_m is another parameter, this shows that for different δ_m we set a different value of κ

Upper FC Limits in terms of K on dm scale



nTMM 2 BG quick analysis

- Need to add errorbars, but relationship is visible



Tried 5 ROI sizes

Smallest: 0.10*0.12

Small: 0.20*0.24

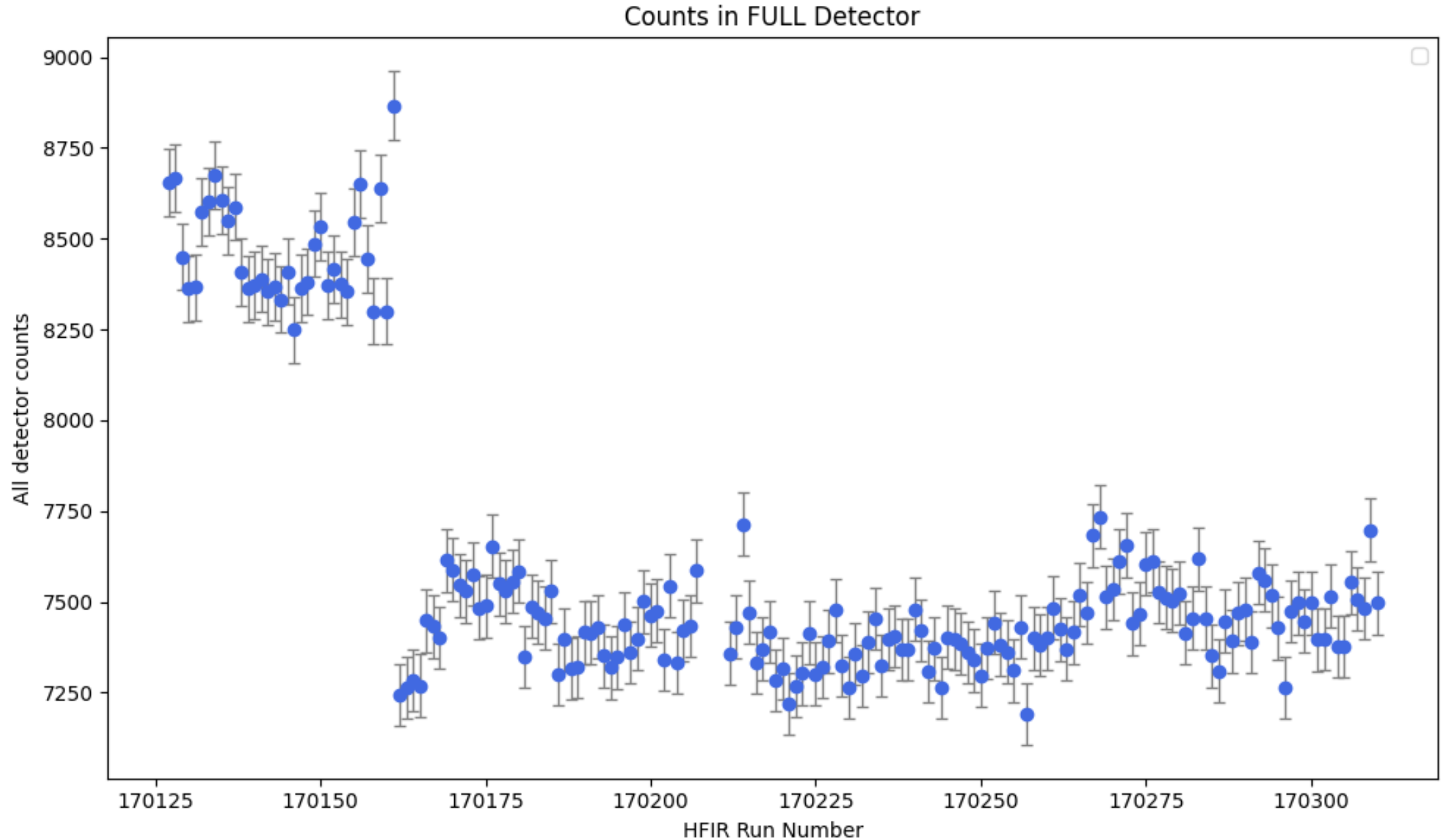
Real: 0.30x0.36

Big: 0.40*0.48

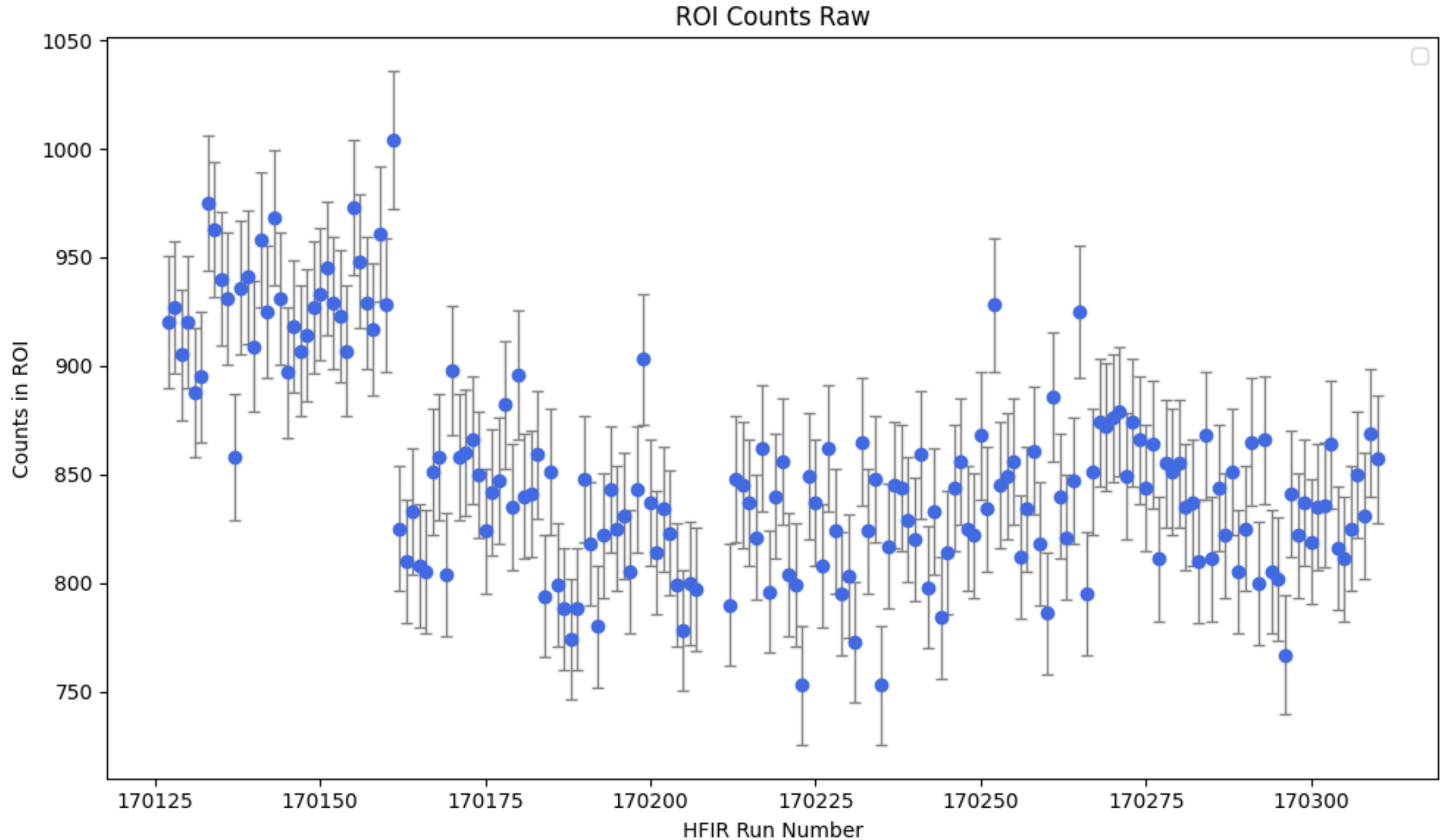
Biggest: 0.5*0.6

Supports that the background is uniform throughout the detector, makes subtraction from ROI counts feasible

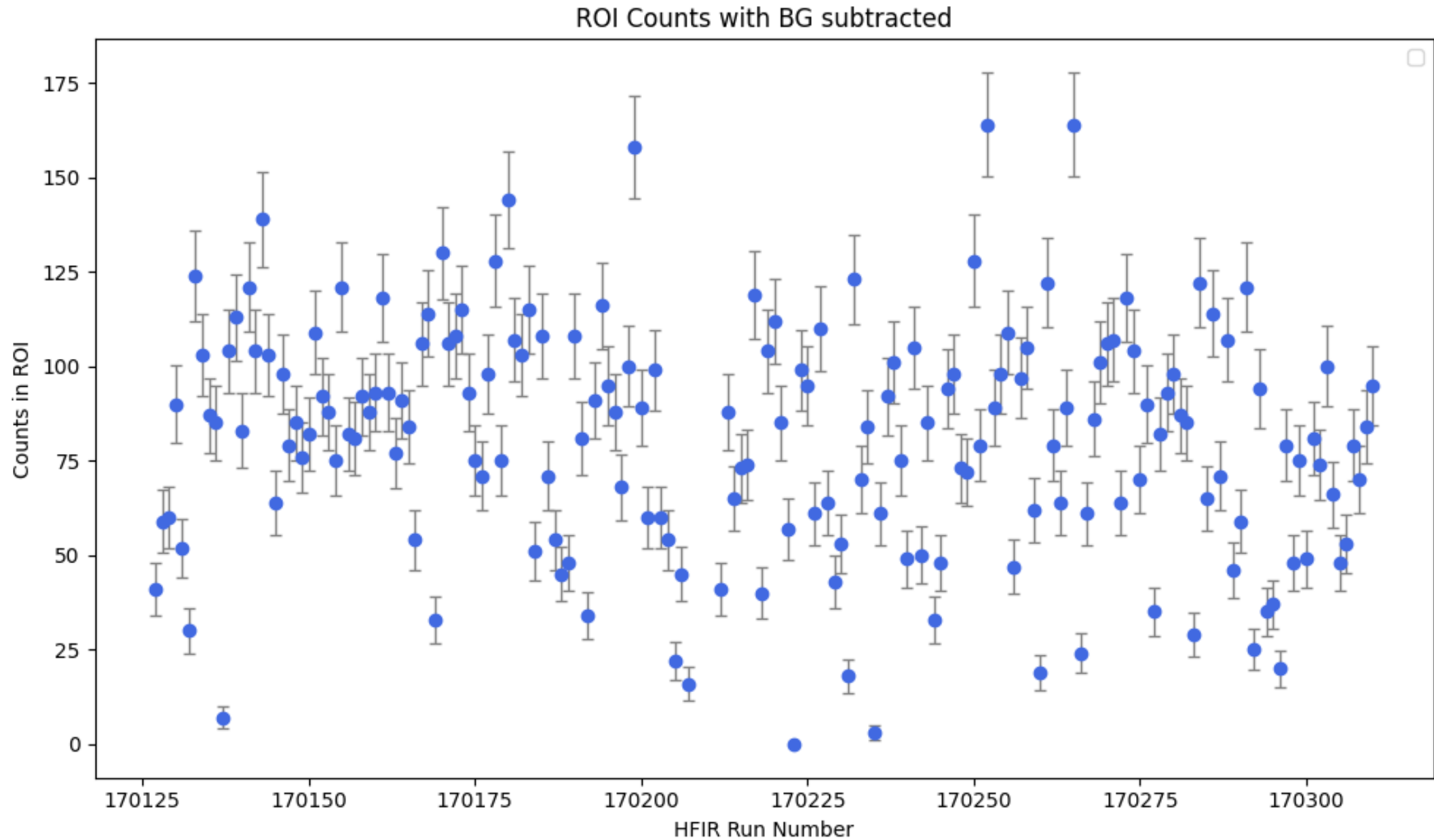
nTMM2 Counts over time (HFIR run num)



nTMM2 Counts over time (HFIR run num) contd.



nTMM2 Counts over time (HFIR run num) contd. 2



Reference Slides

Intensity calculation

$$Intensity = \frac{GPM_{Sec} * DT_{Corr} * Sapp_{Atten}}{GPMEff * 2} * 1800$$

nTMM Intensity Correction Factors	
name	value
GPM_efficiency	3.00E-05
deadtime	1.15
sapphire_window_attn	0.908
al_transmission	0.9
si_attn	0.9315
detector_eff	0.96
reg_ROI_eff	0.90156
reg_filteredROI_eff	0.766972005
si_transmission	0.91122

Counts correction	$\frac{RawROIcounts_i * GPMVariationCorr_i}{AlAtten * SiAtten * DetectorEff * ROIEff}$
Intensity	$\frac{GPMsec * DTCorr * SapphireAtten}{GPMEff * 2}$

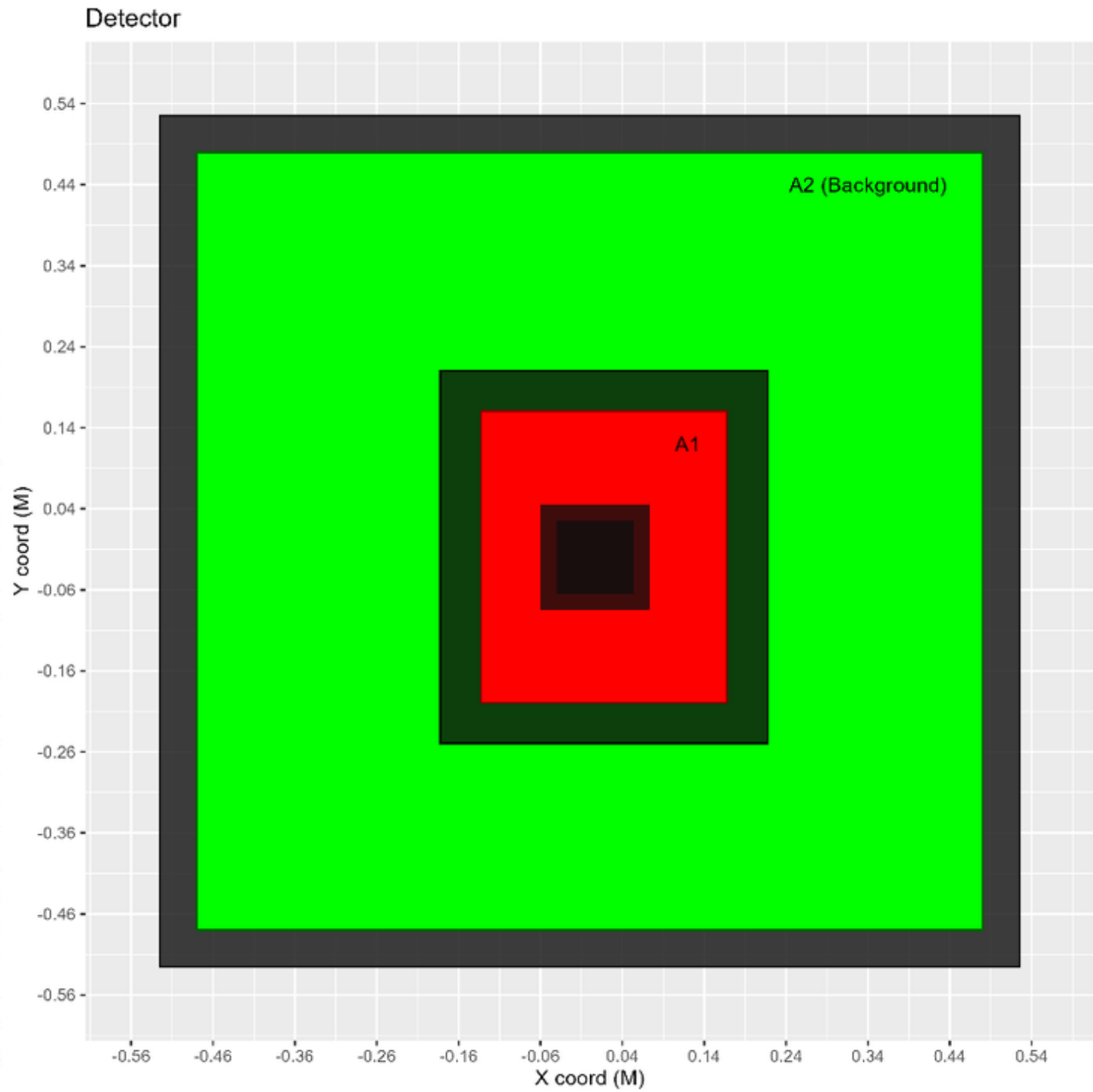
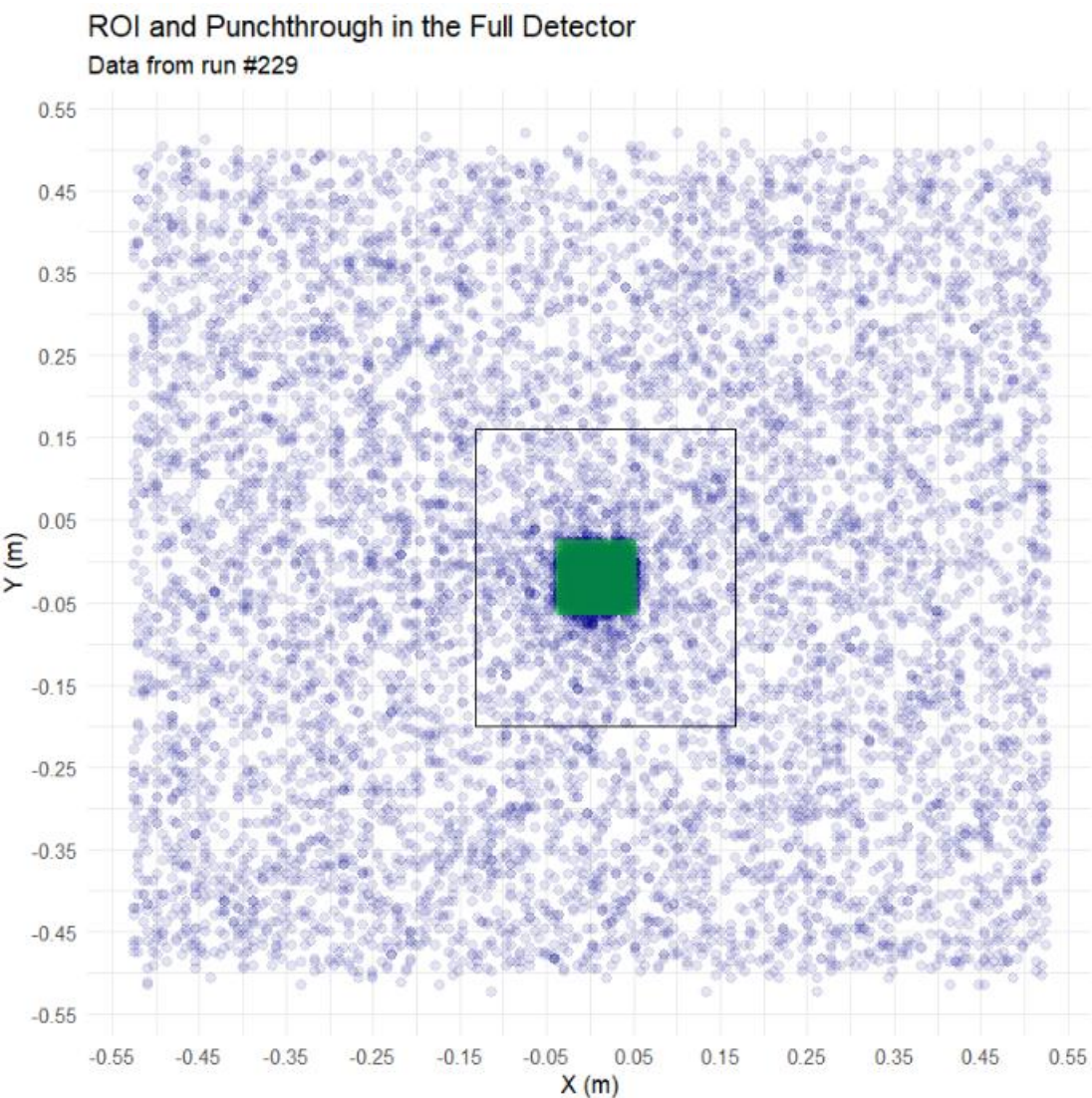
FC Limit method in python

- Clean data
- Define test true μ range 0 to 8 with 10000 points
 - “If μ = the tested value, what x_0 s would be in its acceptance region?”
- Define grid of test x_0 s -9 to 9 with 10000 points
- For each test true μ value calculate likelihood ratio $R(x_0)$ to order the x_0 s for taking them into the acceptance region
- Add x_0 s into the acceptance region in decreasing order of R until the cumulative probability =0.95
- Repeat for every tested μ
- Look at the real x_0 s, find what μ' s acceptance regions they appear in, take min & max of those μ' s to get the limits

- x_0 are the observed a parameter estimate values divided by their error
- Approach from part B of Feldman Cousins paper

$$R(x) = \frac{P(x|\mu)}{P(x|\mu_{\text{best}})} = \begin{cases} \exp(-(x - \mu)^2/2), & x \geq 0 \\ \exp(x\mu - \mu^2/2), & x < 0. \end{cases}$$

ROI and Other Area Selections



ROI and Other Area Coordinates

Area	X start	X end	Y start	Y end
A2 Limits	-0.48	0.48	-0.48	0.48
ROI (A1) Limits	-0.13249	0.16751	-0.19941	0.16059
ROI + 5cm	-0.18249	0.21751	-0.24941	0.21059
<u>Punchthrough</u>	-0.04	0.054	-0.065	0.025
<u>Punchthrough</u> + 2cm	-0.06	0.074	-0.085	0.045
* <u>all</u> measurements in meters				

2025 nTMM-1 Analysis

Carolyn Haviland | nn' Collaboration Meeting | 09/18/2026

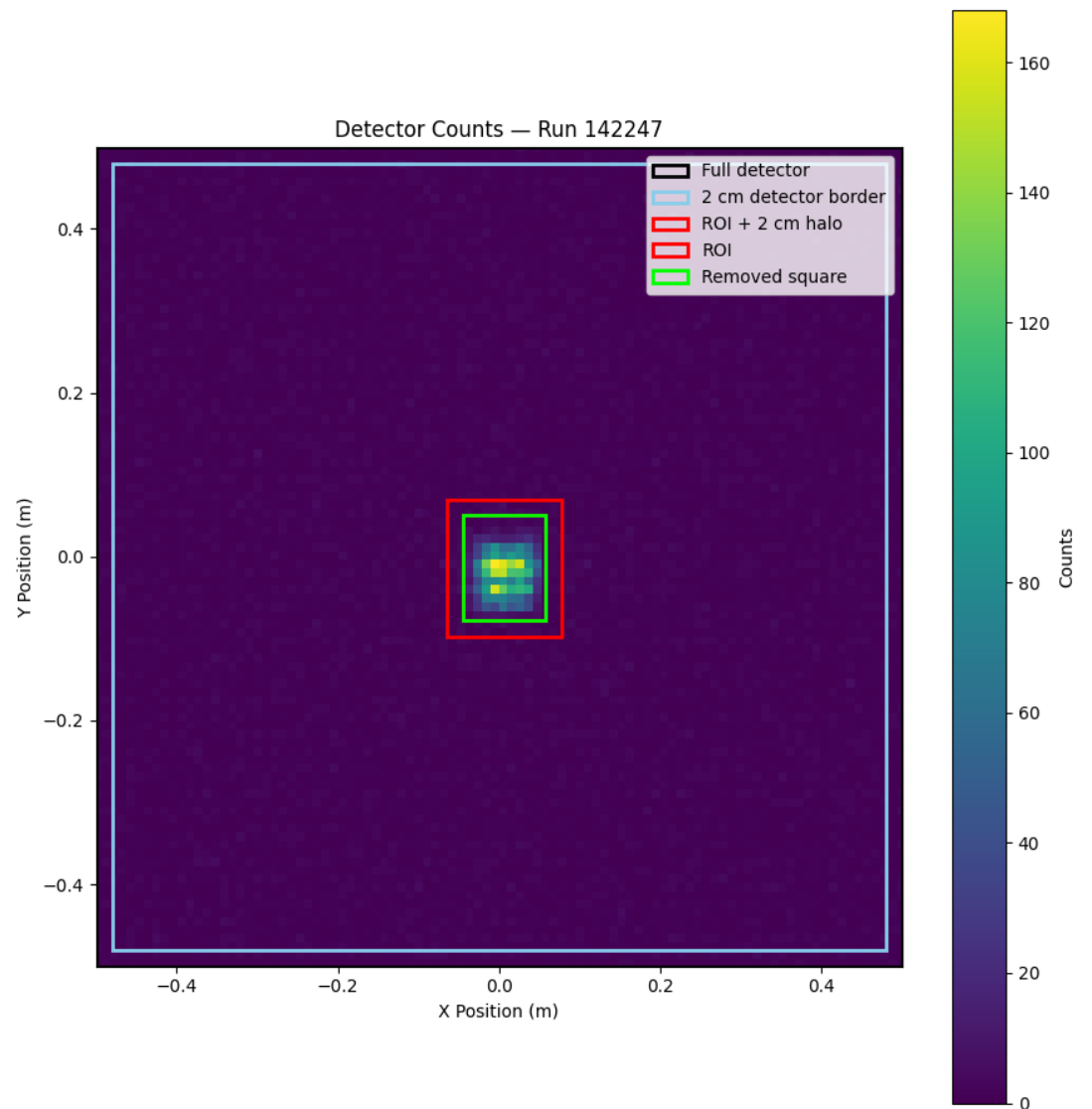
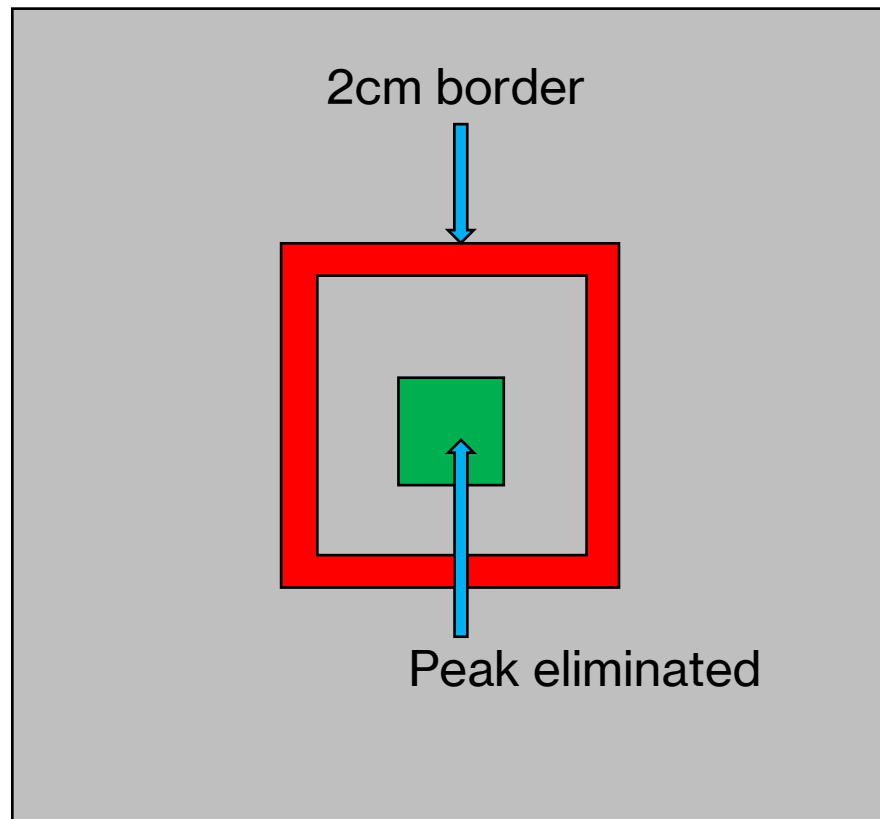
Agenda

- Analysis procedure
 - Baseline counts
 - Fitting
 - α to κ
- Findings
- nTMM-2 analysis

Procedure

1. Plot 2D histogram of each run
 - Make ROI cuts
 - Make peak cuts
2. Output excel with counts for baseline
 - Use area ratios for counts in each section
3. Use counts with fit function
 - Find parameter 'a' at each run
4. Calculate limit assuming no effect
5. Calculate limit with $\kappa = 1e-5$

ROI and Peak Cuts

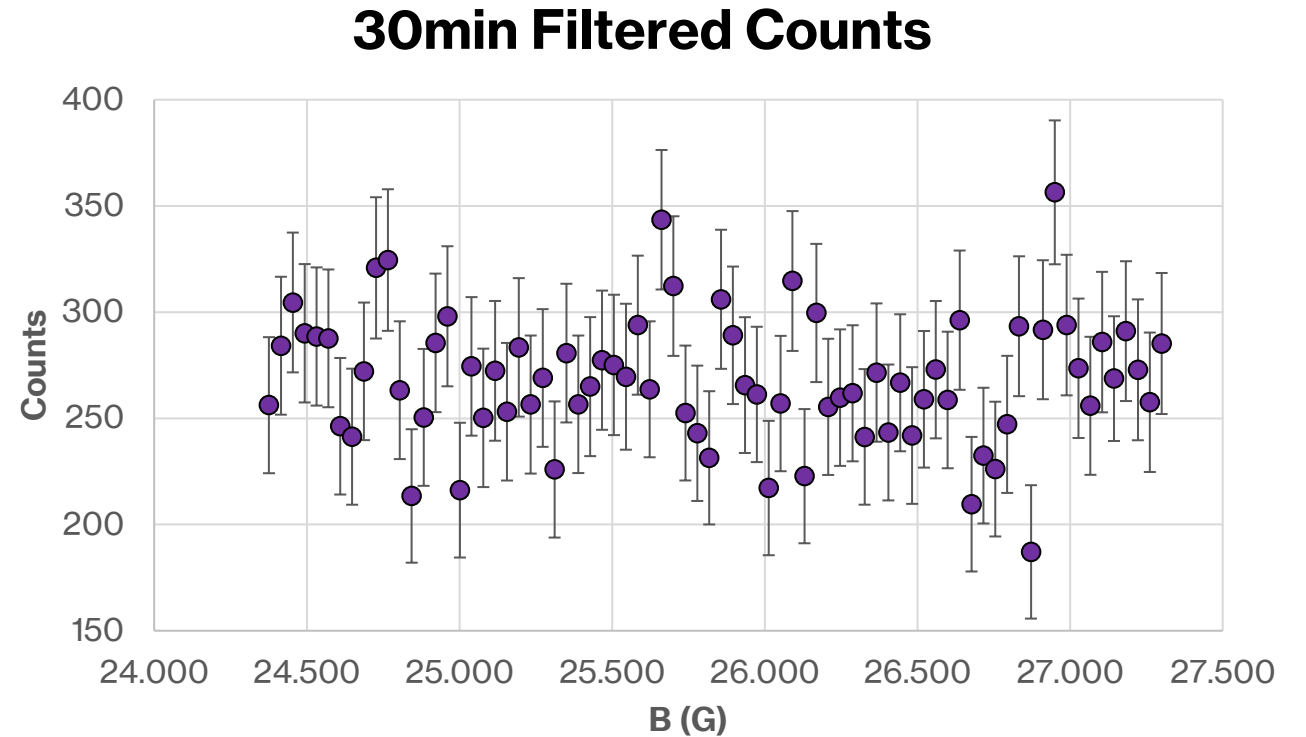


Correction Factors

- Series of correction factors for intensity and counts
- Counts corrected by GPM reference runs (20min 346, 30min 253)
- Intensity corrected by
 - GPM efficiency: $3e-5$
 - Sapphire window: 0.908
 - Aluminum: 0.975
 - Polarization: 0.5
 - Silicon window: 0.911
 - Detector efficiency: 0.96
 - ROI efficiency: 0.7
- Intensity: $5.0336e12$

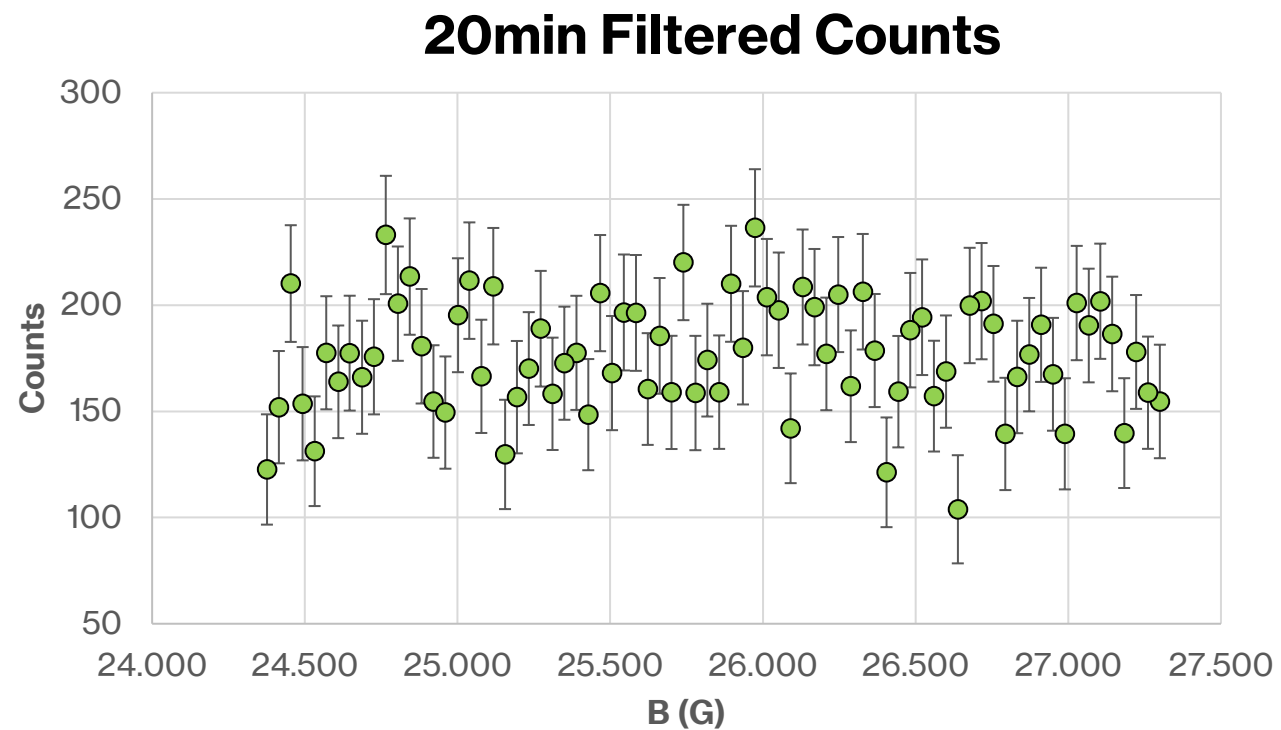
Baseline

- ROI counts – background = baseline
- Average ~267
 - Used when looking for effect from background
- Each point used for fitting procedure



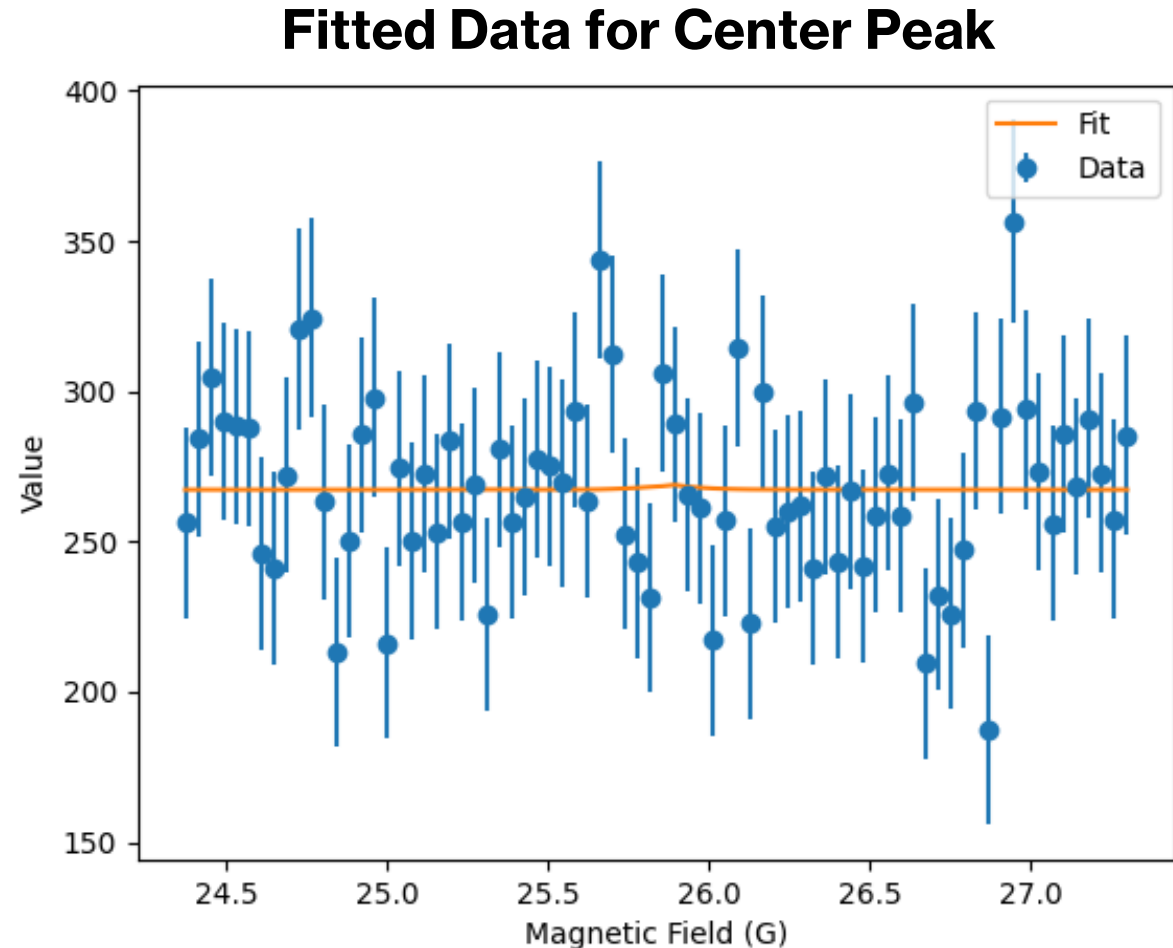
Baseline

- 20min average ~176



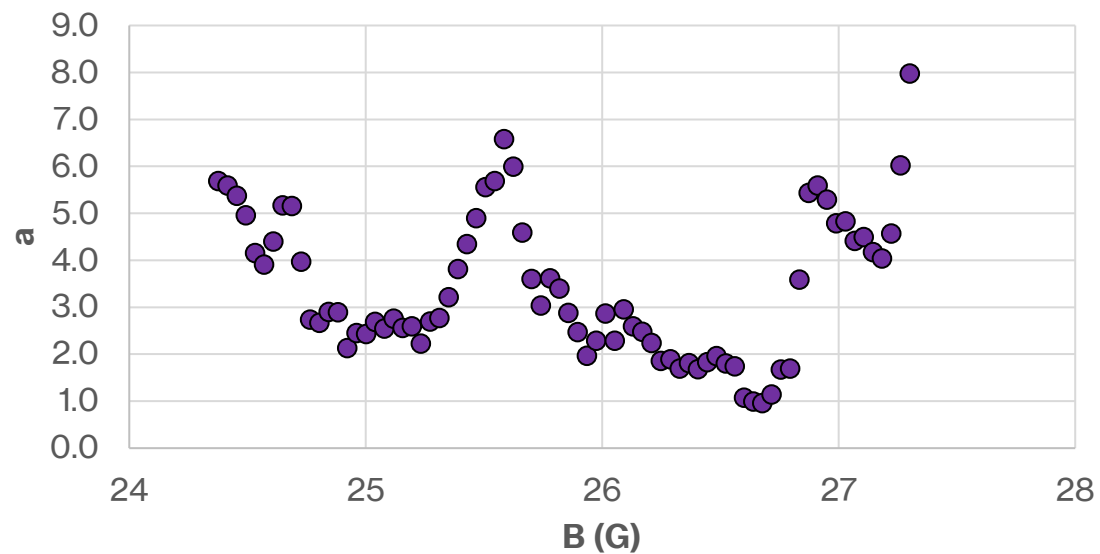
Fitting

- Data fit to $b + a * I * f(a)$
 - b = fitted background
 - a = effect in counts above baseline
 - Intensity = $5.0336e12$
 - $f(a)$ = template function
- $\text{Kappa} = 5e-6$
- Small peak visible
- Take points and slide fit function across data to extract peak (a value) at each point

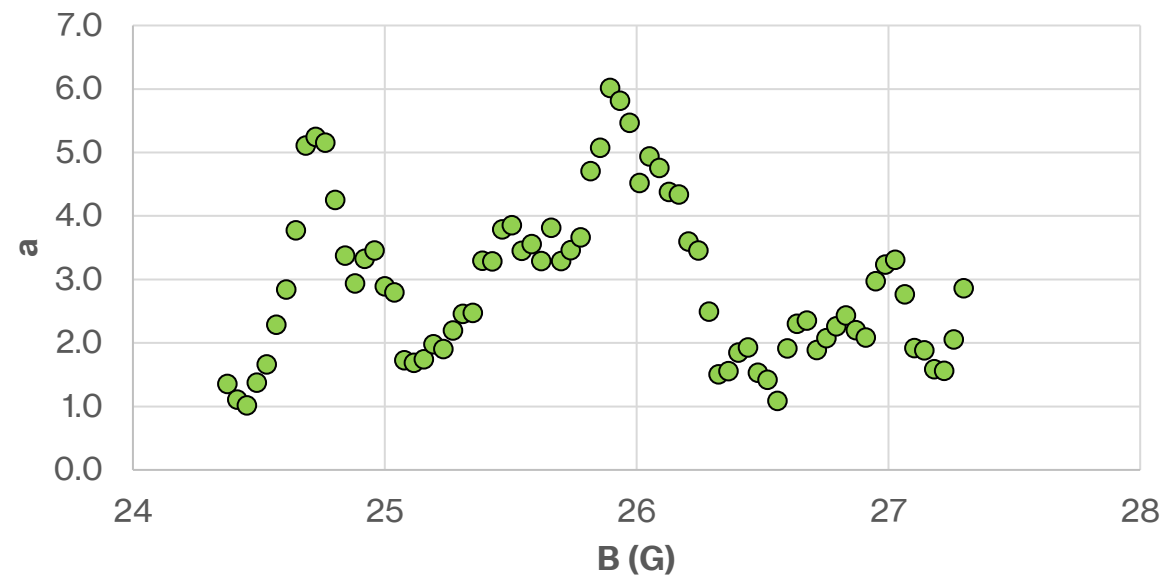


Limits on a

30min Data a Limits

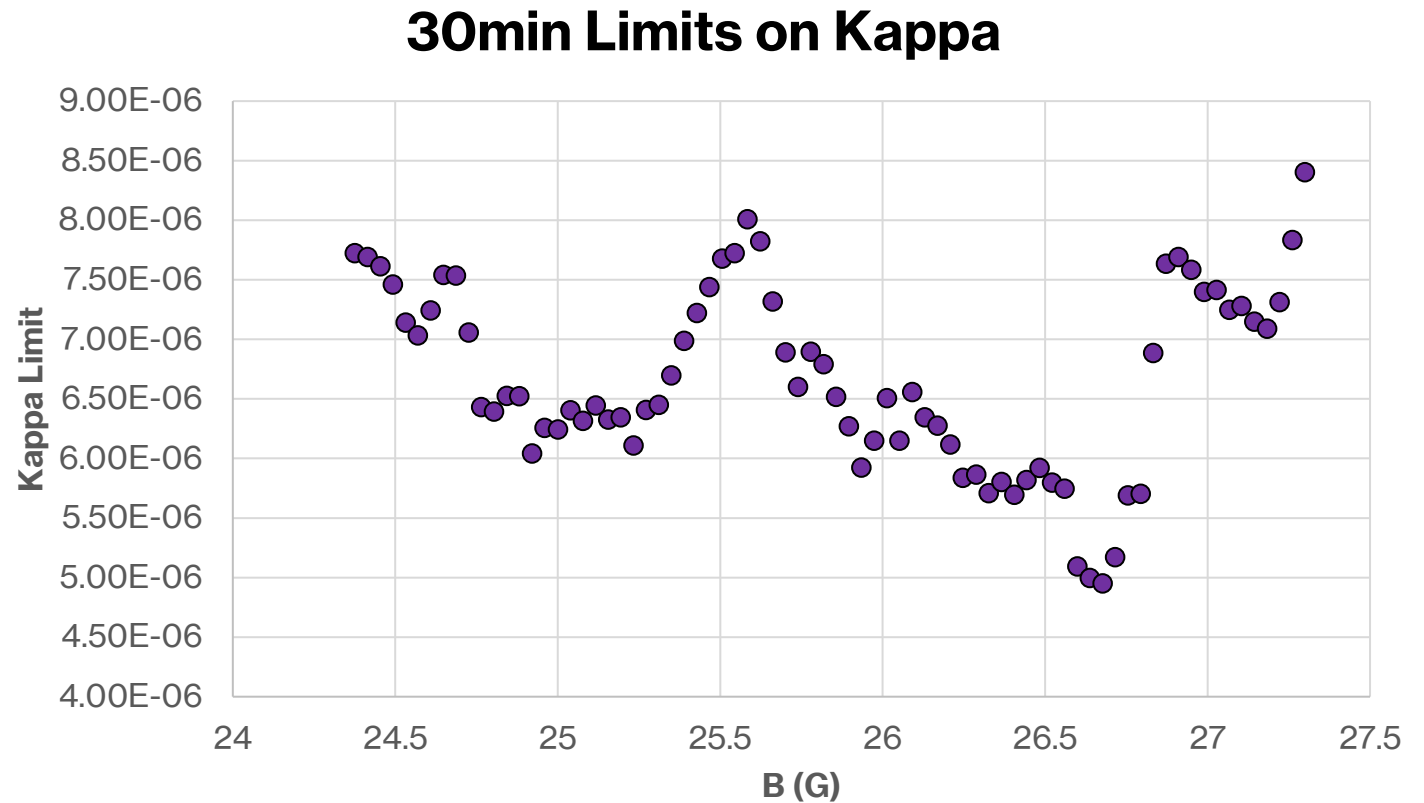


20min Data a Limits



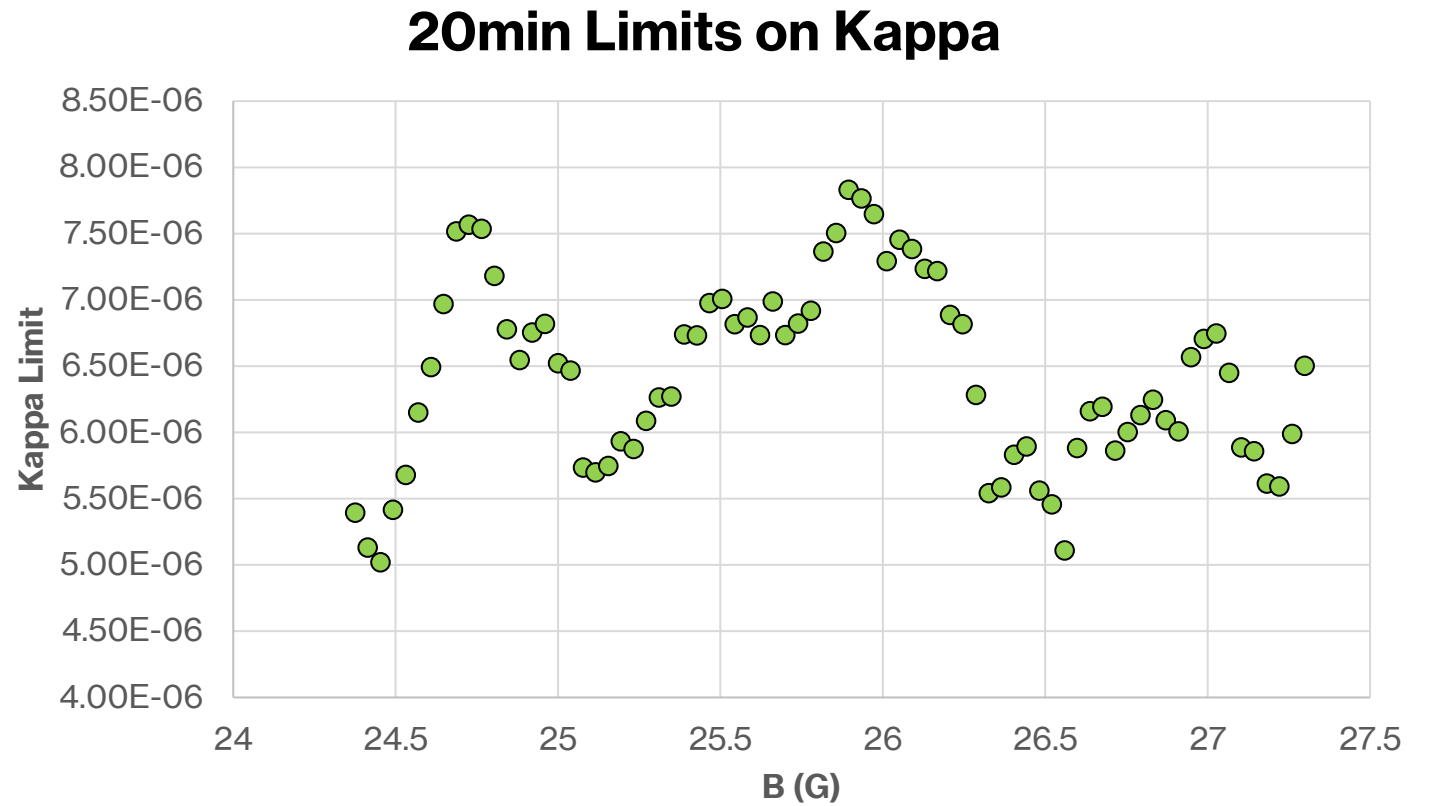
Kappa from a

- $a^* = a$ limit
- Conversion: $5e-6 * (a^{*1/4})$
- $a^*: 2.8856$
- Center limit: $\kappa < 6.27 \times 10^{-6}$



Kappa from a

- 20-minute kappa limits
- a^* : 6.0163
- Center limit: $\kappa < 6.83 \times 10^{-6}$

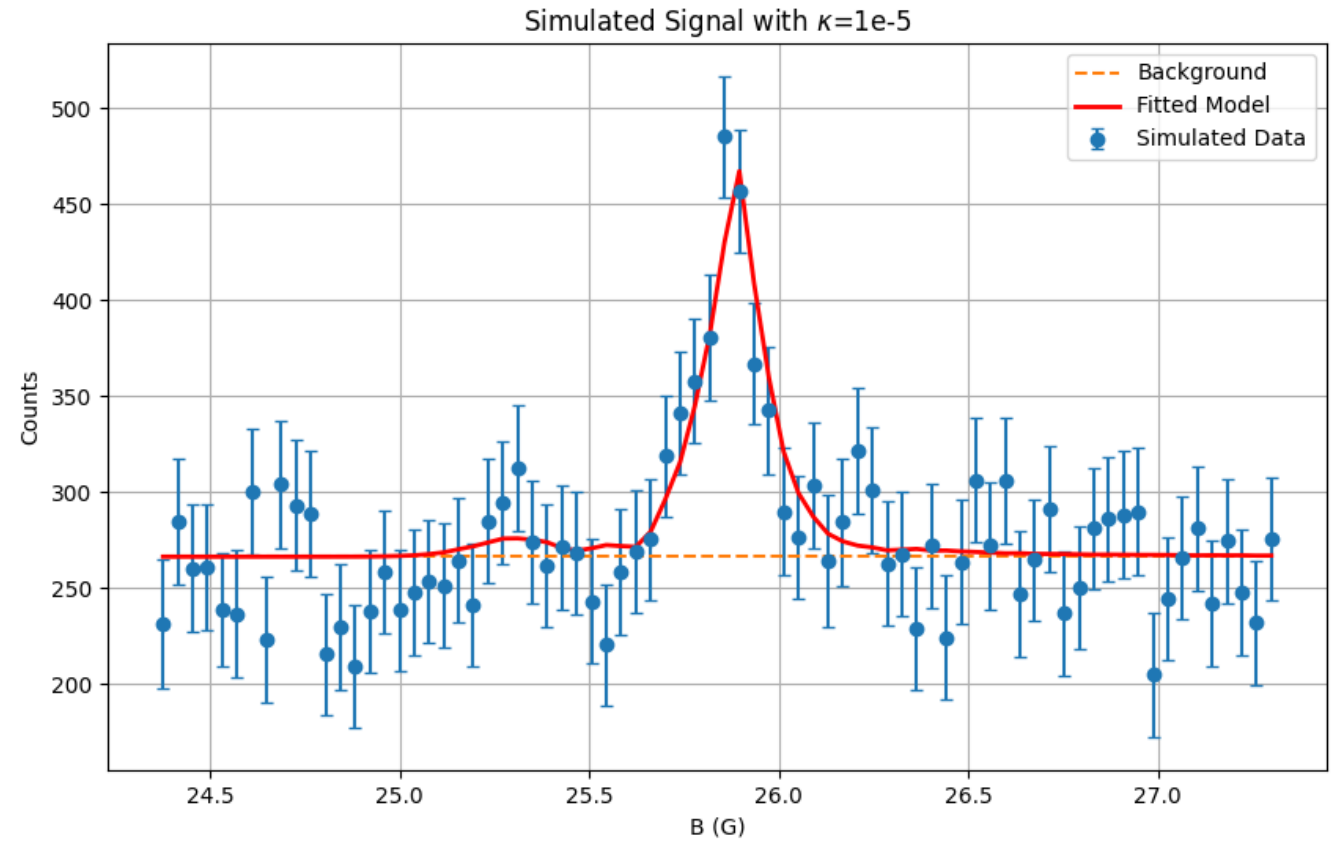


Simulated Limit

- Monte Carlo simulations
- Only 30min runs
- Extract peak from background
 - Same process as before
- Use base values to generate data sets
- Repeat 1e6 times
- a^* : 2.89328
- Kappa limit: $\kappa < 6.27 \times 10^{-6}$

Simulated kappa = 1e-5

- Generated 1e6 experiments
 - Example on right
- Visible peak
- a : 1.0202
- a err: 0.09234
- Calibration:



Next Steps

- Use scripts and process created from this analysis with nTMM-2 data
- Organize procedure to get Michael and Evan started with analysis

Magnetics simulation plan for nTMM data

2026-09-18

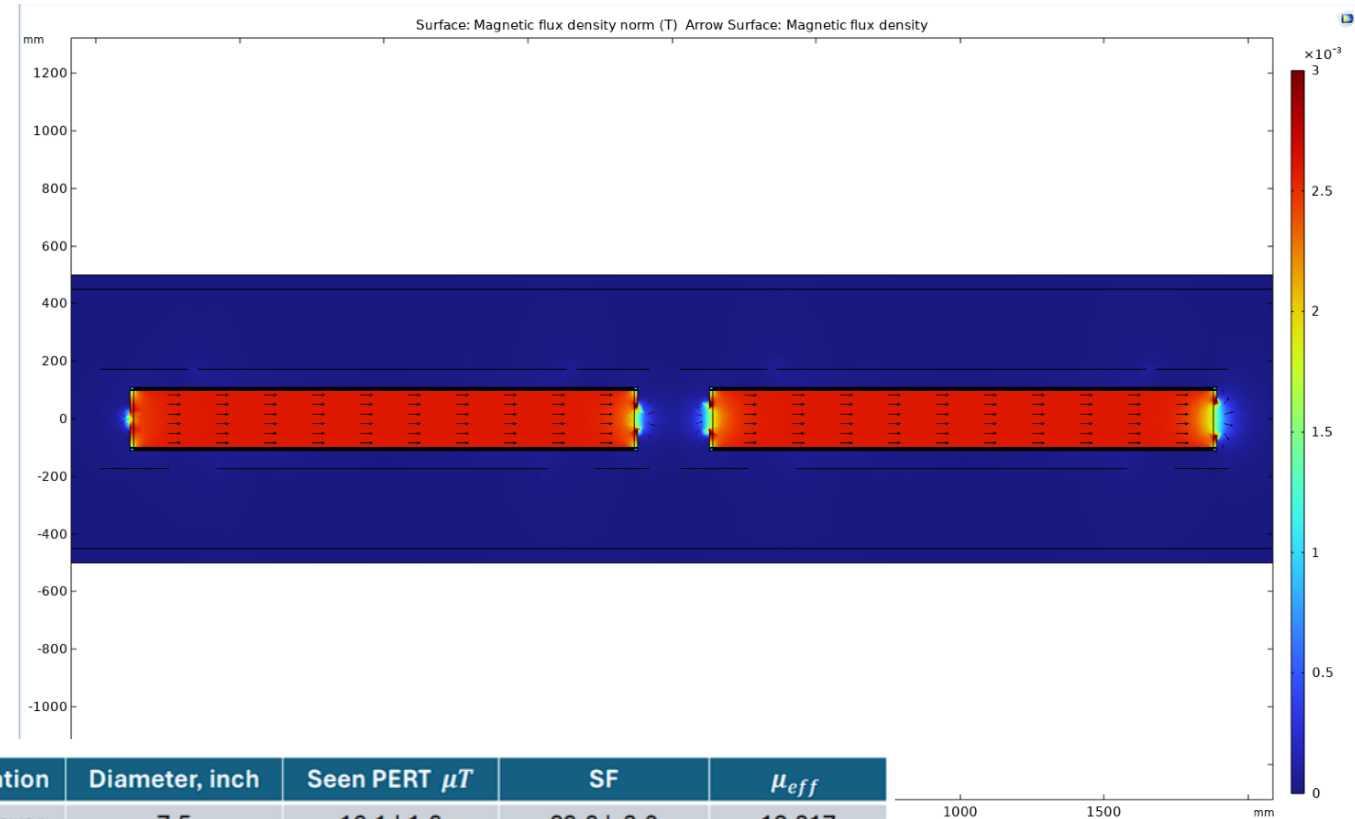
linus.persson@fysik.lu.se

nTMM-2 data analysis

- Imagine-X magnet was turned off during the full nTMM-2 data collection.
- This means the background environment, dominated by the magnetised support under box 8, remains approximately the same as during nTMM-1.
- We need a strategy for how to simulate the backgrounds and incorporate them into the sensitivity analysis.
- Let's start with a brief reminder from before...

Simulation of coil field

- Both inner and outer mumetal with 0.6 mm thickness and effective μ from UT prototype measurements.
- Individually modelled coil layers.
- This model gave a current-to-field conversion factor of 97.91 G/A.



Result from
Mubi (18/8), can
be updated

	Configuration	Diameter, inch	Seen PERT μT	SF	μ_{eff}
only	1 μ inner layer	7.5	16.1 $\pm$ 1.8	32.6 $\pm$ 3.8	13,317
only	2 μ inner layer	7.6	4.51 $\pm$ 0.61	116.2 $\pm$ 16.1	24,050
only	1 μ outer layer	9.0	27.6 $\pm$ 3.7	19.0 $\pm$ 2.6	11,220
only	2 μ outer layer	9.1	7.9 $\pm$ 1.1	66.3 $\pm$ 9.4	19,794

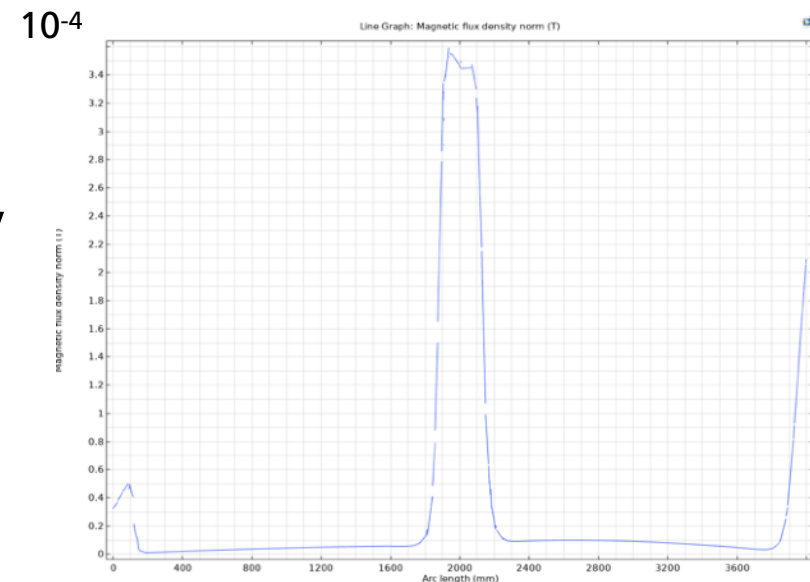
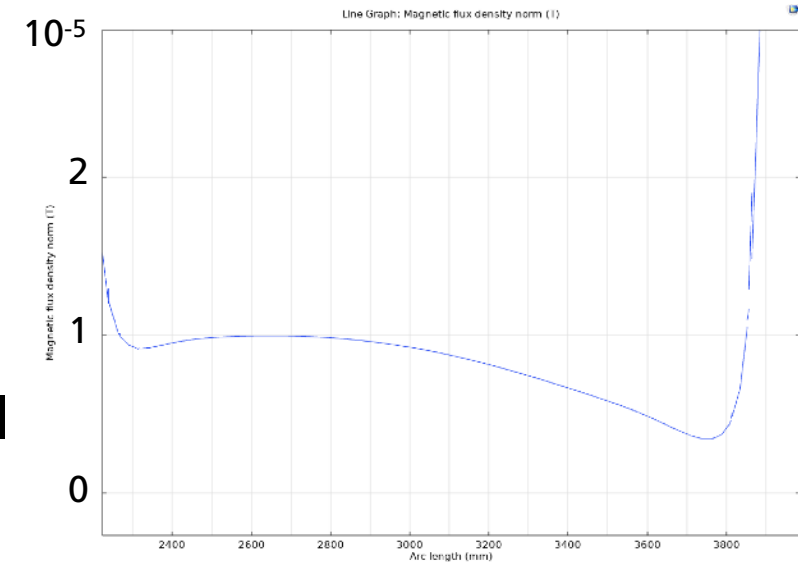
Prototype measurement



LUND
UNIVERSITY

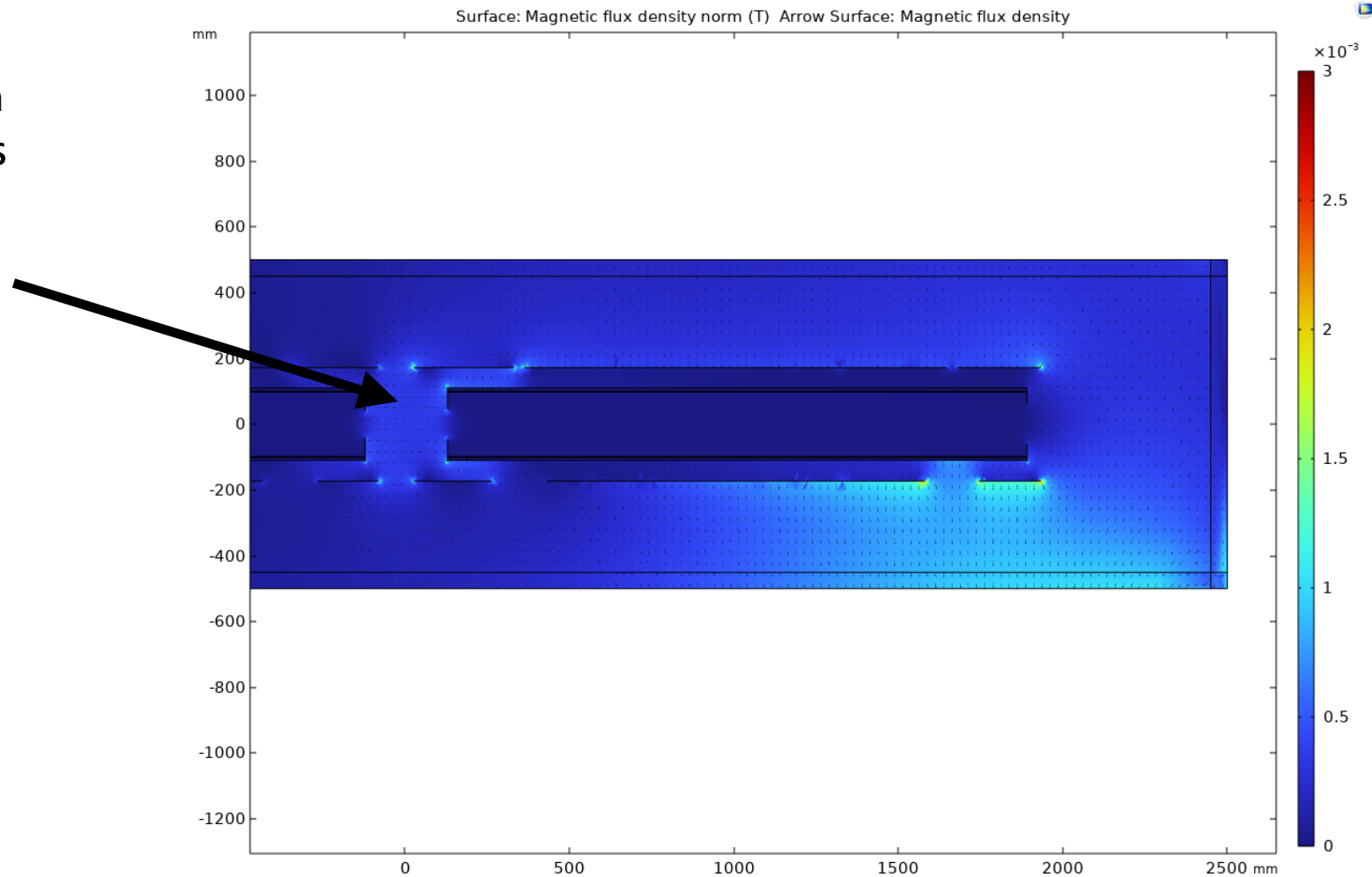
Magnetic scalar potential simulation

- I used a low-order approximation ($L \leq 6$) spherical harmonic fit from German and applied on the new model for the background.
- Infinite element domain with scalar potential boundary condition.
- This is very similar to German's results.
- Field between the magnets is much larger than expected ($\sim 350 \mu\text{T}$), but is supported by simulations in constant longitudinal background field.
- Additional scrutiny needed.



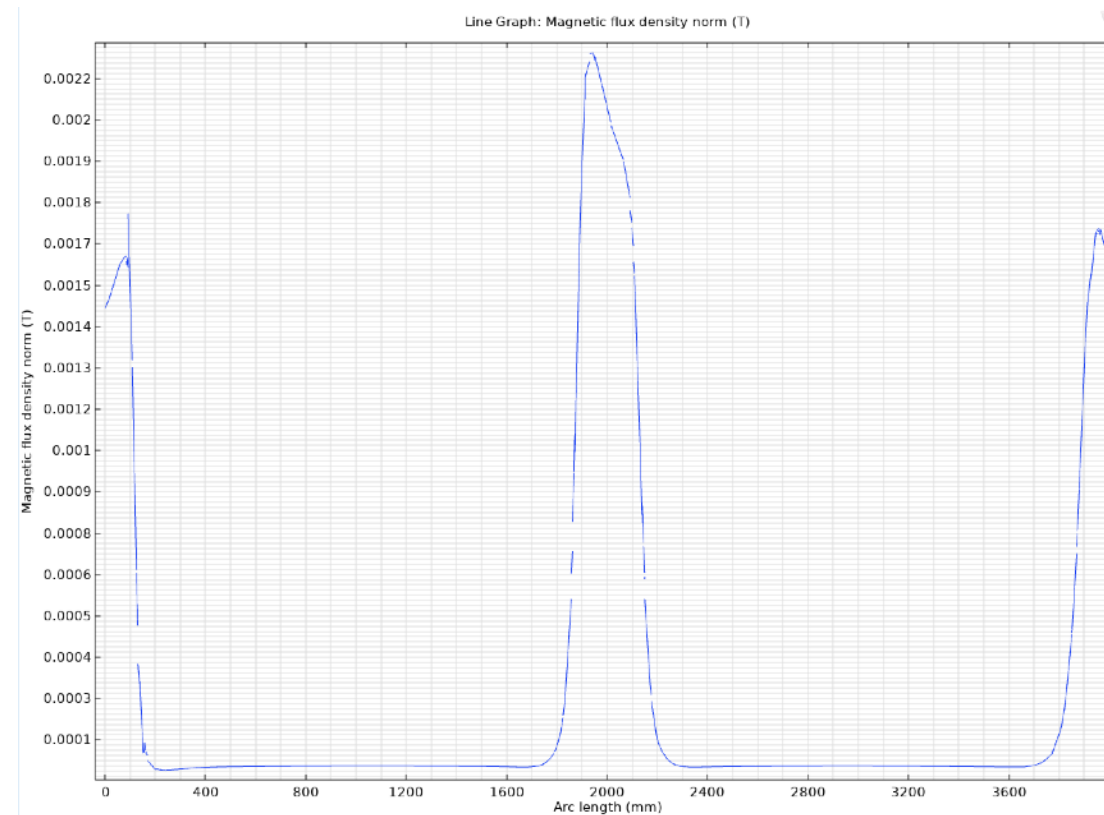
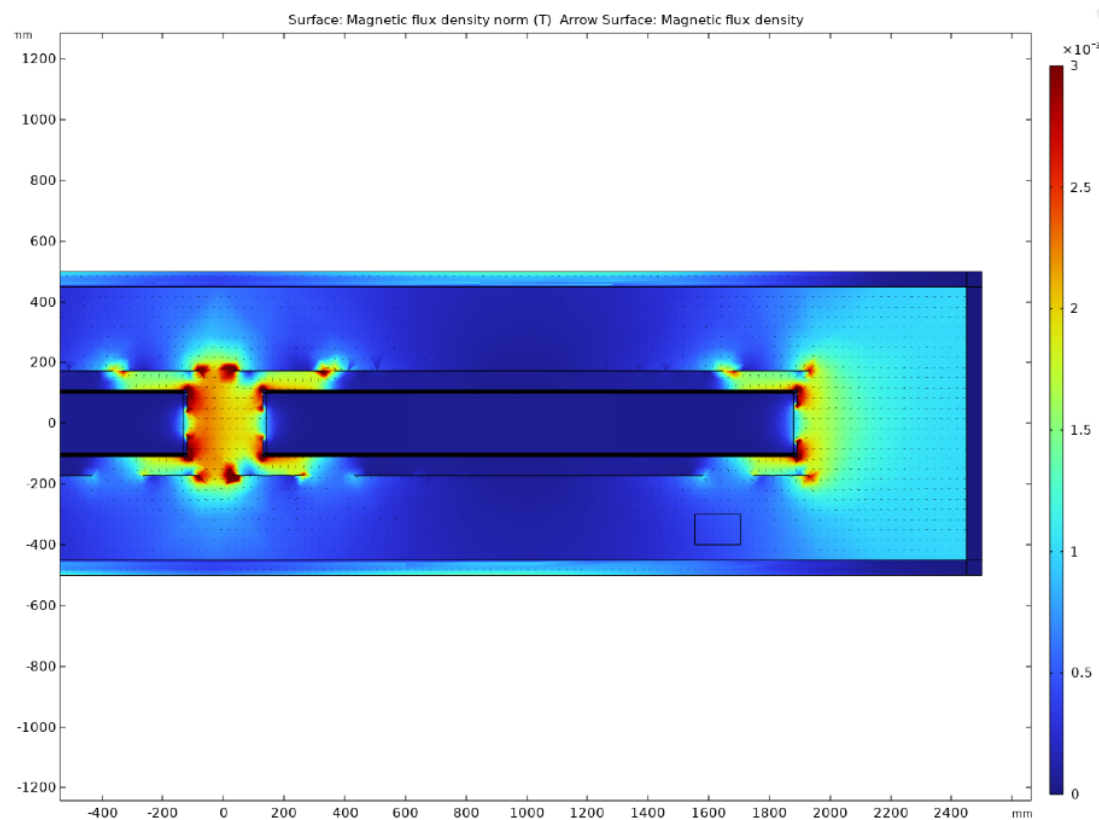
Simulation with magnetic scalar potential

The field between the inner shields is stronger than the field outside



Simulation with horizontal background field

- Notice a large amplification in field strength where field lines jumps from the outer mumetal layer in one magnet to the other.



Suggested analysis strategy

- Simulate full beam profile every 1 mm along the axis by adding (as vectors) the perturbation field and the field generated by coil current.
- Verify transverse homogeneity of the field.
- Verify that higher-order expansion of spherical harmonic results converge for large L .
- Use the new B-field curve to calculate the regeneration probability curve as a function of applied magnetic field.
- Use the input function shape as input to Alina/Carolyn's analysis workflow to extract nTMM sensitivity.

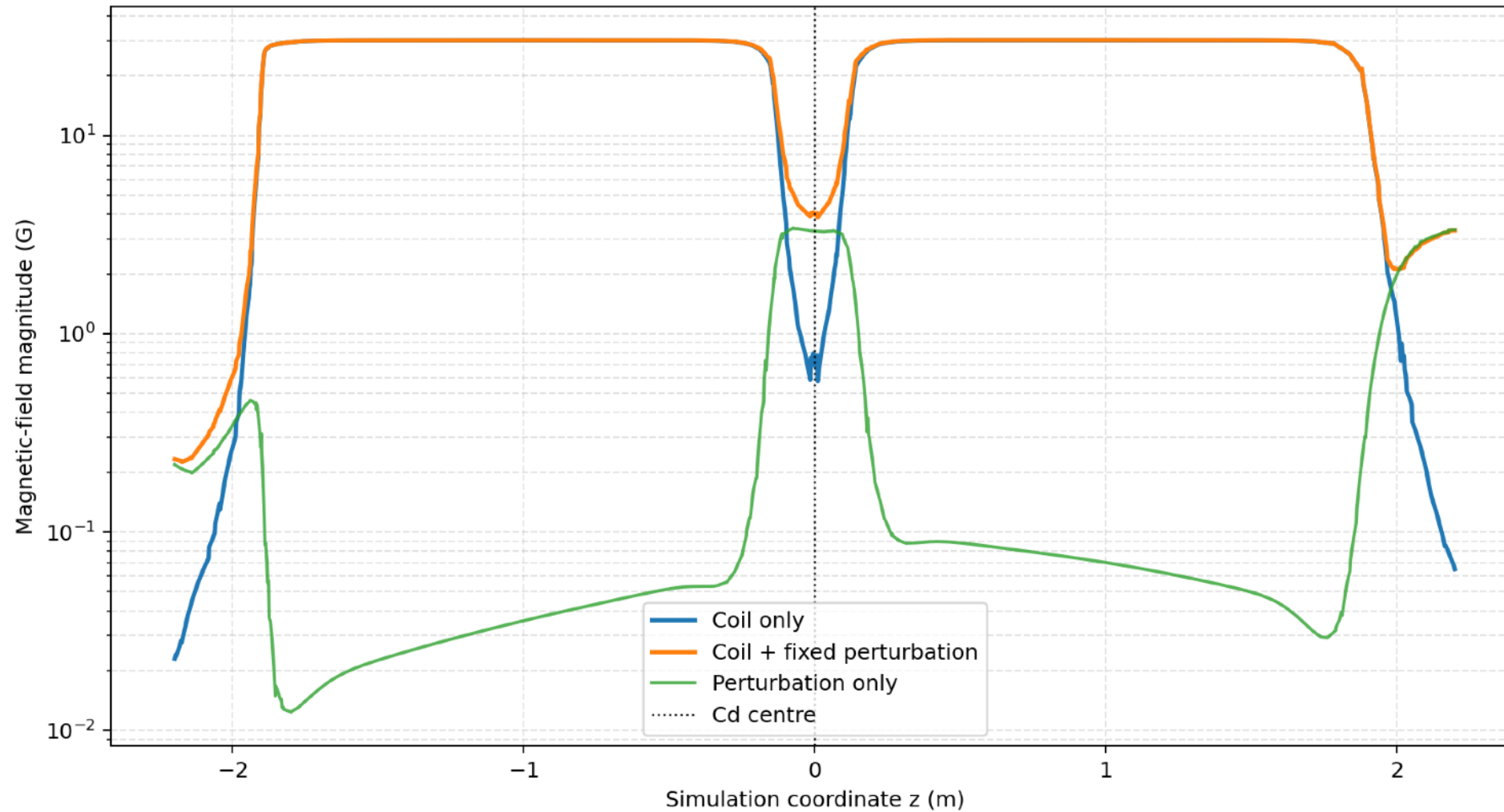
Disclaimer: Highly preliminary results coming up

(i.e done in an airport or the back row of a conference)

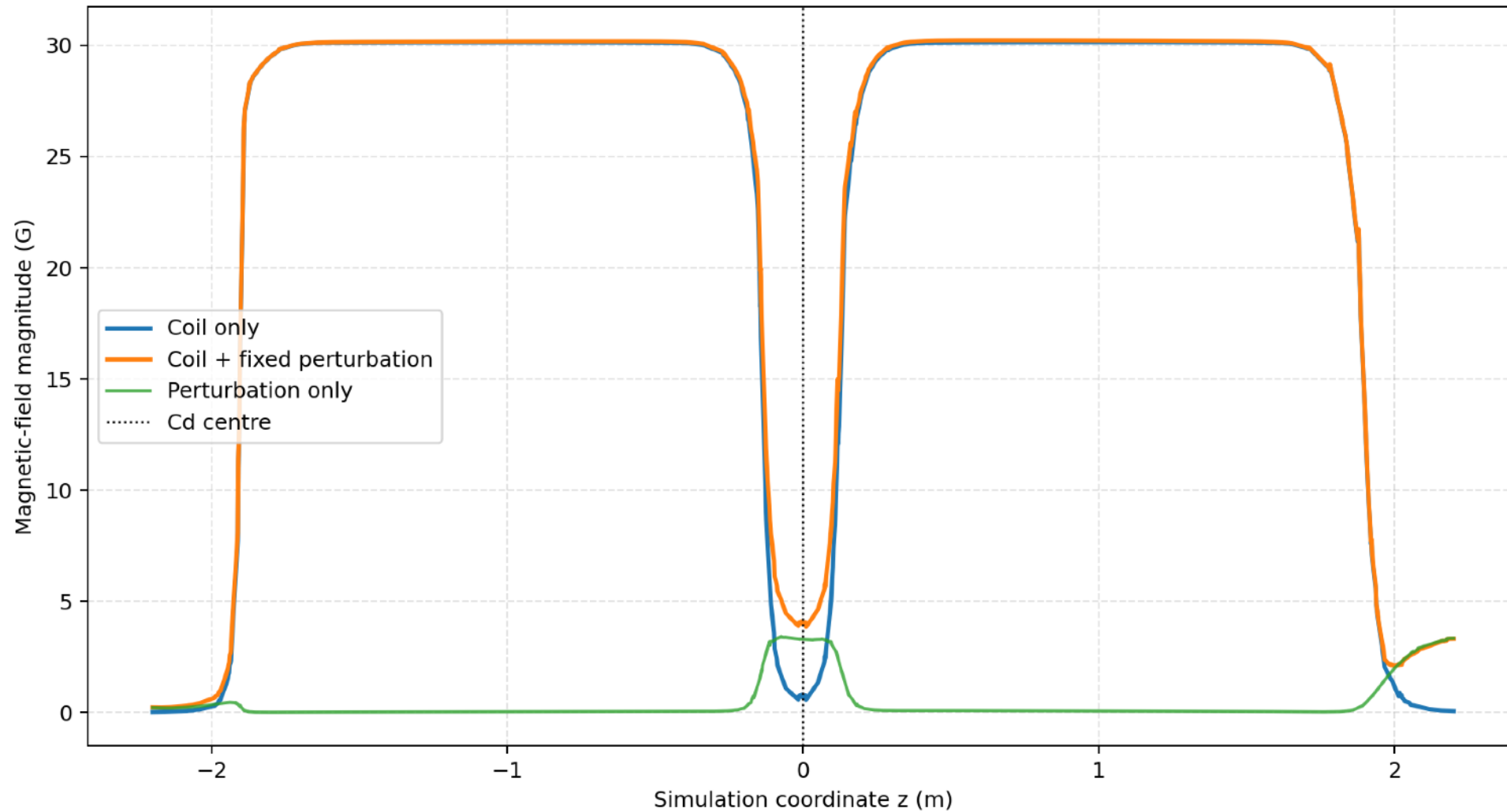
Input to updated sensitivity analysis

- I used the coil simulation that determined the current scan interval for the generated field.
- I used the low-order scalar potential fit ($L \leq 6$) discussed previously for the background field.
- These were added as vectors and the total magnitude was used as input to the Python version that I made of Yuri's Fortran script (based on Berezhiani's formula).
- Simulations were performed for relative nTMM $\kappa = 5 \times 10^{-6}$.
- Note: For now I only included the cadmium absorber, not the borated polyethylene (although this should not have a major effect), what is the exact thickness and placement of all materials relative to the beam direction?

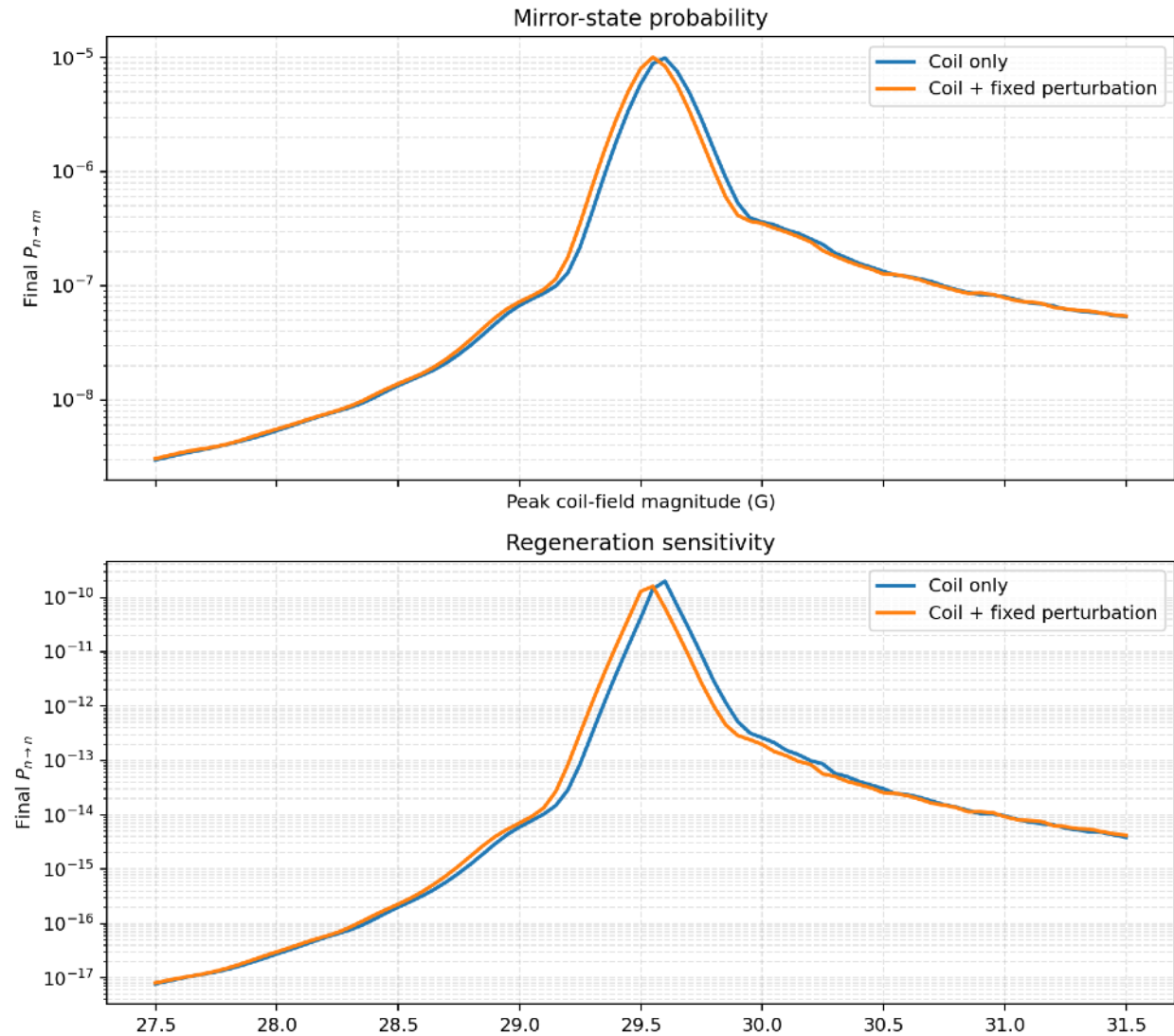
Field profile (log scale)



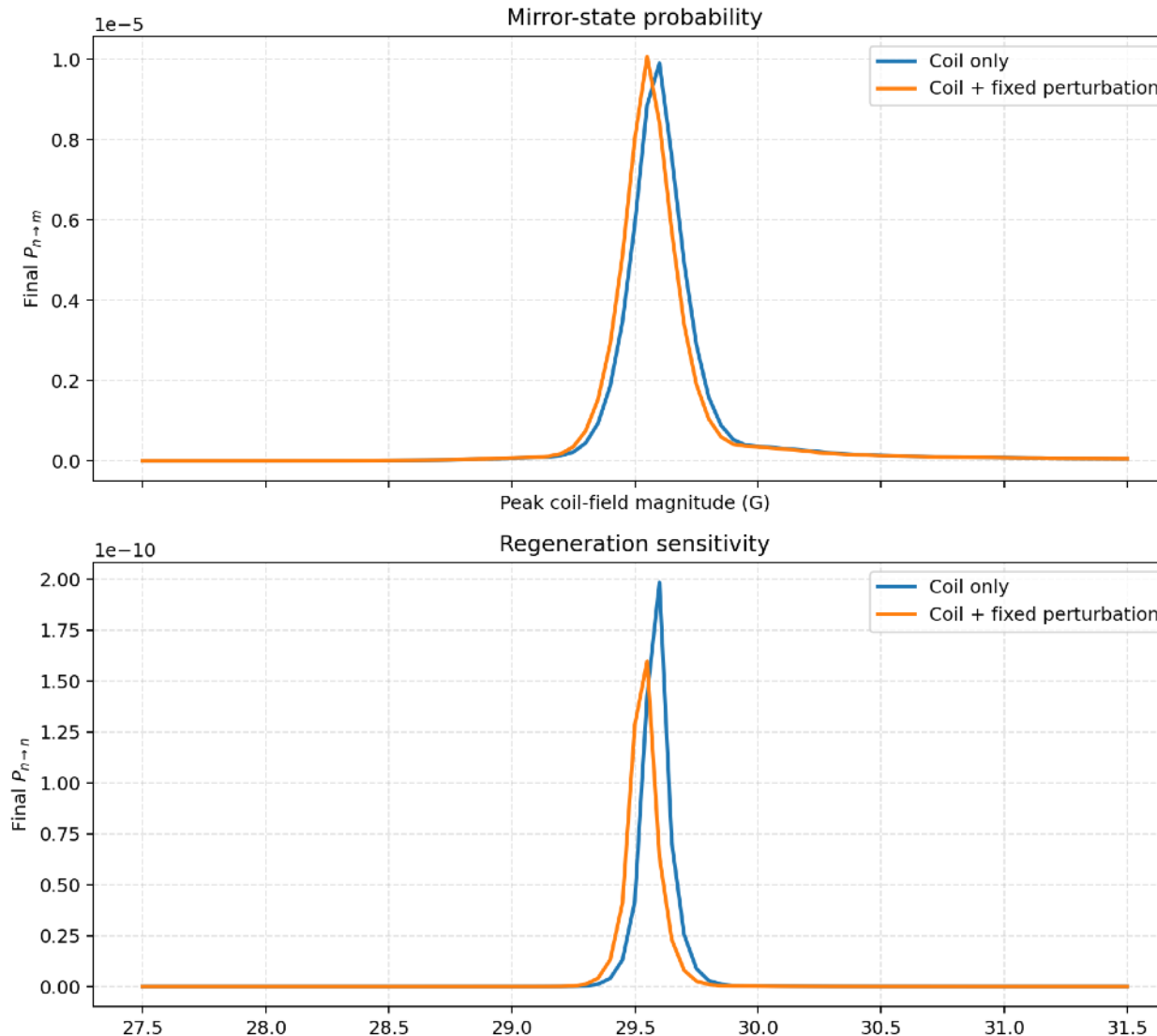
Field profile (linear scale)



Resonance shape (log scale)



Resonance shape (linear scale)



Conclusions

- The peak appears at roughly 29.5 G, consistent with expectations and the chosen current scan range. The “coil only” case gives approximately the same peak regeneration probability as the calculations from Yuri.
- When including the background fields, the peak probability drops by about 20% and shifts to lower coil fields (0.05 G). This makes sense since the background field makes the total field slightly larger.
- A small broadening in FWHM of about 8% is also seen. This can be understood intuitively as the region with correct field is smaller at a given current but reachable from more current settings.
- This is encouraging, but several tests remain before being publication-ready, including full understanding of the effective μ , understanding of scalar potential convergence and transverse field homogeneity.

Optical Character Recognition for nTMM-2

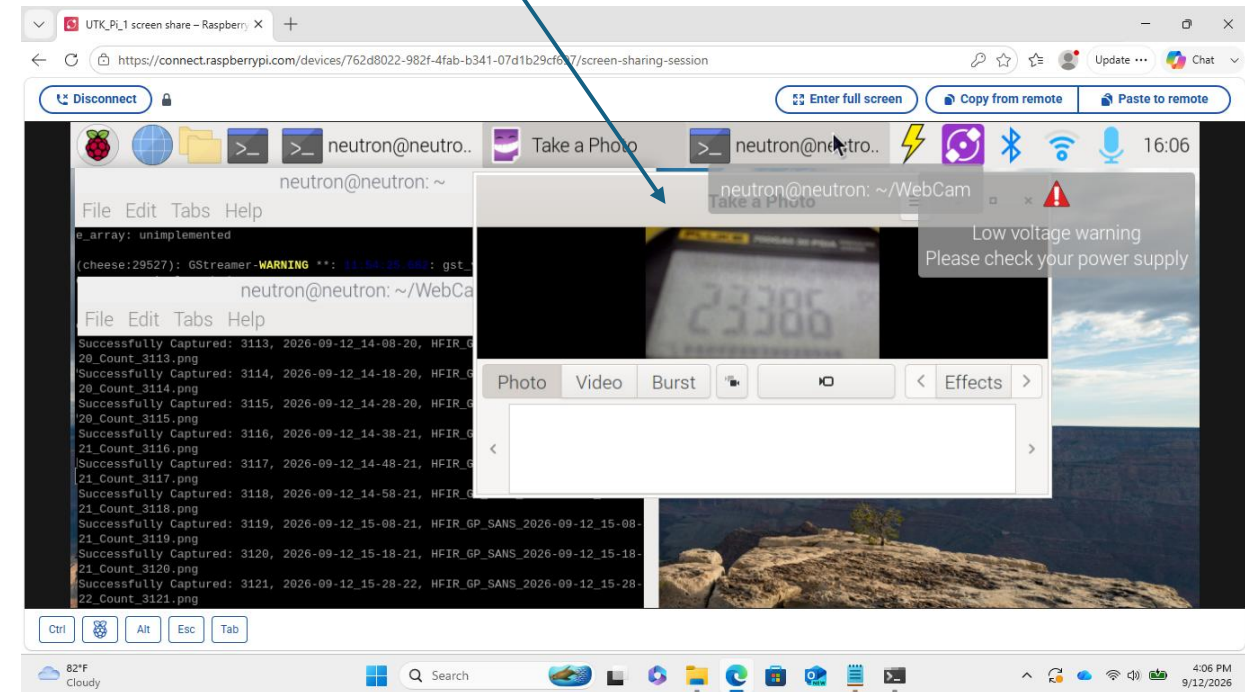
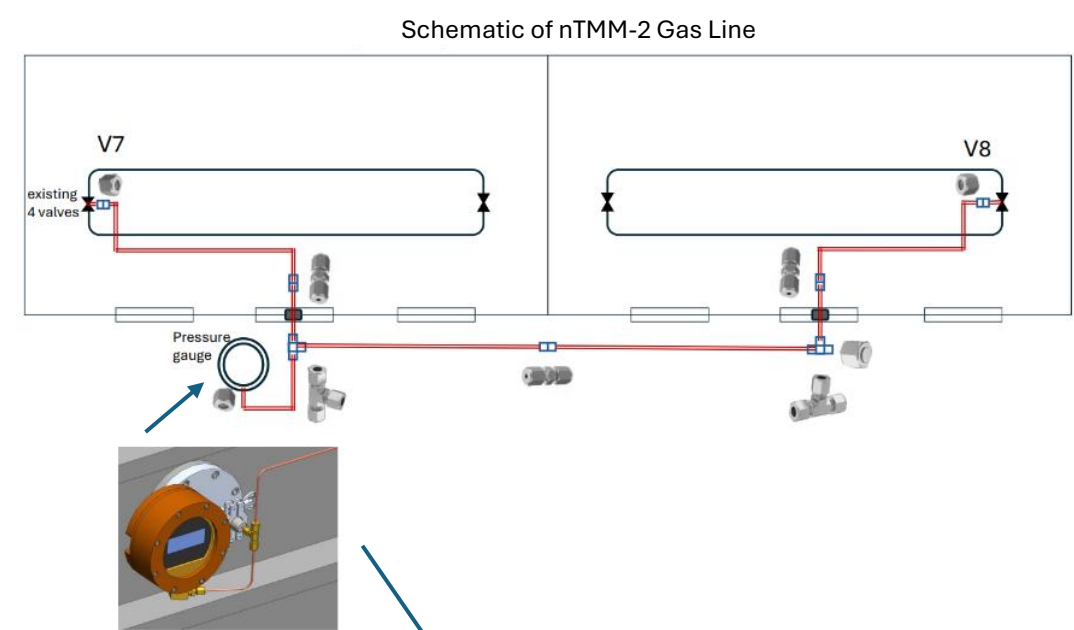
Evan Michael - University of Tennessee
- nn' Collaboration Meeting -
9/18/2026

Overview

- Changes in electrical current throughout the experiment heated and cooled the magnets through resistive heating, causing CO₂ gas pressure changes throughout the runs.
- The time of day also lead to notable fluctuations in the temperature of the magnet.
- CO₂ gas pressure readings and temperature readings must be reconciled to understand their correlation throughout the runs.
- This correlation will allow for a correction to the density of the CO₂ gas and thus a further correction to the calculated Fermi optical potential.

Pressure Gauge Readout

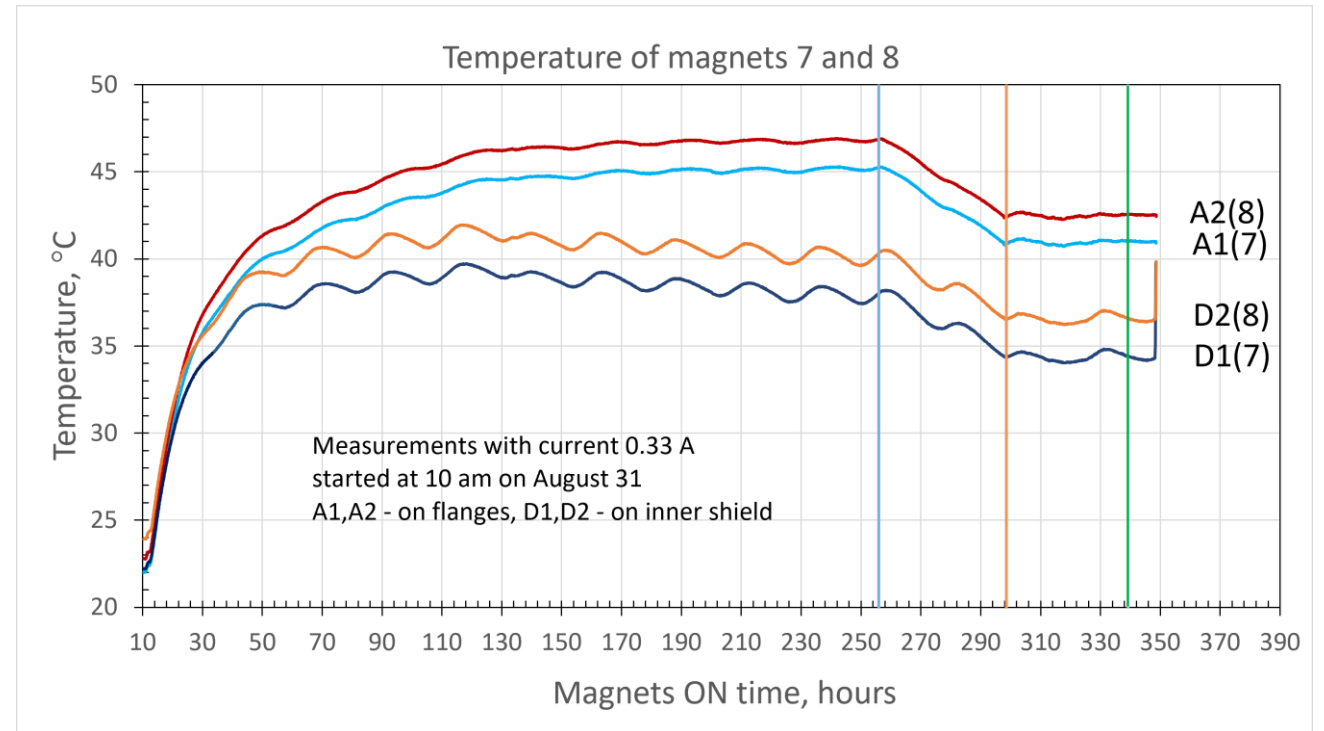
- Analysis of nTMM-1 (2025) found that there was a 2% difference in CO₂ gas pressure.
- A gas line was added to the magnets for nTMM-2 (2026) to maintain equal pressure in both vessels.
- The connected system was monitored by a single gas pressure gauge, whose reading was recorded via camera in the 2026 runs.
- The pressure readout from the gauge was captured every 10 minutes.



Raw Pressure Capture (Uncropped Image)

Thermocouple Readout

- Four thermocouples are attached to each nTMM-2 magnet to monitor temperature changes throughout the runs.
- Two thermocouples on the outer flanges, two on the inner Mu Metal shielding layers.
- Temperature changes are due to resistive heating and cooling as the electrical current changes through the runs.
- Temperature fluctuations from daily environment are also present.
- The readouts of the thermocouples are recorded digitally every 5 seconds.



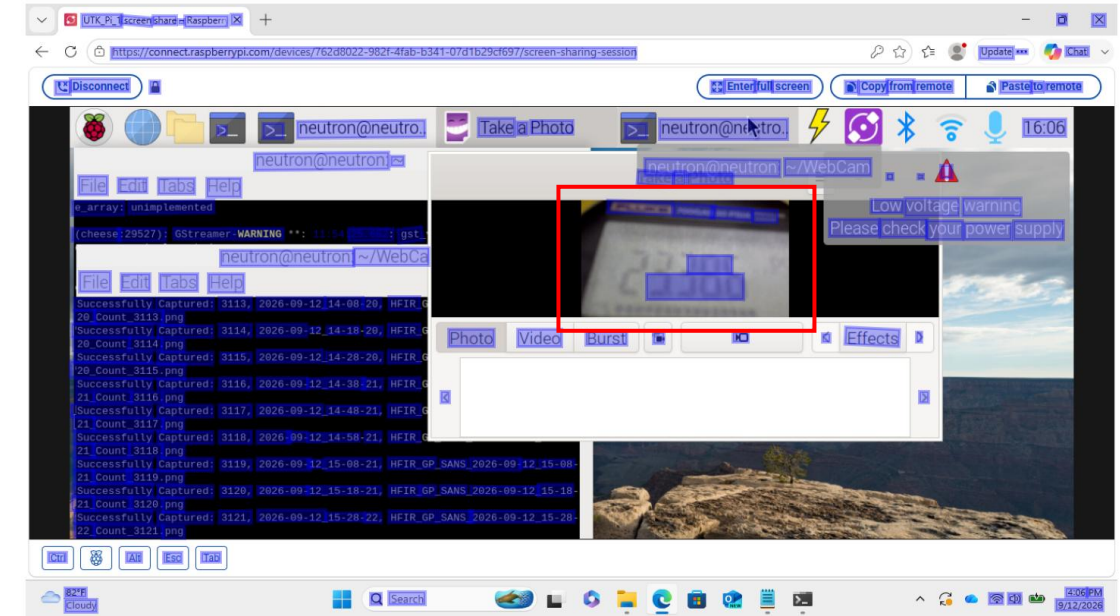
Experiment Time

Temperature-Pressure Relation

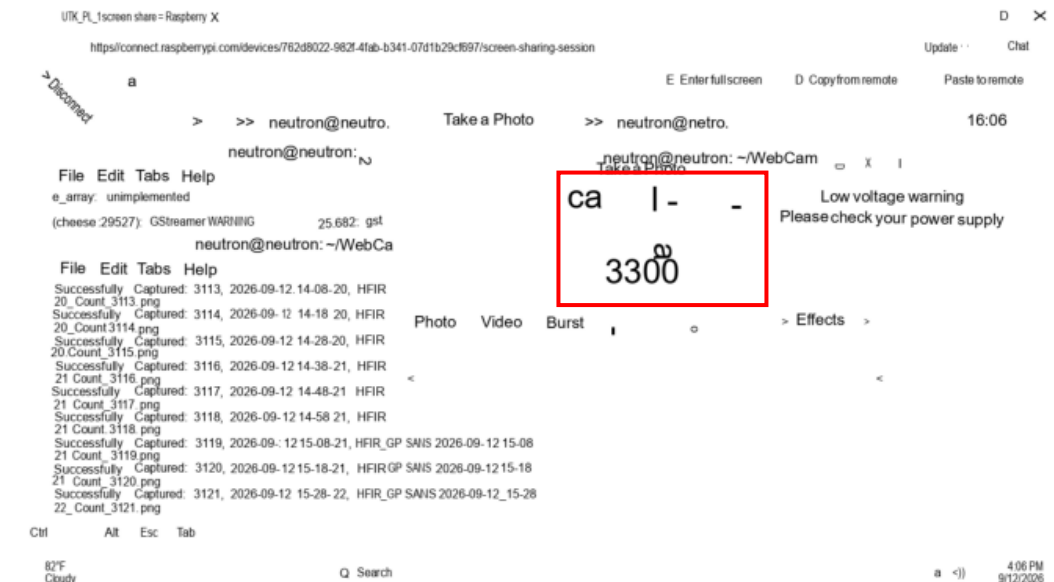
- A temperature-pressure relationship of the CO₂ gas throughout the runs should be established calculate gas density throughout the runs via a NIST calculator.
- Knowing the density allows for corrections to calculations of the Fermi optical potential in the experiment.
- This can be accomplished by pairing coinciding pressure and temperature data throughout the runs.
- Challenge: 500+ pressure captures are difficult to manually catalog.
- Solution: Use an Optical Character Recognition (OCR) machine learning model to detect (locate the text), recognize (read the text), and catalog pressure readings.

OCR Implementation

- Method 1: Utilize a pre-trained OCR deep learning model and immediately deploy on pressure captures.
 - First attempted with the PyTorch-based docTR model. Unsuccessful in both detection and recognition.
 - The model is not trained on the seven-segment number font on the gauge display.
- Method 2: Use a standard deep learning OCR model and train it on datasets of seven-segment digital display images, then deploy
 - Example: train a YOLOv8 model (also PyTorch-based) on a large dataset, then deploy to extract pressure data from captures.
 - Train the model on images of various lighting, orientation, and resolution to account for grain and tilt of display on pressure capture.
 - Require manual input if confidence score falls below set percentage.
 - Work in progress

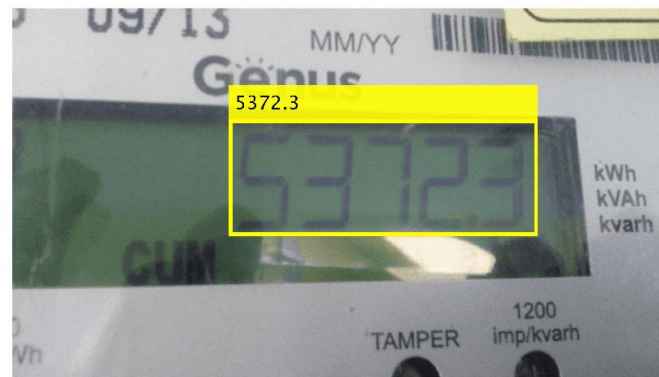
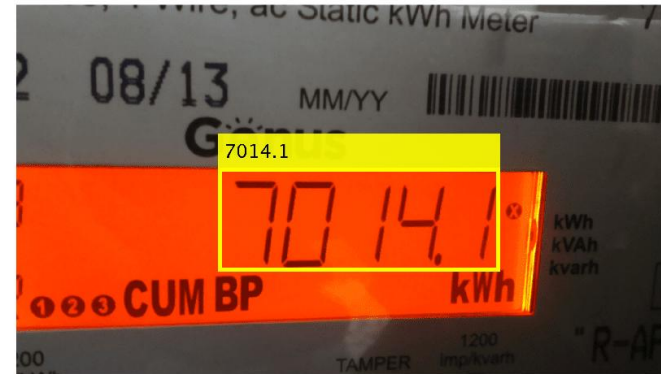
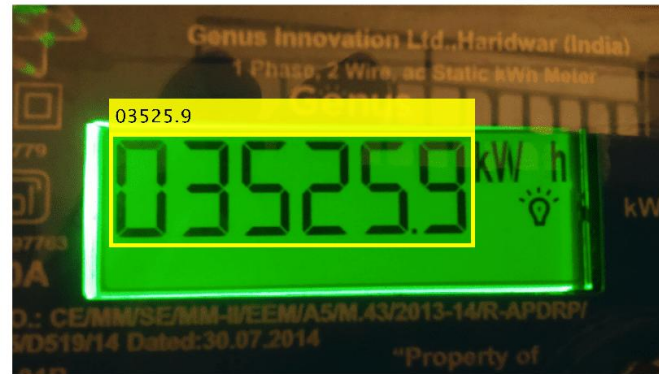


docTR Recognition Results (what the model sees)



Seven-Segment Training Dataset

Example: YUVA EB dataset (image credit: MathWorks)



- After recording pressure, pair with average temperature of coinciding time (via a Numpy array) for use in calculations

Plans and projects for 2026 data analysis

Yuri Kamyshev / UTK

1. Recover CO₂ pressure from camera shots (every 10') (Evan)
2. Recover and sync temperature recording from Data Logger
3. Correlate Pressure and average Vessel Temperature. Functional dependence.
4. Use this function for Optical Potential calculation for each run
5. Estimate error or Optical Potential reconstruction
6. Understand spatial distribution of background in the detector. Use it for subtraction in ROI
7. Understand reason for different μ from single & double layer in mu-prototype (Mubi + Linus)
8. Using μ for inner and outer double shield suppression factor measurements
calculate 3D profile of magnetic field in magnets with external perturbations (Mubi + Linus)
9. Estimate error in magnetic field reconstruction
10. Using corrected O.P. and $B_{\text{total}}(z)$, calculate nTMM expectation function (used by fit) 2000 velocities averaged
11. Estimate effect of neglecting $B(x)$ and $B(y)$ in nTMM expectation function calculations, Spin rotation?
12. Check time uniformity of GP monitor
13. Find GPM dead time correction
14. Check that there is no anomalies in magnet currents
14. Update on Scattering and absorption in beam windows for nTMM configuration (Mubi)
15. Estimate errors in intensity

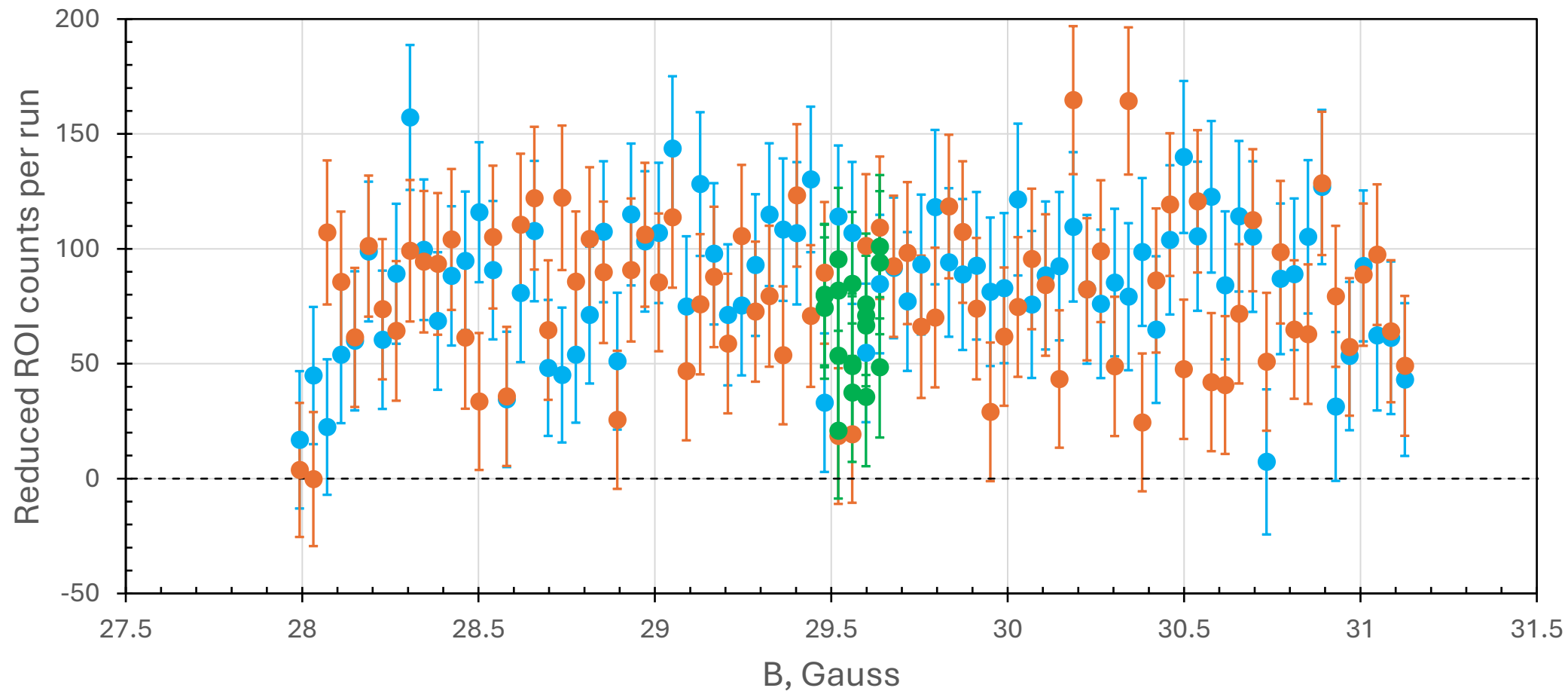
Publication scenarios for nTMM data:

Yuri Kamyshev / UTK

0. Can we combine all 180 points (all 30 min) into single data set for theory fit?
1. For TMM-related peak should appear around 29.56 Gauss $\pm$ error (to be calculated).
We can set a limit on κ (n TMM). Limiting value of $\kappa\mu B$ is $\gg$ larger than ϵ measured in all $n \rightarrow n'$ experiment, so, it will be pure case of measurement of κ . We can use only ~ 5 points centered around 29.56 Gauss and use all other 81-5 points as background. Experience of 2025 data analysis can be used.
2. For following PRD: Using same data we can set 2-dimensional 95% CL parameter exclusion in terms of $(\epsilon, \Delta m)$ or $(\kappa, \Delta m)$. With more complicated effort we can set 2-dim limit in terms of (ϵ, B') or (κ, B') . Recent PSI published two independent papers (and 2 PhD) Δm and B' analysis. We can do it within one PRD paper and Mubi and Linus might include that in their PhD works. UT students can use it for undergraduate theses.
3. Our prototype measurement of shielding factors for mu-metal with corresponding COMSOL analysis can be described in the paper that we submit to NIM. (Evan can lead this effort)
- MM model, in principle, for n has many unknown parameter: $[\epsilon, \kappa(nTMM), \Delta m, B', V_F', V_{B-L}]$
In principle, suppressing effect of these parameters can be compensated by our magnetic field as μB

PRELIMINARY

SCAN1,SCAN2,SCAN3 bkgr subtracted, intensity corrected

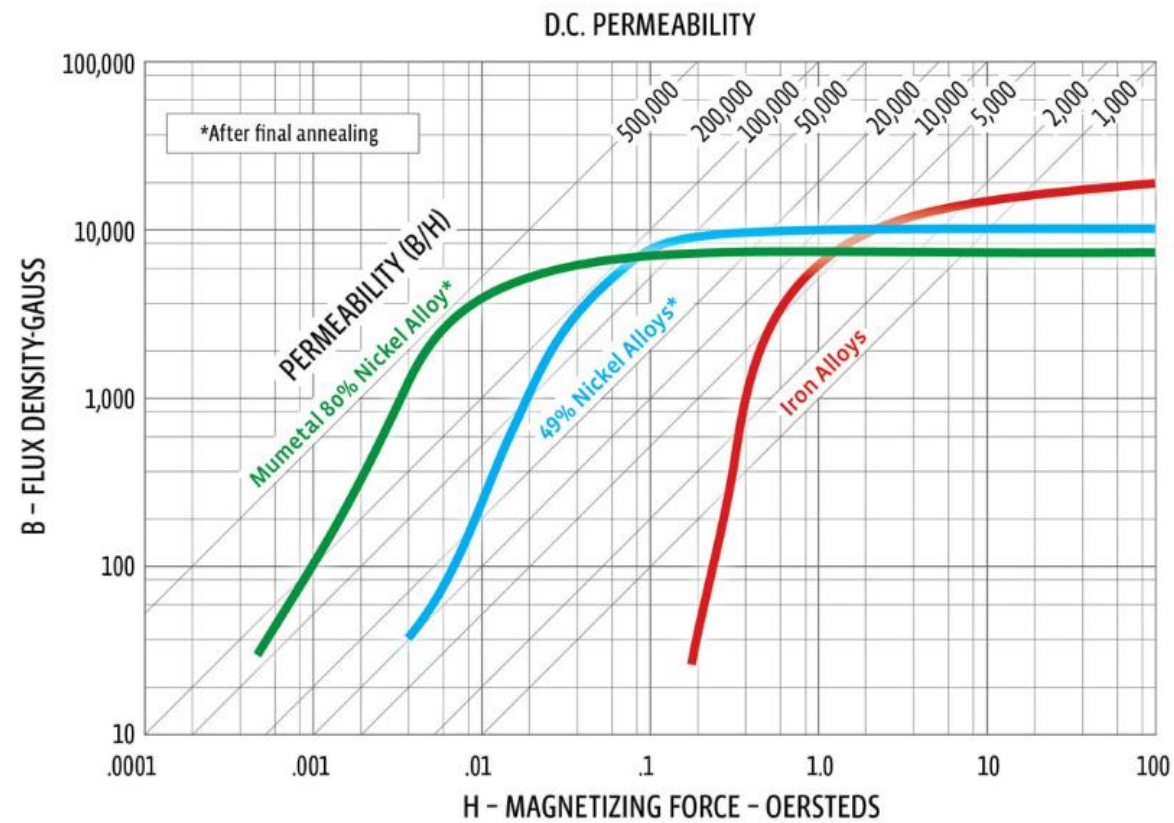
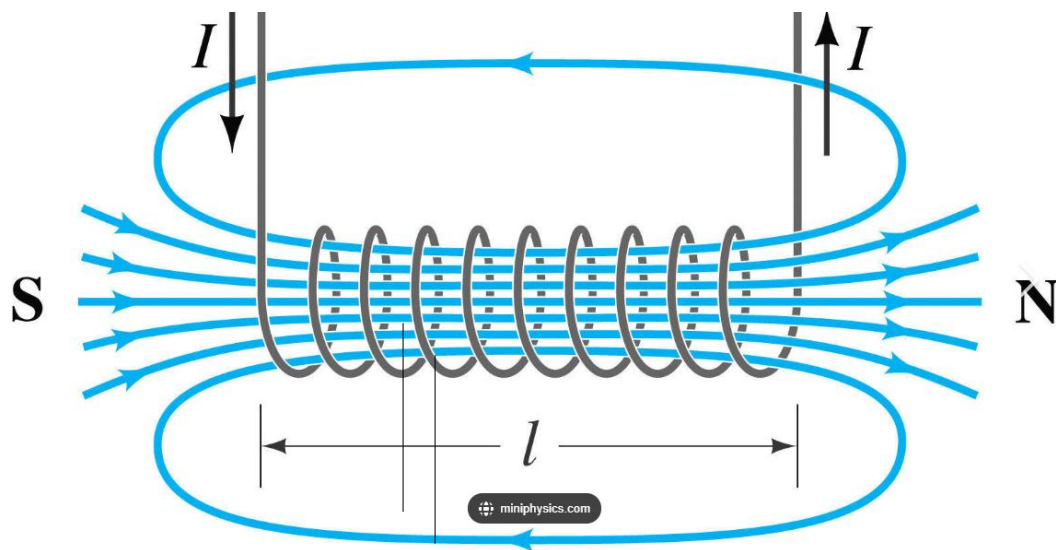
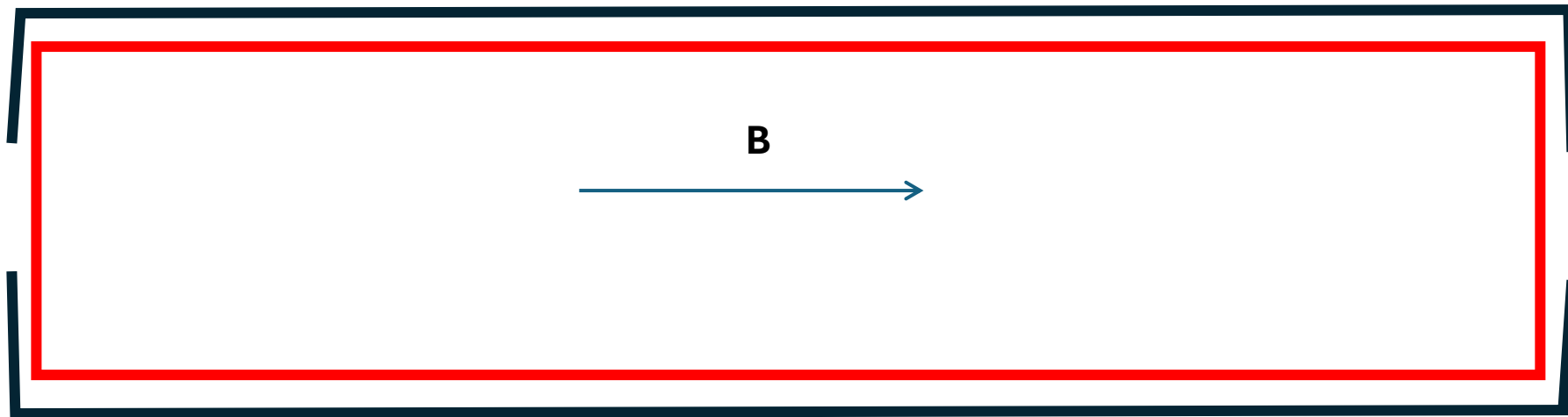


Measurements with μ prototype at UT

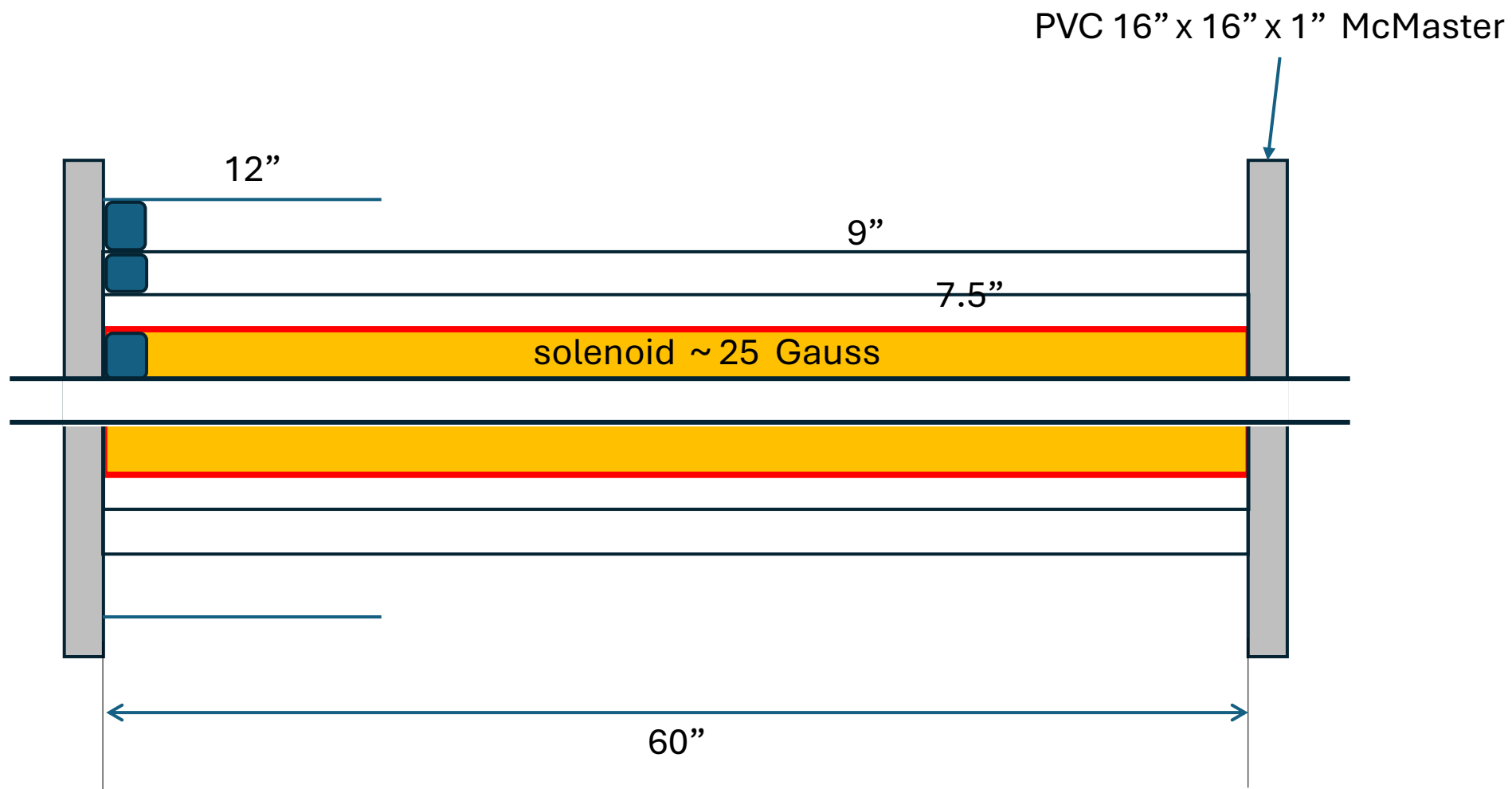
Yuri Kamyshev

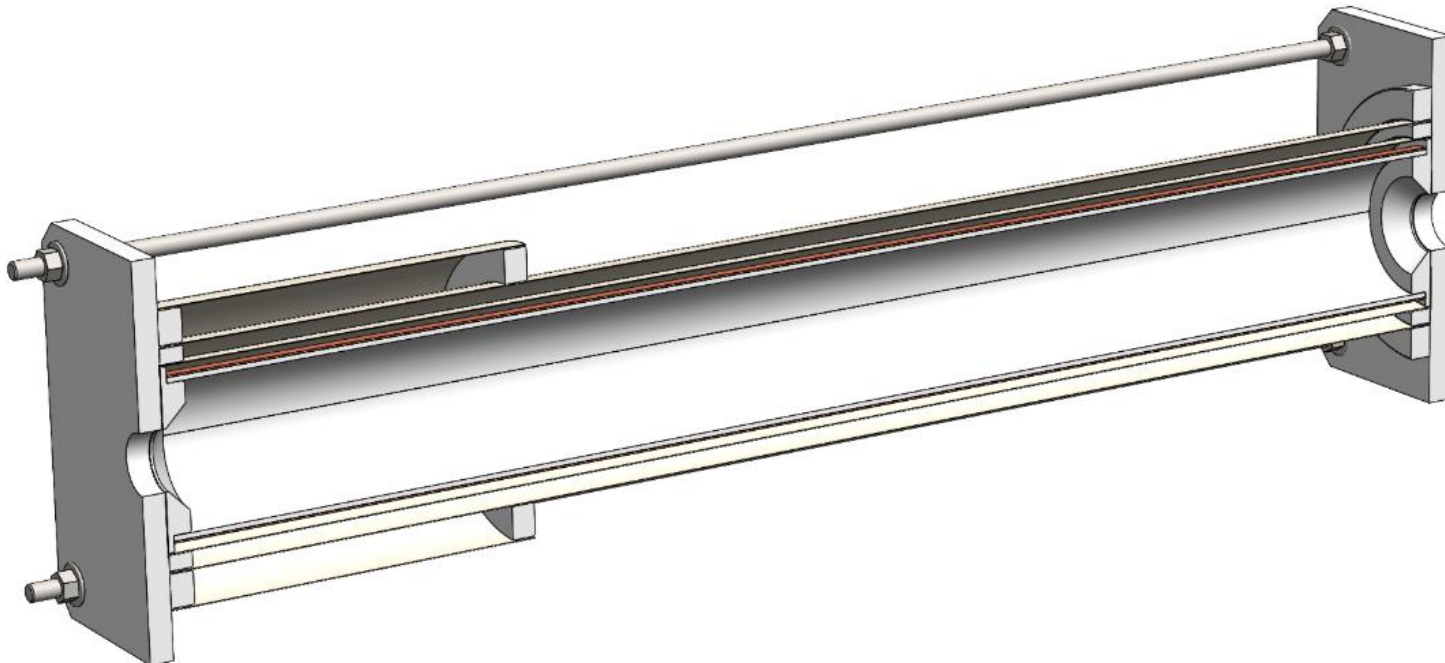
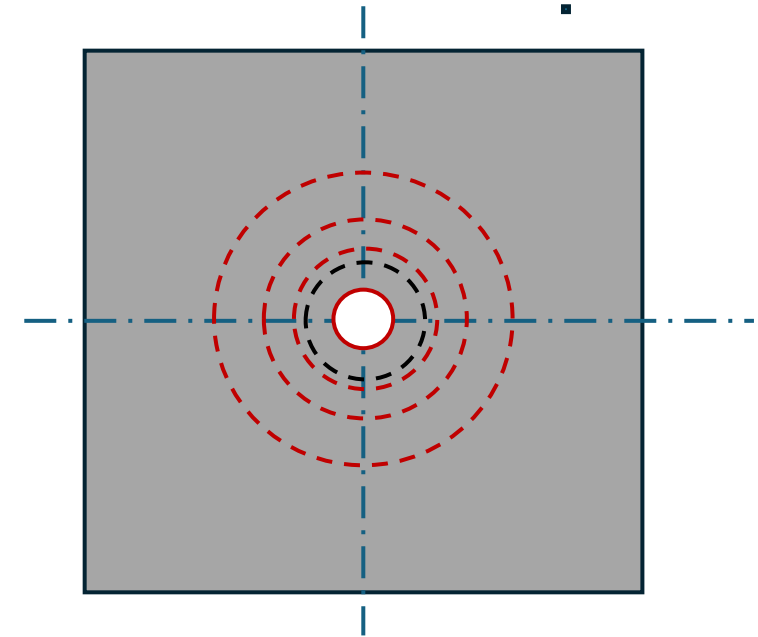
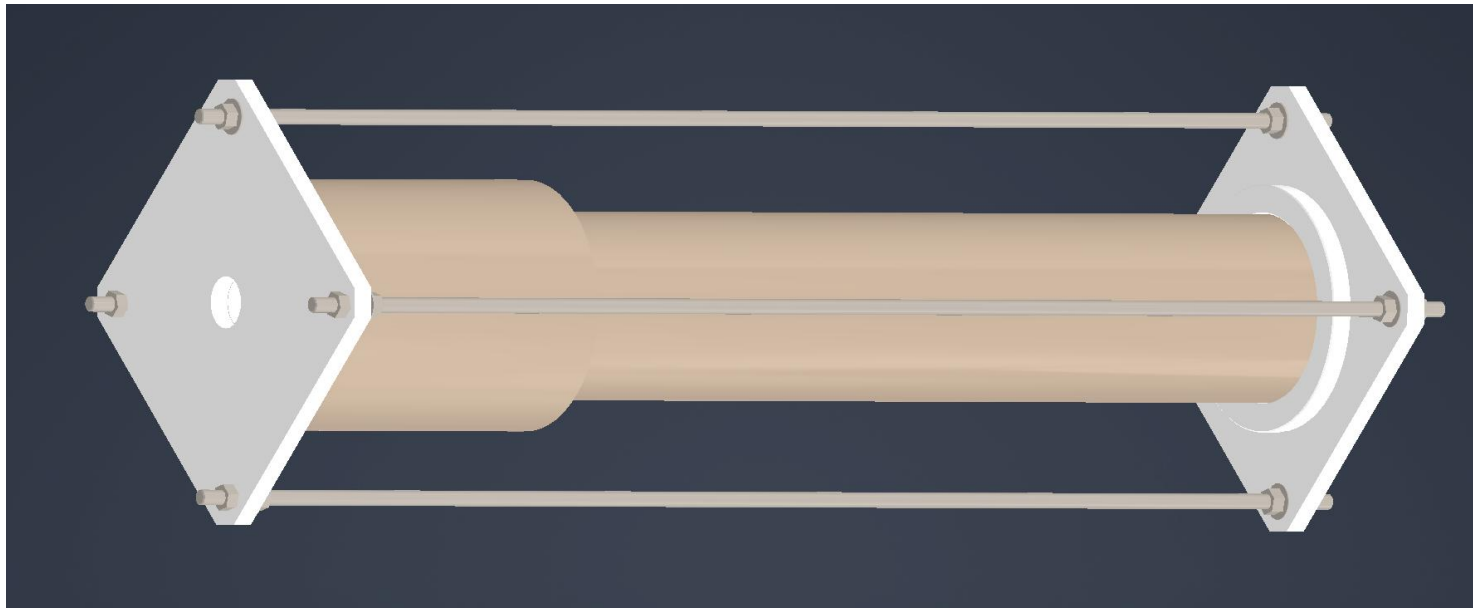
Goals of μ Prototype for n TMM experiment

- Confirm that in n TMM experiment the magnetic shielding is sufficient for obtaining adjustable UNIFORM magnetic field in both magnets in the presence of external perturbations
- Width of n TMM resonance peak (@ 1.5 atm CO₂ $\approx$ 29.5 Gauss) is $\sim$ 0.15 Gauss; field scanning step is 0.04 Gauss.
Goal of variation of uniformity under perturbation is 0.01 Gauss
- Thus, max perturbation at GP-SANS 6 Gauss should be shielded with shielding factor S.F. $\gtrsim$ 600
- Lower SF will imply corrections lowering the sensitivity of n TMM
- Effective SF in n TMM can be function of solenoidal field due to saturation of the flux in mu-metal of internal shield. This should be addressed in prototype tests.

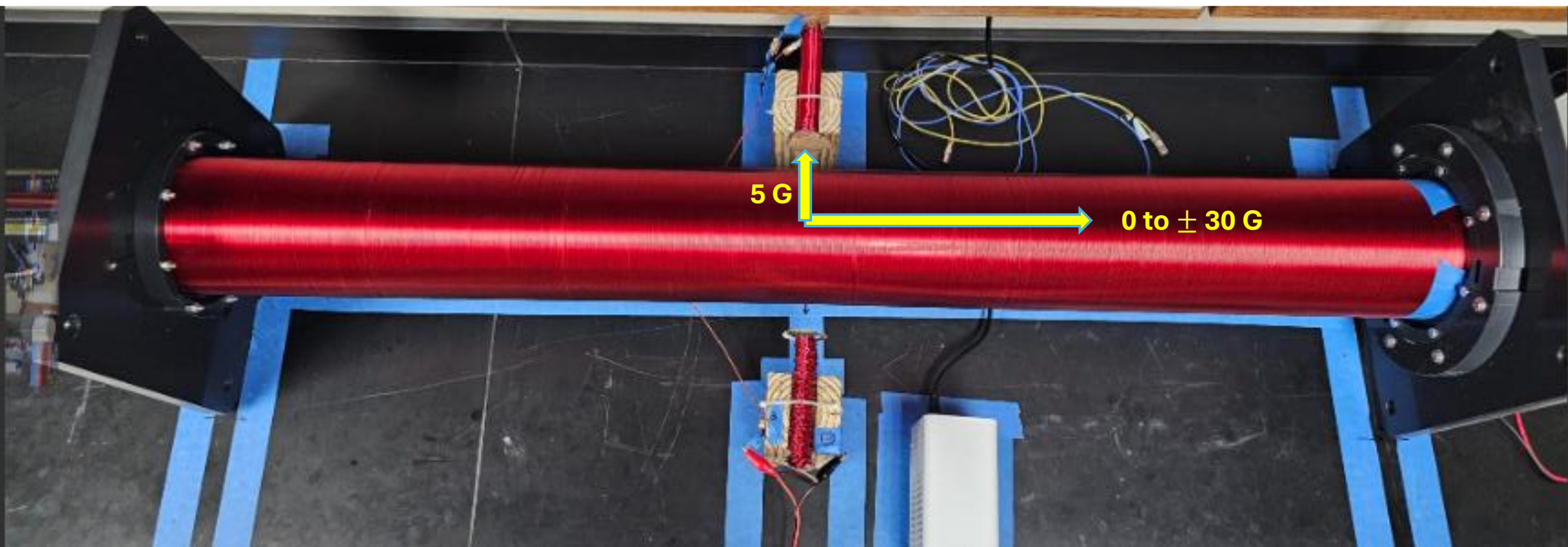


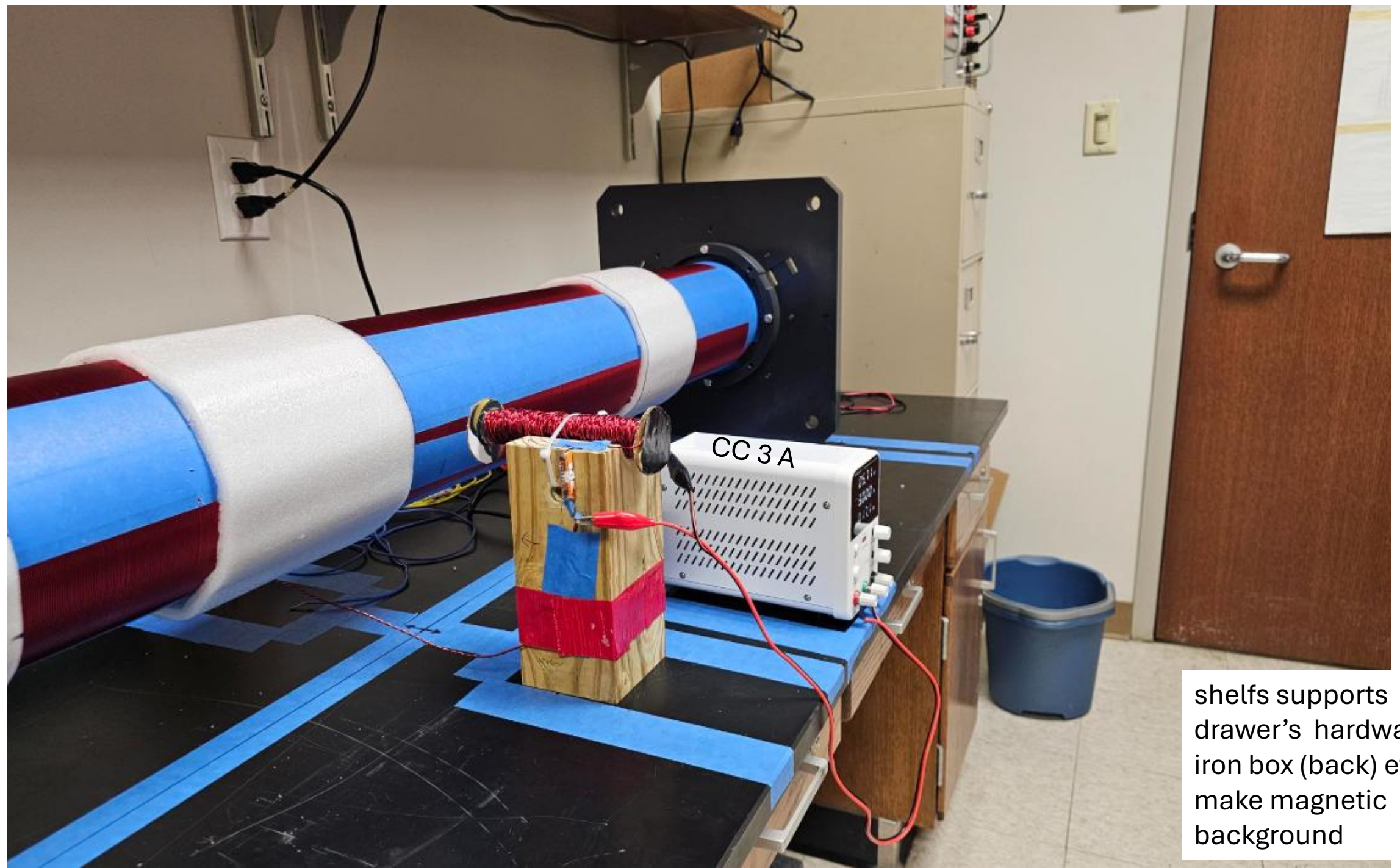
Schematic of μ -prototype



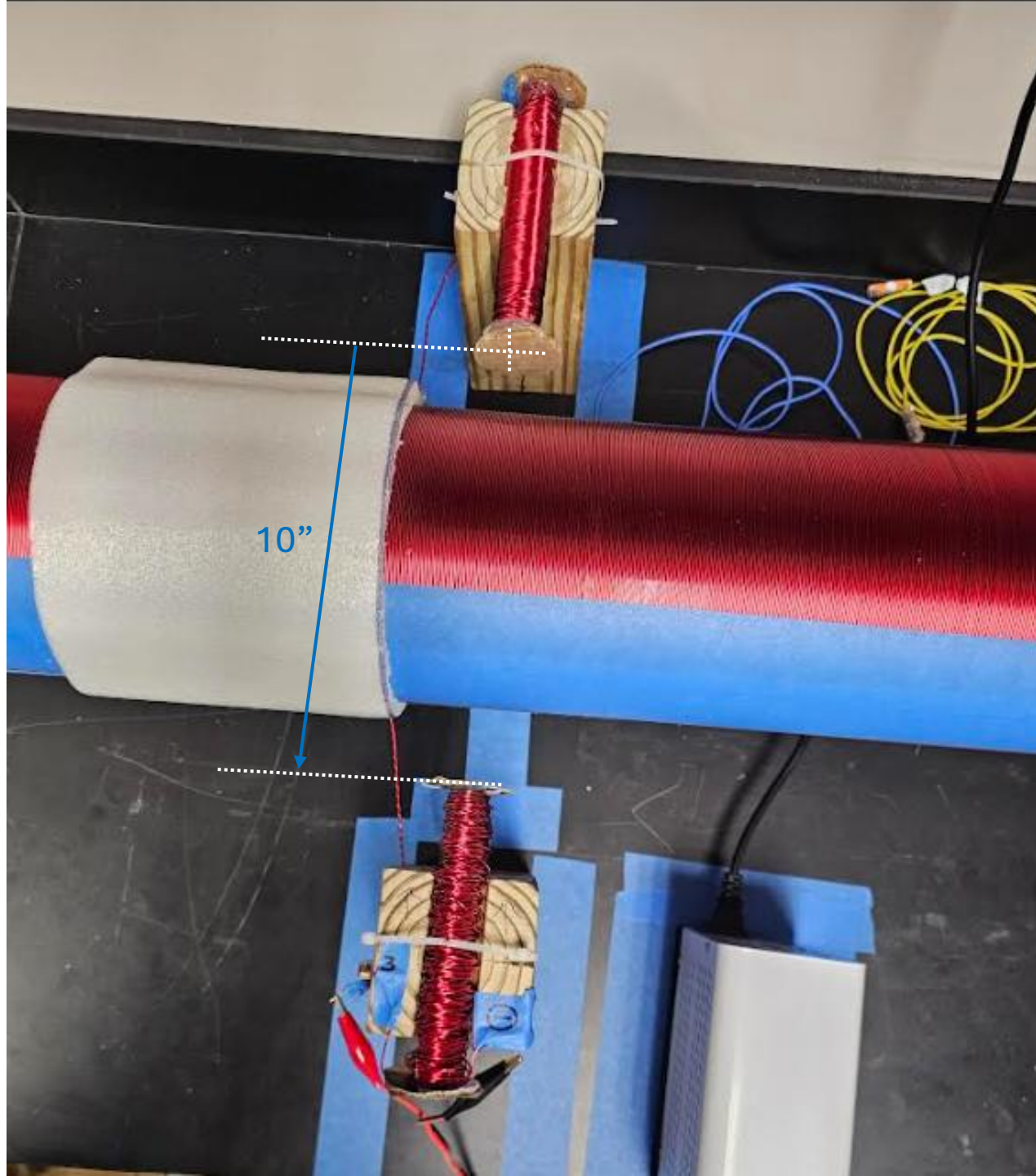


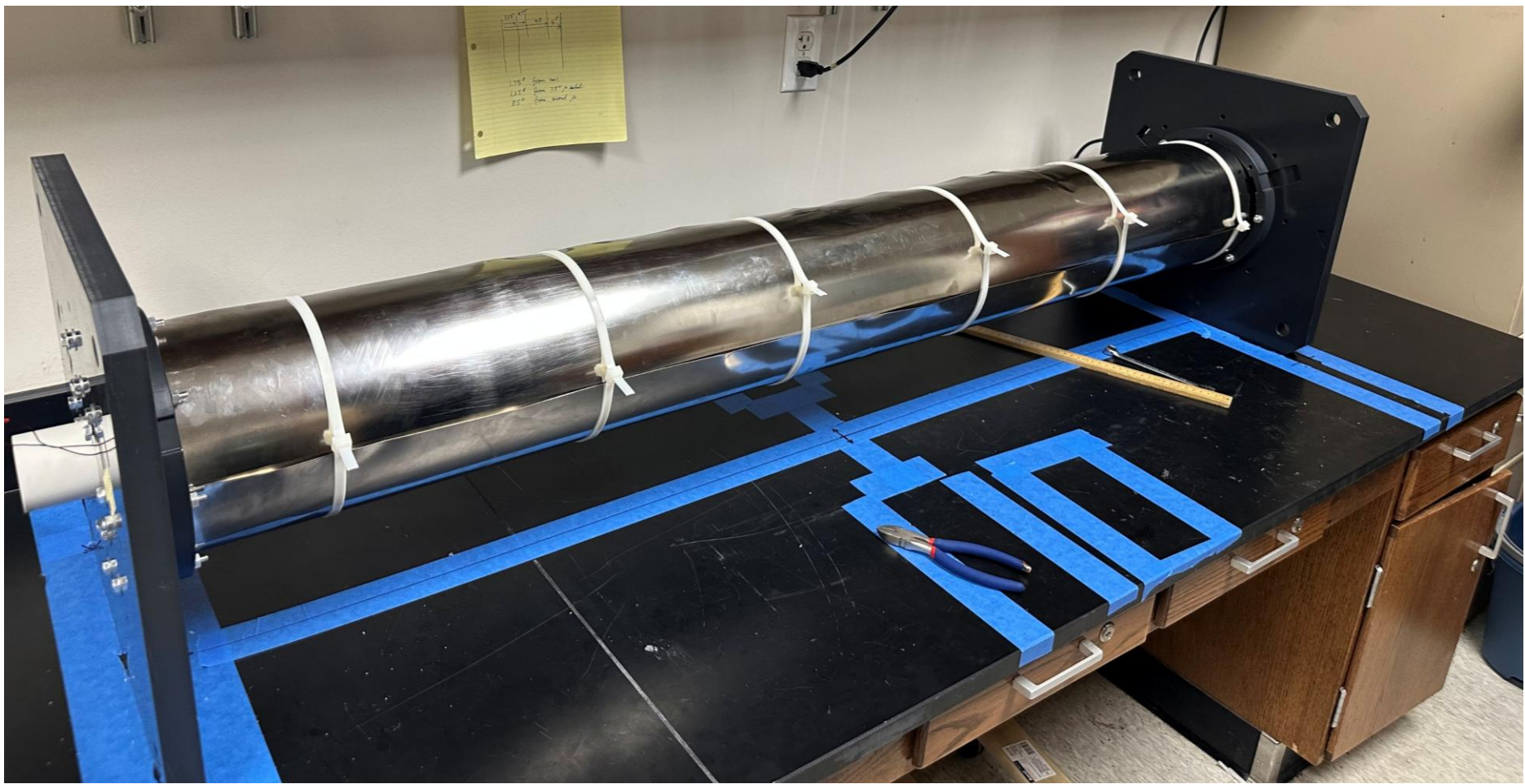
June 15, 2026



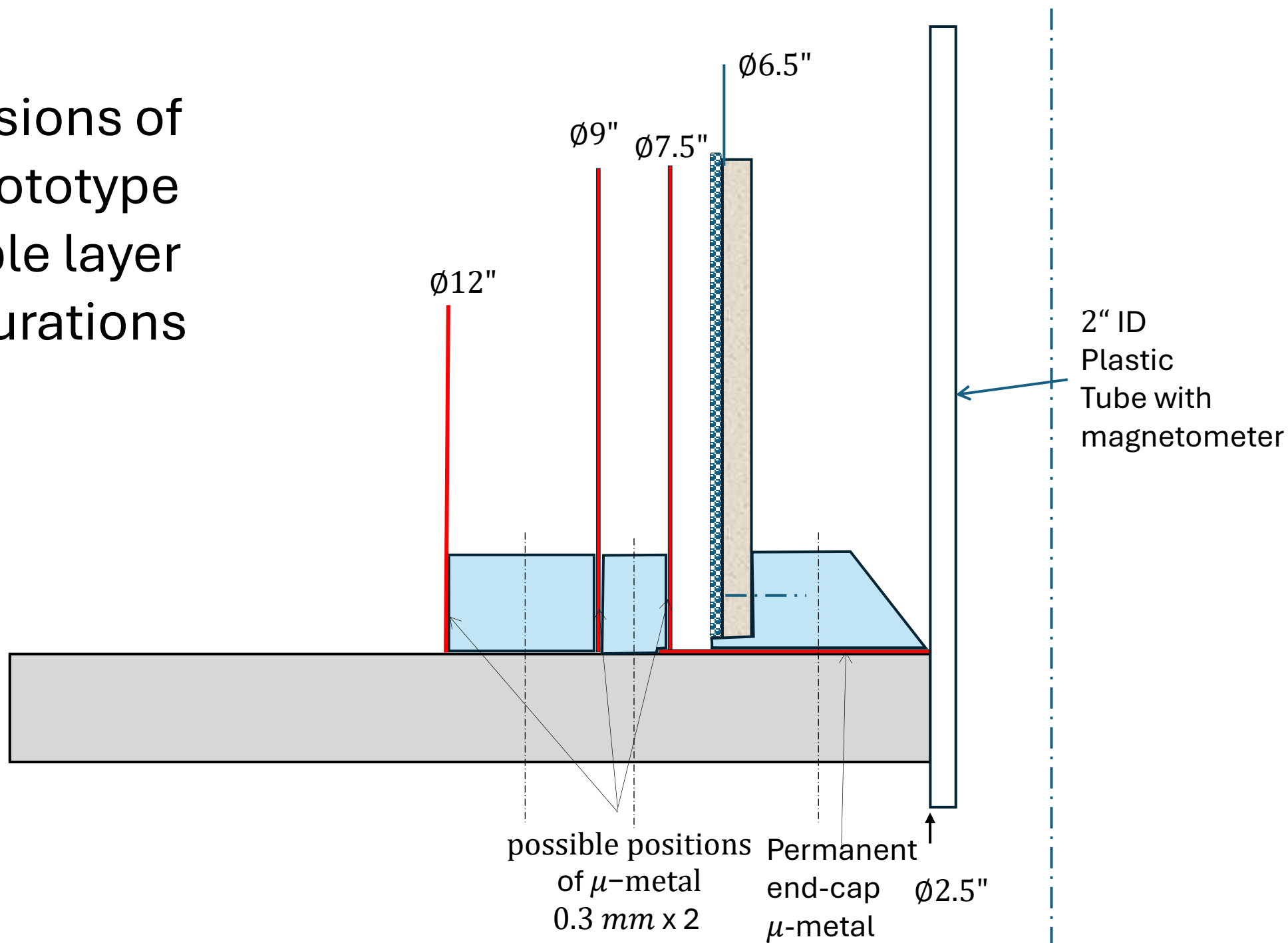


shelves supports
drawer's hardware,
iron box (back) etc.
make magnetic
background

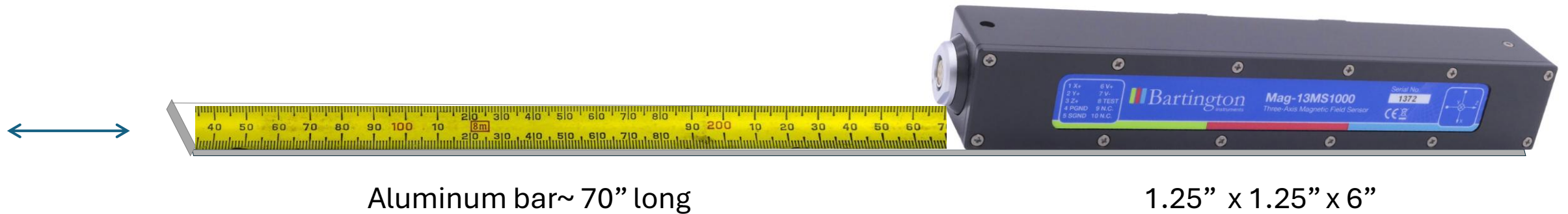




Dimensions of mu-prototype possible layer configurations



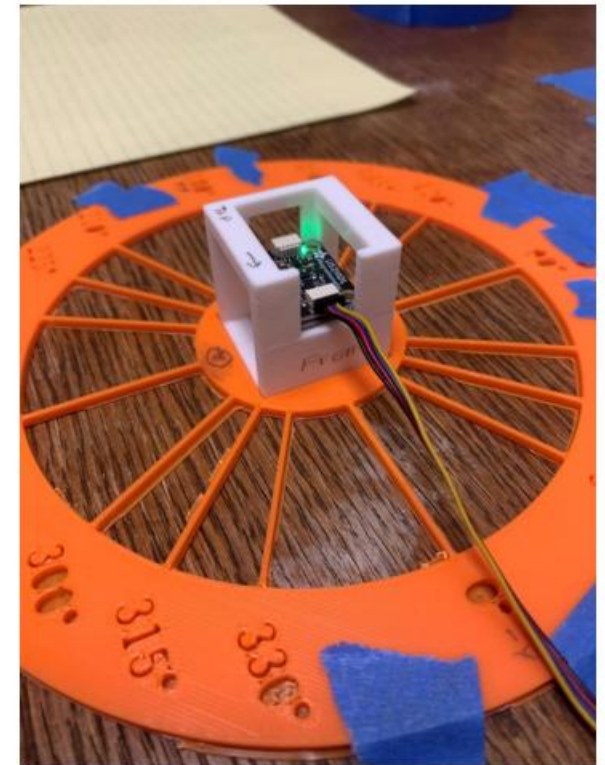
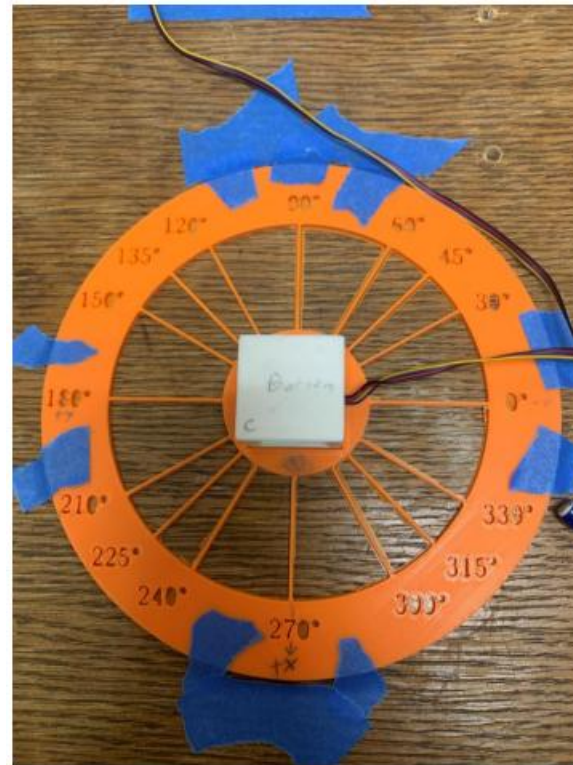
It was designed for
Bartington Flux-gate 3D sensor with Spectromag-6 DAQ system
Mag-13MS1000 with range $\pm 1000\mu T$ (± 10 Gauss)



Limitations: Short range of measurement < 10 G, distortion of B-readings if one of sensors is saturated at high field; no temperature control

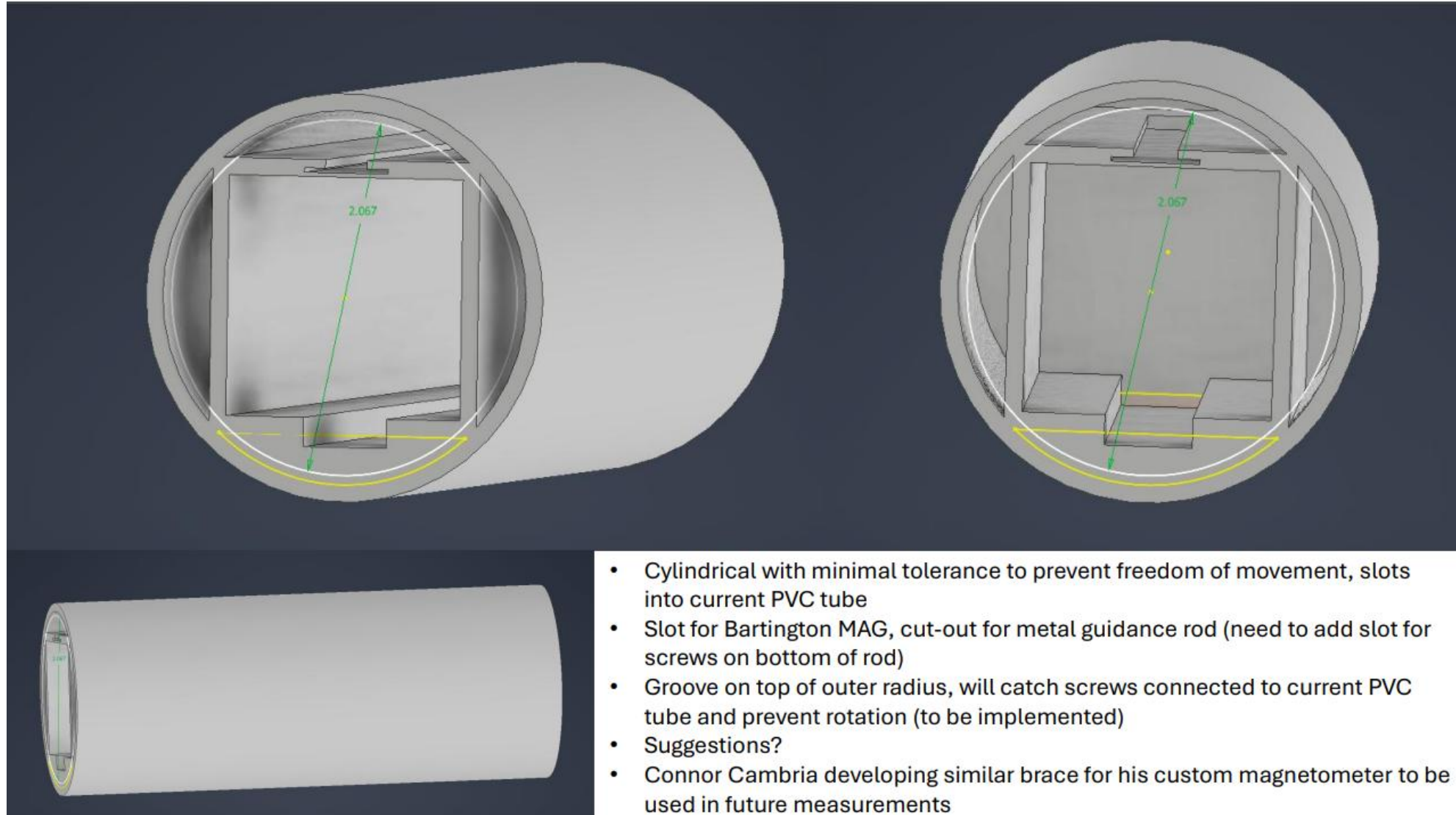
Alternative chip magnetometer LIS2MDL with range ± 50 Gauss

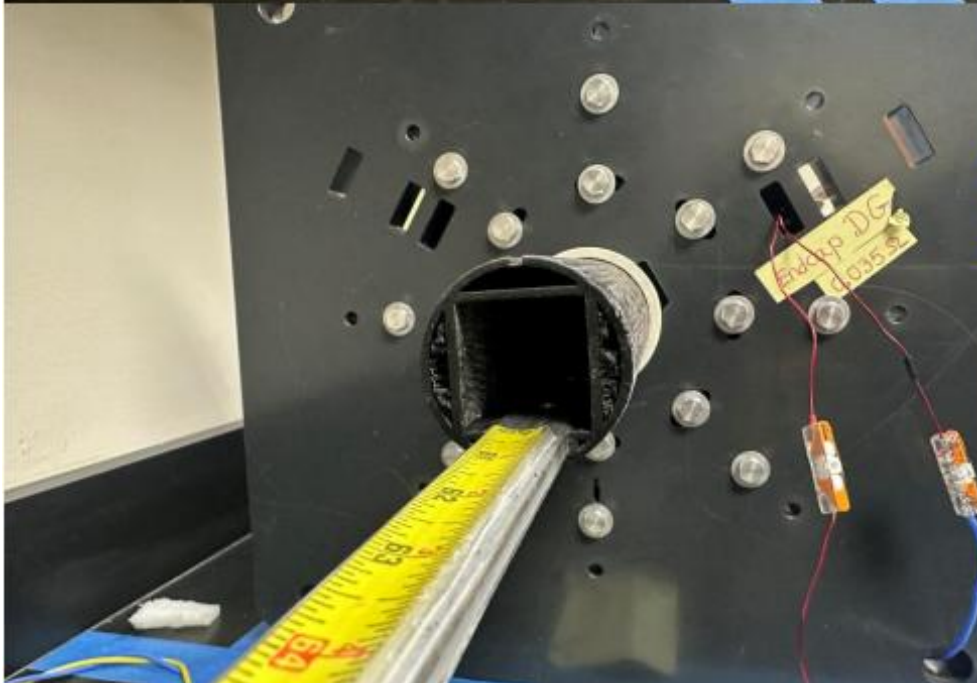
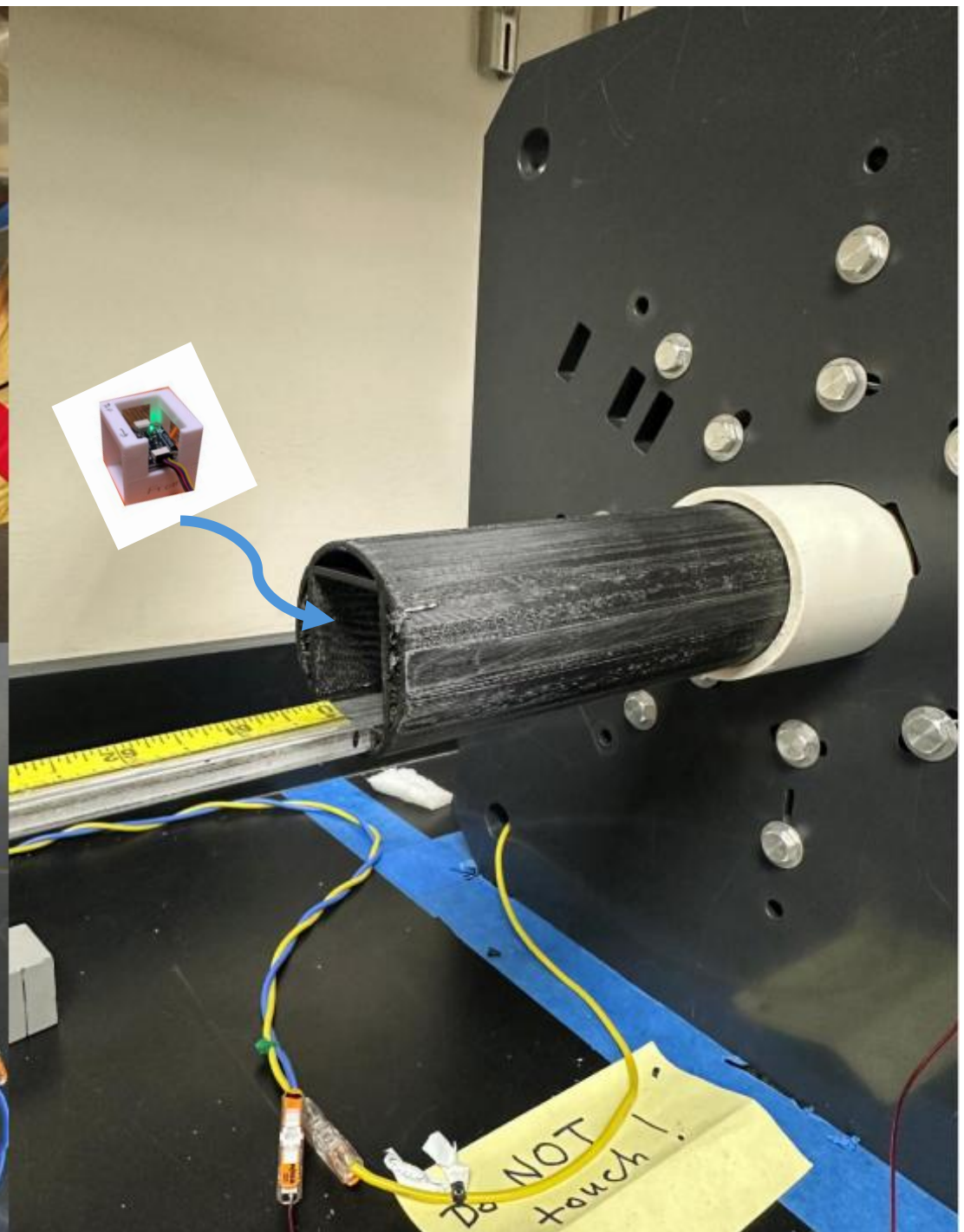
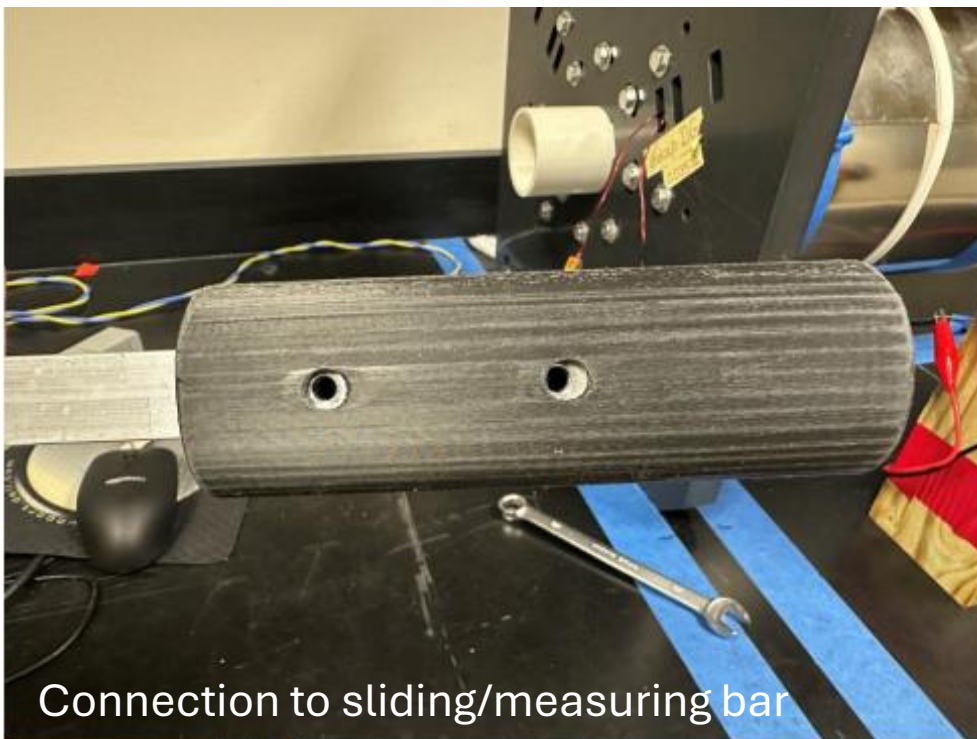
Challenges: in control of alignment (angle B_x/B_z). **Essential temperature dependence.**
Introduced thermocouples and compressed air flow in the guide tube to maintain thermal equilibrium in field measurements.

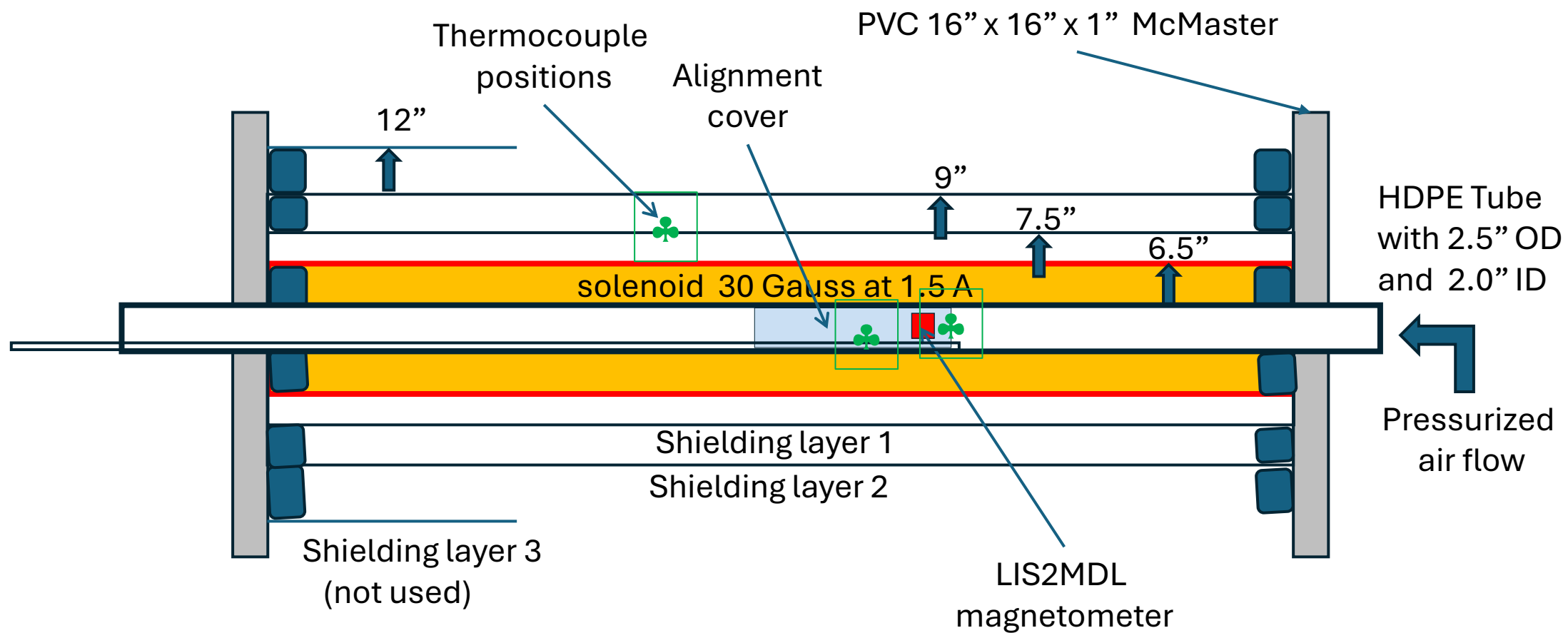


LIS2MDL inside the box with the side 1.25"

3D printer model of magnetometer housing



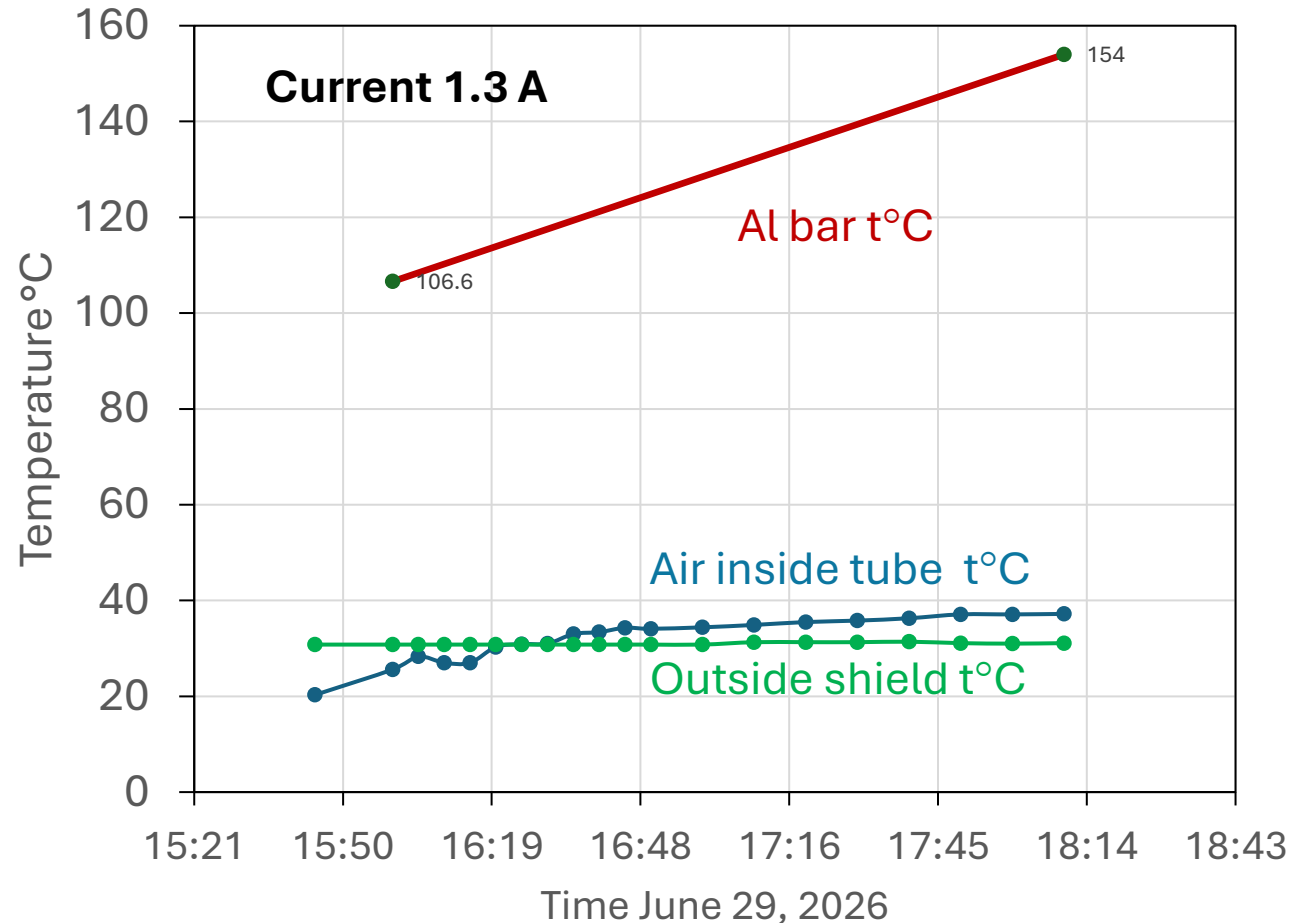




At high coil currents, 1.3A (~ 26 G field) dissipated power is ~ 47 W.

Inside the plastic measurement tube temperature can substantially increased

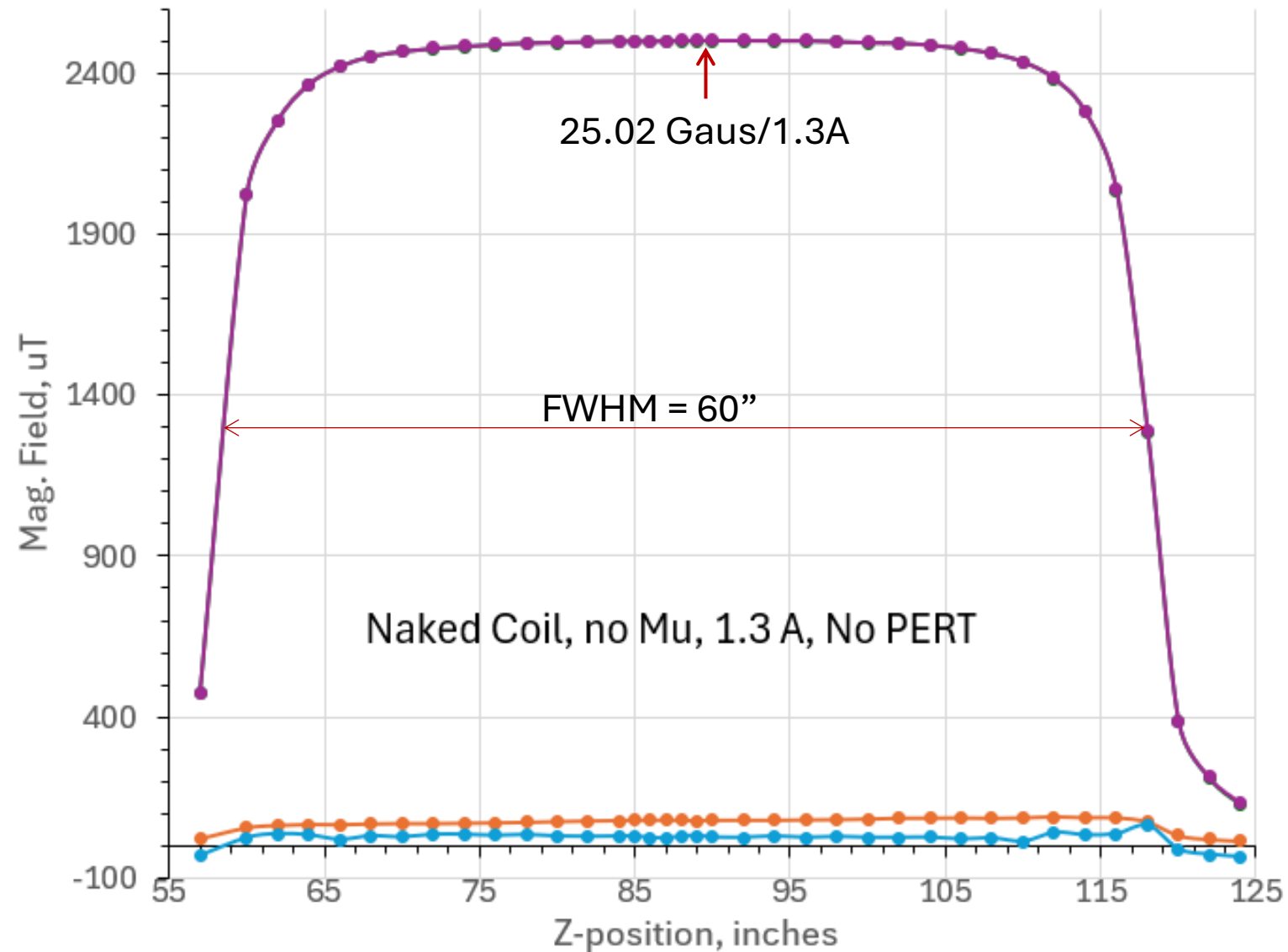
Configuration : inner mu-shielding is 0.6 mm, chip magnetometer in the center of coil z=88”



**Forced air colling flow
inside the tube
is mandatory**

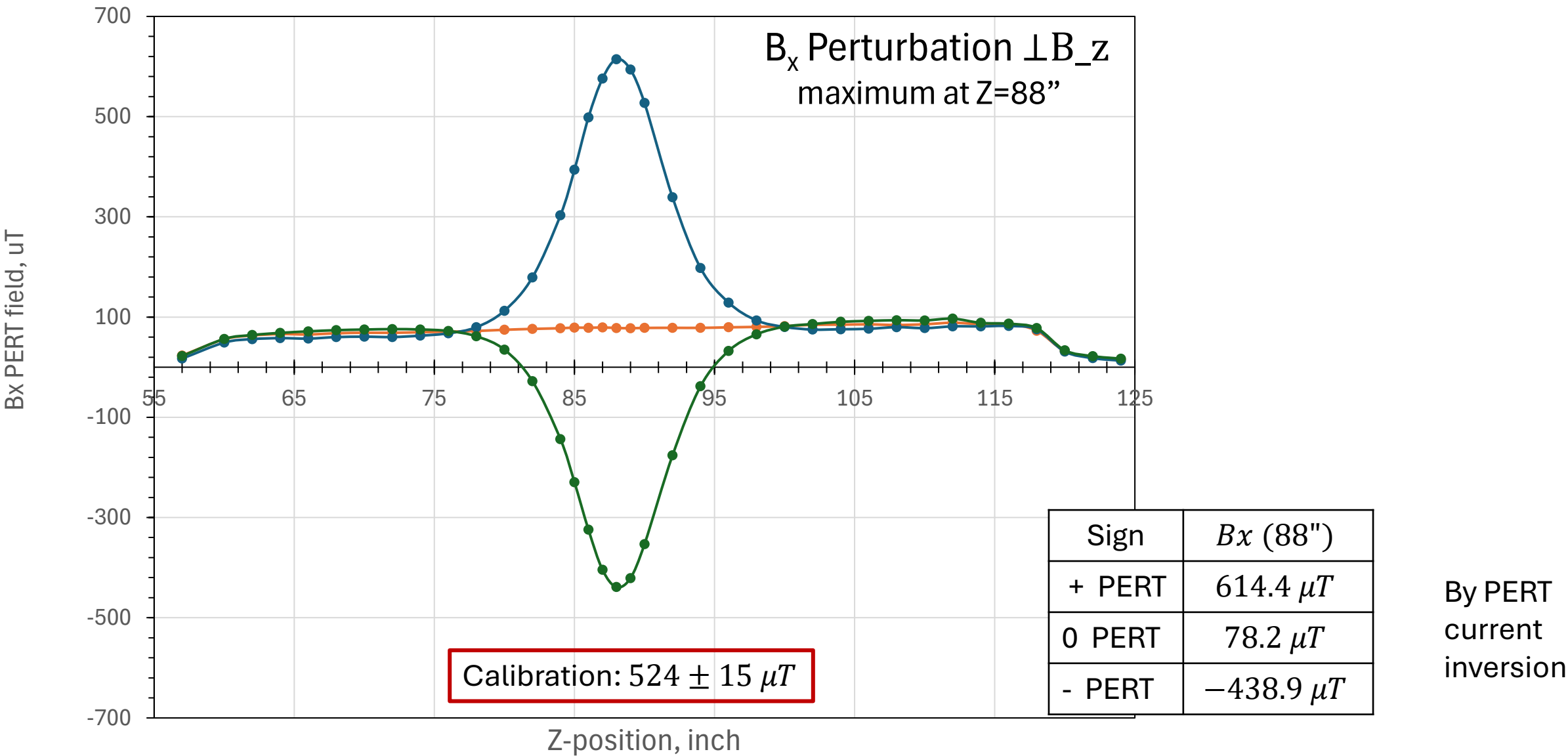
- When previously we used 10-Gauss Bartington magnetometer (0.5 A) power dissipation was ~ 10 times lower. Now we perform all measurements with chip magnetometer and field 25 Gauss.
- With high current operation longitudinal position scan are not reliable even with air flow cooling due to temperature variation along z

Naked Coil 3 field components, no mu-metal, no PERT

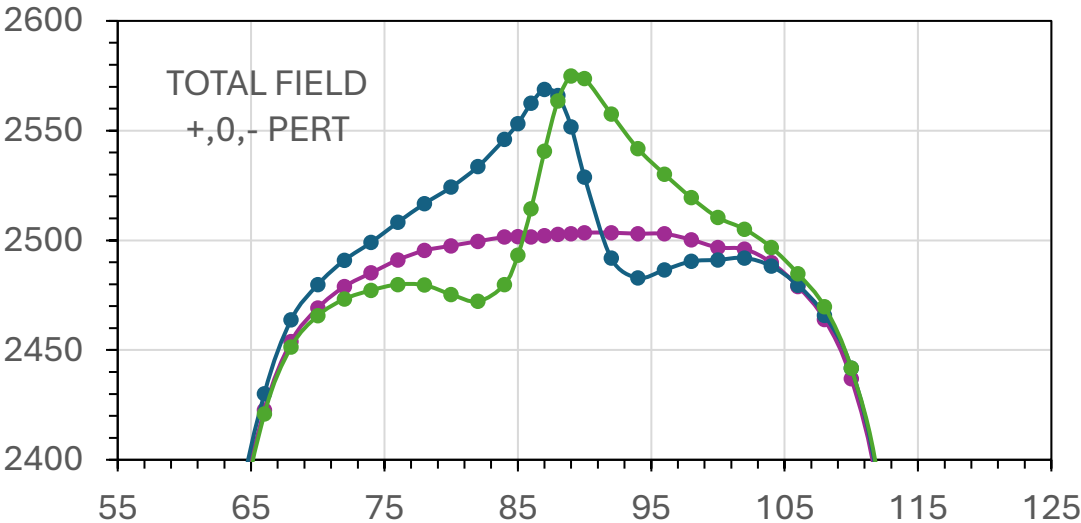
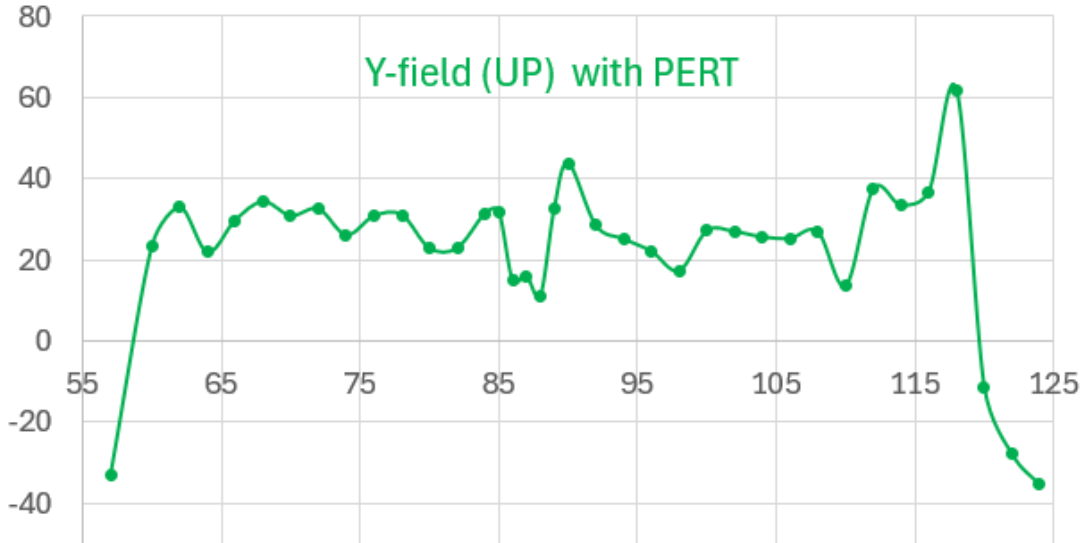
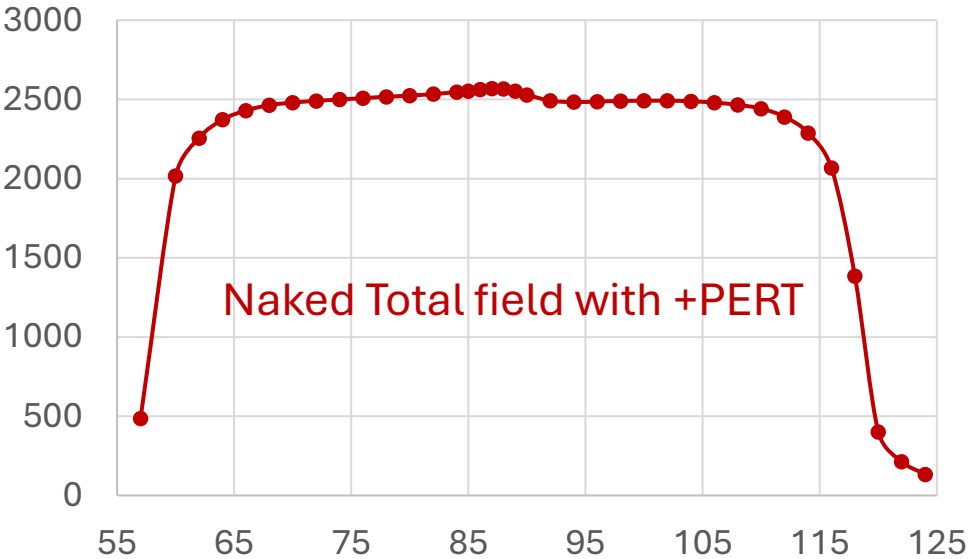


mu-metal endcaps
are installed

Naked Coil, with PERT, Coil current 1.3 A



Naked Coil with current 1.3 A and PERT

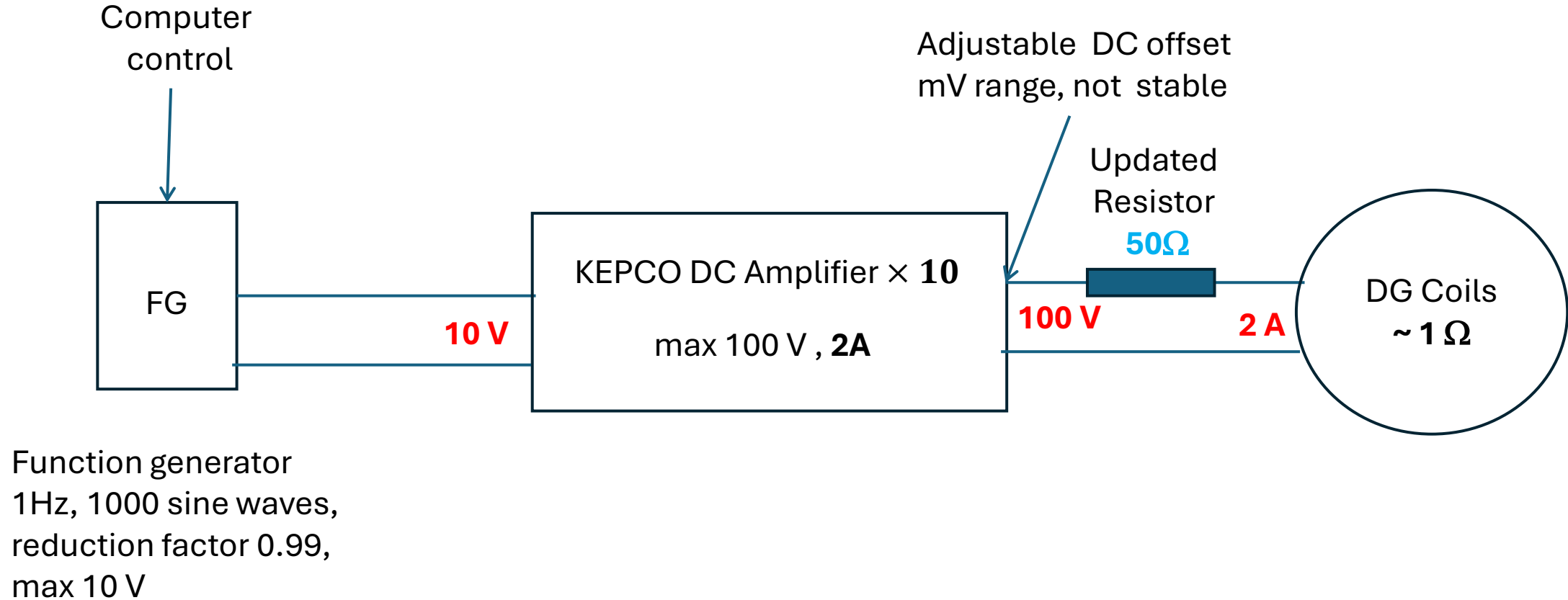


Two perturbation magnets are slightly misaligned

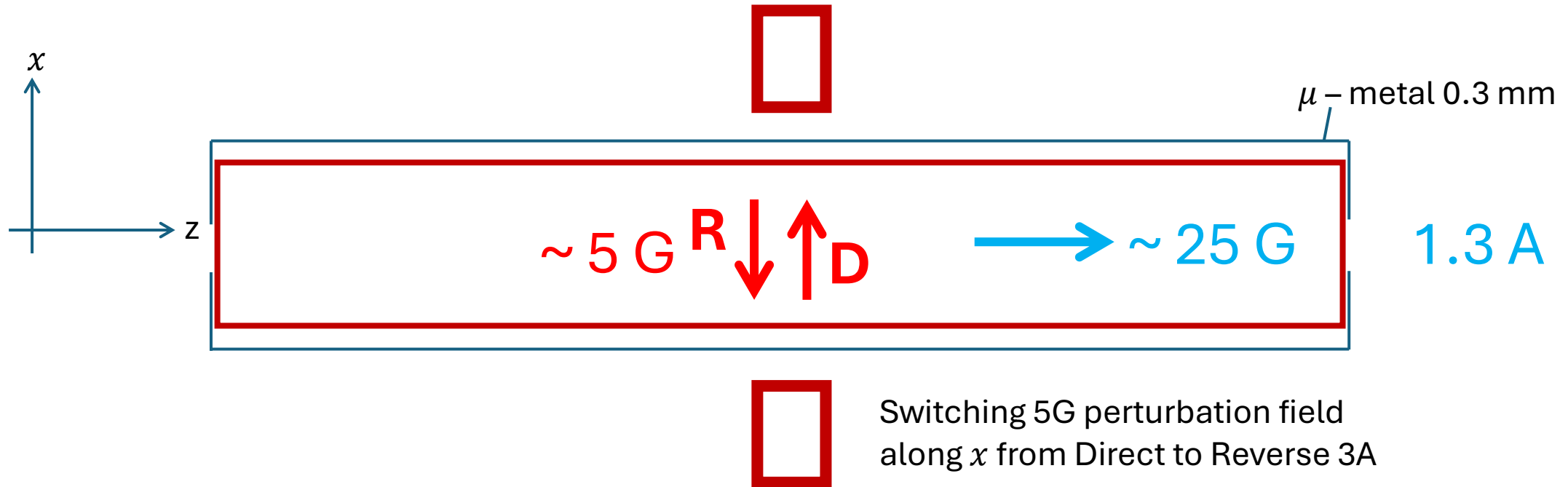
In General:

- Longitudinal scans along z with magnet ON are subject of temperature variation of readings
- Air blowing through the guide tube improves the t-variation, however, needs time for thermalization
- Best compensation of t-dependence can be reached when magnetometer is at fixed z-position (88" – peak position) and air-cooled for long time for equilibrium setting.
- In this case, if coil current variation/measurements are quick; effect of temperature variation is small.

Scheme of degaussing DG

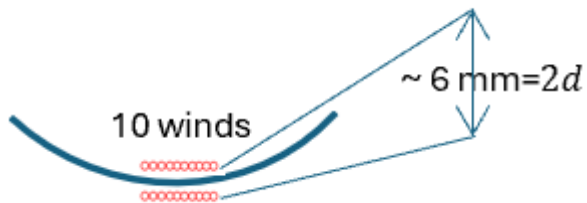
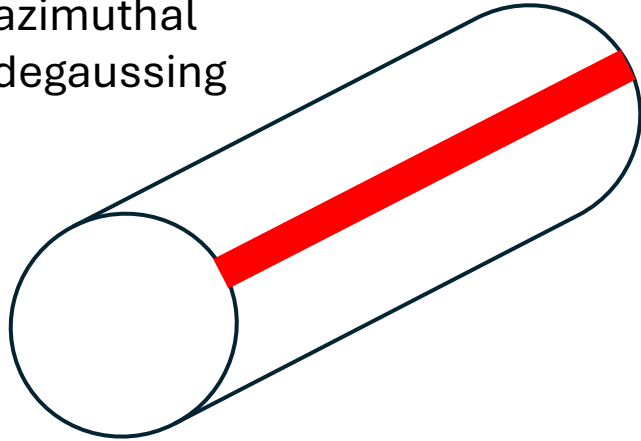


3 degaussing circuits (cylinder and endcaps) serially connected with total resistance $\sim 1 \Omega$



Degaussing in mu-prototype at UT

azimuthal
degaussing

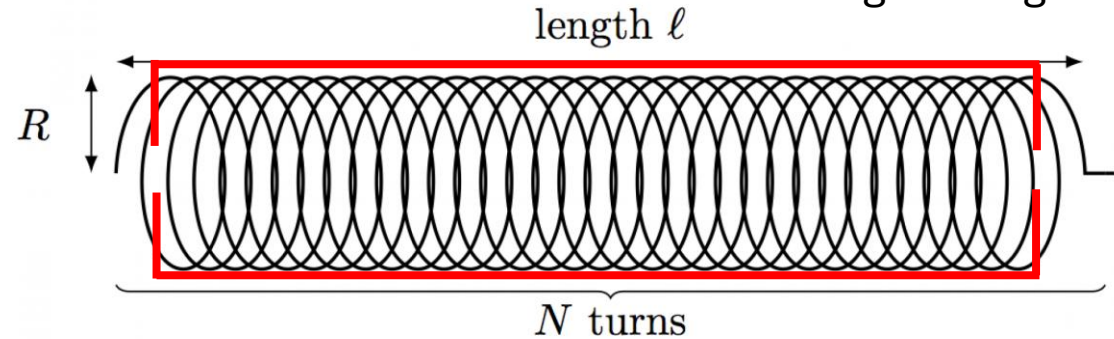


$$\text{Max degaussing current} = \frac{100V}{50\Omega} = 2A$$

$$B = 10 \times \mu_0 \frac{I}{\pi d} = 26.7 \times 10^{-4} T = 26.7 \text{ Oe} \rightarrow 0.76 \text{ T}$$

↑
Max azimuthal
field in mu-metal

longitudinal
degaussing



$$R = 6.5"/2 = 82.6 \text{ mm}$$

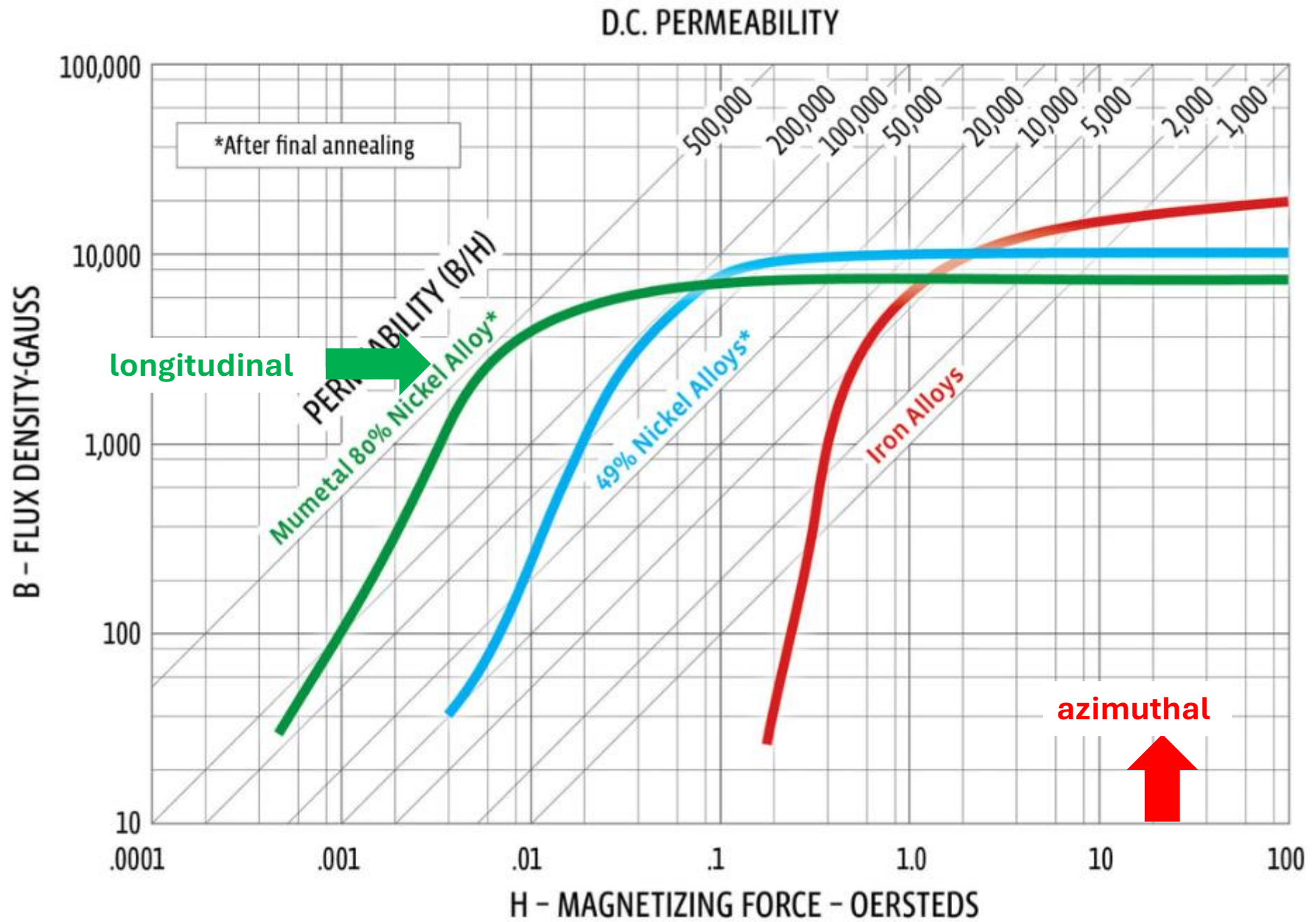
$$\Phi = B \cdot A = 25G \cdot \pi \cdot (82.6)^2 \cong$$

$$\cong B_\mu \cdot 2\pi \cdot (82.6) \cdot 0.46$$

Magnetic field inside mu-metal

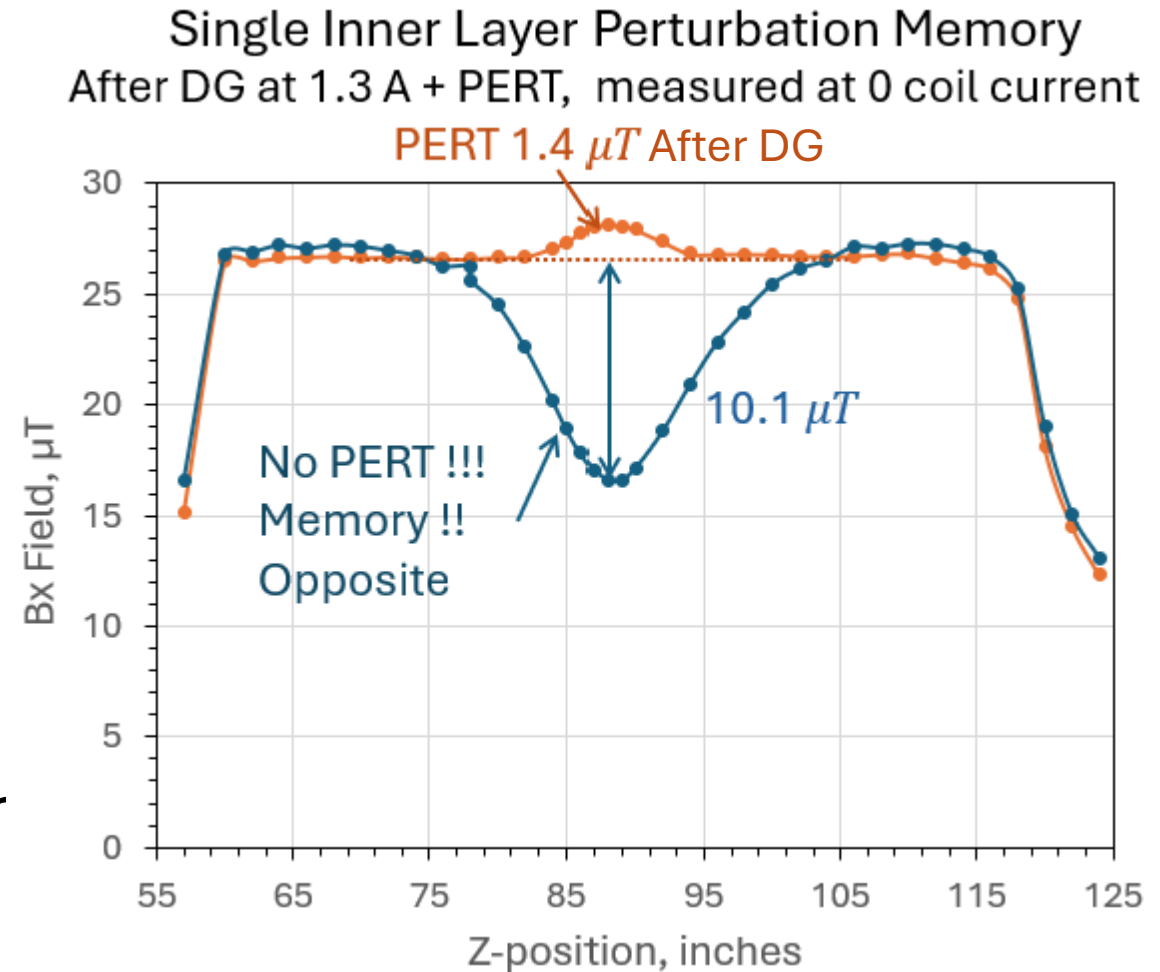
$$B_\mu(\text{max}) = 0.23 \text{ T}$$

↑
Max longitudinal
field in mu-metal



- Degaussing of mu-metal should be performed with the magnet and perturbation state close to operational !

- Our standard mode of degaussing was with + PERT and coil 1.3 A
- By degaussing the shielding remembers the perturbation with opposite sign →
- For Shielding Factor measurement it was important to set magnetometer position on the peak of PERT

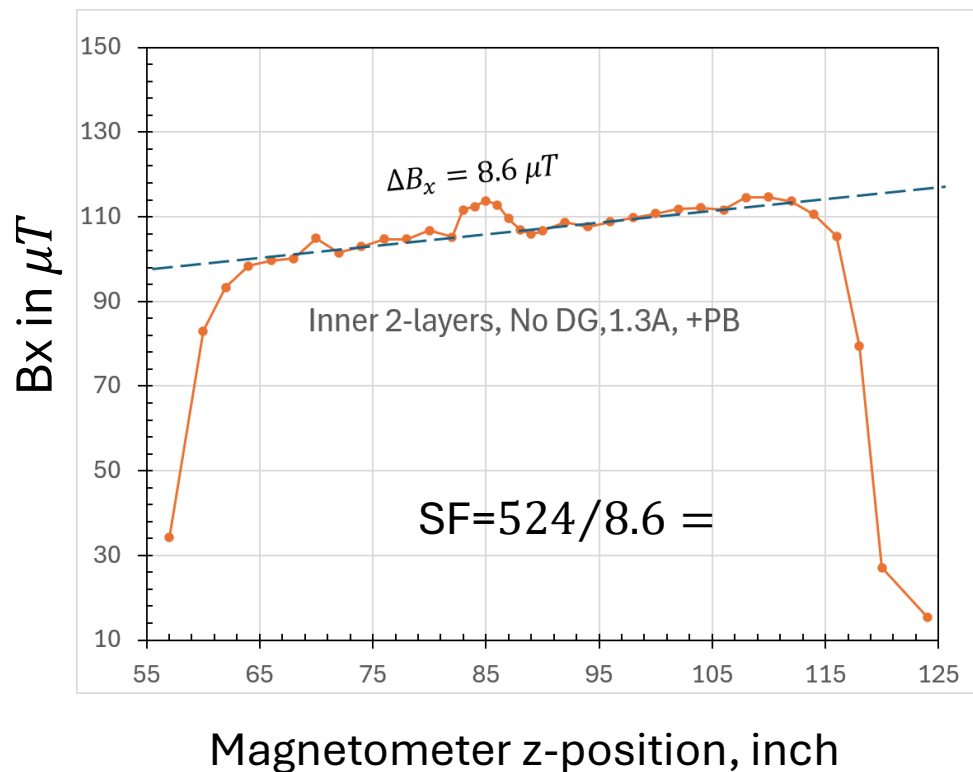


Two methods of perturbation measurement in the presence of μ shield:

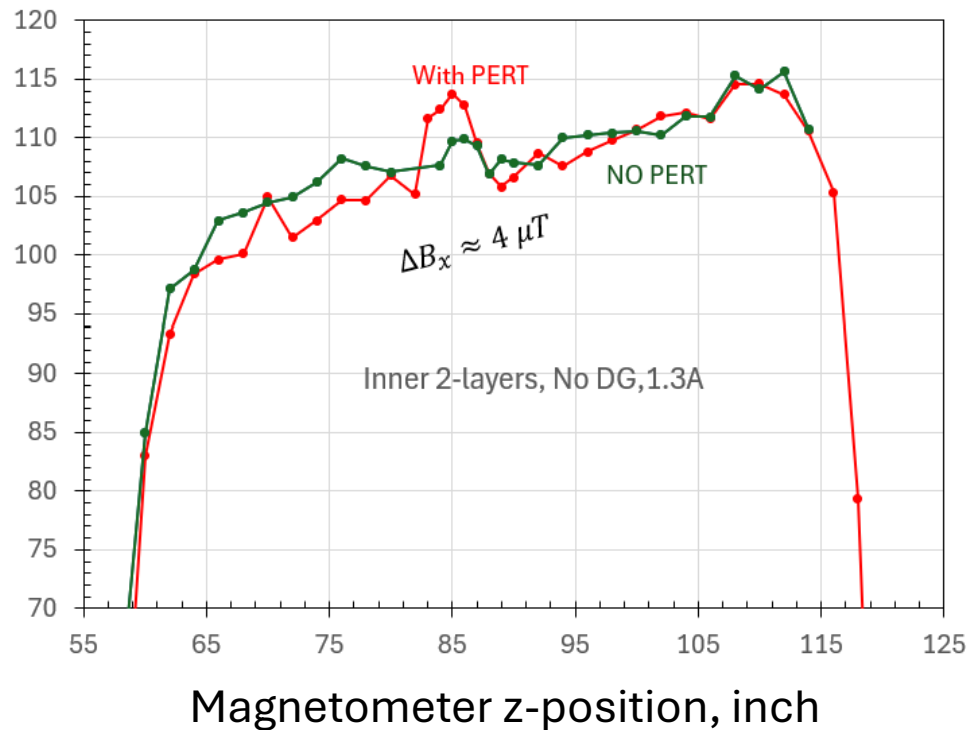
1

Based on z-scan

1a



1b



- Slope due to temperature distribution along the tube
- With Air flow present

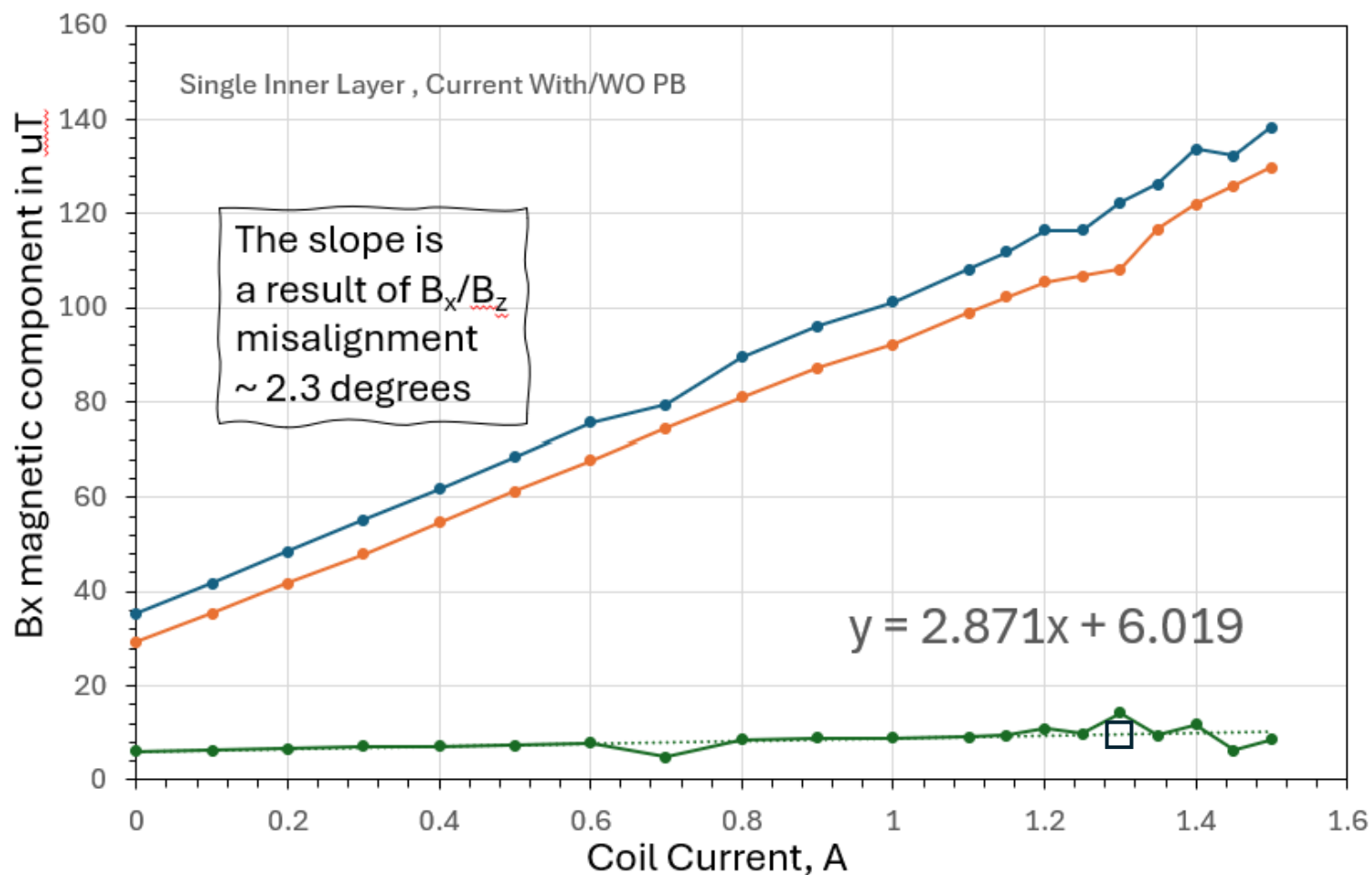
2

Magnetometer z-positioned at the peak of perturbation.

Thermal equilibrium reached with Air flow at coil current 1.3 A

Quick variation of coil current from 0 A to 1.5 A with and without PERT.

Perturbation signal is measured at 1.3 A



Shielding factor
at $I=1.3$ A

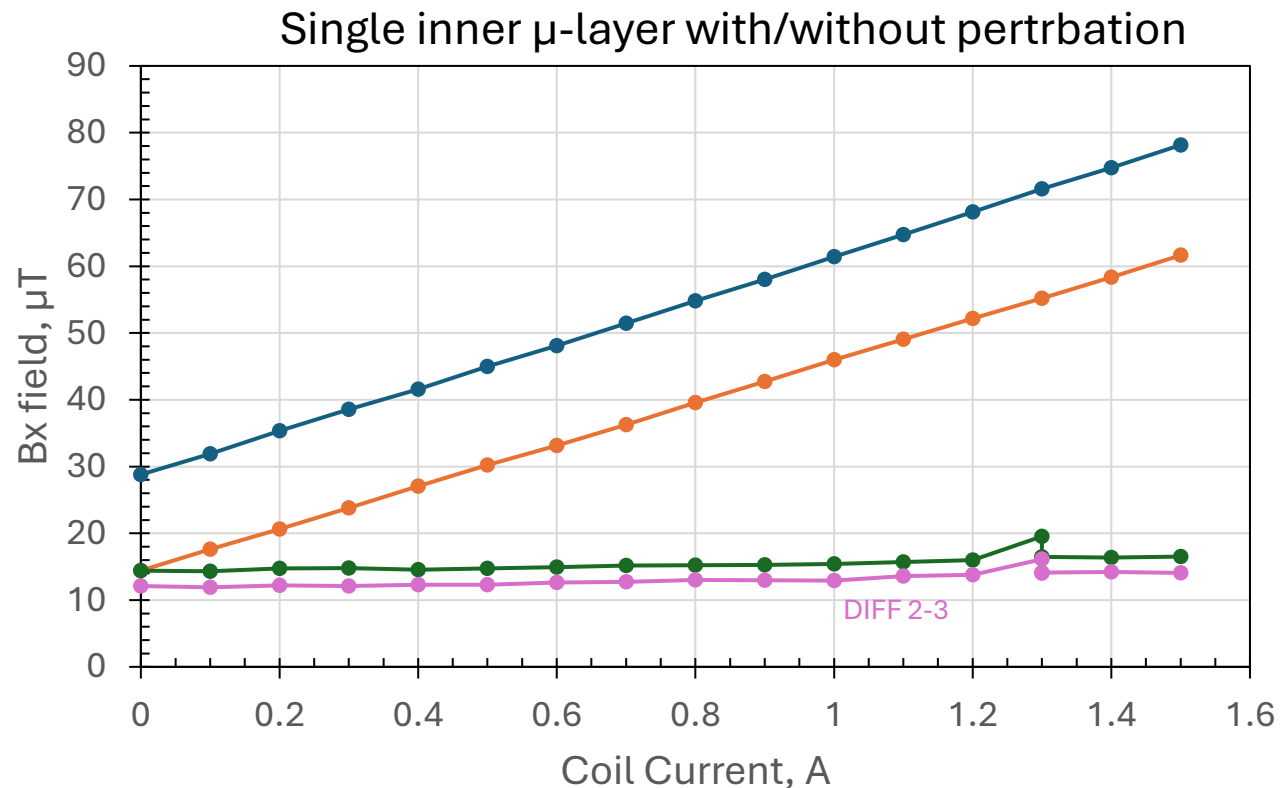
$$SF_1 = 524/9.75 = 53.7$$

Method 2 seems
more reliable

1 μ - inner layers

- Major question was how shielding factor will change with the saturation of mu-metal serving as flux return yoke of the solenoid.

Standard degaussing, B_x field with/without perturbation at z peak position 88"

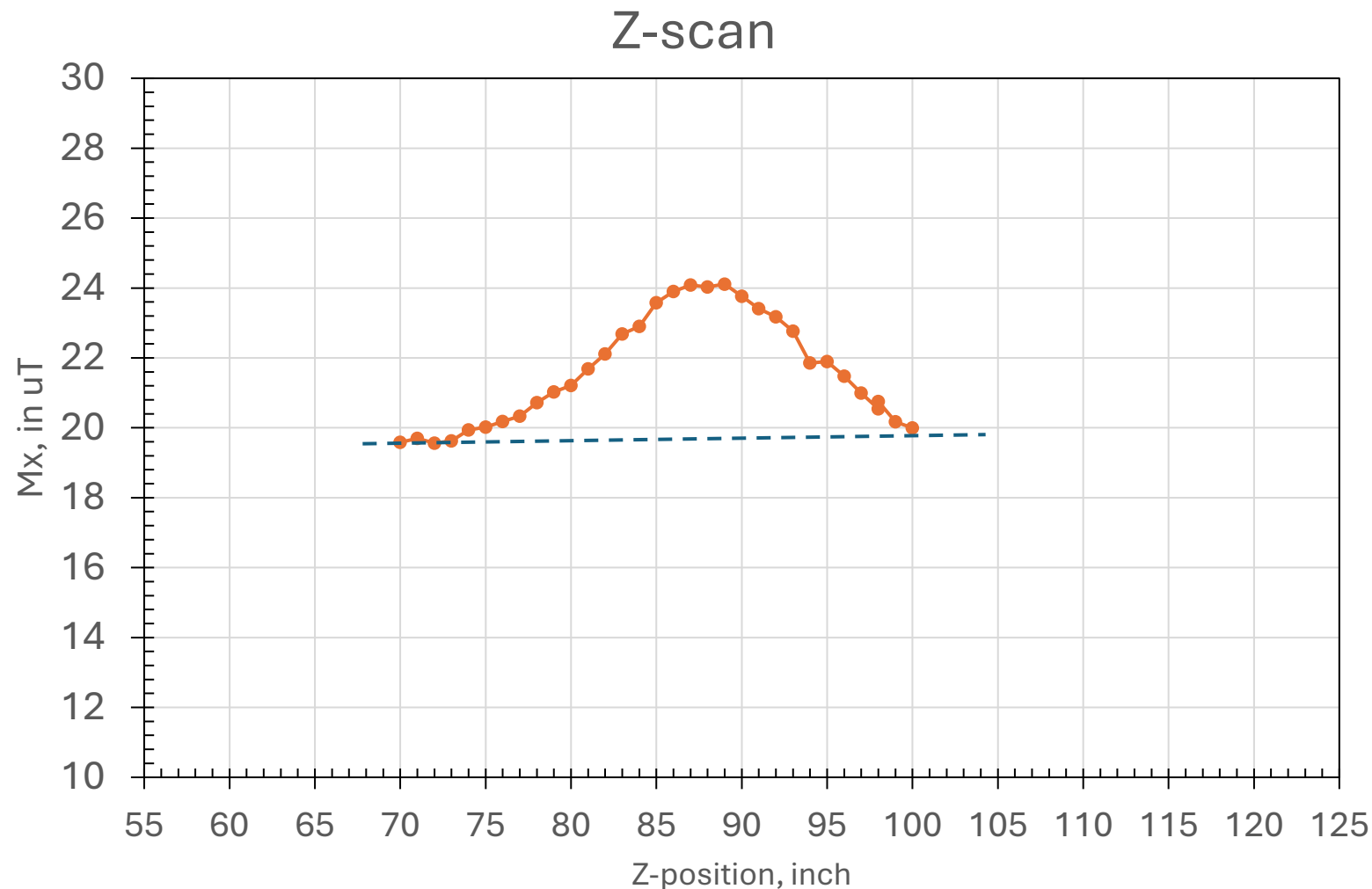


Single inner mu-layer
Seen PERT at 1.3 A
 $\overline{\text{PERT}} = 16.1 \pm 1.8 \mu T$

DIFF dependence on current
 $\sim 14.18 + 1.47 \cdot i(A) \mu T$

2 μ - inner layers

Double shielding is good. Difficult to see the perturbation peak after 2- μ shield
Special degaussing with 0A and NO PERT. Then, Z-scan with PERT, current 0 A



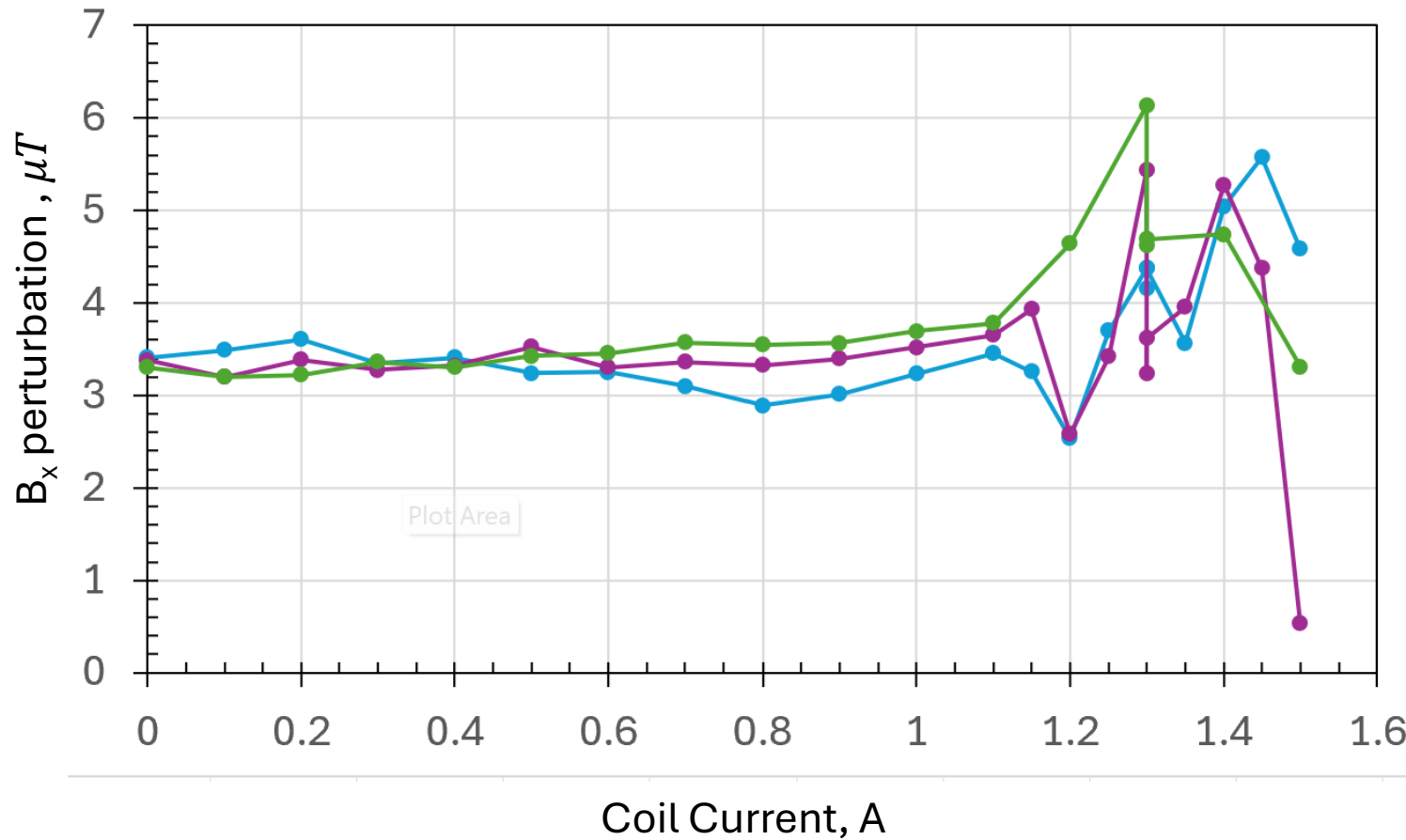
Peak position 88"
Perturbation $\sim 4.5 \mu T$

2 μ - inner layers

After Standard Degaussing, Z=88". DIFF=PERT-NOPERT, with Air

At smallest PERT signals (double layer) with large coil current temp. regime is not stable

Perturbation Difference vs Coil Current



Double inner mu-layer

Seen PERT at 1.3 A

$$\overline{\text{PERT}} = 4.51 \pm 0.61 \mu T$$

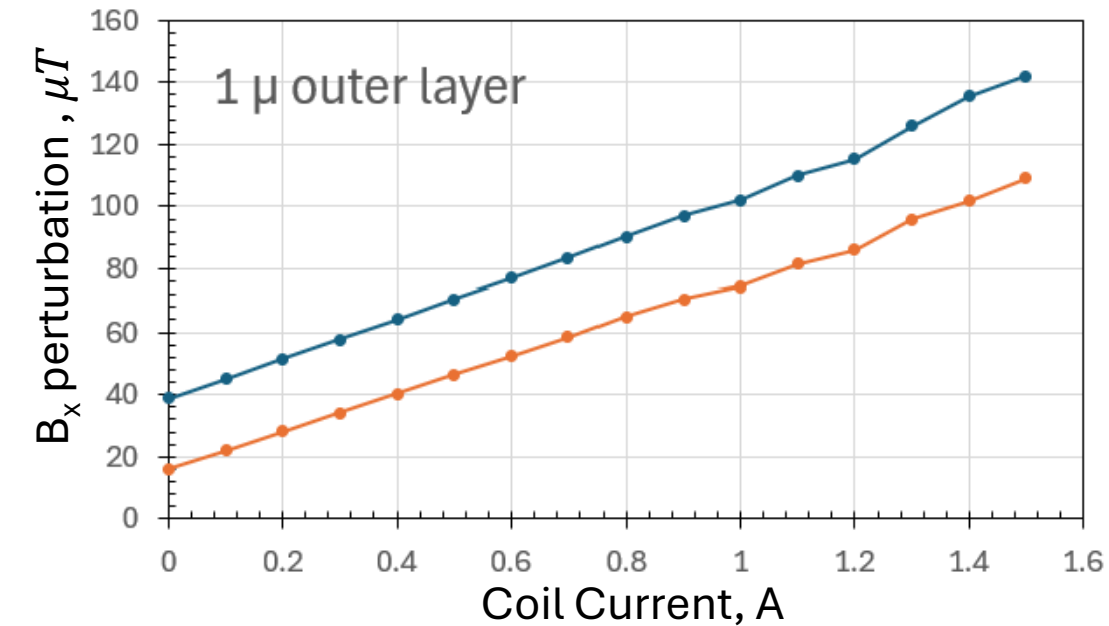
DIFF dependence on current

$$\sim 1.2965i^2 - 1.2367i + 3.5084 \mu T$$

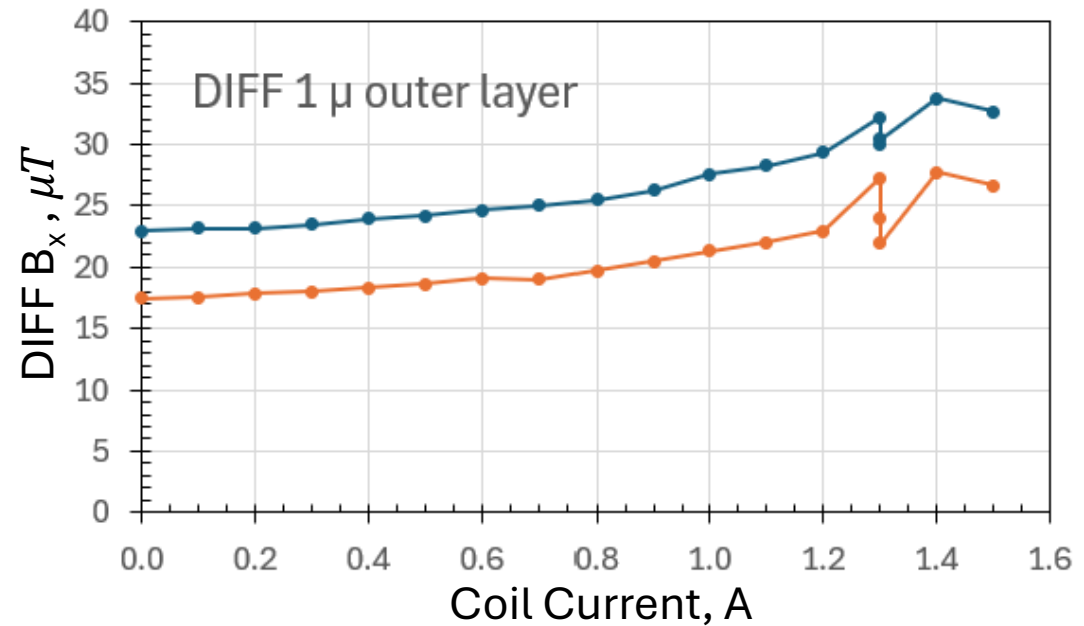
$$P = I^2 R$$

1 μ - OUTER layers

Standard degaussing, B_x field with/without perturbation at z peak position 88"



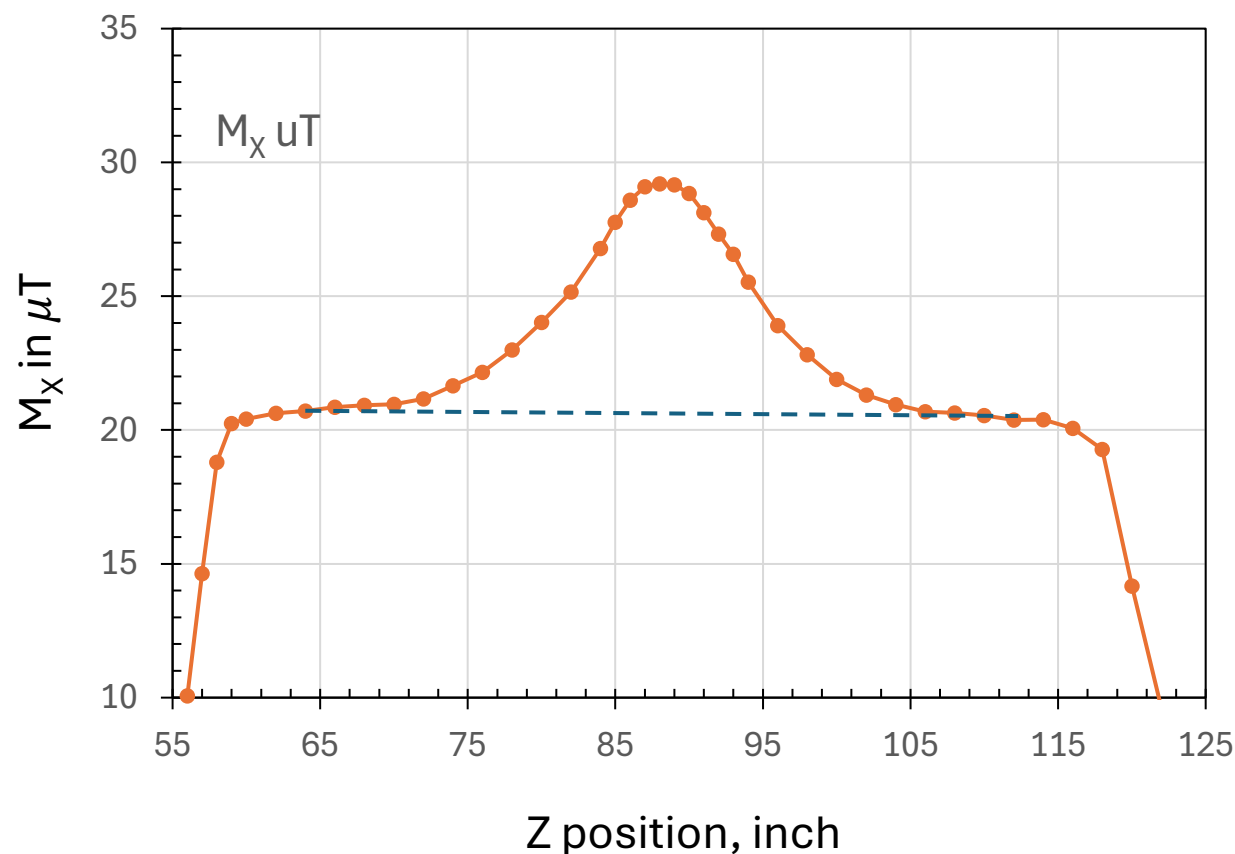
Single outer mu-layer
seen PERT at 1.3 A
 $\overline{\text{PERT}} = 27.6 \pm 3.7 \mu T$



DIFF dependence on current
 $\sim 5.27i^2 - 1.26i + 20.46 \mu T$

2 μ - OUTER layers

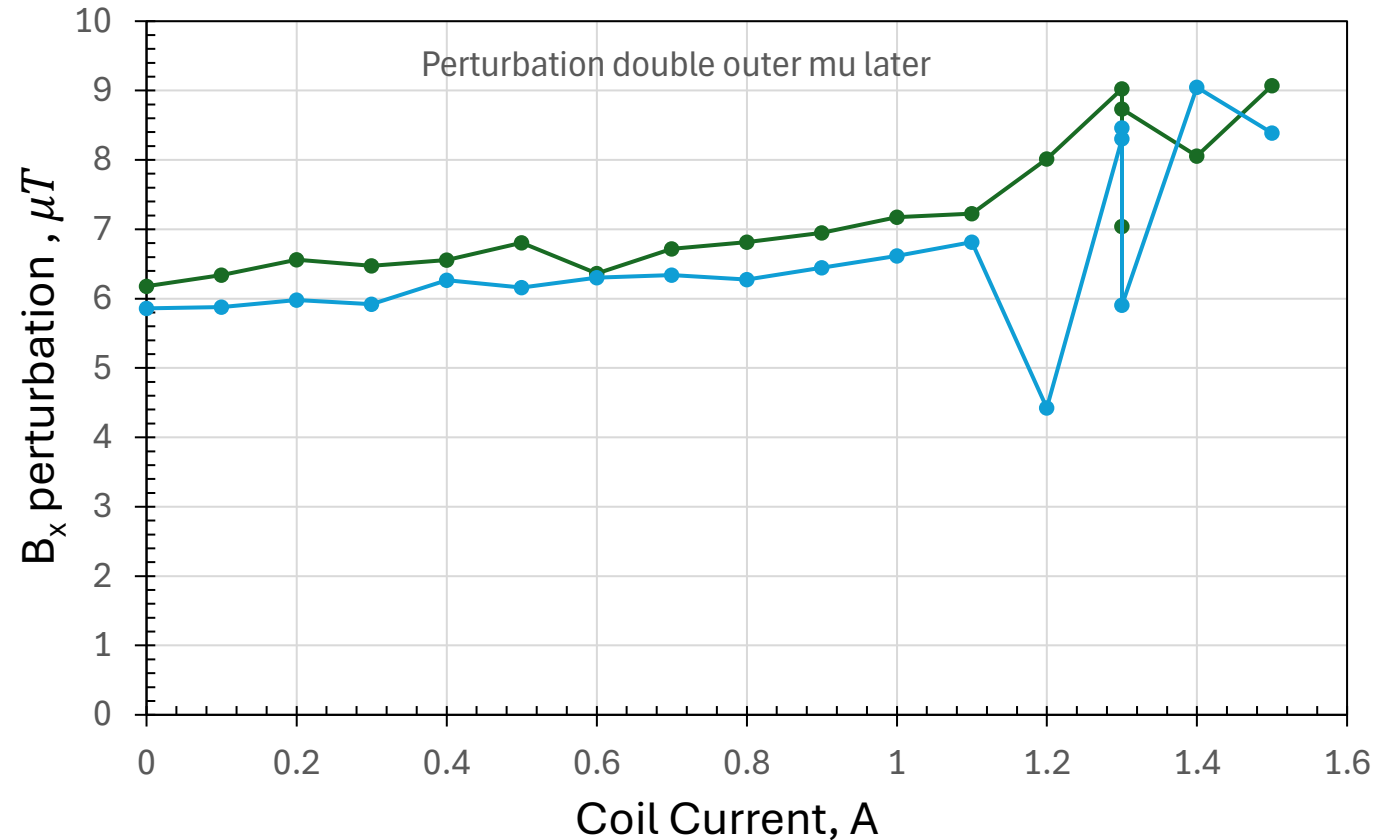
Special degaussing with 0A and NO PERT. Then, Z-scan with PERT, current 0 A



Peak position 88"
Perturbation $\sim 8.6 \mu T$

2 μ - OUTER layers

After Standard Degaussing, Z=88". DIFF=PERT-NOPERT, with Air



Double outer mu-layer
Seen PERT at 1.3 A
 $\overline{\text{PERT}} = 7.9 \pm 1.1 \mu T$

DIFF dependence on current
 $\sim 1.713i^2 - 1.088i + 6.311 \mu T$

Summary of shielding factors

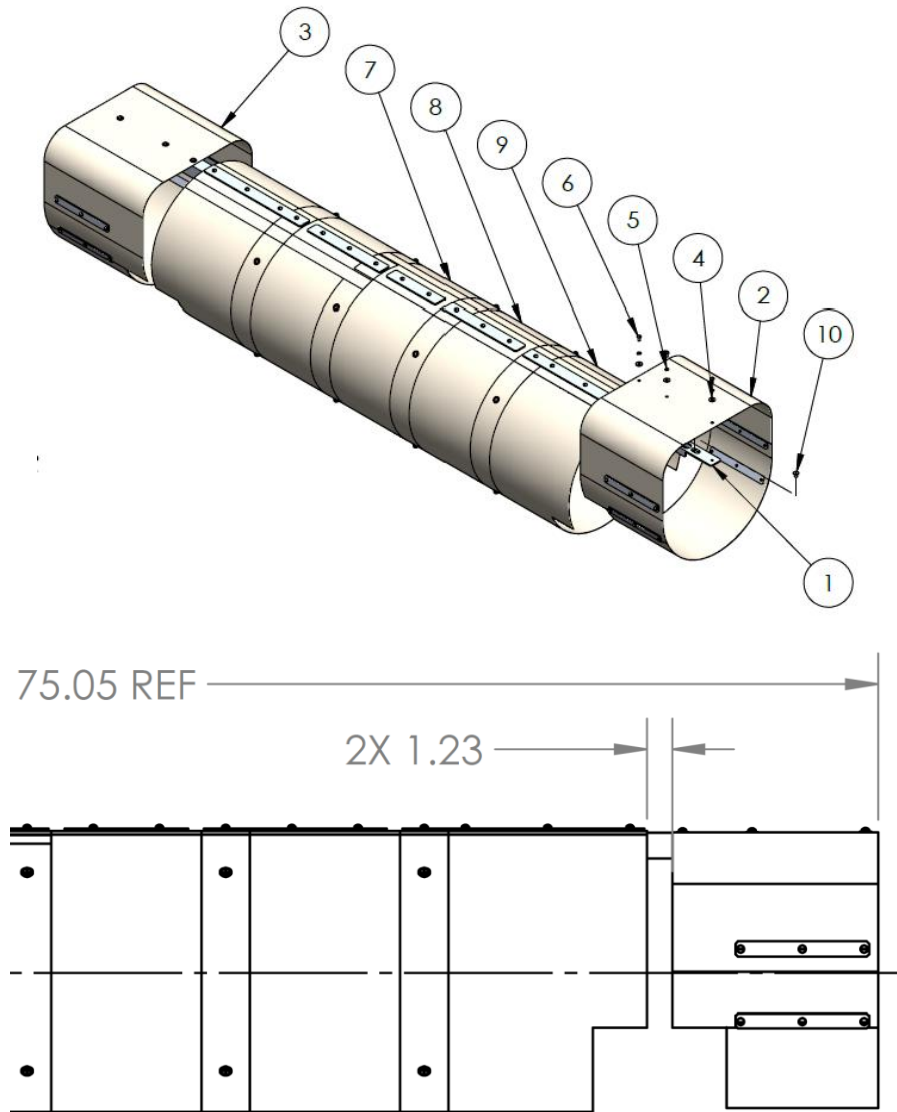
All measurements are after standard degaussing with coil current 1.3 A (~ 25 Gauss)

Configuration	Diameter, inch	Seen PERT μT	SF	μ_{eff}
1 μ inner layer	7.5 (0.46mm)	16.1 ± 1.8	32.6 ± 3.8	13,317
2 μ inner layer	7.6 (0.90mm)	4.51 ± 0.61	116.2 ± 16.1	24,050
1 μ outer layer	9.0 (0.38 mm)	27.6 ± 3.7	19.0 ± 2.6	11,220
2 μ outer layer	9.1(0.76 mm)	7.9 ± 1.1	66.3 ± 9.4	19,794

↑
Effective
thickness

- We did not have degaussing of the outer layer – maybe we should
- Double layer seems have higher μ probably due to effective space between two layers
- Layer thickness calculated effectively accounting for sheets overlay
- It was not possible measure SF with both inner and outer shields

Gap in nTMM outer shield



Gap simulated in prototype

